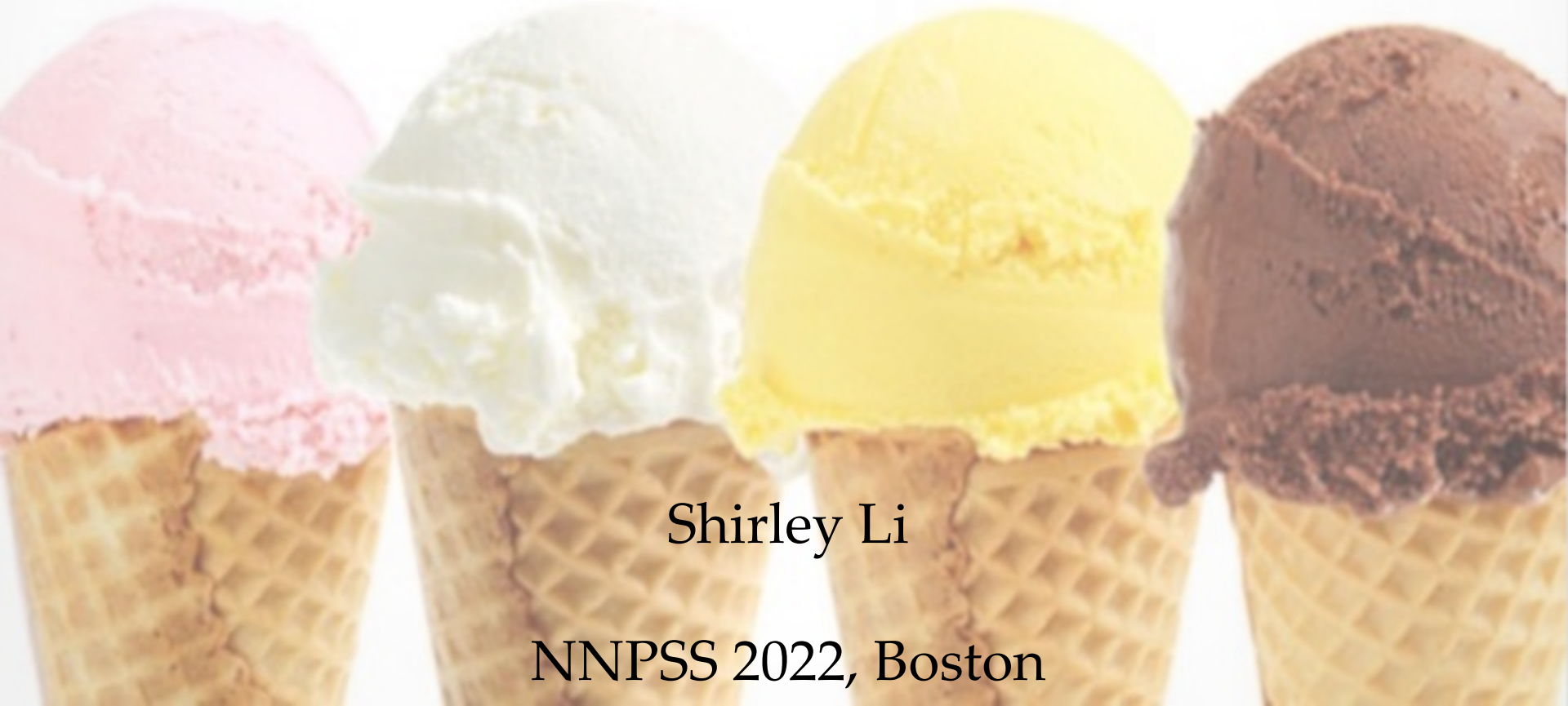


How do neutrinos oscillate?

What we know and do not know about neutrino oscillation



Shirley Li

NNPSS 2022, Boston

Outline

- Standard three-flavor neutrino oscillation
 - Two-level system
 - Matter effect
 - Three-flavor oscillation

- Sterile neutrinos

3-Flavor Neutrino Oscillations

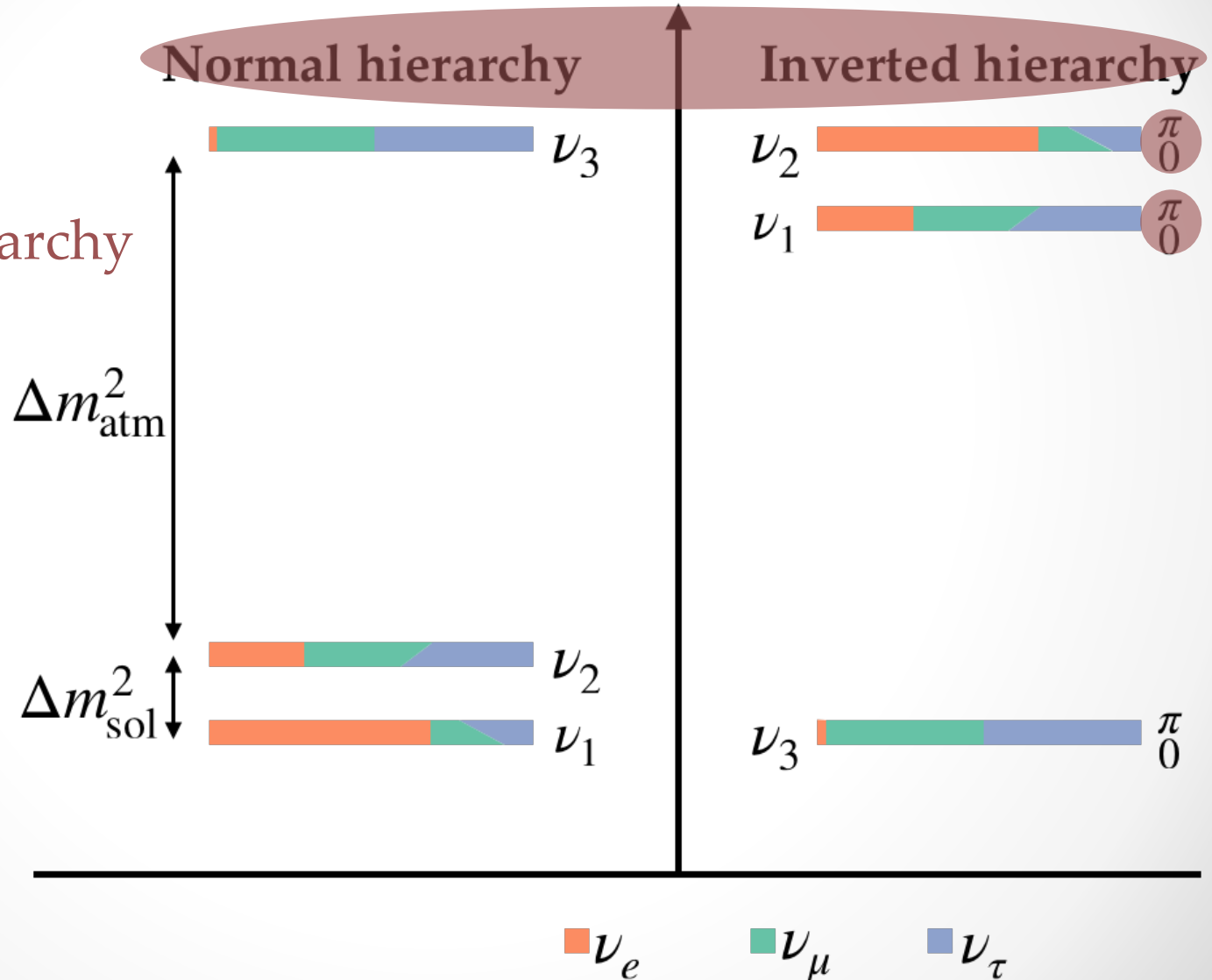
Unknown:

δ_{CP} , mass hierarchy

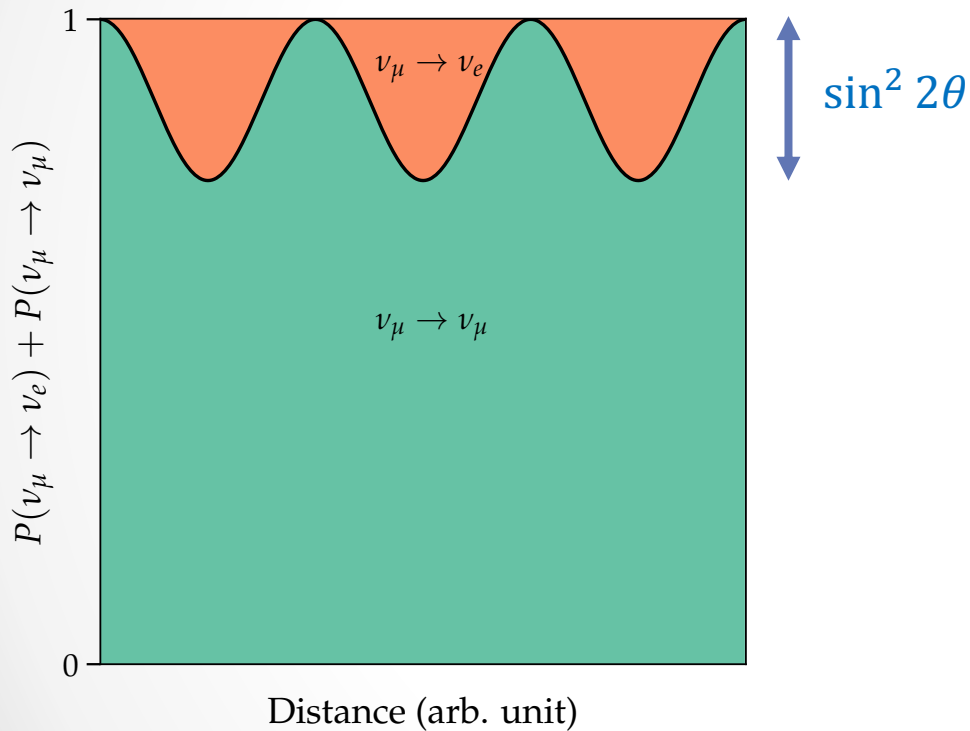
A Few %:

$\theta_{12}, \theta_{13}, \theta_{23}$,

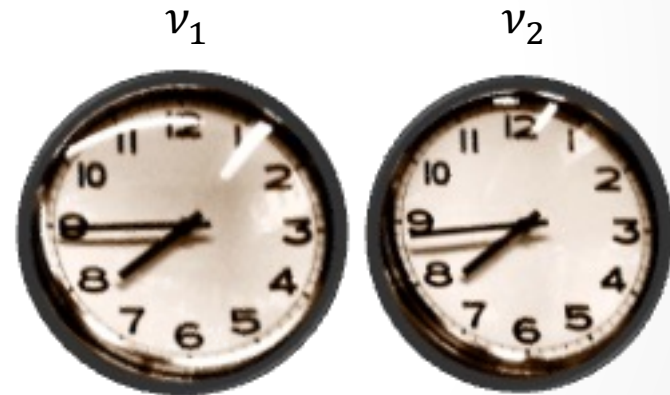
$\Delta m_{21}^2, \Delta m_{32}^2$



Conceptual Idea



Two Out-of-Sync Clocks:



Vacuum Oscillation

Production

$$\underset{\text{Flavor}}{\nu_\alpha(x)} = \sum_{k=1}^3 U_{\alpha k} \underset{\text{Mass}}{\nu_k(x)}$$

3-Flavor Mixing Parameterized by:

$$U_{\text{PMNS}} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Atmospheric

