



Illinois Center for Advanced Studies of the Universe



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Lecture 1 on Hot QCD Matter: Lattice QCD at finite T

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National Nuclear Physics Summer School
MIT 2022

Maximum temperature of matter?

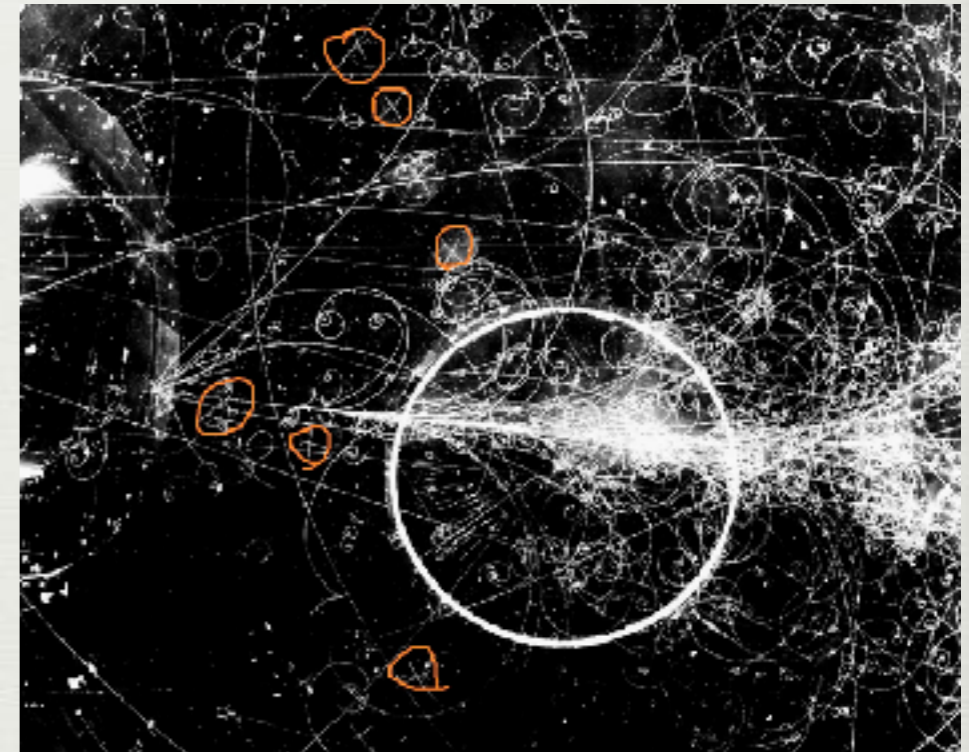
- In the 1960's Hagedorn suggested that matter has a maximum temperature, now known as the Hagedorn Temperature
- Instead of \uparrow temperature, heavier particles are created
- Hagedorn States: "fireballs consist of fireballs, which consist of fireballs..."



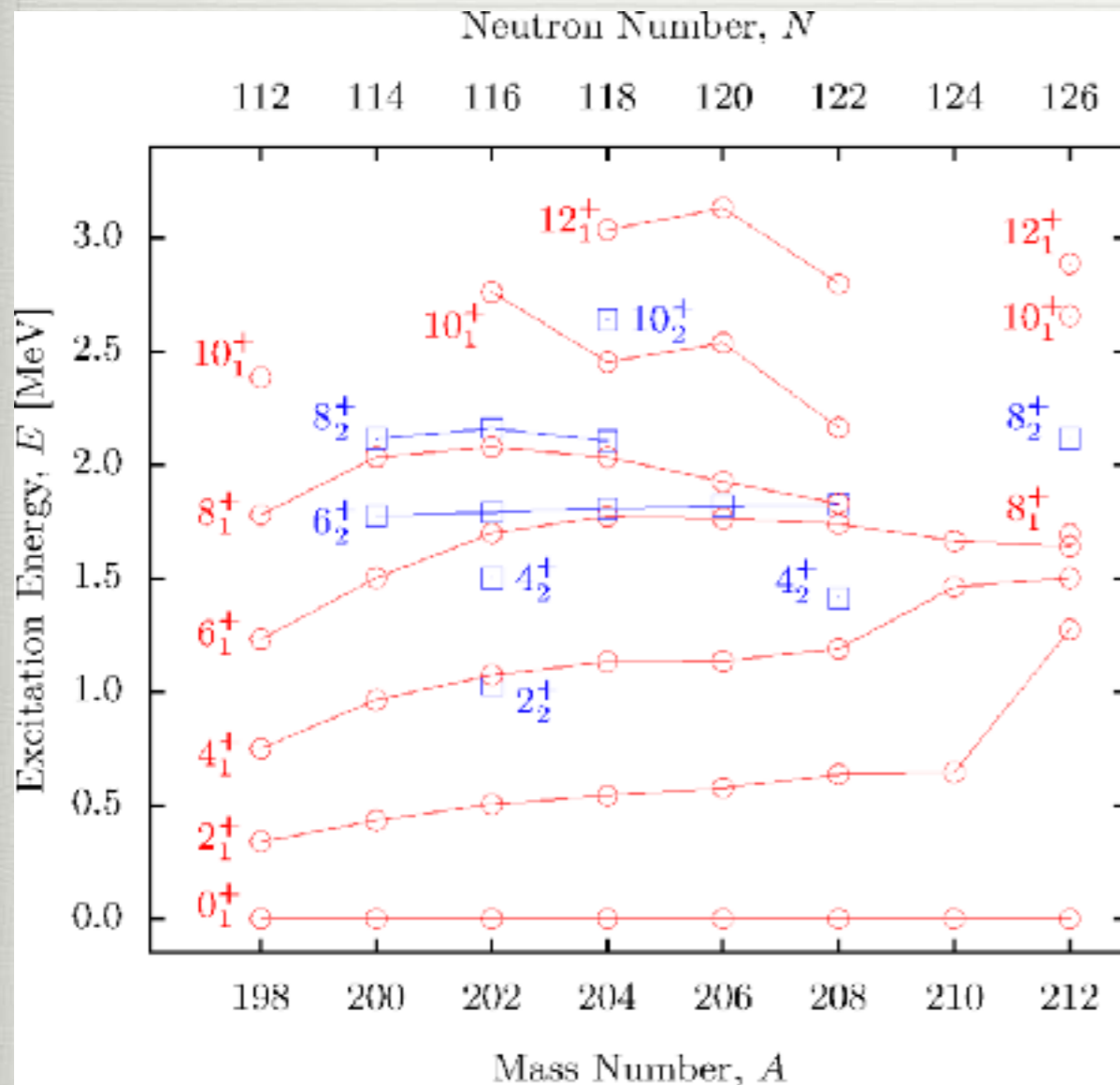
Rolf Hagedorn

Hadrons: “Stable” vs. resonances

- Light mesons: π (stable), ρ , ...
- Light baryons: p (super stable), n (stable), \mathbf{N} , Δ , ...
- Strange mesons: \mathbf{K} (stable), ...
- $S = \pm 1$ baryons: Λ , Σ (stable), ...
- $S = \pm 2$ baryons: Ξ (stable), ...
- $S = \pm 3$ baryons: Ω (stable), ...



Nuclear excited states vs. resonances



$m_\rho \sim 770 \text{ MeV}$

$m_\pi \sim 140 \text{ MeV}$

Mass number ~ 200

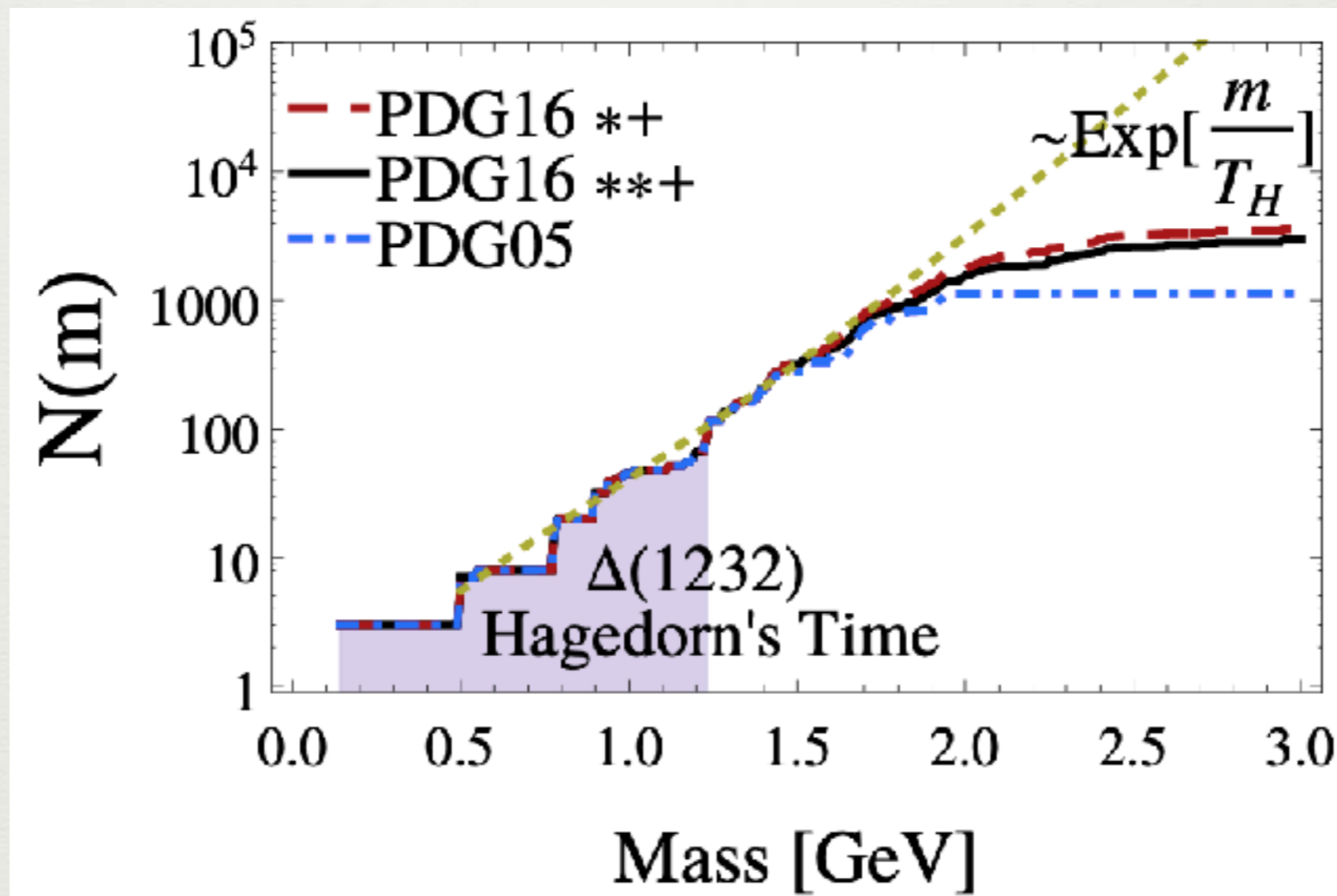
Mass $m_\pi \sim 140 \text{ MeV}$

Excitation energy $\mathcal{O}(1 \text{ MeV})$

Mass difference $\Delta m \sim 630 \text{ MeV}$

Limiting temperature: 1960's to today

Count up all known particles, fit to exponential mass spectrum



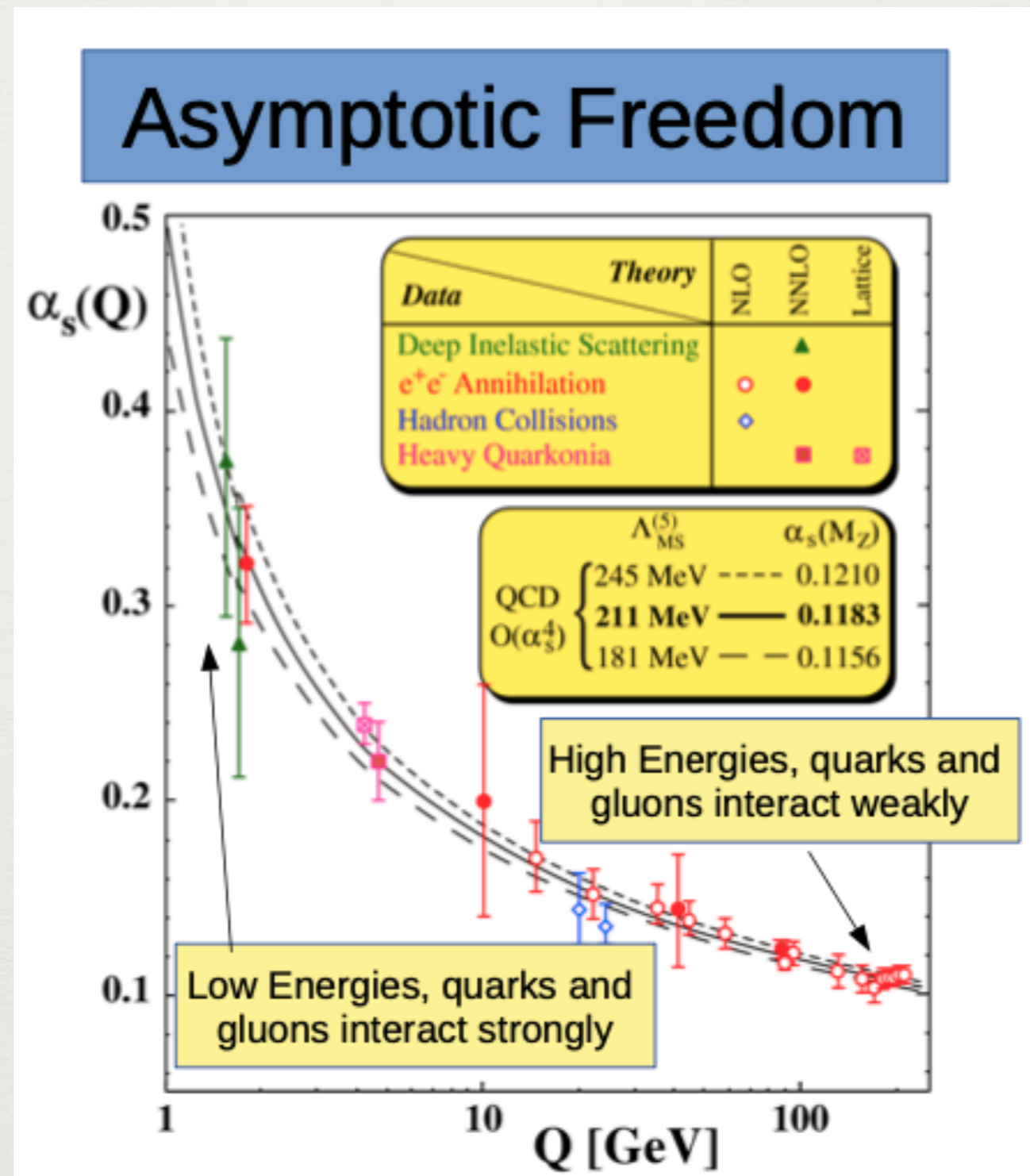
Particle Data Book 2021 gives $T_H \sim 170 \text{ MeV}$

What happens beyond this temperature?

- Normally, you add energy to the system, the system heats up
- Hagedorn: if there's a limiting temperature, you open new degrees of freedom
- What happens to matter at the highest temperatures we can reach on Earth?

