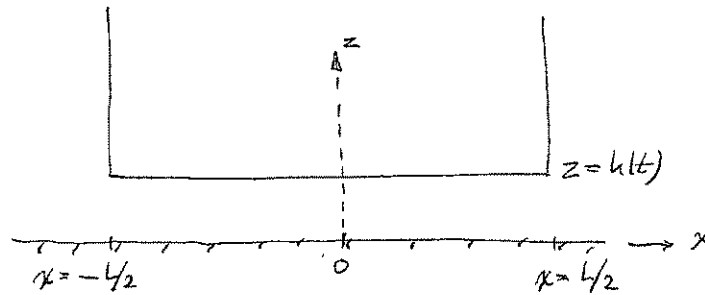


Problem Set No. 2 ; Problem 1 Solution

Separating a flat object from a smooth surface



\* Mass conservation:  $\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0$   $\frac{h}{L} \ll 1$

$u \sim U, w \sim W \Rightarrow W \sim U \frac{h}{L} \sim \frac{h}{T} \Rightarrow T \sim \frac{L}{U}$

\* x-momentum eqn:  $\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + w \frac{\partial u}{\partial z} = -\frac{1}{\rho} p_x + \mu \left( \frac{\partial^2 u}{\partial x^2} + \frac{\partial^2 u}{\partial z^2} \right)$

typical inertia terms  $\frac{\partial u}{\partial t} \sim \frac{U^2}{L} \sim u \frac{\partial u}{\partial x}, w \frac{\partial u}{\partial z}$

typical viscous term  $\mu \frac{\partial^2 u}{\partial z^2} \sim \frac{\mu U}{h^2}$

Neglect inertia terms if  $\frac{\mu U}{\rho h^2} \gg \frac{U^2}{L} \Rightarrow \frac{h^2 U}{L \nu} \ll 1$

ie, slow, viscous flow in a narrow gap.

$\therefore p_x = \mu \frac{\partial^2 u}{\partial z^2} \Rightarrow p \sim \mu \frac{U L}{h^2}$

\* z-momentum eqn Neglecting inertia terms again,

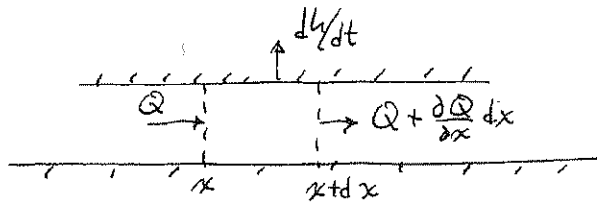
$p_z = \mu \frac{\partial^2 w}{\partial z^2}$   
 $p_z \sim \frac{p}{h} \sim \frac{\mu U L}{h^2} \frac{1}{h}, \mu \frac{\partial^2 w}{\partial z^2} \sim \mu \frac{W}{h^2} \sim \mu \frac{U h / L}{h^2} \sim \mu \frac{U}{h L}$   
 $\frac{p_z}{\mu \frac{\partial^2 w}{\partial z^2}} \sim \frac{L^2}{h^2} \gg 1 \Rightarrow \underline{p_z \approx 0} \quad \text{ie} \quad \underline{p \approx p(x)}$

\* Since  $p_x = G(x)$ , may solve for  $u(z, x)$ :

$$u(z, x) = \frac{G(x)}{2\mu} z(z-h) \quad (u = 0 \text{ at } z=0, h)$$

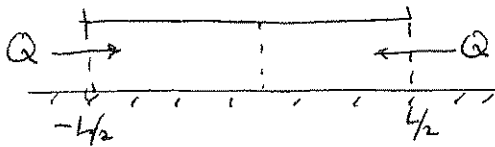
so locally we have Poiseuille flow

\* Impose mass conservation:



$$\frac{\partial}{\partial t} (\rho h dx) = - \frac{\partial Q}{\partial x} dx \Rightarrow \rho \frac{dh}{dt} = - \frac{\partial Q}{\partial x} \Rightarrow$$