$$\mu \rightarrow \tau$$

500 km/GeV

Reactor/Interference

$$\mu \leftrightarrow e$$

500 km/GeV

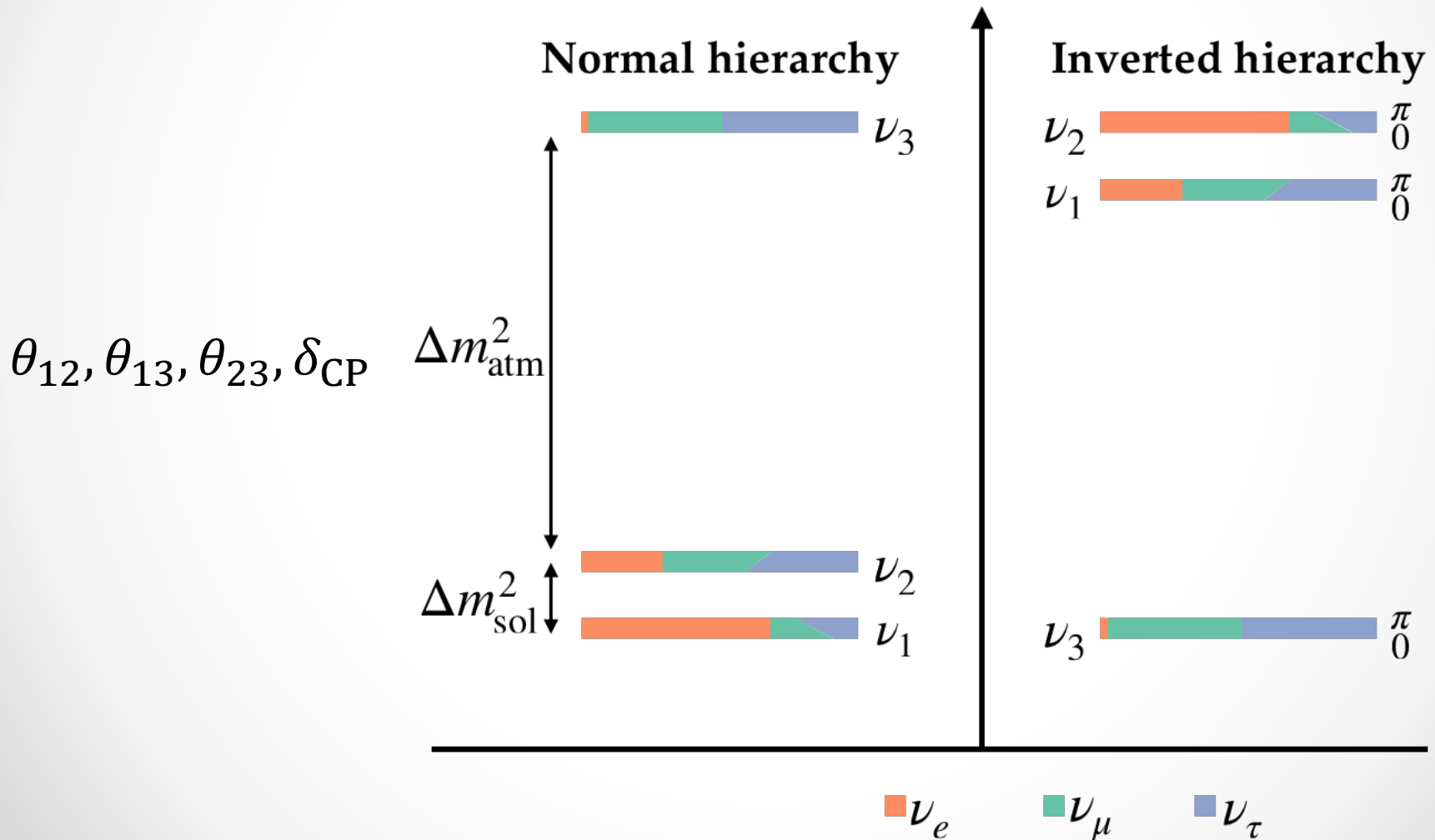
Solar/Reactor

$$e \rightarrow e$$

15000 km/GeV

Vacuum Oscillation

Production



Vacuum Oscillation

Propagation

For a two-level system, the Schrodinger equation in the mass basis:

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} \varphi_1(t) \\ \varphi_2(t) \end{pmatrix} = H_{mass} \begin{pmatrix} \varphi_1(t) \\ \varphi_2(t) \end{pmatrix} = \begin{pmatrix} E_1 & 0 \\ 0 & E_2 \end{pmatrix} \begin{pmatrix} \varphi_1(t) \\ \varphi_2(t) \end{pmatrix}$$

with

$$E_i = \sqrt{p_i^2 + m_i^2} \cong E + \frac{m_i^2}{2E} = E + \frac{m_1^2 + m_2^2}{4E} + \frac{m_i^2 - (m_1^2 + m_2^2)/2}{2E}$$

Dropping identities:

$$H_{mass} = \frac{\delta m^2}{4E} \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix}, \quad \delta m^2 = m_2^2 - m_1^2$$

Vacuum Oscillation

Propagation

In the flavor basis

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} \varphi_e(t) \\ \varphi_\mu(t) \end{pmatrix} = H_{\text{flavor}} \begin{pmatrix} \varphi_e(t) \\ \varphi_\mu(t) \end{pmatrix}$$

with

$$H_{\text{flavor}} = U^\dagger H_{\text{mass}} U = \frac{\delta m^2}{4E} \begin{pmatrix} -\cos 2\theta & \sin 2\theta \\ \sin 2\theta & \cos 2\theta \end{pmatrix}$$

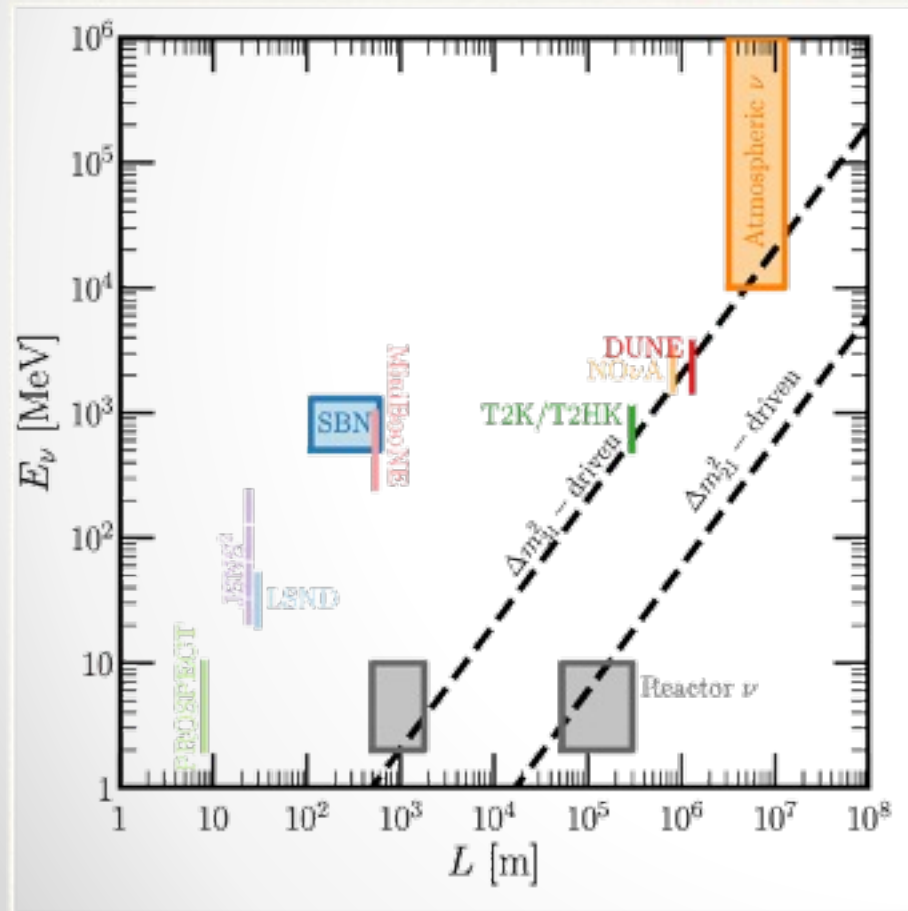
As a result:

$$P(\nu_e \rightarrow \nu_e) = |\varphi_e(t)|^2 = 1 - \sin^2 2\theta \sin^2 \left(\frac{\delta m^2 L}{4E} \right)$$