Asymptotic Freedom: hitting QCD matter at high energies

Hadrons

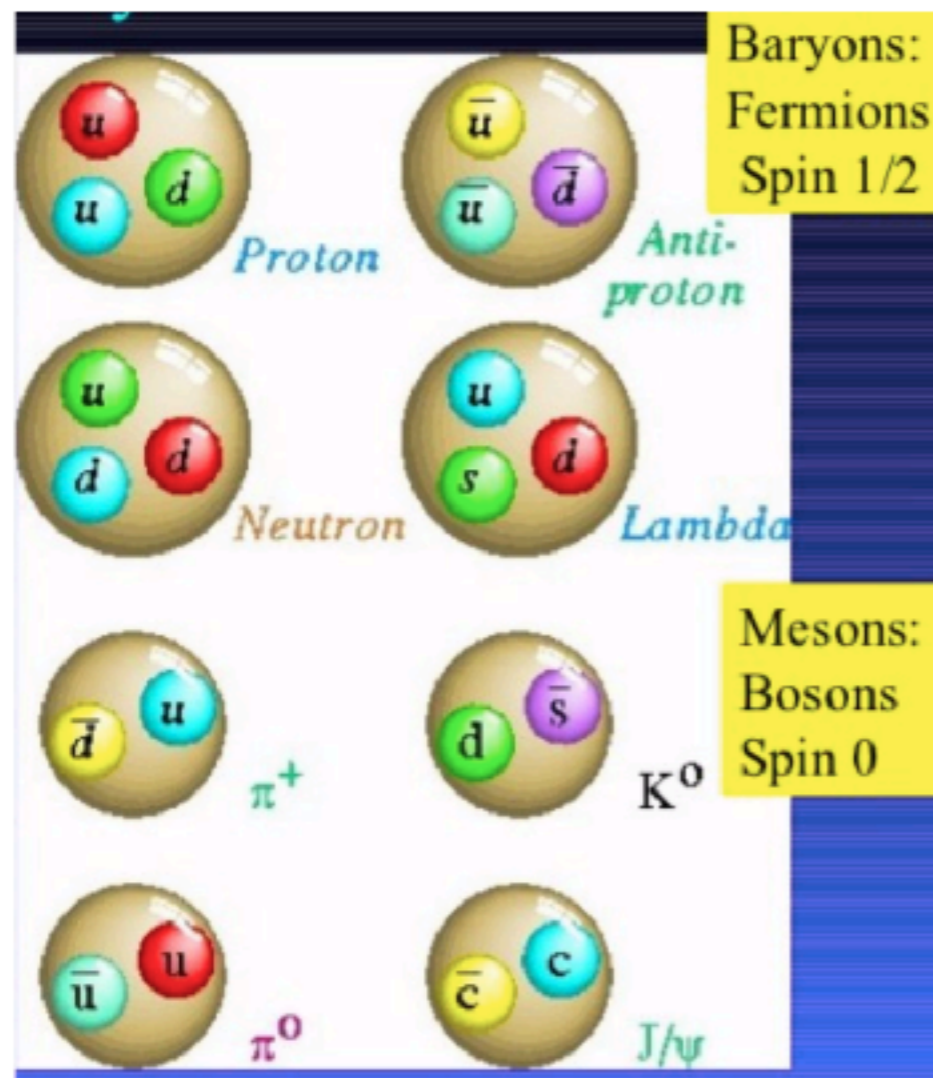


Quarks and gluons

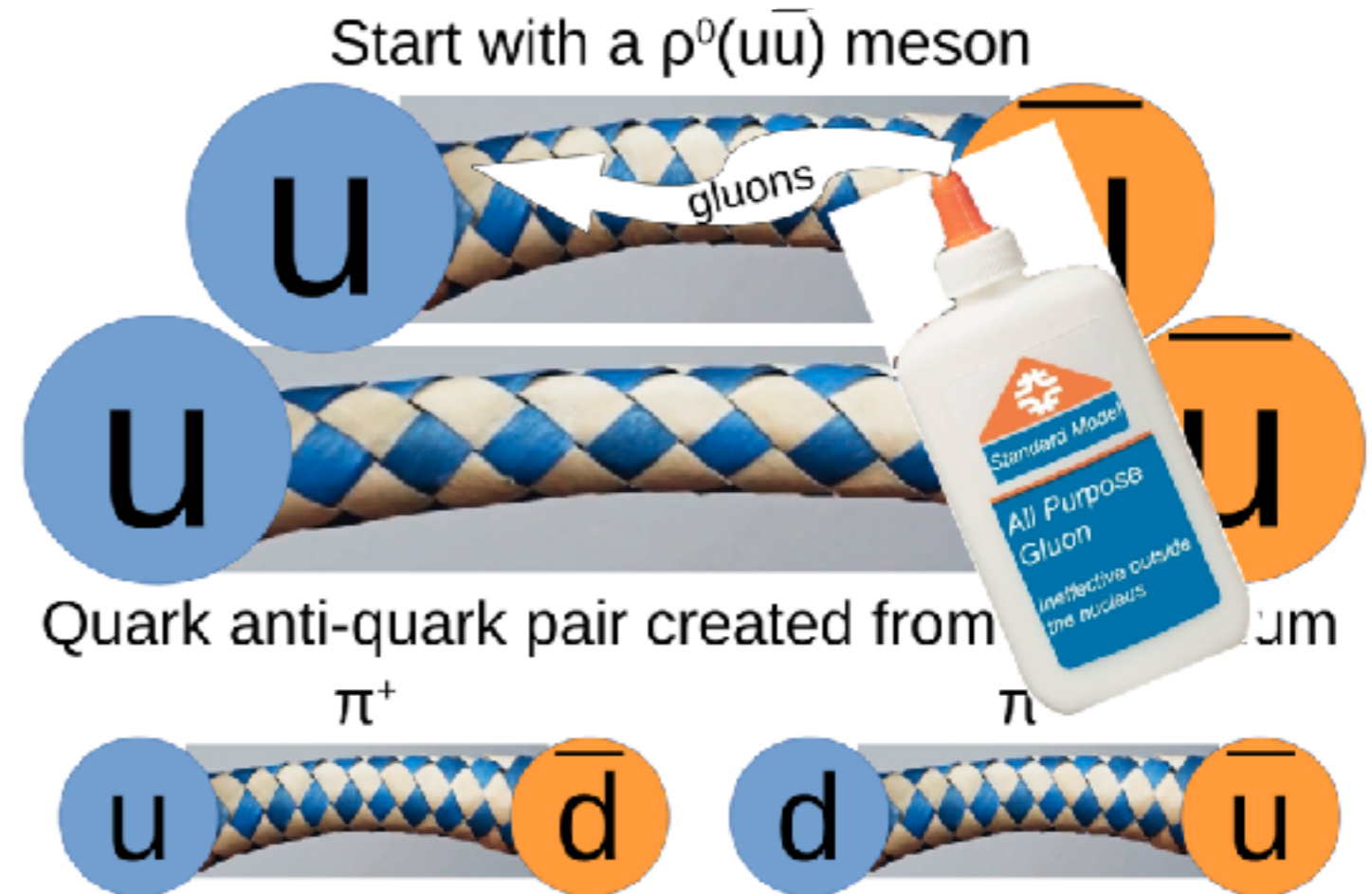
Theory of the strong force: Quantum Chromodynamics (QCD)

Quantum Chromodynamics (QCD)- David Politzer, Frank Wilczek and David Gross 1973 (2004 Nobel Prize for Physics)

Bound states=Hadrons

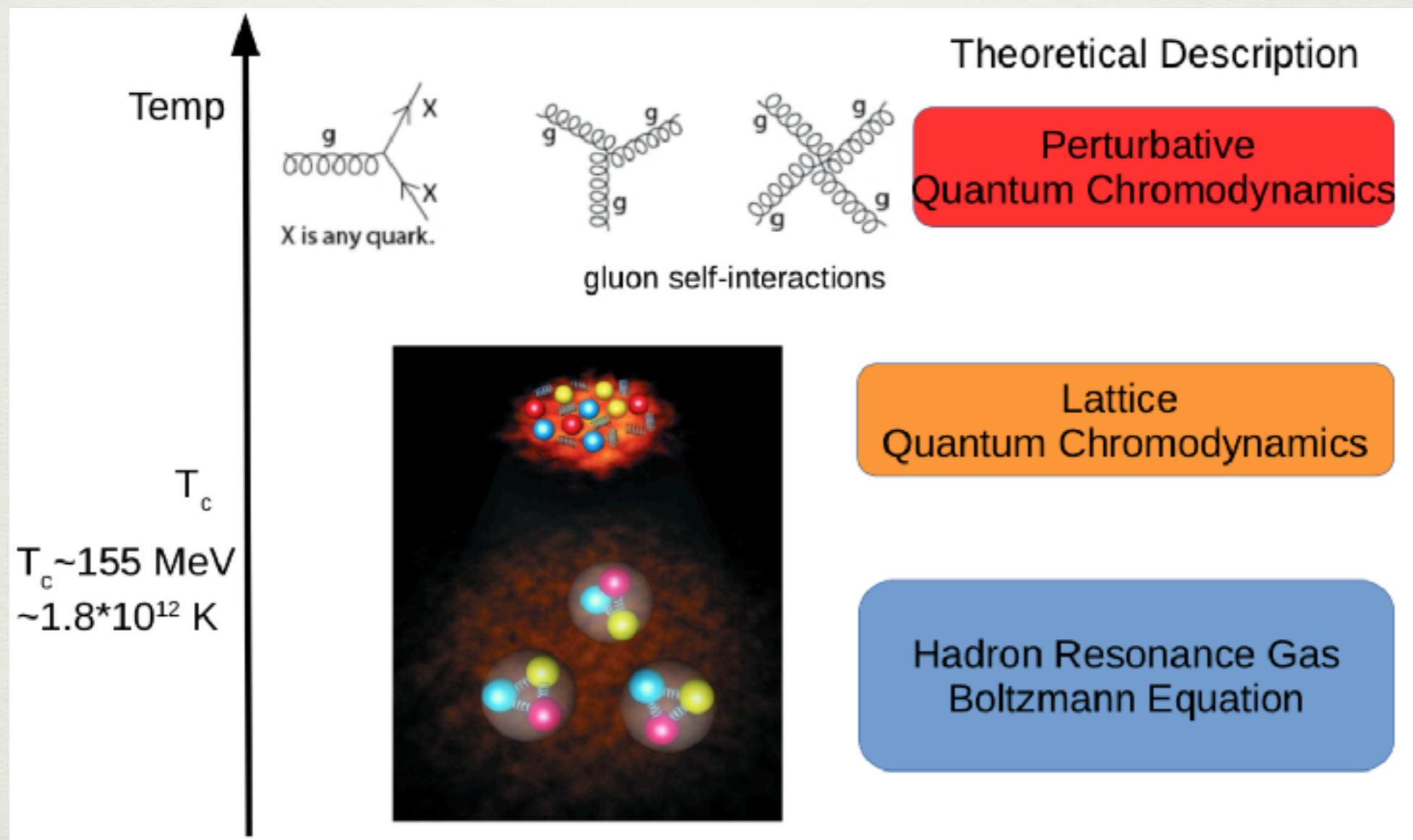


Confinement - no free quarks



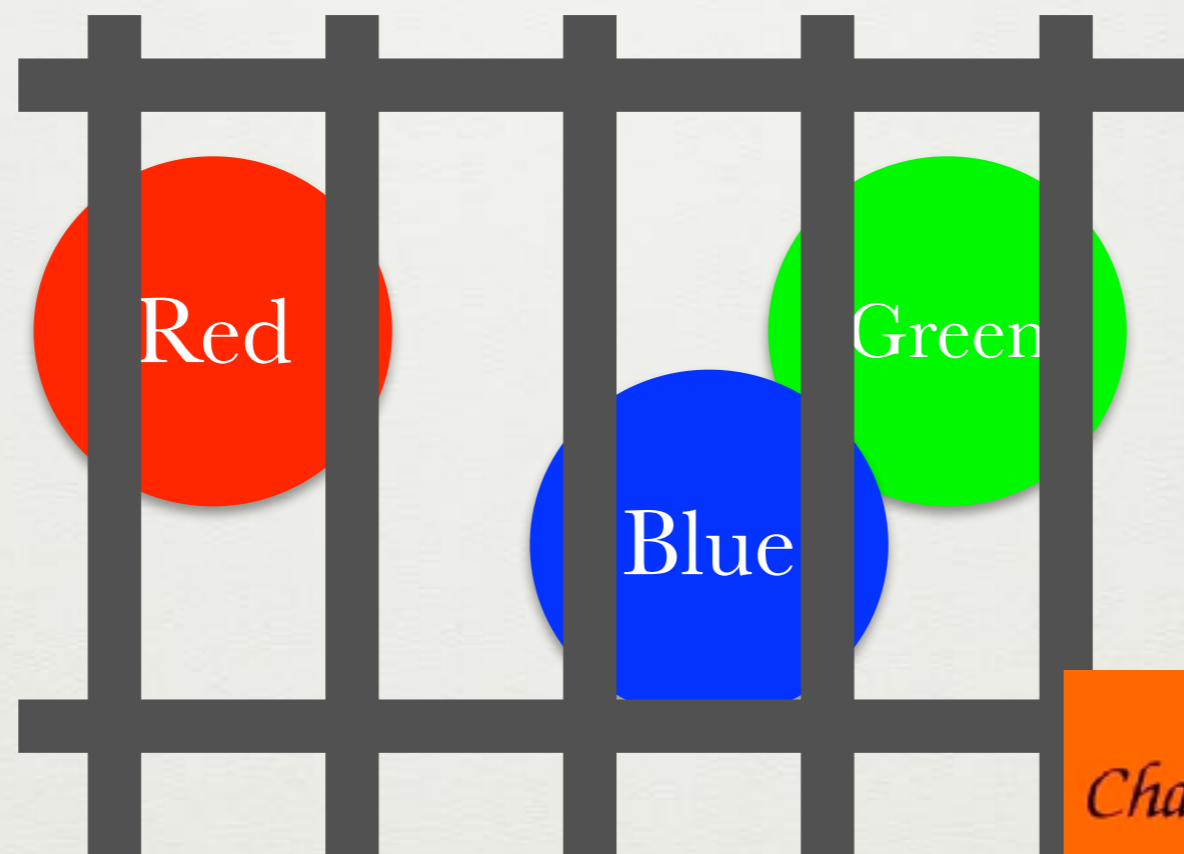
Solving Quantum Chromodynamics

$$L_{QCD} = \underbrace{-\frac{1}{4} F_{\mu\nu}^a F_a^{\mu\nu}}_{\text{Gluon Interactions}} + \underbrace{\bar{\psi}^q \left(i\gamma_\mu D^\mu - m_q \right) \psi^q}_{\text{Quark Interactions}} \quad \text{where } D^\mu = \partial^\mu - ig \underbrace{A^\mu(x)}_{\text{Gluons}}$$



Deconfinement of QCD matter

High enough temperatures/large enough energies can deconfine quarks and gluons



Chance

THIS CARD MAY BE KEPT
UNTIL NEEDED, OR SOLD

GET OUT OF JAIL
~~FREE~~

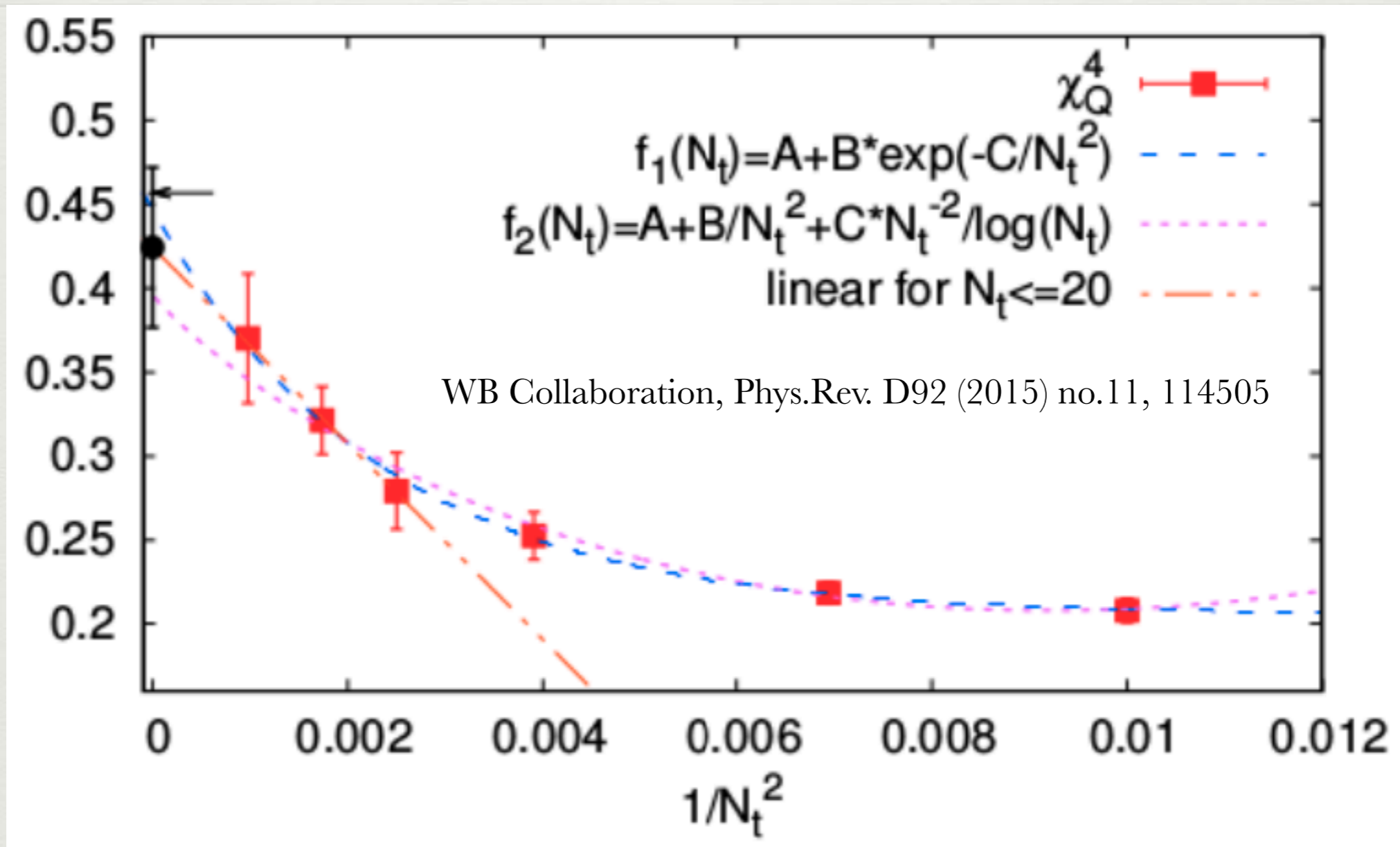
~Billion dollars



Finite Temperature QCD

- Analytical solutions are not possible, however, solving Quantum Chromodynamics on a lattice makes calculations possible.
- Limits of QCD \mathcal{L}
 - $m_q \rightarrow \infty$ pure gluon theory (quenched lattice)
 - $m_q \rightarrow 0$ chiral limit
 - $m_q \neq 0$ but u,d,s stronger interactions (non-perturbative)

Caution: Lattice Spacing and finite T observables



Too coarse of lattice spacing can lead to misleading results!

