

Illinois Center for Advanced Studies of the Universe





Lecture 2 on Hot QCD Matter: Heavy-Ion Collisions

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Deconfined Quarks and Gluons in the Early Universe

~ 10^{-6} s after the Big Bang \rightarrow Quark-Gluon Plasma (1975 Collins and Perry)



How far back in time can we see?

Cosmic Microwave Background ~ 10^5 years after Big Bang

Quark-Gluon Plasma existed ~ 10^{-6} seconds

Little Bangs in the Lab

The Large Hadron Collider and RHIC create "little bangs": deconfined quarks and gluons in the lab



T=4 trillion degrees Celsius



10⁸ times hotter than the sun!



Evolution of a heavy-ion collision

Smashing two gold ions at the speed of light

Nuclear collisions and the QGP expansion



Big Bang vs. Heavy-Ion Collisions

| | Quark Epoch | Little Bang |
|--------------|-------------------------------------|--|
| Similarities | Pressure, entropy, energy density | |
| | Quark/gluons vs. Hadrons | |
| | Temperature when hadrons are formed | |
| | Strong Force | |
| | Nearly perfect fluid | |
| Differences | System Size | |
| | Entire Universe | 10^-14 m |
| | | Finite volume effects |
| | Equilibrated | Out-of-Equilibrium |
| | | Viscosity matters! |
| | | Expansion rate large |
| | 1 data point | Billions of events |
| | | Initial Conditions (many different shapes) |

Collisions of fluids vs. solids



The Quark Gluon Plasma is the

Hottest GUINNESS S O P D RECOR

Most Perfect

SCIENTIFIC **DNA** Computers AMERICAN MAY 2006 WWW SCIAM COM

Bringing

to Life

Duark Soup

SICISTS RE-CREATE THE LIQUID STUFF OF THE EARLIEST UNIVERSE

Stopping

Birth of the Amazon

Future **Giant Telesco** nature UNE 2017 VOL 13 NO 6

Strange

Stranger and stranger says ALICE

ELECTRON GASES Spin and charge part ways

QUANTUM SIMULATION Hamiltonian learning

TOPOLOGICAL PHOTONICS Optical Weyl points and Fermi arcs S U B A T O W C S W F R L S YOUTI FUL Second of the second SUMMER PARIS AGREEMENT SELECTION

namallest

vsics

Most Vortical

nature

of fluid vortices ormed by heavy

The geometry of a

quark-gluon plasma

Kinematics of Heavy-Ion collisions

How do heavy-ions collide?



Pb nucleus A = 208Center of mass energy *per nucleon* $\sqrt{s_{NN}} = 5.02$ TeV Collider mode: $s = (E + E) = 4E^2$ Radius along beam $E = \gamma m_N c^2$ $L = L_0/\gamma$

 $R_{beam} = \frac{2R_0m_N}{\sqrt{s}}$

How do heavy-ions collide?

Beam



Heavy-ions: Beam direction



Schenke IP-Glasma+MUSIC

Boost-invariance: 3+1 vs 2+1D

Heavy-ion collisions are **boost-invariant**- identical slices along longitudinal direction (z)

$$v_z = \frac{z}{t}$$

Thus, we talk about the proper time τ and space-time rapidity y

 $t = \tau \cosh y,$

 $z = \tau \sinh y,$ $\tanh y = \frac{z}{t}$

Pseudo-rapidity η

When a particle is traveling at the speed of light (massless) $y \sim \eta$

For massive particles, use **pseudo-rapidity** η



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Heavy-ions: Transverse plane



Transverse plan vs. longitudinal



Initial conditions



How 2 heavy-ions collide

Assuming Gold ions are spheres...



Cold Atoms



Standard Model of HIC



Initial Conditions

Wood Saxon

Sample nucleons over a Wood Saxon density distribution

$$\rho = \rho_0 \left[1 + \exp\left(\frac{r - R_0}{a}\right) \right]^{-1}$$

Radius of nucleus R_0 a- surface diffusion parameter ρ_0 - normal nuclear density

We also consider deformed nuclei where $R_0 \neq const$

Alternative: nucleon configurations from low-energy nuclear structure calculations

Wood-Saxon continued



Ann.Rev.Nucl.Part.Sci. 57 (2007) 205-243

The spherical cow has bumps: From optical Glauber to event-by-event fluctuations





Sub-nucleonic Fluctuations

Centrality circa 2000's



Centrality circa 2000's



Centrality circa 2000's



Eccentricity early 2000's

$$e_n = \frac{\{y^2 - x^2\}}{\{y^2 + x^2\}}$$

$$\{\dots\} \equiv \frac{\int d^3 x e(\vec{x}) \dots}{\int d^3 x e(\vec{x})}$$

Defined with impact parameter

Collision: Lorentz contracted



Collisions impact parameter *b*, and reaction plane angle ψ_R

Participant Plane vs. Reaction Plane



Event plane angle ψ_{EP}

Now Eccentricity had a magnitude AND angle!