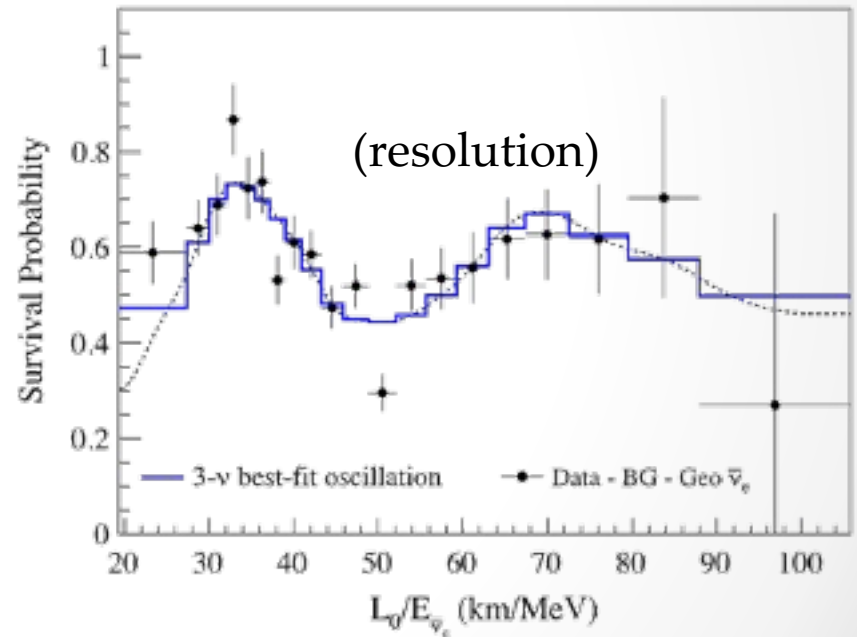
Takeaway Points

L/E dependence

Figure stole from K. Kelly



KamLAND 2016

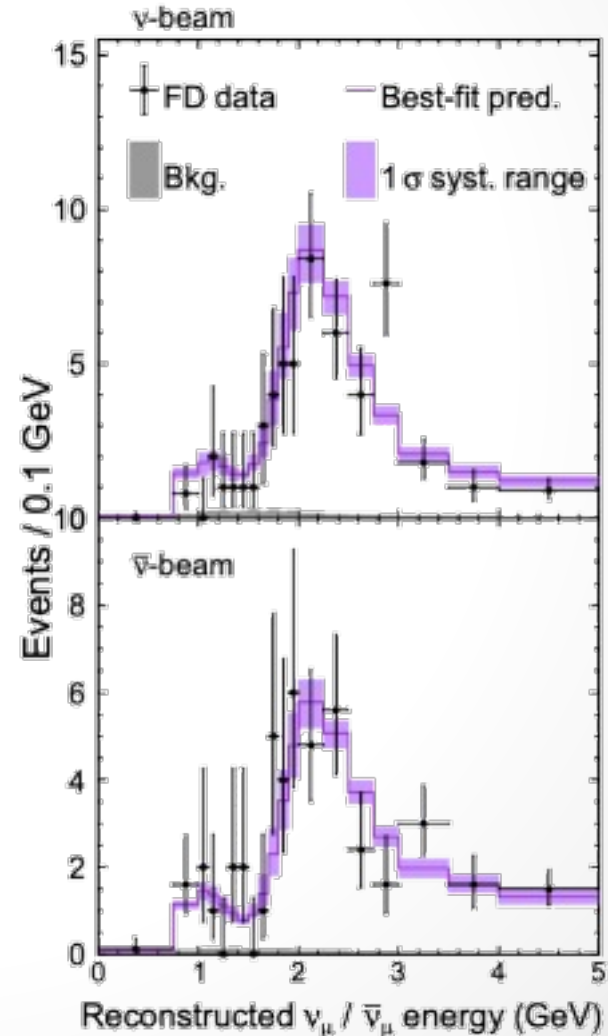
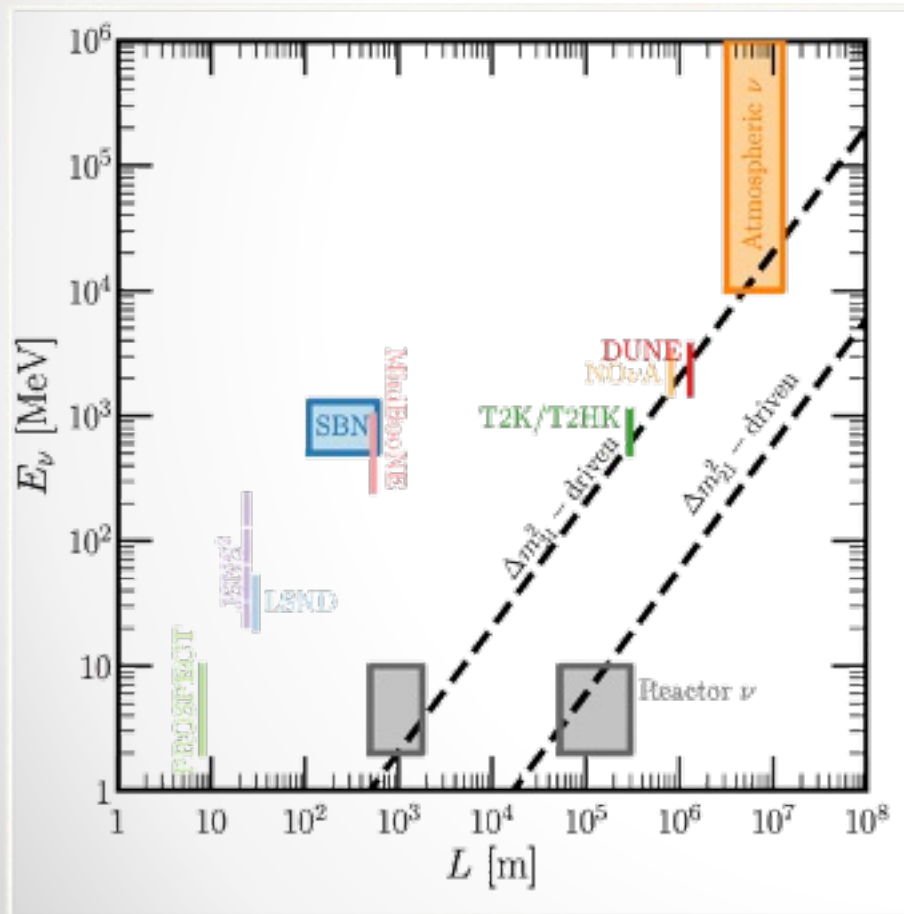