$$Q = - \rho x \frac{dh}{dt}$$



$$Q(-\frac{h}{2}) = \rho \frac{h}{2} \frac{dh}{dt}$$

$$Q(\frac{h}{2}) = -\rho \frac{h}{2} \frac{dh}{dt}$$

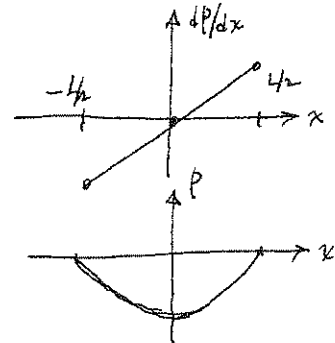
\* Also, from  $u(z, x) = \frac{dp/dx}{2\mu} z(z-h)$ ,

$$Q = \rho \int_0^h u dz = \rho \frac{dp/dx}{2\mu} \int_0^h (z^2 - zh) dz = -\rho \frac{dp/dx}{12\mu} h^3 = Q$$

$$\therefore -\rho \frac{dp/dx}{12\mu} h^3 = -\rho x \frac{dh}{dt} \Rightarrow \frac{dp}{dx} = \frac{12\mu}{h^3} x \frac{dh}{dt}$$

So, taking  $p=0$  at  $x = \pm h/2$ ,

$$p(x) = \frac{12\mu}{h^3} \frac{dh}{dt} \left( \frac{x^2}{2} - \frac{h^2}{8} \right)$$



\* Total force (directed downwards)

$$\begin{aligned}
 F &= \frac{12\mu}{h^3} \frac{dh}{dt} \int_{-4h}^{4h} \left( \frac{x^2}{2} - \frac{L^2}{8} \right) dx \\
 &= \frac{12\mu}{h^3} \frac{dh}{dt} \left( \frac{x^3}{6} - \frac{L^2 x}{8} \right) \Big|_{-4h}^{4h} = \frac{12\mu}{h^3} \frac{dh}{dt} \left( -\frac{h^3}{8} + \frac{L^3}{24} \right) \\
 \Rightarrow F &= \mu \frac{L^3}{h^3} \frac{dh}{dt} \Rightarrow \underline{\underline{\frac{dh}{dt} = \frac{Fh^3}{\mu L^3}}}
 \end{aligned}$$

\* Suppose the block falls under its own weight

$$F = -W$$

Hence, 
$$\frac{dh}{dt} = -\frac{Wh^3}{\mu L^3}$$

Integrating, 
$$\int_{h_0}^h \frac{dh}{h^3} = -\frac{W}{\mu L^3} \int_0^t dt$$

$$\Rightarrow \frac{1}{2} \left( \frac{1}{h_0^2} - \frac{1}{h^2} \right) = -\frac{W}{\mu L^3} t \Rightarrow$$

$$\begin{aligned}
 h^2 &= \frac{1}{\frac{1}{h_0^2} + \frac{2W}{\mu L^3} t} \Rightarrow \underline{\underline{h(t) = \frac{1}{\sqrt{\frac{1}{h_0^2} + \frac{2W}{\mu L^3} t}}}}
 \end{aligned}$$

Problem Set No. 2 ; Problem 2 Solution

Spreading of lava on a horizontal plane

\* From lubrication approximation, local radial velocity

$$u(r, z, t) = - \frac{g}{2\nu} \frac{\partial h}{\partial r} z (2h - z)$$

Conservation of mass requires that

$$\frac{\partial h}{\partial t} + \frac{1}{r} \frac{\partial}{\partial r} \left( r \int_0^h u dz \right) = 0$$

But  $\int_0^h u dz = - \frac{g}{2\nu} \frac{\partial h}{\partial r} \int_0^h z (2h - z) dz =$   
 $= - \frac{g}{2\nu} \frac{\partial h}{\partial r} \left( h z^2 \Big|_0^h - \frac{z^3}{3} \Big|_0^h \right) = - \frac{g}{3\nu} h^3 \frac{\partial h}{\partial r}$

∴ evolution equation for  $h(r, t)$  is

$$\frac{\partial h}{\partial t} = \frac{g}{3\nu} \frac{1}{r} \frac{\partial}{\partial r} \left( r h^3 \frac{\partial h}{\partial r} \right) \quad \left. \begin{array}{l} \text{B.C.} \\ h(R(t), t) = 0 \\ \frac{\partial h}{\partial r}(0, t) = 0 \end{array} \right\}$$