Monte Carlo sampling of ¹⁶O



 d_{min} Minimum distance apart

Thickness function

 $\hat{T}_A(\mathbf{s}) = \int \hat{\rho}_A(\mathbf{s}, z_A) dz_A$, where $\hat{\rho}_A(\mathbf{s}, z_A)$ is the probability per unit volume,

Joint probability per unit area of nucleons being located in the respective overlapping target and projectile flux tubes

Reduced thickness function

$$\hat{T}_{AB}(\vec{s},\vec{b}) \equiv f\left[\hat{T}_A(\vec{s}),\hat{T}_B(\vec{s}-\vec{b})\right]$$

Probability of a collision $p(\vec{s}, \vec{b}) = \sigma_{NN}^{inel} \hat{T}_{AB}(\vec{s}, \vec{b})$

$$\left(\frac{dS}{dy}\right)_{\tau=\tau_0} \propto f\left(T_A, T_B\right)$$

Monte Carlo sampling of collisions



Keep the participating nucleons that collide in the overlap region

Collisions and Participants

$$N_{\text{coll}}\left(b\right) = \sum_{n=1}^{AB} nP\left(n,b\right) = AB\hat{T}_{AB}\left(b\right)\sigma_{\text{inel}}^{\text{NN}}$$

$$egin{aligned} N_{ ext{part}}\left(\mathbf{b}
ight) &= & A \int \hat{T}_A\left(\mathbf{s}
ight) \left\{1 - \left[1 - \hat{T}_B\left(\mathbf{s} - \mathbf{b}
ight) \sigma_{ ext{inel}}^{ ext{NN}}
ight]^B
ight\} d^2s + \ & B \int \hat{T}_B\left(\mathbf{s} - \mathbf{b}
ight) \left\{1 - \left[1 - \hat{T}_A\left(\mathbf{s}
ight) \sigma_{ ext{inel}}^{ ext{NN}}
ight]^A
ight\} d^2s, \end{aligned}$$



Connecting to initial conditions

- Not 100% clear how to go from thickness functions to initial energy density
- Commonly used in the field: $\varepsilon \propto T_A + T_B$ (Glauber), $\varepsilon \propto T_A T_B$ (CGC), or $s \propto \sqrt{T_A T_B}$ (Phenomenological)
- Scaling constant needed i.e. $\varepsilon = CT_A T_B$ where C is fitted to measured multiplicity. Unsolved mystery how to find C from first principles.

Initial state "phase diagram"



Gluon Saturation of the nucleus



Different shape/fluctuations

"Event-by-Event" Holding the number of partons (density) constant for the same types of collisions, different shapes can be formed.



Triangles, squares etc can even appear...

What about other shapes?

 $\varepsilon_{n,m}e^{i\Phi_{n,m}}\equiv-\frac{\{r^me^{in\phi}\}}{\{r^m\}}.$

Fourier series....



Quantify eccentricities in different models

Eccentricities:

$$\varepsilon_n e^{i\Phi_n} = \frac{\int d^2 \mathbf{r} \ r^n \ e^{in\phi} \ \epsilon(\tau_0, \mathbf{r})}{\int d^2 \mathbf{r} \ r^n \ \epsilon(\tau_0, \mathbf{r})}$$

Variance in the Eccentricities

Different initial conditions with varying underlying physical assumptions



Initial conditions out-of-equilibrium



J. Noronha-Hostler

What are people studying today? Initial State

- Measuring deformations in nuclei, eccentric protons
- Full $T^{\mu\nu}$ and finite baryon density (baryon stopping)
- Signatures of the Color Glass Condensate
- Connections to the Electron Ion Collider
- Correct time evolution: free-stream, kinetic theory, AdS/ CFT, classical Yang Mills
- Start time of hydrodynamics?