Takeaway Points

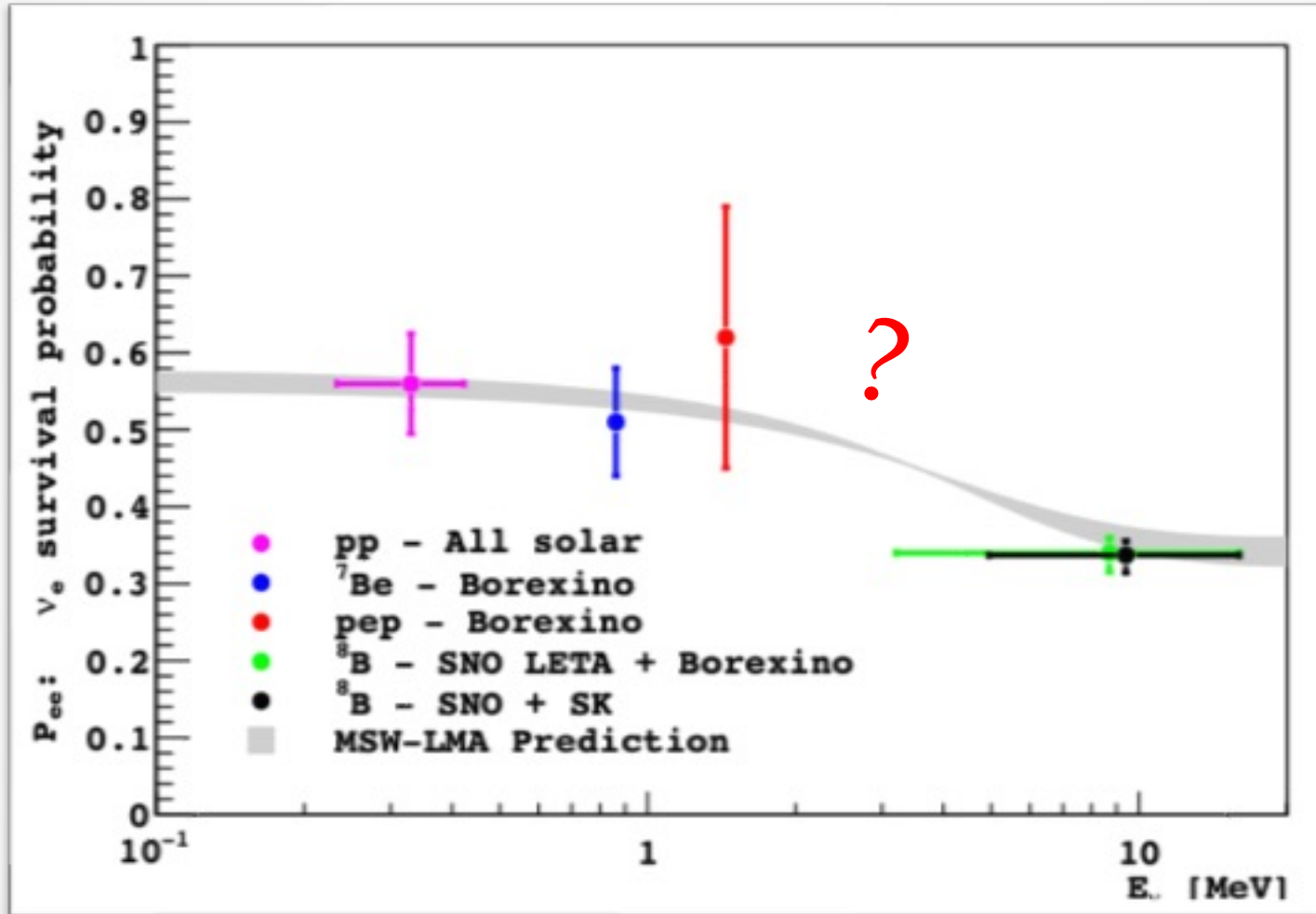
L/E dependence

NOvA 2018

Figure stole from K. Kelly



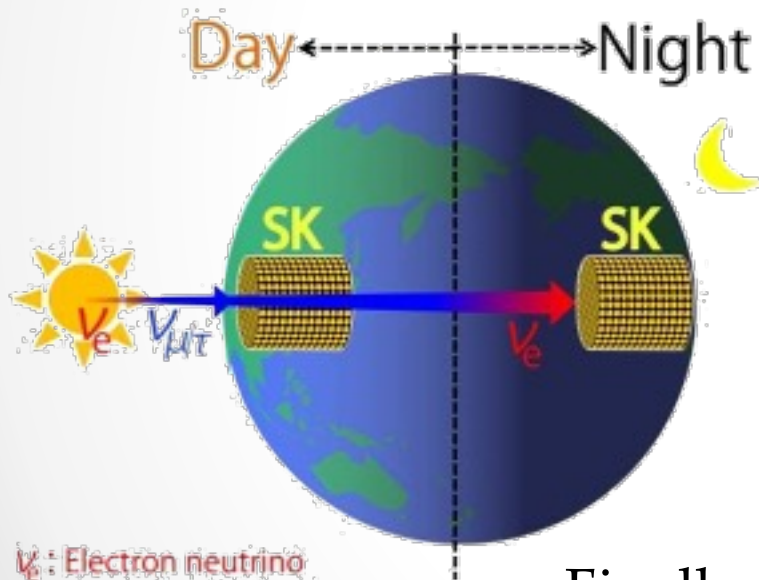
How does this make sense??