\* Consider  $\frac{\partial}{\partial t} \int_0^{R(t)} r h(r, t) dr = \int_0^{R(t)} r \frac{\partial h}{\partial t} dr + \frac{\partial R}{\partial t} R h(R, t)$   
 $= \int_0^R \cancel{\frac{g}{3\nu}} \frac{1}{r} \frac{\partial}{\partial r} \left( r h^3 \frac{\partial h}{\partial r} \right) dr = \frac{g}{3\nu} \cancel{R h^3} \frac{\partial h}{\partial r} \Big|_R = 0$

∴  $\int_0^R r h(r, t) dr = \text{const}$

Let  $\underline{\underline{2\pi \int_0^{R(t)} r h(r, t) dr = V}}$

\* Similarity solution

$$h = \lambda^a h', \quad r = \lambda^b r', \quad t = \lambda^c t'$$

$$\frac{\partial h}{\partial t} = \frac{g}{3\rho} \frac{1}{r} \frac{\partial}{\partial r} \left( r h^3 \frac{\partial h}{\partial r} \right) \Rightarrow \lambda^{a-c} \frac{\partial h'}{\partial t'} = \frac{g}{3\rho} \lambda^{4a-2b} \frac{1}{r'} \frac{\partial}{\partial r'} \left( r' h'^3 \frac{\partial h'}{\partial r'} \right)$$

$$2\pi \int_0^R r h(r,t) dr = V \Rightarrow 2\pi \lambda^{2b+a} \int_0^{R'} r' h(r',t') dr' = V$$

$$\text{Require } \left. \begin{array}{l} a-c = 4a-2b \\ 2b+a = 0 \end{array} \right\} \Rightarrow \begin{array}{l} b = -\frac{a}{2} \\ \underline{\underline{c = -4a}} \end{array}$$

$$\text{So, } h = \lambda^a h', \quad r = \lambda^{-\frac{a}{2}} r', \quad t = \lambda^{-4a} t'$$

$$\text{Note that } \frac{r}{t^{1/8}} = \lambda^{-\frac{a}{2} + \frac{a}{2}} \frac{r'}{t'^{1/8}}$$

Therefore, the similarity variable is  $\eta = A \frac{r}{t^{1/8}}$

$$\text{and } t^{1/4} h = \lambda^{-a} \lambda^a h' \quad \underline{\underline{B f(\eta) = t^{1/4} h}}$$

So, we write

$$\underline{\underline{h = \frac{B f(\eta)}{t^{1/4}}, \quad \eta = A \frac{r}{t^{1/8}}}}$$

$$\frac{\partial h}{\partial r} = \frac{B}{t^{1/4}} \frac{A}{t^{1/8}} f'(\eta), \quad \frac{\partial h}{\partial t} = -\frac{1}{8} \frac{B}{t^{5/4}} (2f(\eta) + \eta f'(\eta))$$

Manipulation

$$\frac{\partial}{\partial r} \left( r h^3 \frac{\partial h}{\partial r} \right) = \frac{\partial}{\partial r} \left( r \frac{B^3}{t^{3/4}} f^3 \frac{BA}{t^{1/8}} f'(\eta) \right)$$

$$= \frac{B^4 A}{t^{9/8}} f^3 f' + \frac{r B^4 A^2}{t^{5/4}} \frac{d}{d\eta} (f^3 f') \quad \left( C \equiv \frac{g}{3\nu} \right)$$

$$\therefore \frac{\partial h}{\partial t} = C \frac{1}{r} \frac{\partial}{\partial r} \left( r h^3 \frac{\partial h}{\partial r} \right) \Rightarrow$$

$$-\frac{1}{8} \frac{B}{t^{5/4}} (2f(\eta) + \eta f'(\eta)) = C \left\{ \frac{B^4 A}{t^{9/8}} \frac{1}{r} f^3 f' + \frac{B^4 A^2}{t^{5/4}} \frac{d}{d\eta} (f^3 f') \right\}$$

$$\text{Since } r = \frac{\eta t^{1/8}}{A},$$

we have

$$-\frac{1}{8} B (2f + \eta f') = C \left\{ \frac{B^4 A^2}{\eta} f^3 f' + B^4 A^2 \frac{d}{d\eta} (f^3 f') \right\}$$

$$\text{Also, } 2\pi \int_0^R r h(\eta, t) dr = V \Rightarrow$$

$$2\pi \left( \frac{t^{1/8}}{A} \right)^2 \frac{B}{t^{1/4}} \int_0^{\eta_R} \eta f(\eta) d\eta = V \Rightarrow \underline{2\pi \frac{B}{A^2} \int_0^{\eta_R} \eta f(\eta) d\eta = V}$$