Lattice QCD basics

$$\text{Temperature: } T = \frac{1}{N_T a}$$

N_T - number of links in time direction

a - lattice spacing

$$\text{Volume: } V = (N_S a)^3$$

Continuum Limit:

$$N_S \rightarrow \infty$$

$$N_t \rightarrow \infty$$

$$a \rightarrow 0$$

While $T = \text{const}$ and $V = \text{const}$

Partition function and thermodynamics

- Sampling over gauge configurations gives the action and from there one calculate the partition function

$$Z = \int \prod_{n,\mu} dU_\mu(n) e^{-S_g} \det(M[U])$$

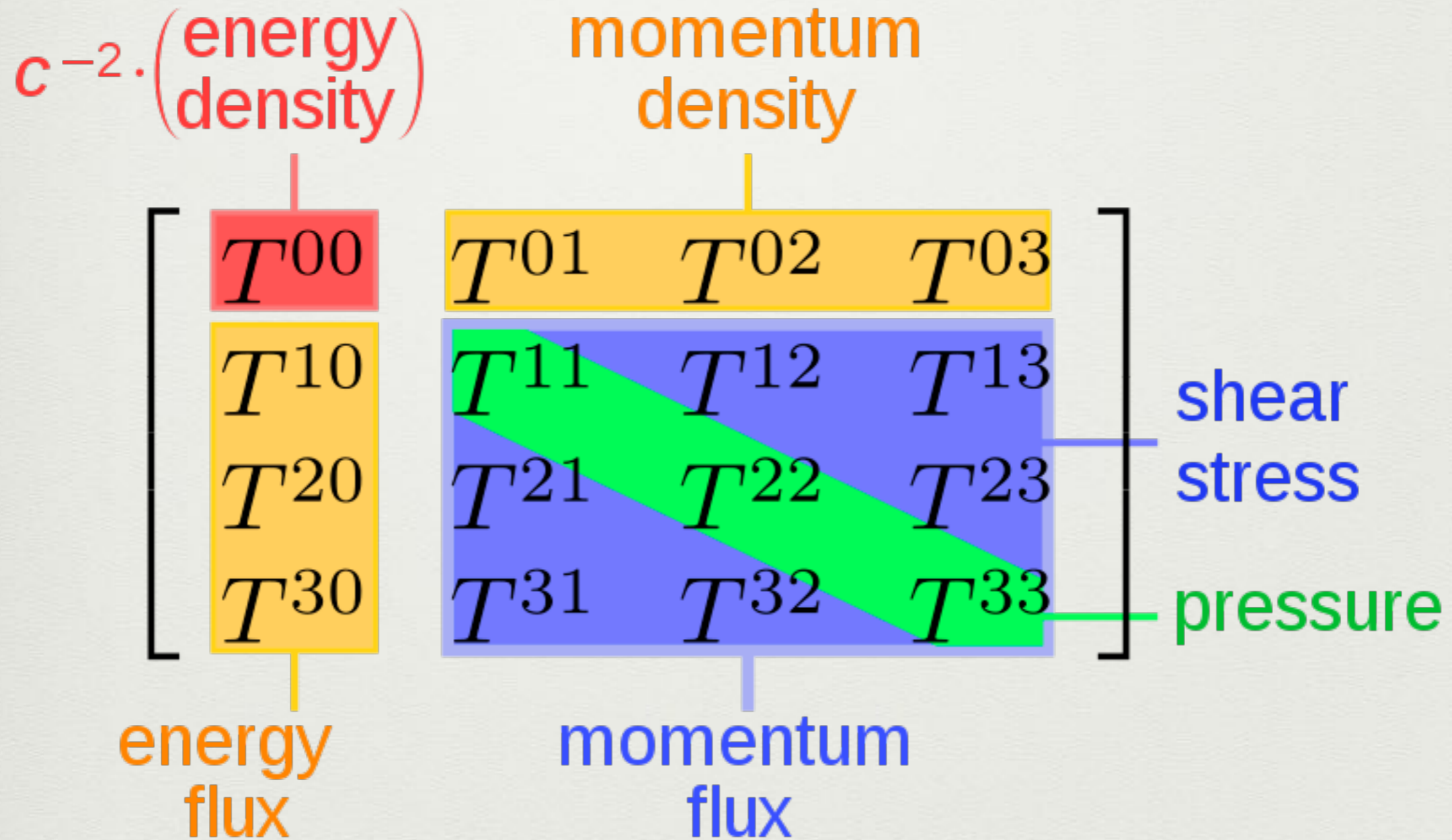
- From Z , one can calculate the pressure ($V \rightarrow \infty$ limit):

$$p = \frac{T}{V} \ln Z$$

or trace anomaly:

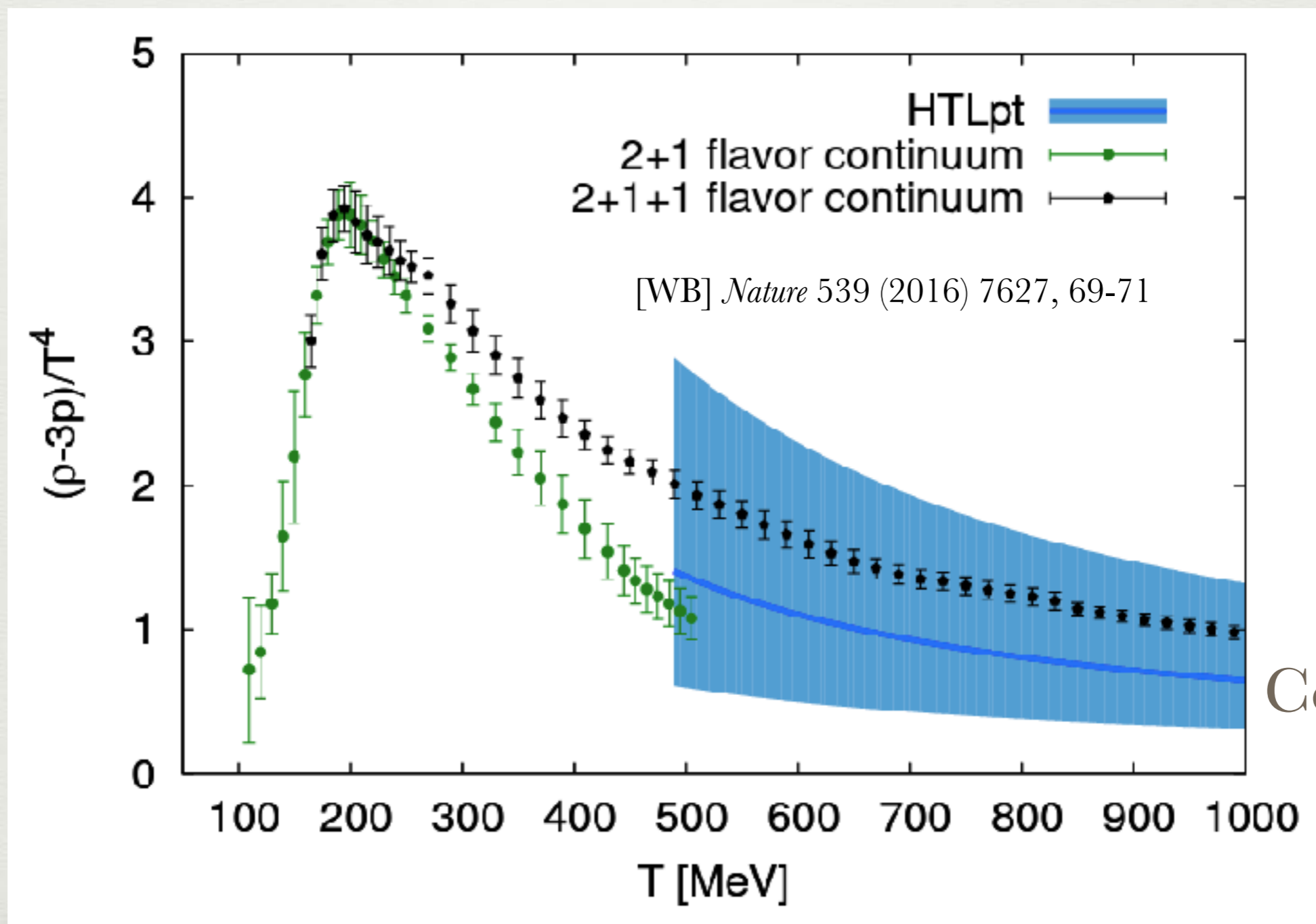
$$\varepsilon - 3p = - \frac{T}{V} \frac{d \ln Z}{d \ln a}$$

Trace anomaly $\varepsilon - 3p$ for
 $\{+, -, -, -\}$



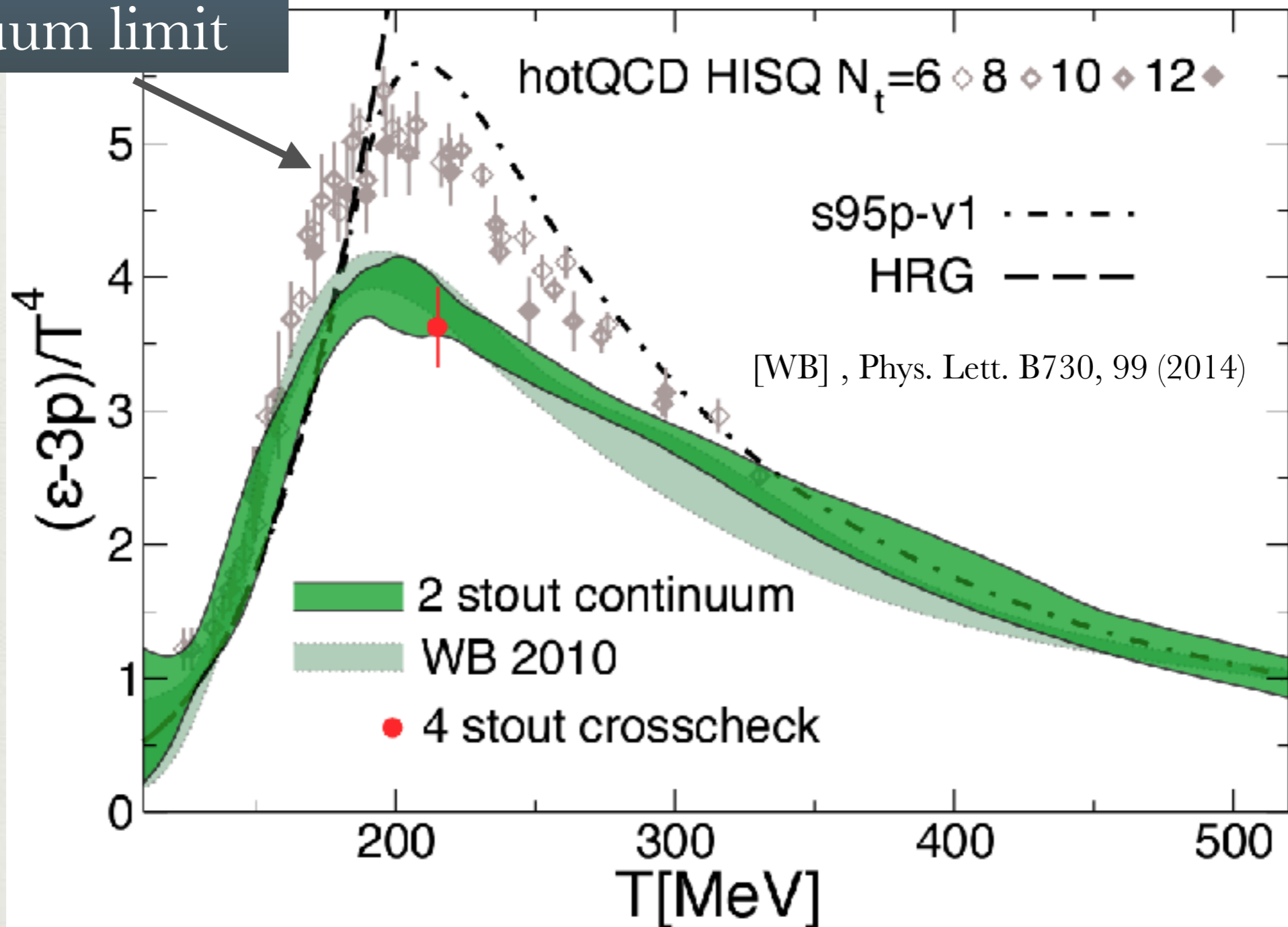
Trace anomaly

Depends on the numbers of quarks



Influence of lattice spacing

Not yet in the
continuum limit



Thermodynamic relations at $\mu_B = 0$

From the trace anomaly ($\varepsilon - 3p$), one can work out the rest

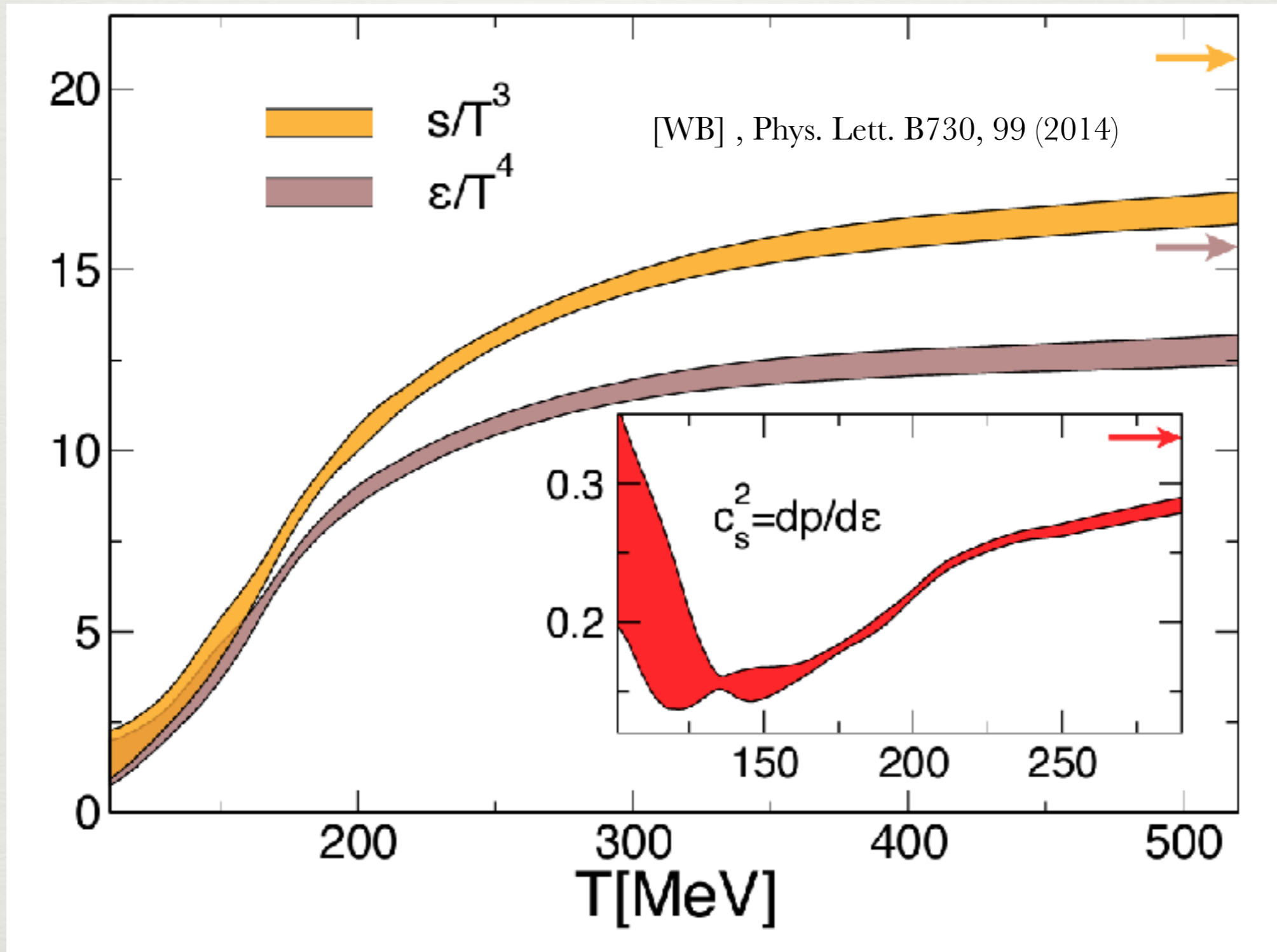
$$\text{Pressure: } \frac{p}{T^4} = \int_0^T dT \frac{1}{T} \left(\frac{\varepsilon - 3p}{T^4} \right)$$

$$\text{Energy density: } \frac{\varepsilon}{T^4} = \left(\frac{\varepsilon - 3p}{T^4} \right) + 3 \frac{p}{T^4}$$

$$\text{Entropy density: } \frac{s}{T^3} = \left(\frac{\varepsilon - 3p}{T^4} \right) + 4 \frac{p}{T^4}$$

$$\text{Speed of sound: } c_s^2 = \frac{dp}{d\varepsilon} = \frac{s}{T} \frac{dT}{ds}$$