Borexino 2014

Matter Effect

Important for solar & supernova neutrinos, and long-baseline oscillation experiments (e.g. DUNE)



ν_e : Electron neutrino
 $\nu_{\mu/\tau}$: Muon/Tau neutrino

sk.icrr.u-tokyo.ac.jp

$$L \supset -2\sqrt{2}G_F(\bar{\nu}_e\gamma_\mu P_L\nu_e)(\bar{e}\gamma^\mu P_L e)$$

Leads to an effective potential:

$$V = \sqrt{2}G_F n_e$$

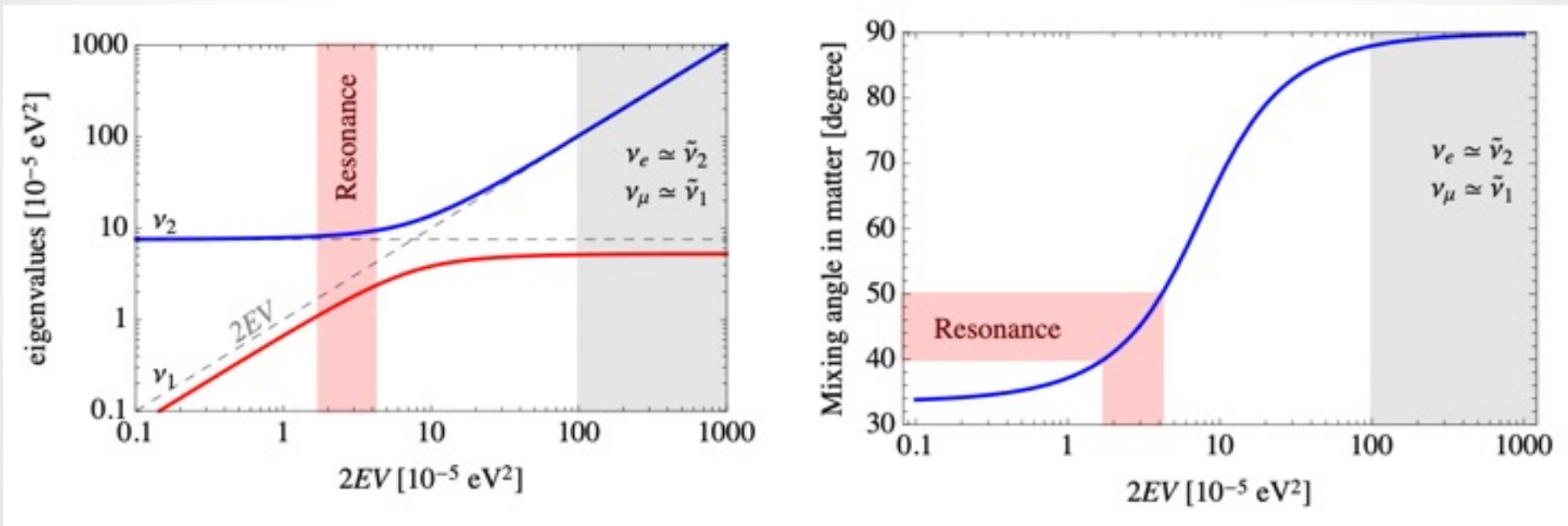
$$H_{\text{matter potential}} = \begin{pmatrix} V & 0 \\ 0 & 0 \end{pmatrix}$$

Finally:

$$H_{\text{flavor}} = \frac{\delta m^2}{4E} \begin{pmatrix} -\cos 2\theta + \frac{2\sqrt{2}G_F n_e}{\delta m^2/E} & \sin 2\theta \\ \sin 2\theta & \cos 2\theta \end{pmatrix}$$