\* Choose (for simplicity)

$$\underline{B^3 A^2 = \frac{1}{8C}} \Rightarrow A^2 = \frac{1}{8CB^3}$$

$$2\pi \frac{B}{A^2} \frac{1}{V} = 2\pi \frac{B^4 C}{V} = 1 \Rightarrow$$

$$\underline{B^4 = \frac{V}{2\pi C}}$$

\* Solution

$$-2f - \eta \frac{df}{d\eta} = \frac{1}{\eta} f^3 \frac{df}{d\eta} + \frac{d}{d\eta} \left( f^3 \frac{df}{d\eta} \right)$$

$$\Rightarrow 2\eta f + \eta^2 \frac{df}{d\eta} + f^3 \frac{df}{d\eta} + \eta \frac{d}{d\eta} \left( f^3 \frac{df}{d\eta} \right) = 0$$

$$\Rightarrow \frac{d}{d\eta} \left\{ \eta^2 f + \eta f^3 \frac{df}{d\eta} \right\} = 0$$

$$\Rightarrow \eta^2 f + \eta f^3 \frac{df}{d\eta} = 0$$

$$\Rightarrow \underline{\underline{\eta + f^2 \frac{df}{d\eta} = 0}} \quad \text{with } \underline{\underline{BC}} \quad \frac{df}{d\eta} = 0 \quad (\eta = 0)$$

$$f(\eta_R) = 0$$

May integrate analytically:

$$f^2 df = -\eta d\eta \Rightarrow \frac{1}{3} d(f^3) = -\frac{1}{2} d\eta^2$$

$$\frac{f^3}{3} = -\frac{\eta^2}{2} + \text{const} = -\frac{1}{2} (\eta^2 - \eta_R^2)$$

$$\Rightarrow \underline{\underline{f(\eta) = \left\{ \frac{3}{2} (\eta_R^2 - \eta^2) \right\}^{1/3}}}$$

\* To determine  $\eta_R$  impose  $\int_0^{\eta_R} \eta f(\eta) d\eta = 1$

$$\Rightarrow \left(\frac{3}{2}\right)^{1/3} \int_0^{\eta_R} \eta (\eta_R^2 - \eta^2)^{1/3} d\eta = 1$$

\* Since  $\eta_R = A \frac{R}{t^{1/8}} \Rightarrow R \propto t^{1/8}$  as  $t$  increases

$h = \frac{f(\eta)}{t^{1/4}} \Rightarrow h \sim \frac{f(0)}{t^{1/4}}$  as  $t$  increases

## Problem Set No. 2 ; Problem 3 Solution

### Rigid sphere rotating in a viscous fluid at low Re

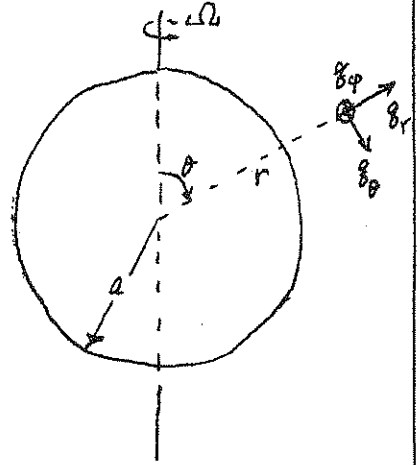
\* It is convenient to use spherical polar coordinates

$r$  (radial),  $\theta$  (polar),  $\phi$  (azimuthal)  
(see Appendix 2 of Batchelor's book)

\* Owing to symmetry, the velocity field

$$\underline{q} = (q_r, q_\theta, q_\phi) \text{ and the pressure } p$$

are independent of  $\phi$



\* At low Reynolds number,  $\frac{\Omega a^2}{\nu} \ll 1$ ,

inertia terms are expected to be negligible; this in turn implies that  $\underline{q} = q_\phi \hat{e}_\phi$ , with  $q_\phi = q_\phi(r, \theta)$