Speed of sound

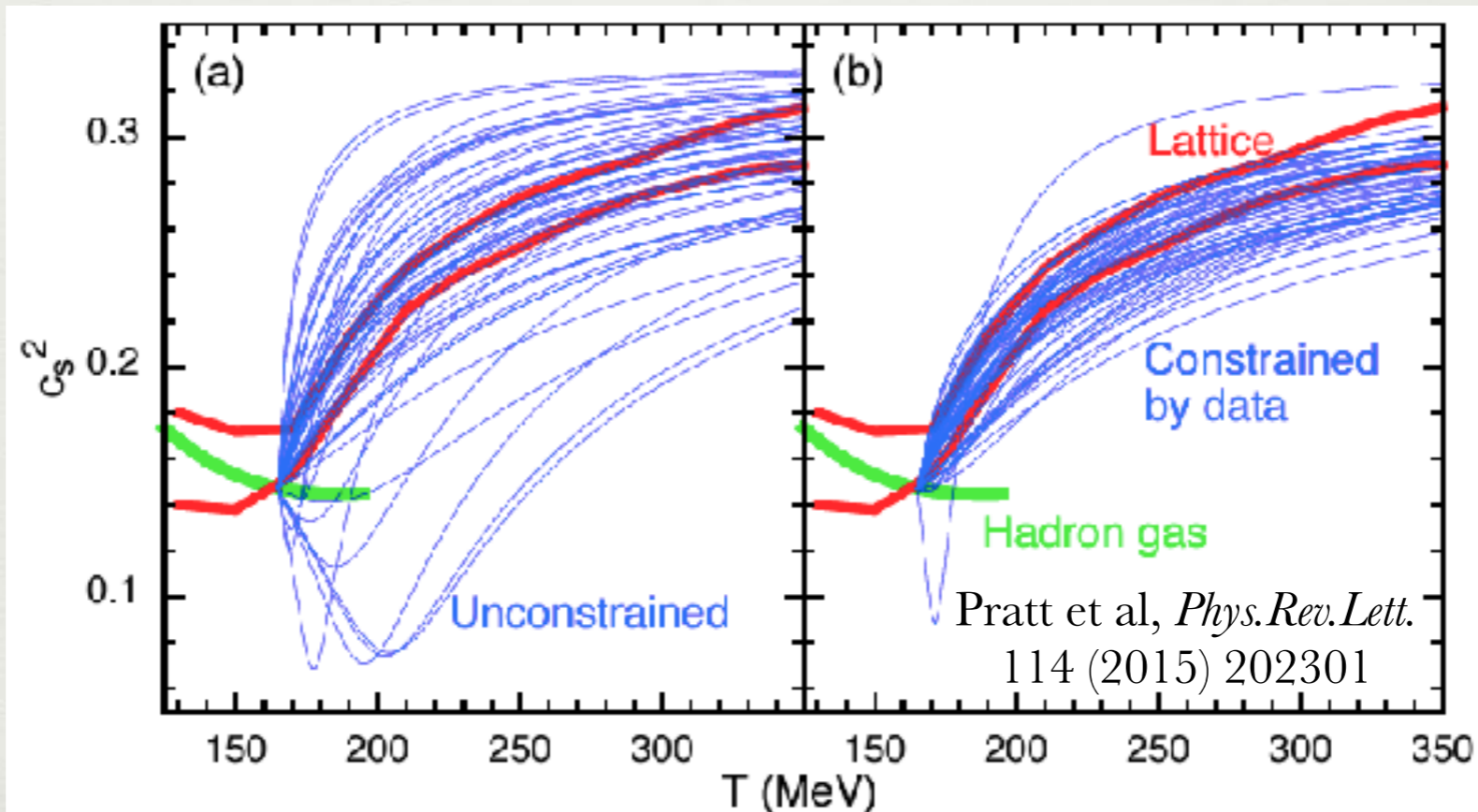


Equation of State

Equation of state: $p(\epsilon) +$ derivatives

Unfortunately, we cannot directly probe $p(\epsilon)$ experimentally

However, use $p(\epsilon)$ as input for relativistic viscous hydrodynamic simulations, compare to heavy-ion data (tomorrow's lecture)



Low T

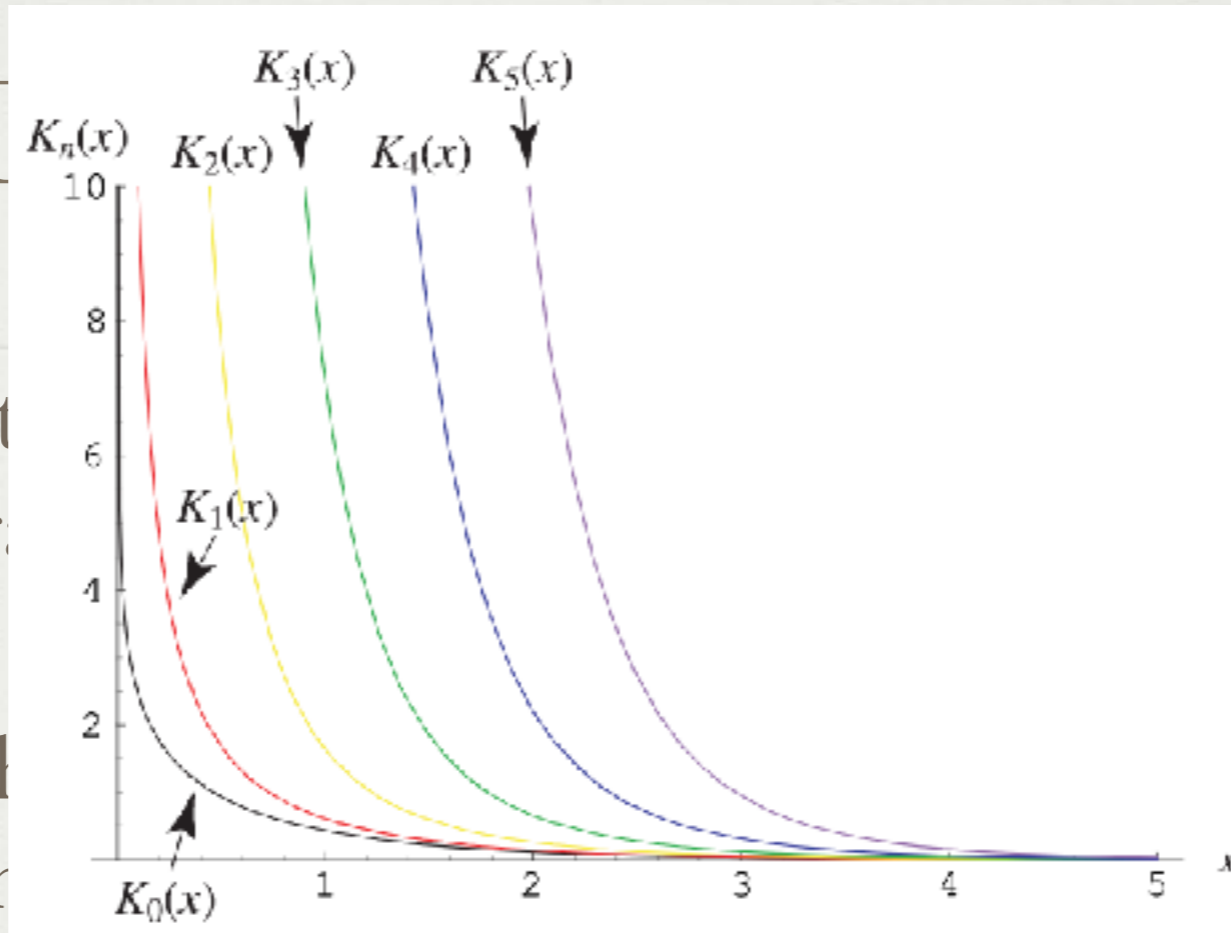
gas

At “low” temperature
a gas

near matter to be

We call this the
many, r

the ideal gas of
nces



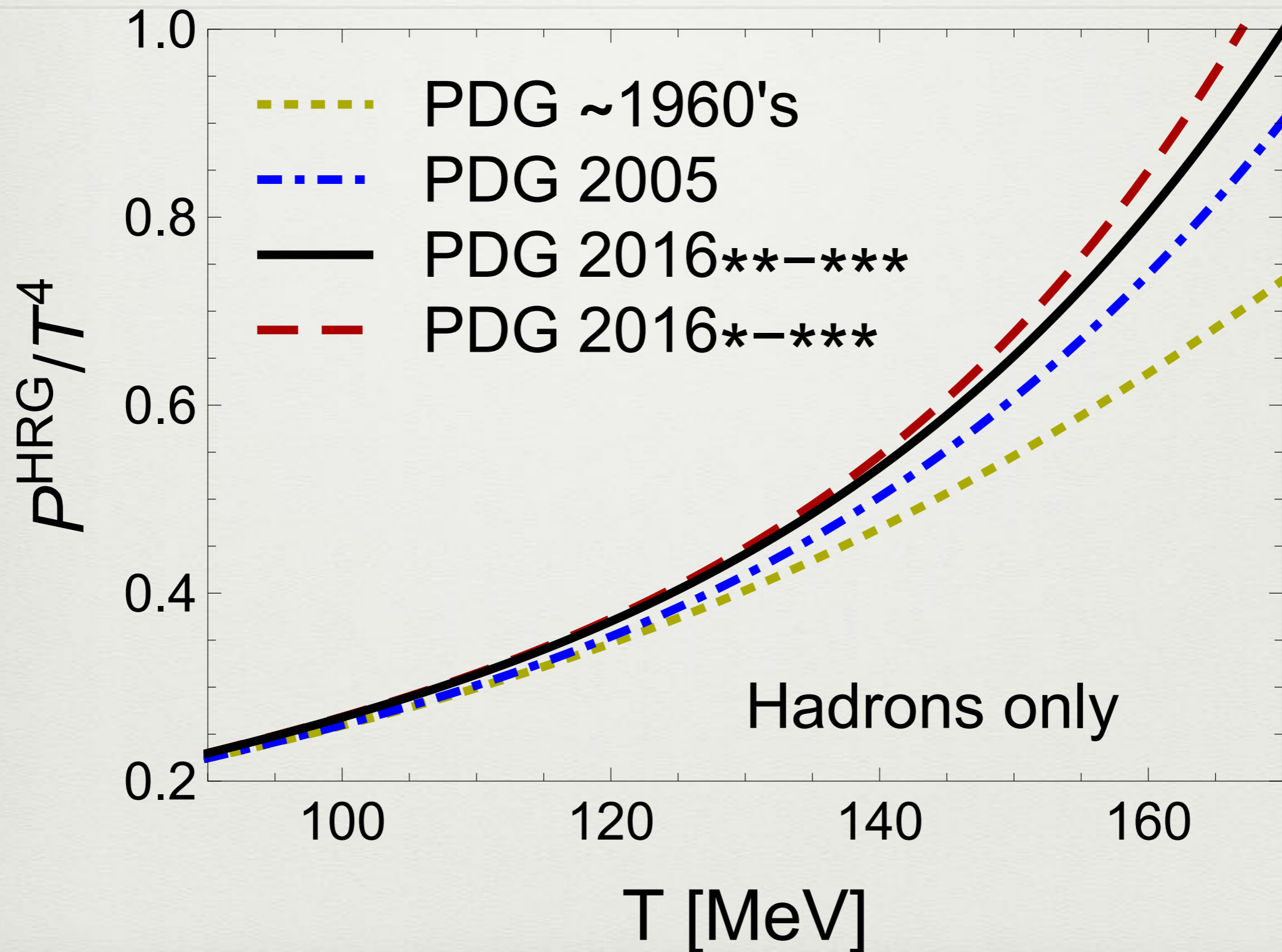
$$\ln Z_i(T, \vec{\mu}) \sim \frac{d_i}{2\pi} \left(\frac{m_i}{T} \right)^2 \sum_{k=1}^{\infty} \frac{(-1)^{(|B_i|-1)(k+1)}}{k^2} K_2 \left(\frac{km_i}{T} \right) \cosh [k \vec{\mu}_i / T]$$

$$\text{Where } \vec{\mu}_i = B_i \mu_B + S_i \mu_S + Q_i \mu_Q$$

In principle, one should describe their interactions. Instead, using all the resonance states does a reasonable job.

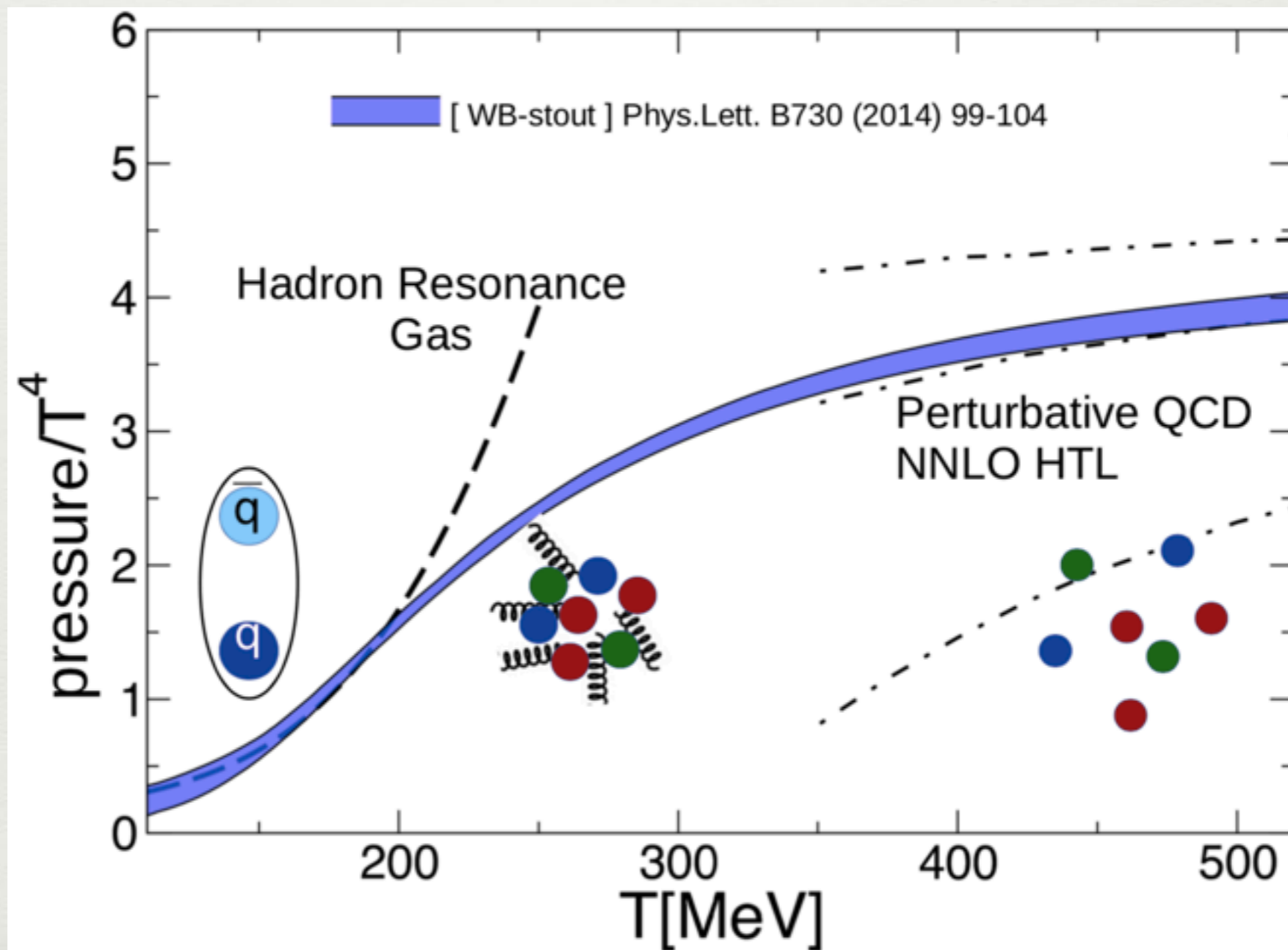
BUT at large μ_B , need interactions!