Fluid Dynamics Dynamics of Quark Gluon Plasma

Non-relativistic Fluid Dynamics

Navier Stokes Equation:

 $\rho \frac{\partial v}{\partial t} = -\nabla p - \eta \nabla^2 v$

How quick the fluid changes its velocity



 $\eta \nabla^2 v$

 $\rho \frac{\partial v}{\partial t}$

 ∇p

Viscosity- how sticky a substance is

Ideal relativistic fluid dynamics

Conservation of Energy and Momentum $\partial_{\mu}T^{\mu\nu} = 0$ and $\partial_{\mu}N^{\mu} = 0$

The *ideal* energy-moment tensor is $T^{\mu\nu} = (\varepsilon + p)u^{\nu}u^{\nu} - pg^{\mu\nu} \text{ where } u^{\mu} = \gamma (1, \vec{v})$ We also write: $\Delta^{\mu\nu} = g^{\mu\nu} - u^{\mu}u^{\nu}$ $T^{\mu\nu} = \varepsilon u^{\nu}u^{\nu} - p\Delta^{\mu\nu}$

Coordinate System: $x^{\mu} = (\tau, x, y, \eta)$ where $\tau = \sqrt{t^2 - z^2}$ and $\eta = 0.5 \ln\left(\frac{t+z}{t-z}\right)$

Solving idea fluid dynamics

• Project the $T^{\mu\nu}$ in the parallel $u_{\nu}\partial_{\mu}T^{\mu\nu} = 0$ $(\varepsilon + p) \partial_{\mu}u^{\mu} + u^{\mu}\partial_{\mu}\varepsilon = 0$

and the perpendicular directions $\Delta^{\alpha}_{\nu}\partial_{\mu}T^{\mu\nu} = 0$ $(\varepsilon + p) u^{\mu}\partial_{\mu}u^{\alpha} - \Delta^{\mu\alpha}\partial_{\mu}p = 0$

- Solve in terms of ε , u^{μ} (hydro variables)
- Find pressure from equation of state $p(\varepsilon)$

Shear viscosity - η/s



Navier-Stokes equations are acausal and unstable relativistically

Perturb the system and the equations will break down



Requires a finite time for the system to return to equilibrium Relxation time τ_{π}

Relativistic Fluid Dynamics

Israel Stewart Annals Phys. 118 (1979) 341-372

Conservation of Energy and Momentum $\partial_{\mu}T^{\mu\nu} = 0$ and $\partial_{\mu}N^{\mu} = 0$ The energy-moment tensor contains a bulk dissipative term Π and the shear stress tensor $\pi^{\mu\nu}$ is Equation of State $T^{\mu\nu} = \underbrace{\epsilon}_{\mu\nu} \nu u^{\nu} - \underbrace{\rho}_{\mu\nu} + \Pi \underbrace{\Delta^{\mu\nu} + \pi^{\mu\nu}}_{\tau_{\pi}} \begin{pmatrix} Transport \\ Coefficients \end{pmatrix}$ $\underbrace{\tau_{\pi}}_{\pi} \begin{pmatrix} \Delta_{\mu\nu\alpha\beta}D\pi^{\alpha\beta} + \frac{4}{3}\pi_{\mu\nu}\theta \\ +\Pi, N^{\mu} \dots \end{pmatrix} = \underbrace{2\eta e_{\mu\nu}}_{\mu\nu} - \pi_{\mu\nu}$

Nearly perfect fluidity



Attempts from QCD





Inversion Problem: correlation functions from Lattice QCD in Euclidean time, must convert to Minkowski time.

For overview, see Moore 2010.15704

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Calculations in certain limits

Pure Yang Mills Theory

Perturbative QCD



Transport coefficients/viscosities

Transport coefficient: Perturb the fluid from equilibrium- how quickly does it return to equilibrium?



Viscosity - resistance to deformation or "thickness" of liquid

Experimental probes of $\eta/s(T)$



Theoretical calculations of viscosity



See references in JNH arXiv:1512.06315 **Dip expected:** Phys.Rev.Lett. 97 (2006) 152303, Nucl.Phys. A769 (2006) 71-94, Phys.Rev.Lett. 103 (2009) 172302

Bayesian analysis (agnostic $\eta/s \& \zeta/s$)

Shear viscosity

Bulk viscosity



Bernhard, Moreland, Bass Nature Phys. 15 (2019) no.11, 1113-1117