Indeed, the continuity equation

$$\nabla \cdot \underline{q} = \frac{1}{r^2} \left( \frac{\partial}{\partial r} (r^2 q_r) \right) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta q_\theta) + \frac{1}{r \sin \theta} \frac{\partial q_\phi}{\partial \phi} = 0$$

is automatically satisfied;

moreover, the only surviving inertia terms in the  $r$ - and  $\theta$ -mom. eqns

$$\frac{q_\phi^2}{r} \sim \frac{\Omega^2 a^4}{r} \sim \Omega^2 a^3,$$

are much smaller than the typical viscous term

$$\nu \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial q_\phi}{\partial r} \right) \sim \frac{\nu q_\phi}{a^2} \sim \frac{\nu \Omega a}{\phi a}$$

kept in the  $\phi$ -mom. equation, since  $\frac{\nu \Omega}{a} \gg \Omega^2 a$

$$\Rightarrow \frac{\Omega a^2}{\nu} \ll 1$$



Therefore,  $\varphi$ -momentum eqn<sup>s</sup>  $\Rightarrow$

$$\nu \left\{ \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial \varphi}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \varphi}{\partial \theta} \right) - \frac{\varphi}{r^2 \sin^2 \theta} \right\} = 0 \quad (*)$$

\* This (linear) equation for  $\varphi$  has a separable solution:

$$\underline{\varphi = f(r) \sin \theta} \quad (**)$$

(This is motivated by the fact that on the sphere ( $r=a$ ),

$$\varphi \hat{e}_\varphi = \underline{\underline{\Omega}} \times a \hat{e}_r = \Omega \cdot a \sin \theta \hat{e}_\varphi)$$

Upon substitution of  $\varphi$   $(**)$  into  $(*)$ , we find that  $f(r)$  satisfies:

$$\frac{d}{dr} \left( r^2 \frac{df}{dr} \right) - 2f(r) = 0$$

This ode has solutions of the form  $r^n$  for  $n = -2, 1$   
Since  $\varphi \rightarrow 0$  as  $r \rightarrow \infty$ , only  $n = -2$  is acceptable

Hence,  $\varphi = \frac{A}{r^2} \sin \theta$

No slip at  $r=a \Rightarrow \Omega a = \frac{A}{a^2} \Rightarrow A = \Omega a^3$

So, finally  $\underline{\underline{\varphi = \frac{\Omega a^3}{r^2} \sin \theta}}$

\* On the surface of the sphere

$$\tau_{r\varphi} \Big|_{r=a} = \mu \left( \frac{\partial \varphi}{\partial r} - \frac{\varphi}{r} \right) \Big|_{r=a} = -3\mu \Omega \sin \theta$$

Therefore, the moment exerted by the fluid on the sphere is given by

$$M = \int_0^\pi \tau_{r\theta} \Big|_r=a \sin\theta \cdot 2\pi a^2 \sin\theta \, d\theta = -8\pi\mu a^3\Omega$$

\* An alternative way to deduce  $M$  is by using the fact that

$$\left( \begin{array}{l} \text{power required to} \\ \text{maintain steady} \\ \text{rotation of sphere} \end{array} \right) = (-M) \cdot \Omega = \left( \begin{array}{l} \text{rate of viscous} \\ \text{dissipation in the} \\ \text{fluid} \end{array} \right)$$

From problem 2 of PS # 1, we know that

$$\left( \begin{array}{l} \text{rate of viscous} \\ \text{dissipation} \end{array} \right) = \iiint_V \Phi \, dV \quad \text{where } \Phi = \frac{1}{2}\mu \left( \frac{\partial \delta_j}{\partial x_j} + \frac{\partial \delta_i}{\partial x_i} \right)^2$$

$$\begin{aligned} \text{Here, } \Phi &= 4\mu e_{r\theta}^2 = \mu r^2 \left( \frac{\partial}{\partial r} \left( \frac{\delta_\theta}{r} \right) \right)^2 = \mu \left( \frac{\partial \delta_\theta}{\partial r} - \frac{\delta_\theta}{r} \right)^2 \\ &= 9\mu \Omega^2 a^6 \frac{\sin^2\theta}{r^6} \end{aligned}$$

Now,

$$\begin{aligned} \iiint_V \Phi \, dV &= 9\mu \Omega^2 a^6 2\pi \int_0^\pi d\theta \int_a^\infty dr \frac{\sin^2\theta}{r^6} r^2 \sin\theta \\ &= 8\pi\mu a^3 \Omega^2 \end{aligned}$$

$$\text{Hence } M = -8\pi\mu a^3 \Omega \quad (\text{as before})$$