Influence of heavy resonances

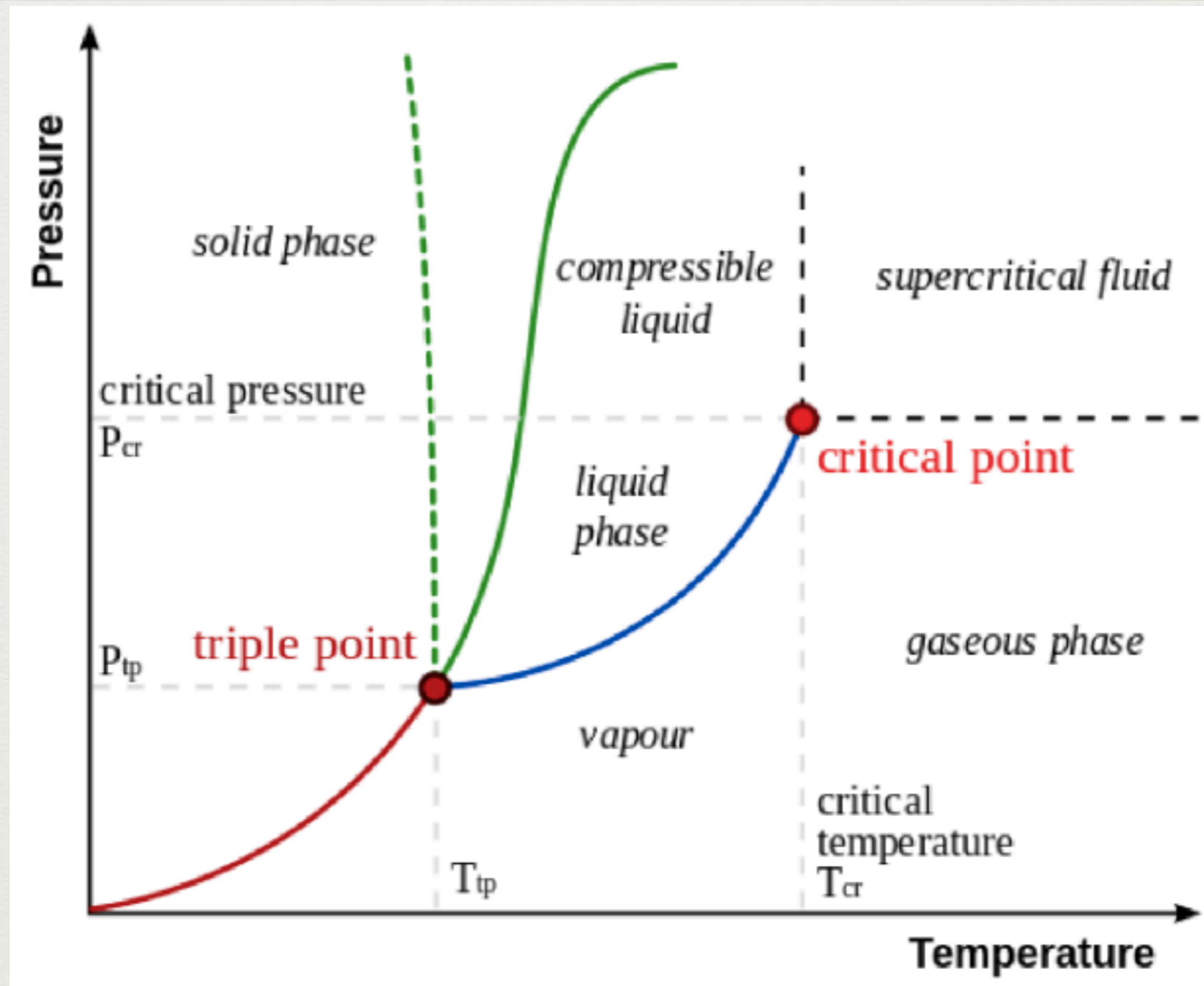


Lattice QCD: Phase Transition

Cross-over phase transition $T \sim 155 \text{ MeV}$



Water phase diagram



Water at atmospheric pressure has 2 first-order phase transitions

Finding a phase transition: order parameter

Order parameter - thermodynamic quantity that goes from 0 to 1 for different phases

Example: compressibility $K_T = -\frac{1}{V} \left(\frac{\partial V}{\partial p} \right)_T$

Solid or liquid - can't compress further with added pressure $K_T \sim 0$

Gas - inversely scales with pressure $K_T = p^{-1} \neq 0$

What is the order parameter in QCD?

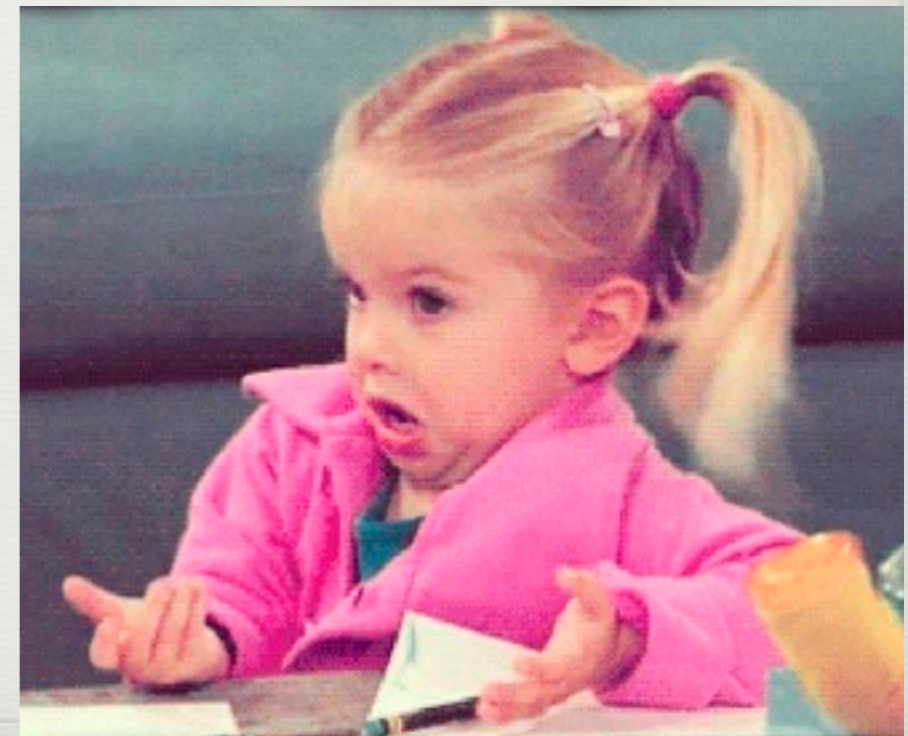
Baryon density n_B

$$\text{Chiral condensate } \langle \psi \bar{\psi} \rangle = \frac{T}{V} \frac{\partial \ln Z}{\partial m_0}$$

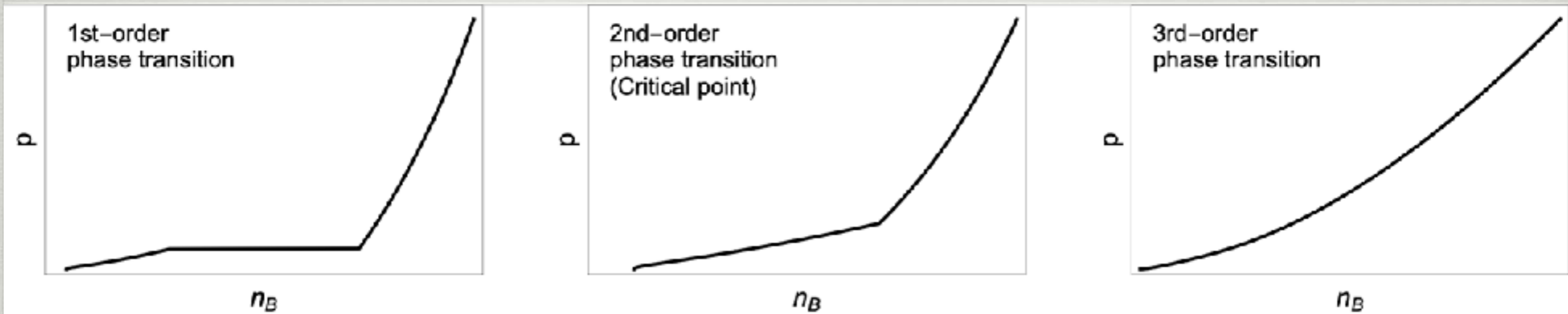
Proxy for deconfinement

Bazavov A et al. 2012 Phys. Rev. D85 054503

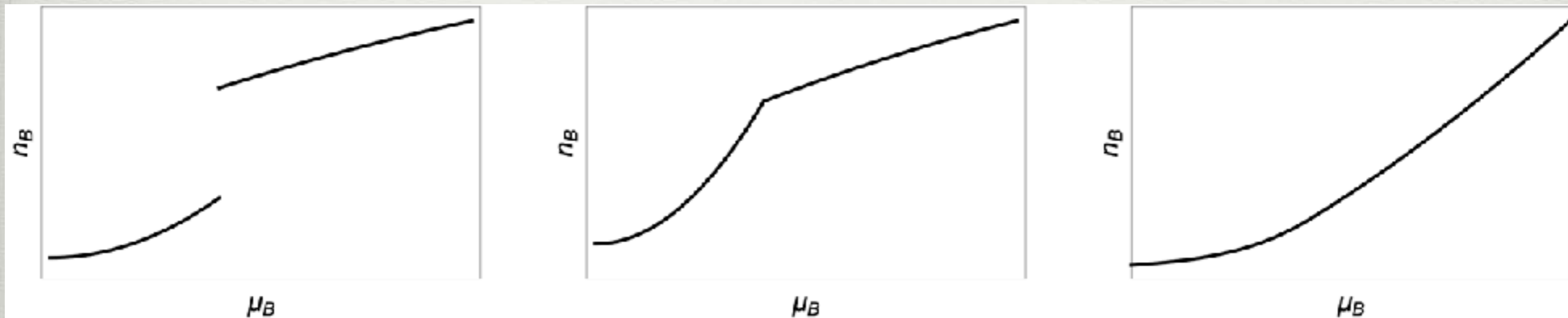
Other charge densities:
strangeness n_s or electric n_Q



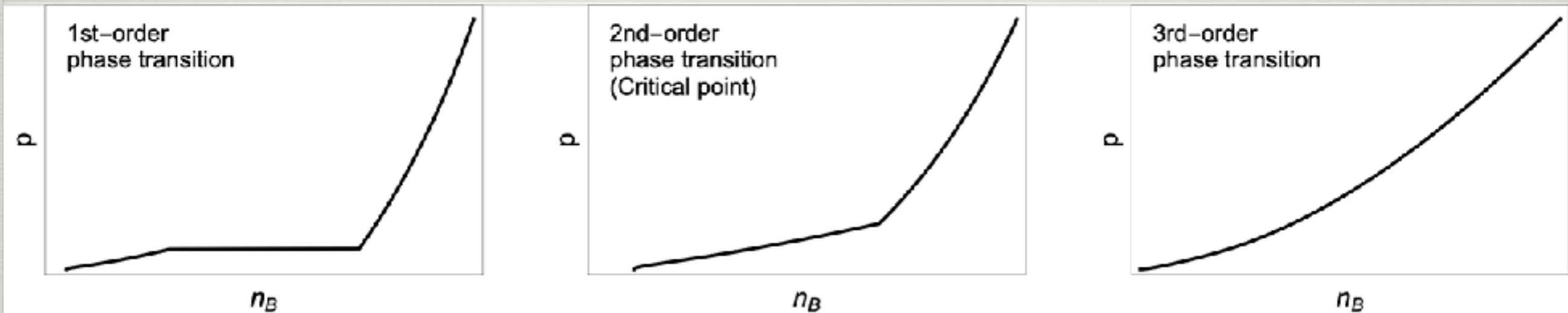
Order of the phase transition



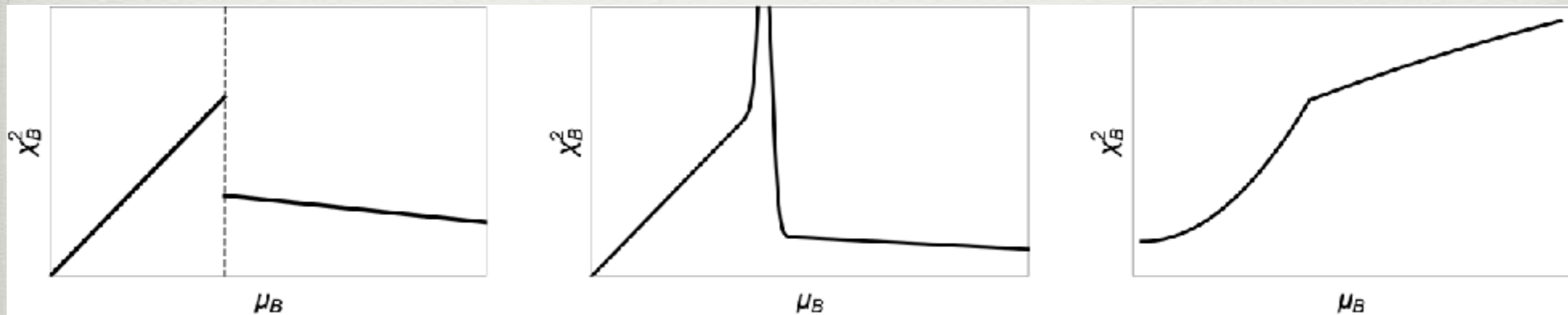
$$\text{Baryon density } n_B = \frac{\partial p}{\partial \mu_B}$$



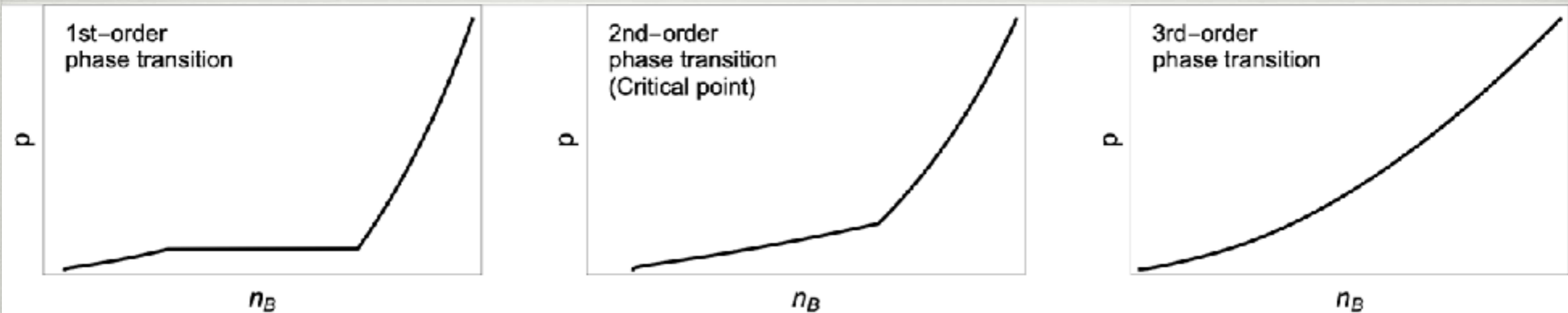
Order of the phase transition



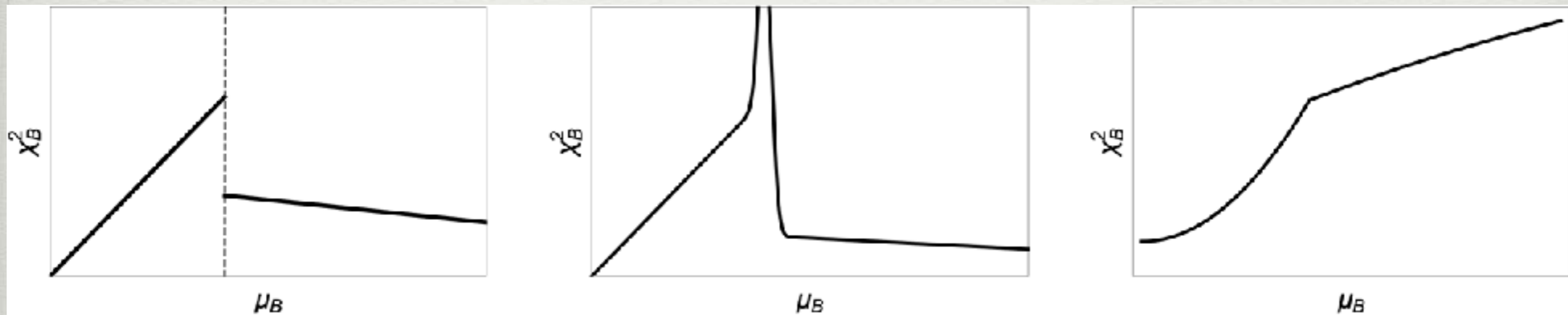
$$\text{Susceptibilities } \chi_B^n = \frac{\partial^n (p/T^4)}{\partial (\mu_B/T)^n}$$



Order of the phase transition



Susceptibilities $c_s^2 = \left(\frac{dp}{d\varepsilon} \right)_{s/n_B}$



What is the order of the QCD deconfinement transition?

Let's review its history to understand...

1981 Pure Glue SU(2)

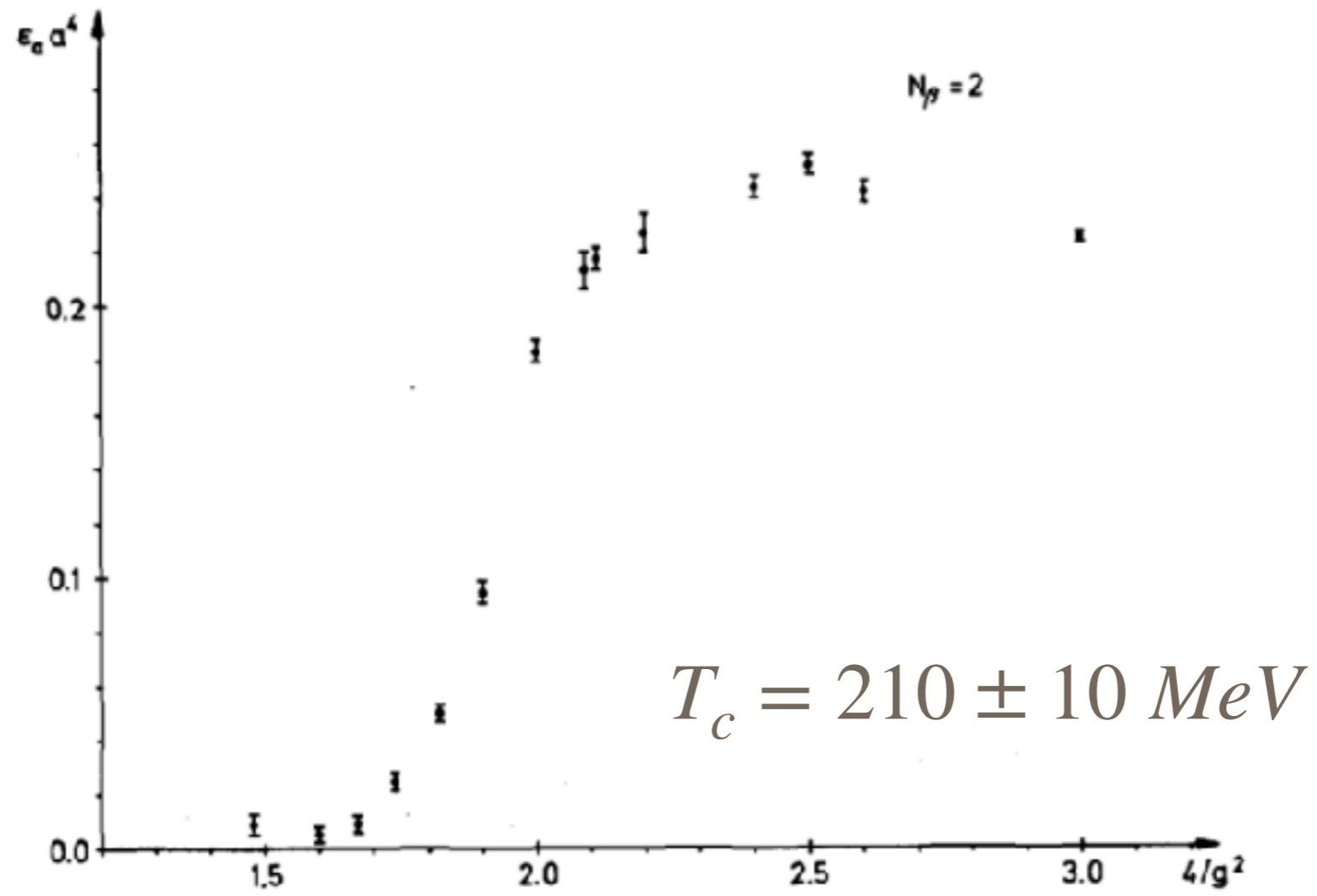


Fig. 3. Energy density of gluon matter versus $4/g^2$, at fixed lattice size $N_\beta = 2$, after about 500 iterations.

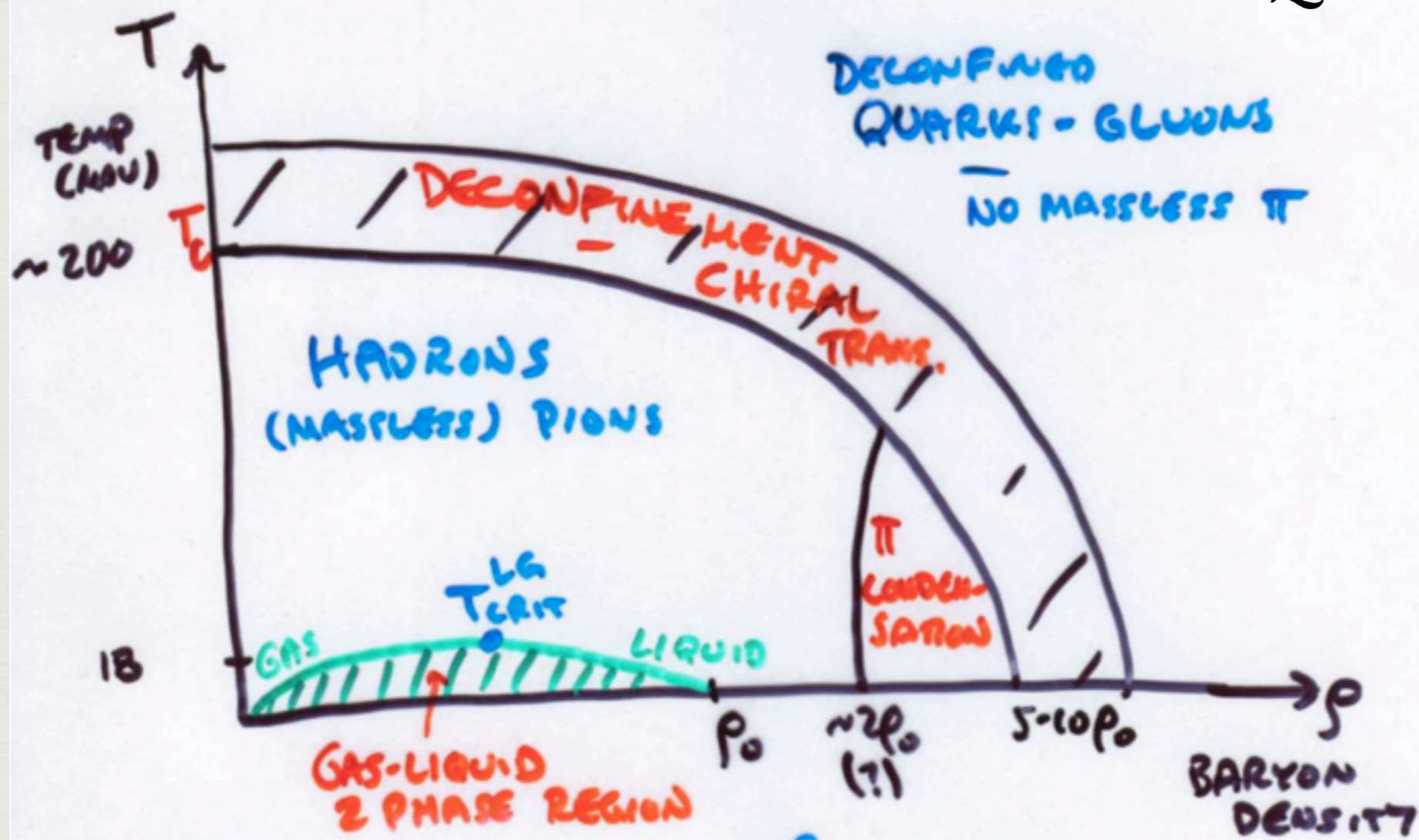
Phys.Lett. B101 (1981) 89

2nd order phase transition (Polyakov Loop McLerran and Svetitsky Phys.Rev. D24 (1981) 450)

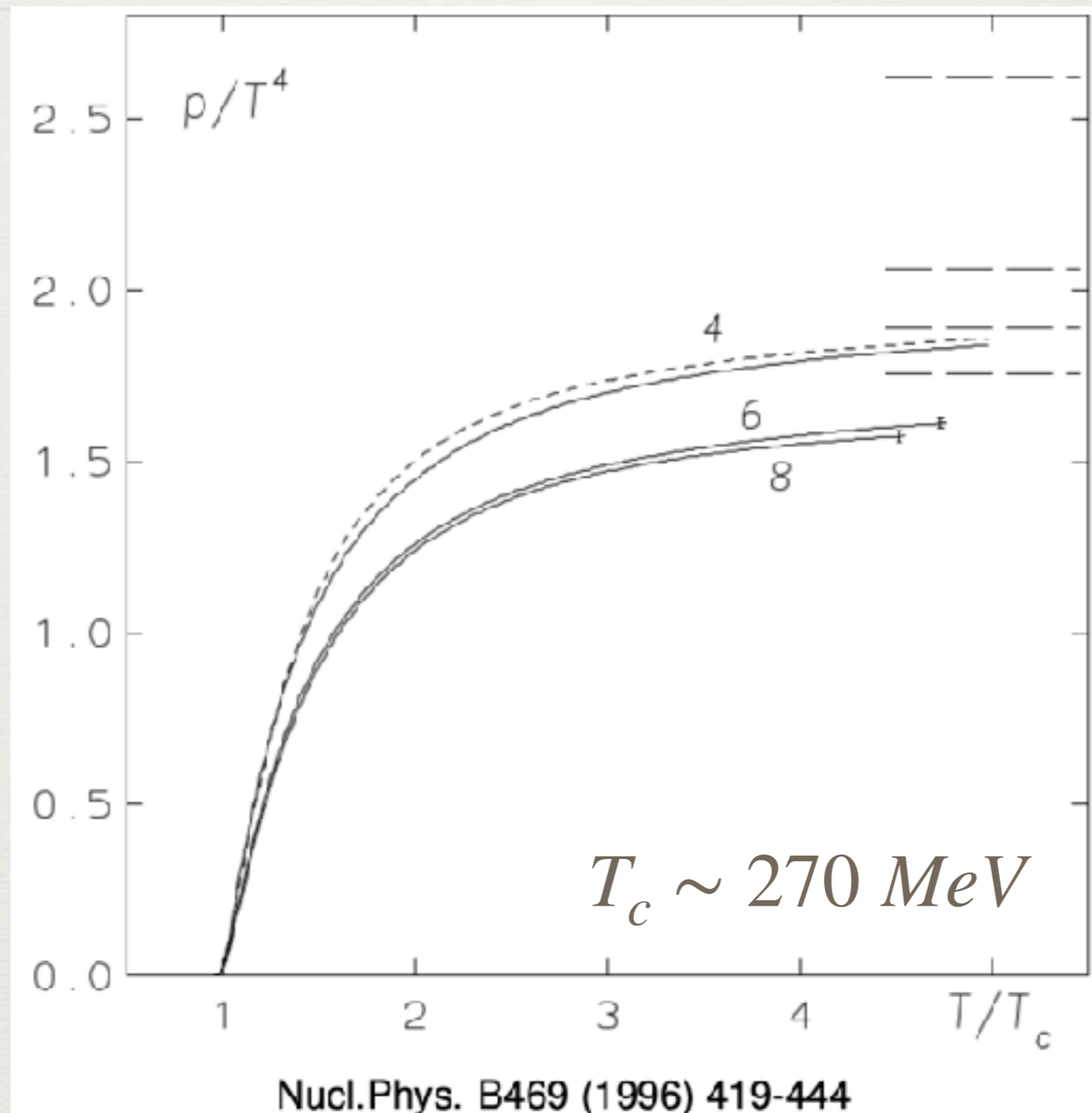
~ 1980's QCD phase diagram

EXPECTED PHASE DIAGRAM
⇒ MANY OPPORTUNITIES

Gordon Baym
Shown at QM15

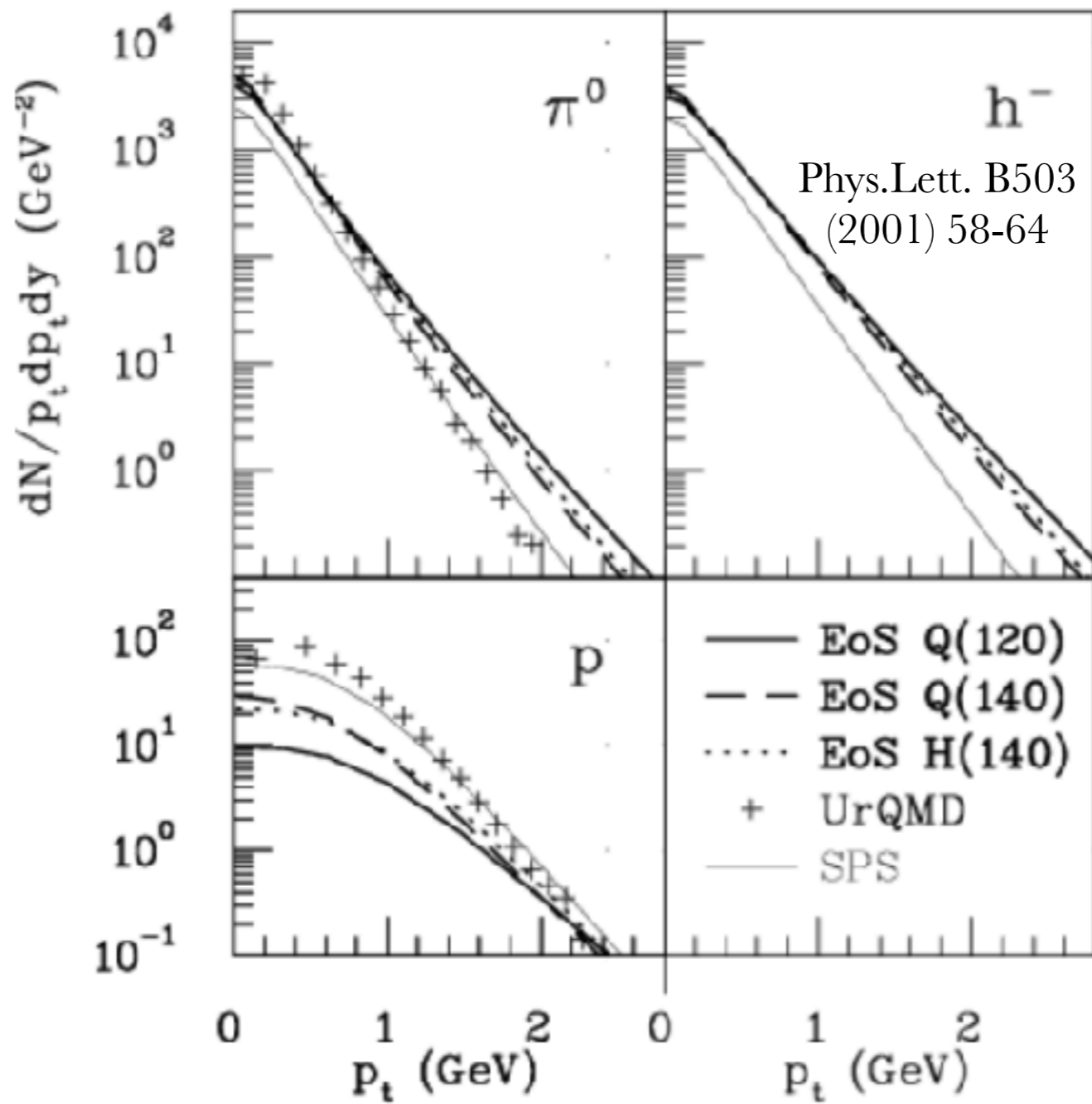


1996 Pure Glue SU(3)



1st order phase transition

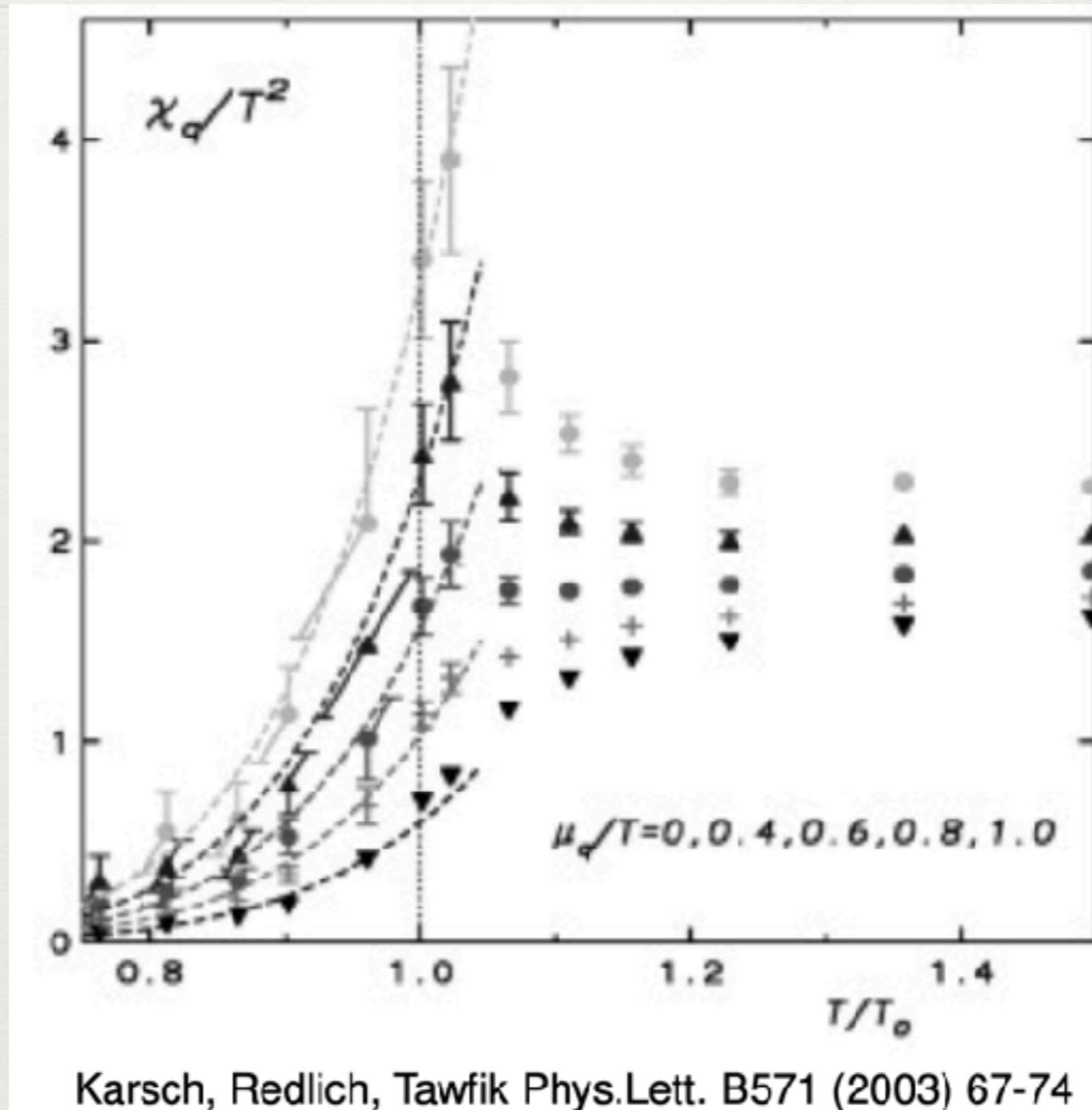
2001 Comparisons to heavy-ion data



See tomorrow's
talk

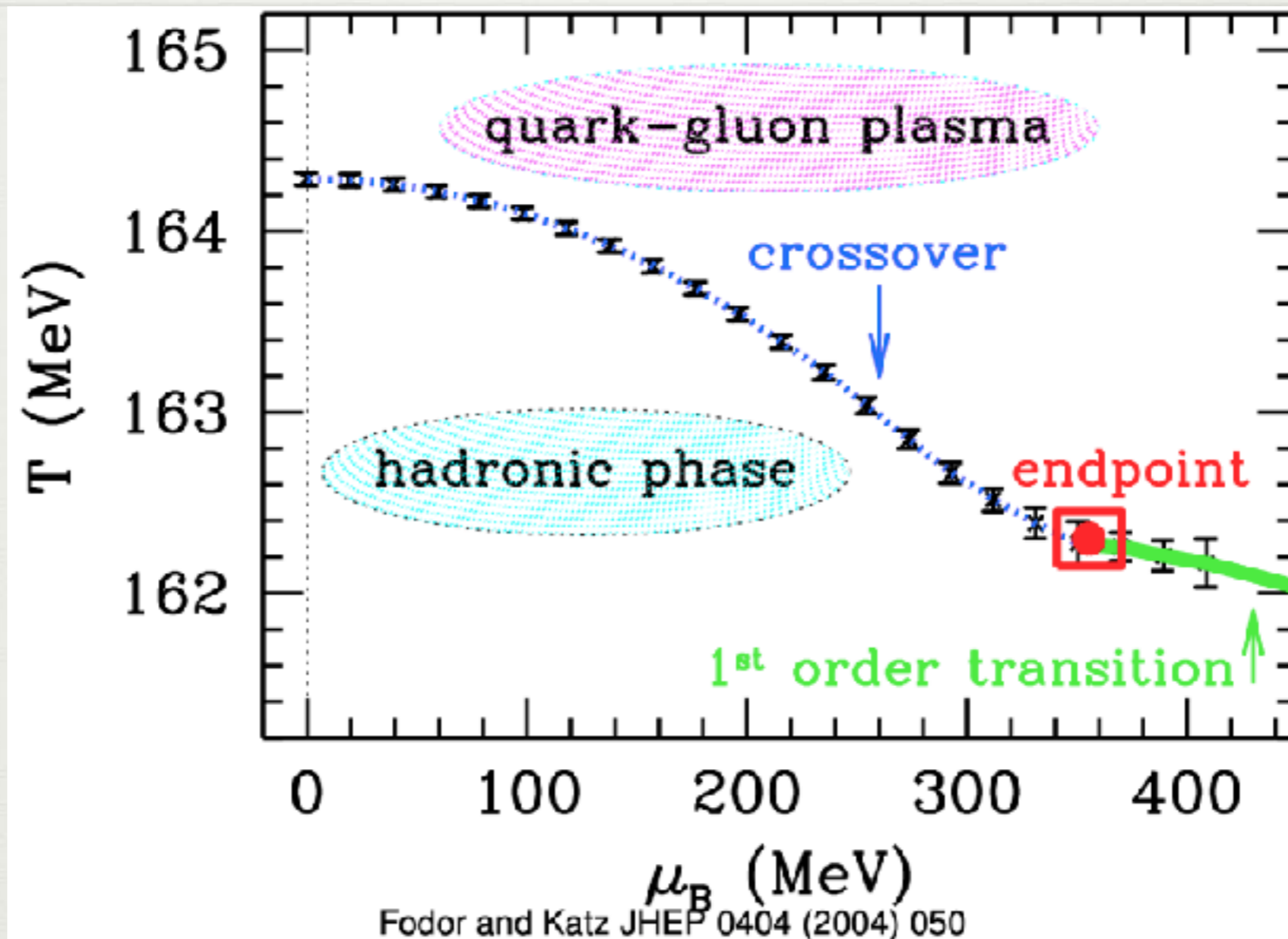
First-order phase transition equation of state compared to data

2003 Matching lattice QCD to hadron resonance gas



Hadron masses readjusted to smoothly match lattice QCD results

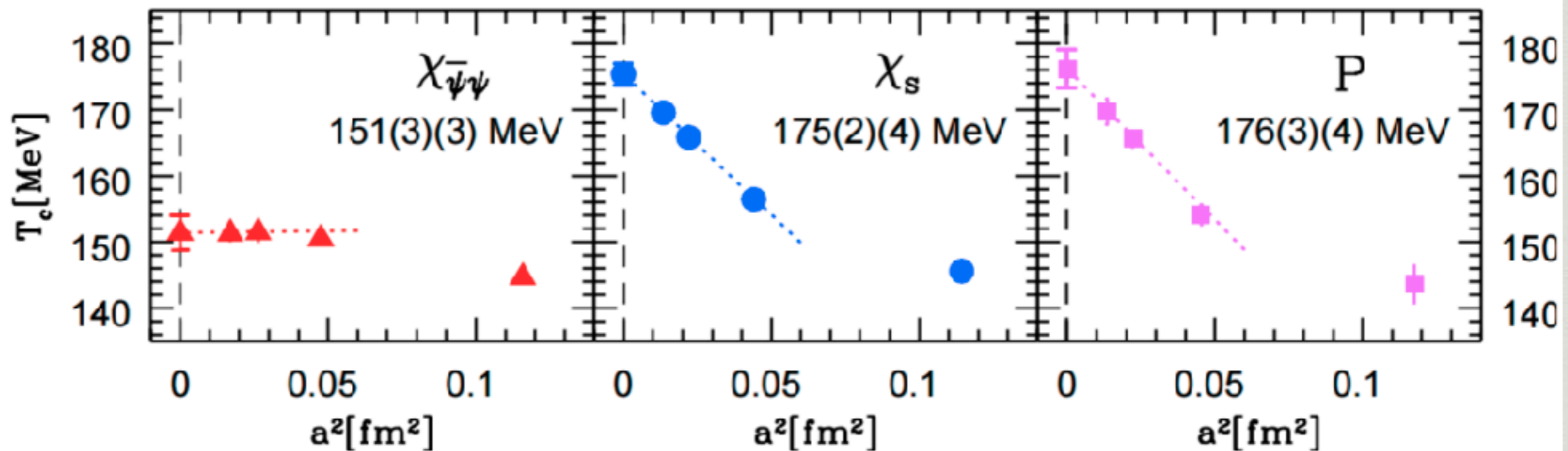
2004 Cross-over at $\mu_B = 0$ and critical point



Reweighting scheme with a coarse lattice

Modern results: *[WB] Phys.Rev.D 105 (2022) 5, L051506*

2006 Cross-over Phase Transition



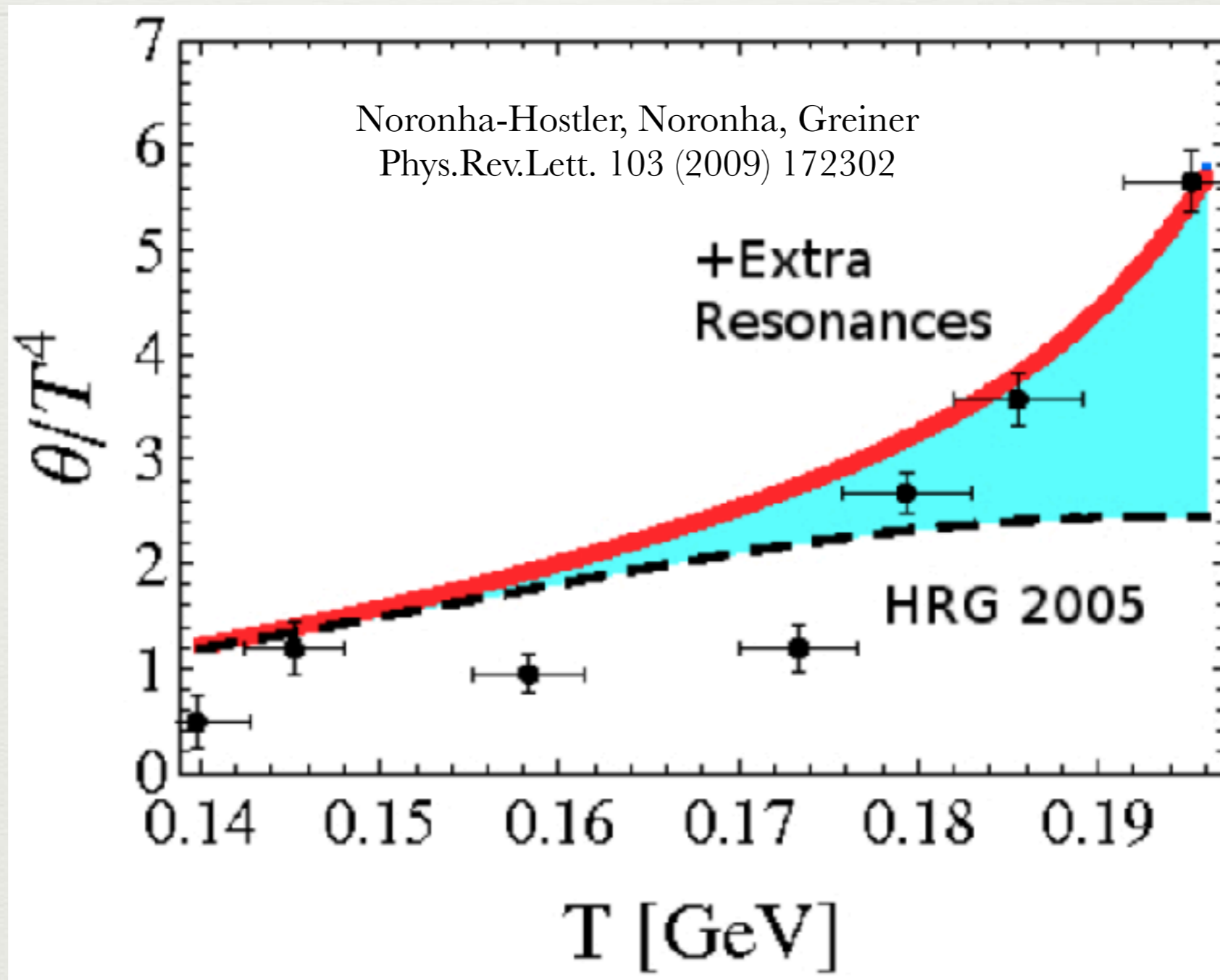
[WB] Phys.Lett. B643 (2006) 46-54; Nature 443 (2006) 675-678

Figure 4: Continuum limit of the transition temperatures obtained from the renormalized chiral susceptibility ($m^2 \Delta \chi_{\bar{\psi}\psi} / T^4$), strange quark number susceptibility (χ_s / T^2) and renormalized Polyakov-loop (P_R).

Checked 3 different “order parameters”, all lead to different T_c

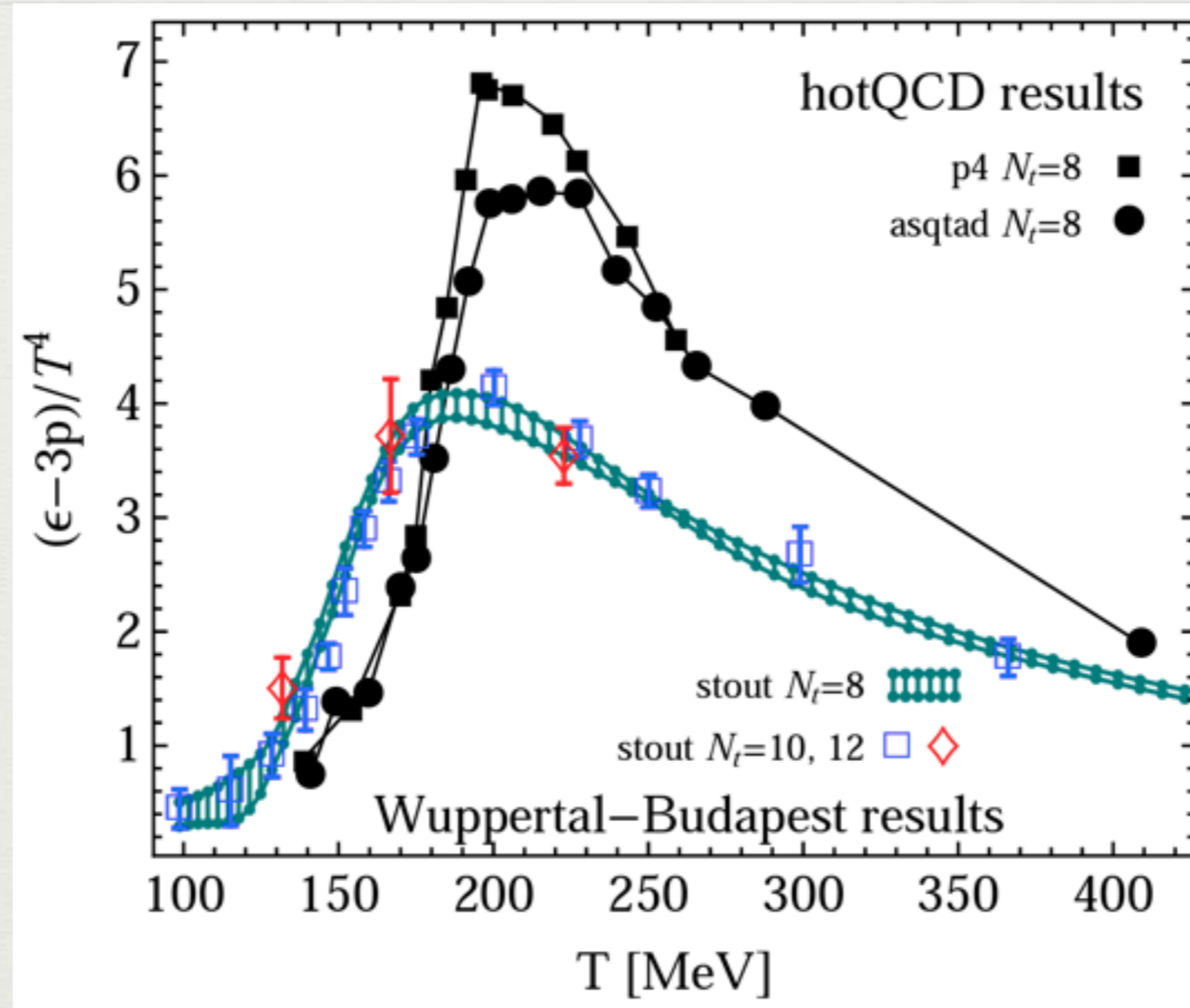
Smooth cross-over phase transition from hadrons to quarks & gluons!

2009 Extra Resonance needed for Cross-over



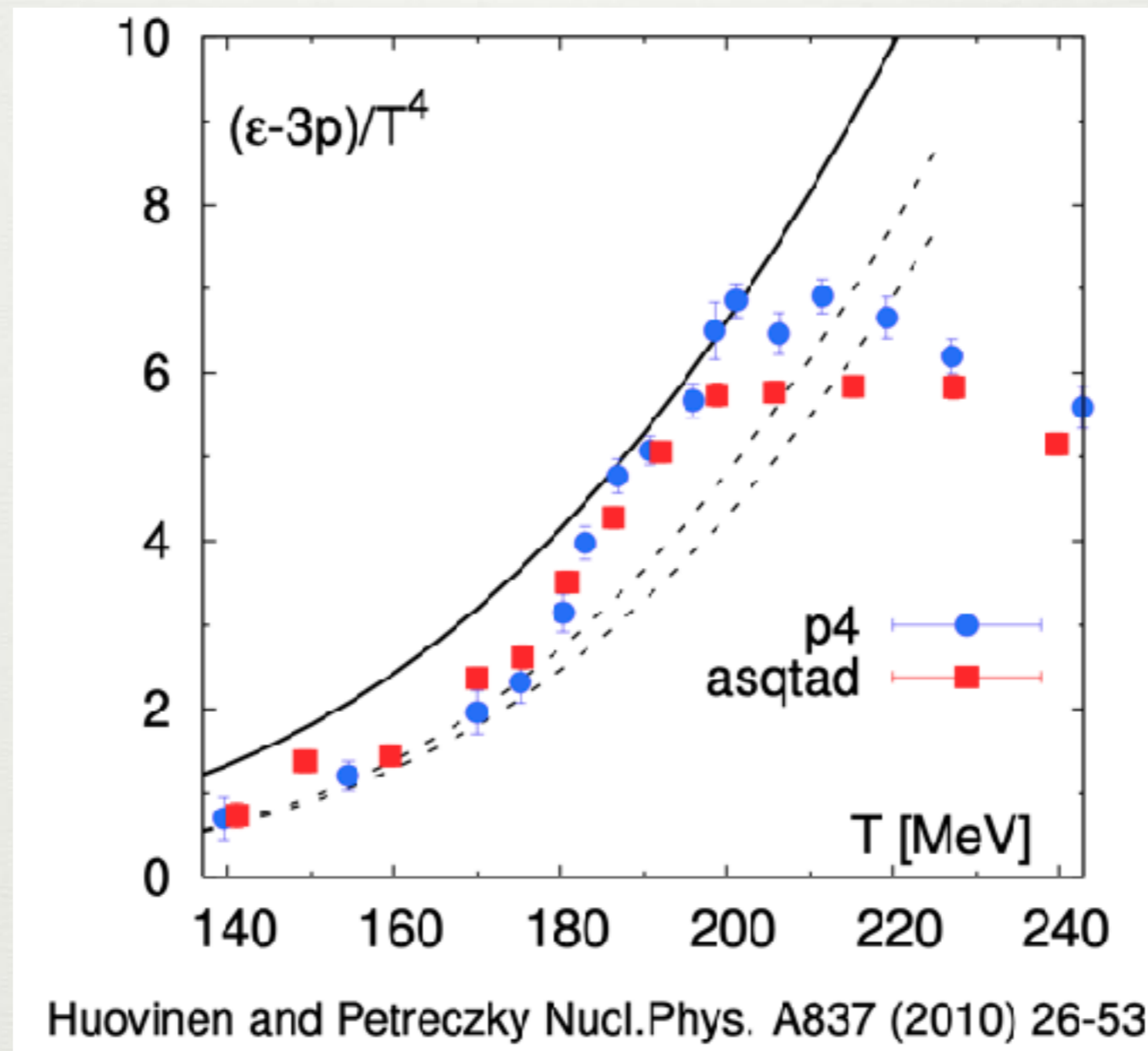
Extend mass spectrum with e^{m/T_H}

~2010 Lattice QCD did not agree on T_c



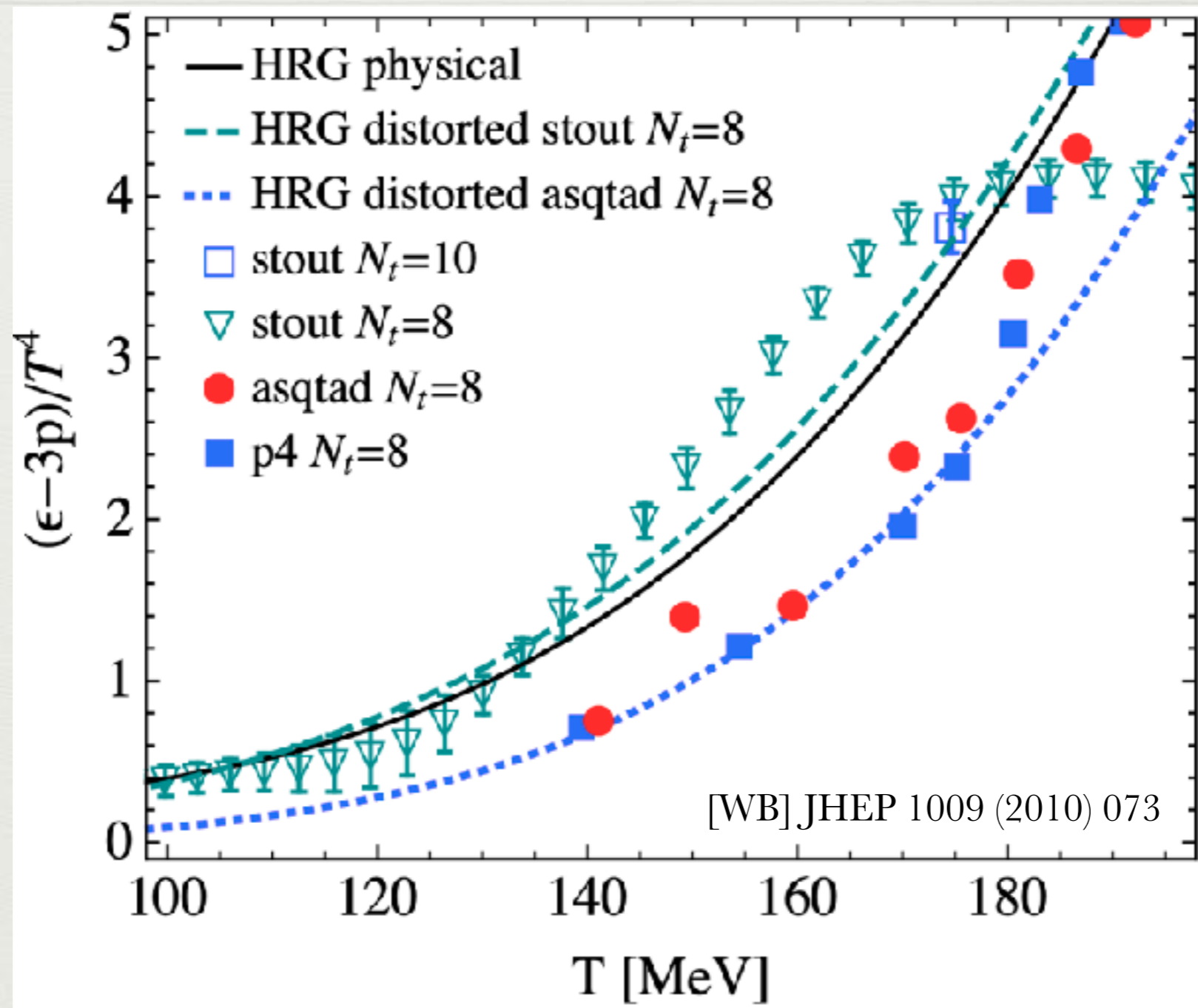
hotQCD $T_c \sim 196$ MeV
WB $T_c \sim 176$ MeV

~2010 rescaling hadron masses



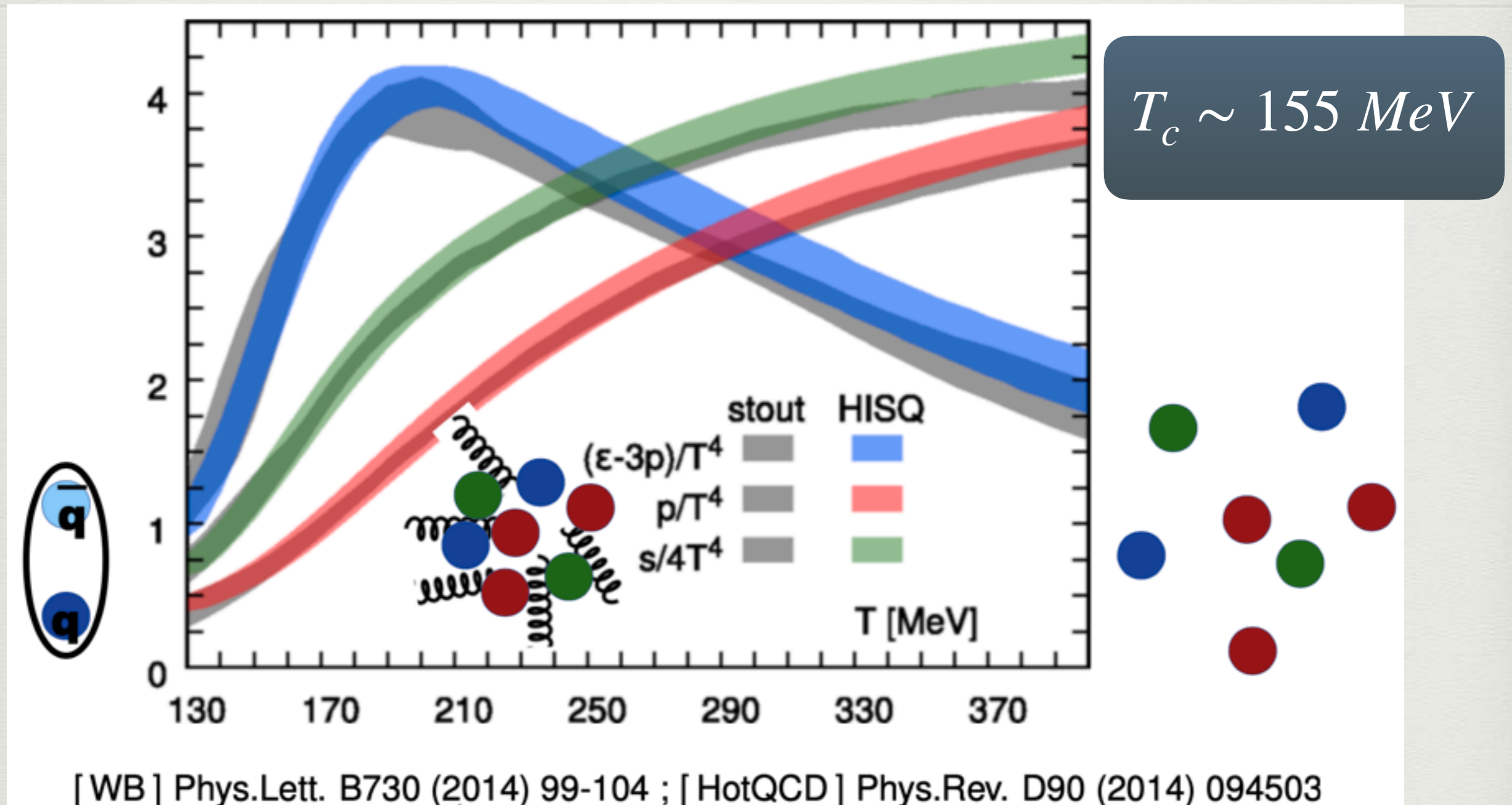
Before reaching the continuum limit (physical m_π), was difficult for HRG to match lattice. Not all hadrons included

2010 Hadron Resonance Gas almost matches Lattice QCD



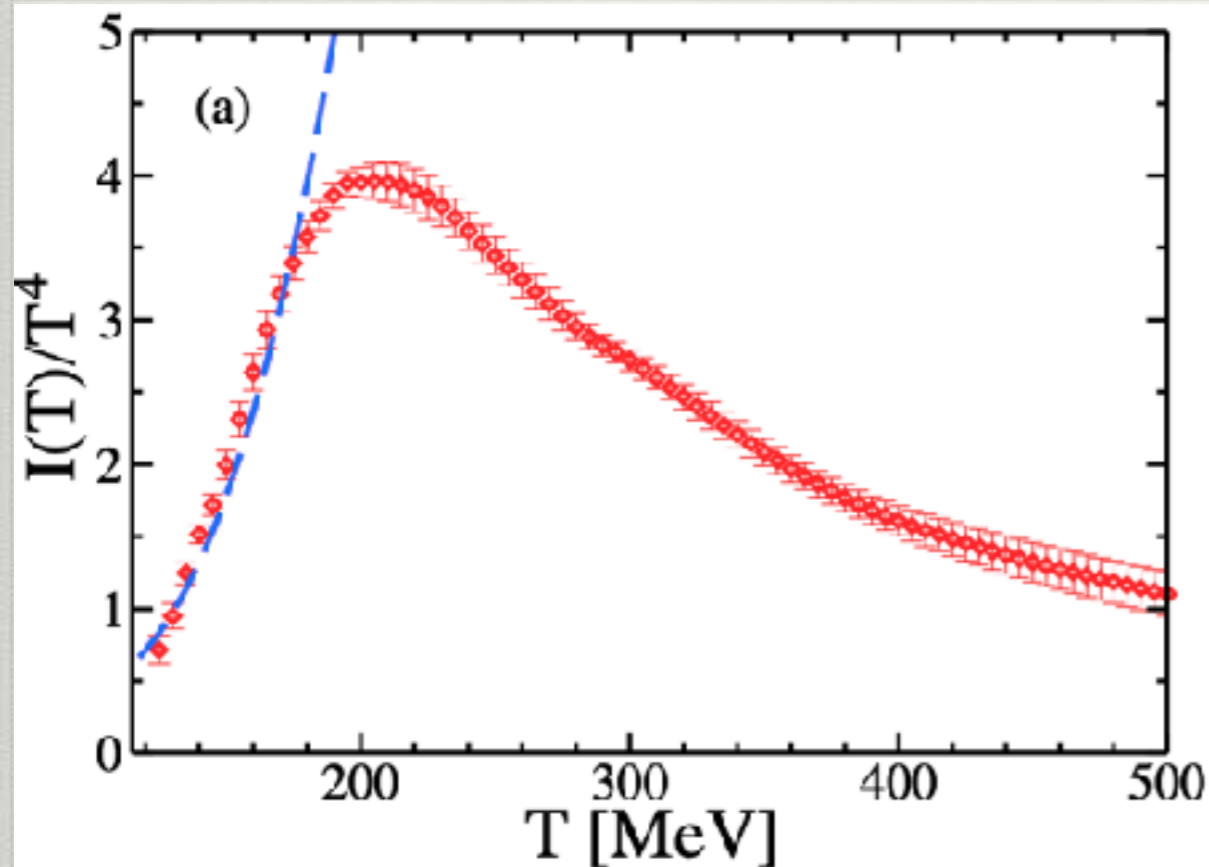
Continuum limit significantly improved match with HRG

2014 Finally, the actions agree

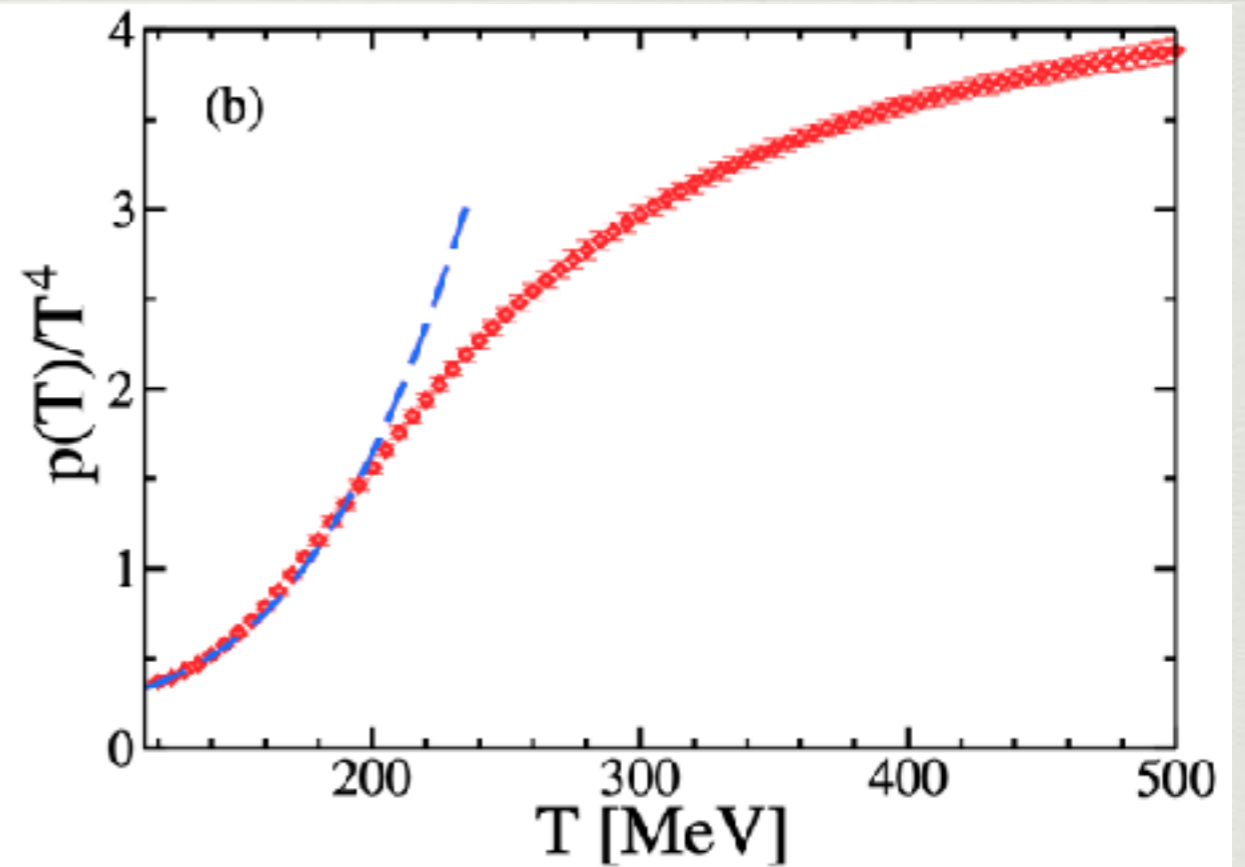


Agreement between different collaborations/actions (WB-stout and HotQCD-HISQ)

2014 Hadron resonance gas and lattice QCD finally fit?



Nucl.Phys. A929 (2014) 157-168



CAUTION

**PROCEED WITH
CAUTION**

Once the continuum limit was reached, wealth of new observables were possible.

Does the HRG actually fit?

What's next?

- Tomorrow: Connecting the equation of state directly to data (out-of-equilibrium effects)
- Weds: Today we focused on $\mu_B = 0$ results, what happens at large baryon densities?
- What can't we get from lattice QCD?
- Tuesday & Wednesday: Open Questions

Summary

Lattice QCD at finite T , $\mu_B = 0$

- We now know precisely the equation of state for the QCD deconfinement phase transition
- This equation of state matches smoothly to a hadron resonance gas with a large number of particles
- The QCD phase transition is a smooth cross-over
- Light and strange particles appear to hadronize/freeze-out at different temperatures