MSW Effect

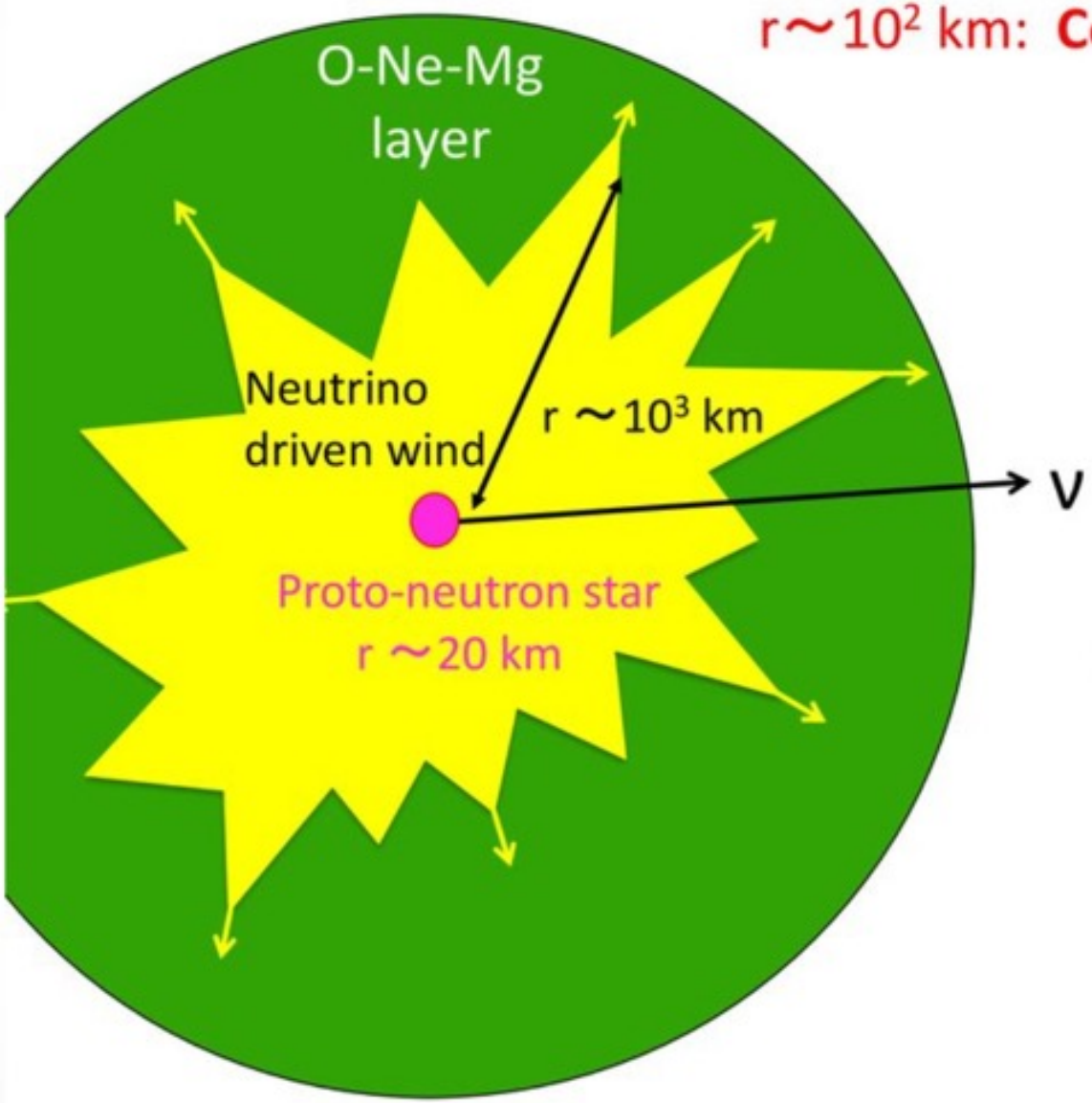
Mikheyev–Smirnov–Wolfenstein effect



Machado 2022

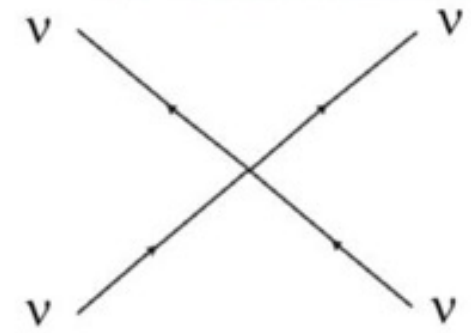
Crucial for establishing solar neutrino oscillation
& determining the sign of Δm_{21}^2

Neutrino oscillations in core-collapse supernovae



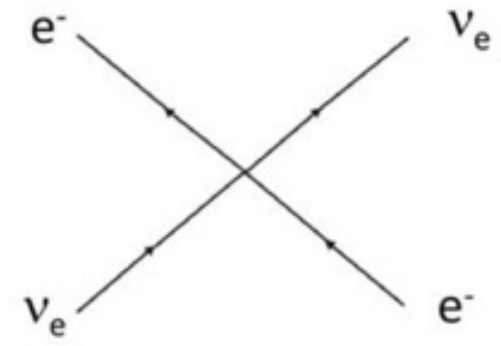
$r \sim 10^2$ km: **Collective neutrino oscillations**

(ν - ν coherent scatterings)



Duan, et al., PRD74, 105014, 2006.

$r \sim 10^3$ km: MSW effect



L. Wolfenstein, PRD17, 2369, 1978.

3-Flavor Neutrino Oscillations

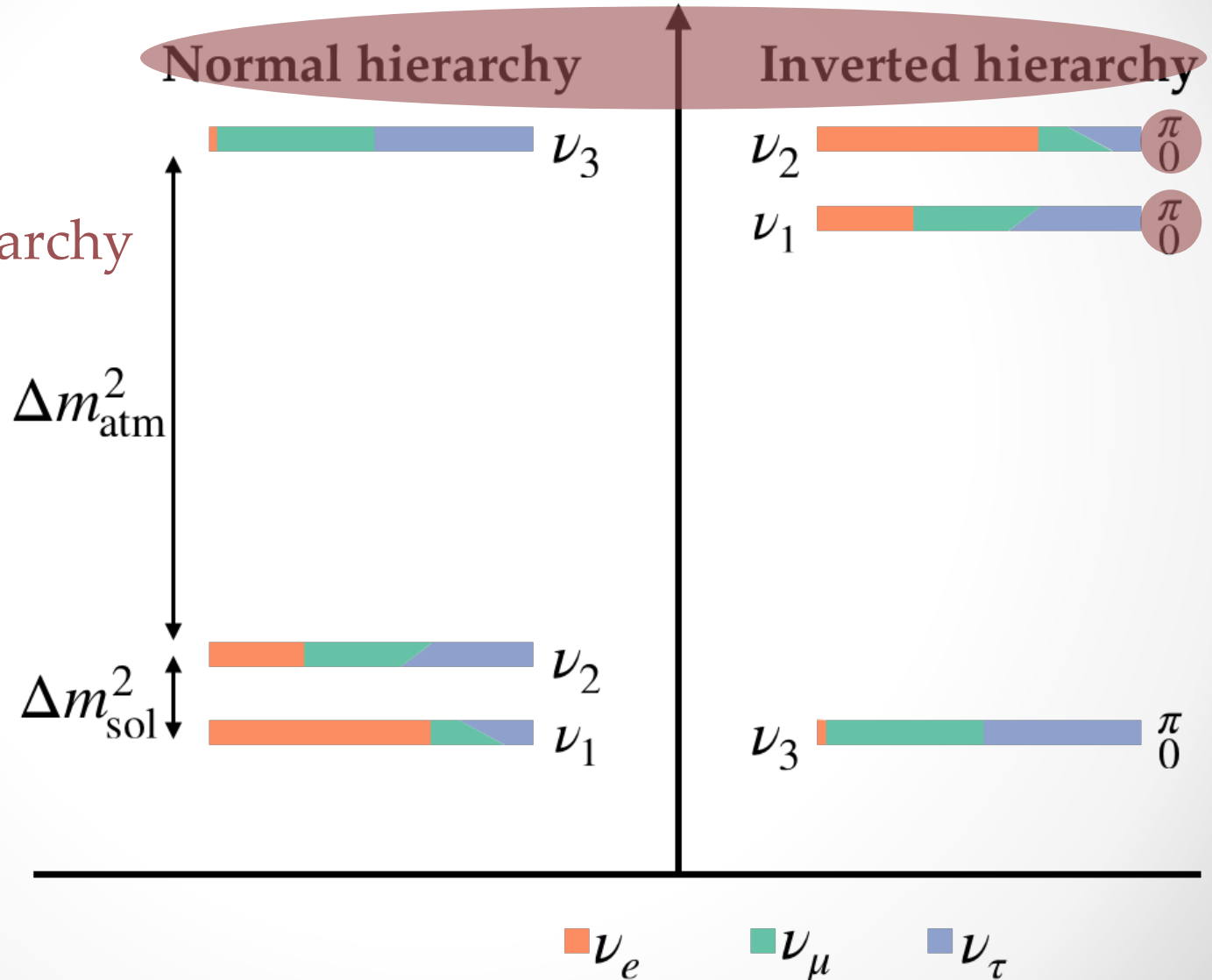
Unknown:

δ_{CP} , mass hierarchy

A Few %:

$\theta_{12}, \theta_{13}, \theta_{23}$,

$\Delta m_{21}^2, \Delta m_{32}^2$



3-Flavor Neutrino Oscillations

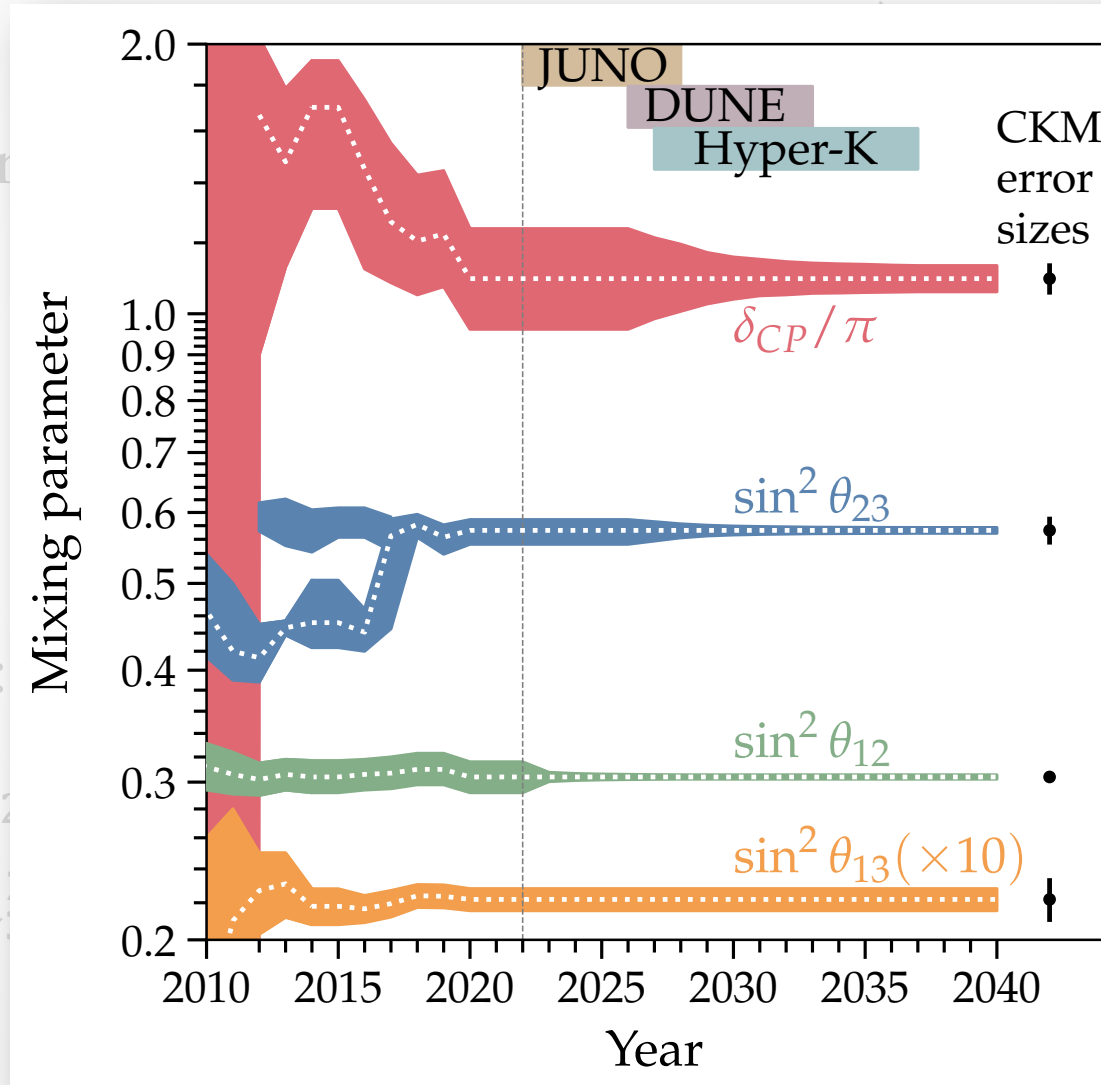
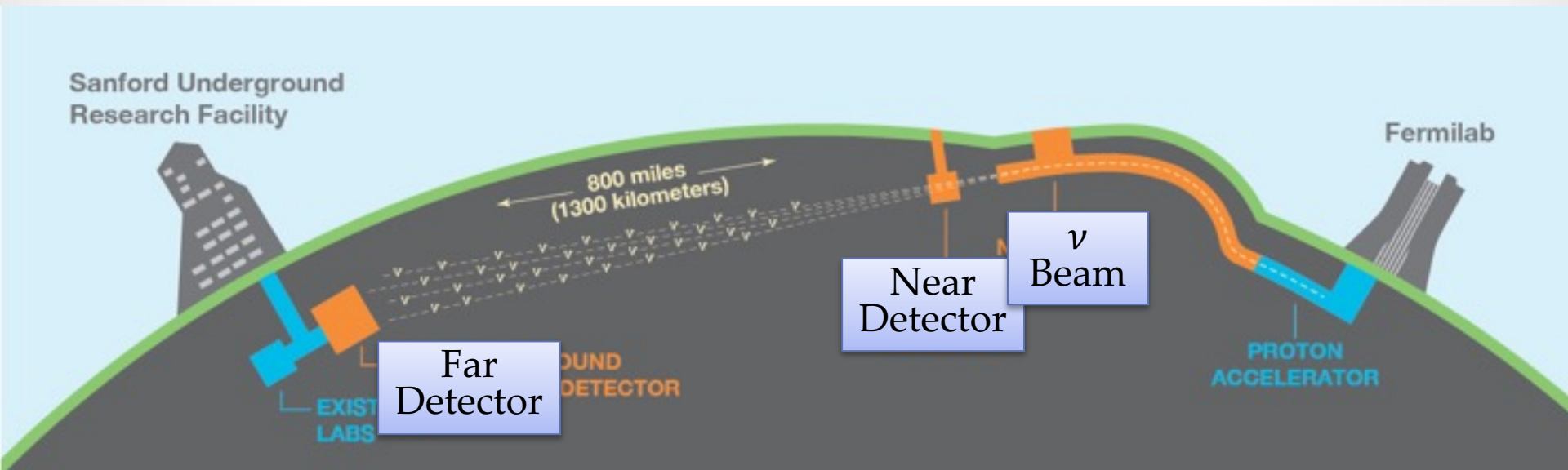


Figure modified from Song et al., 20

Long-Baseline Accelerator Experiments

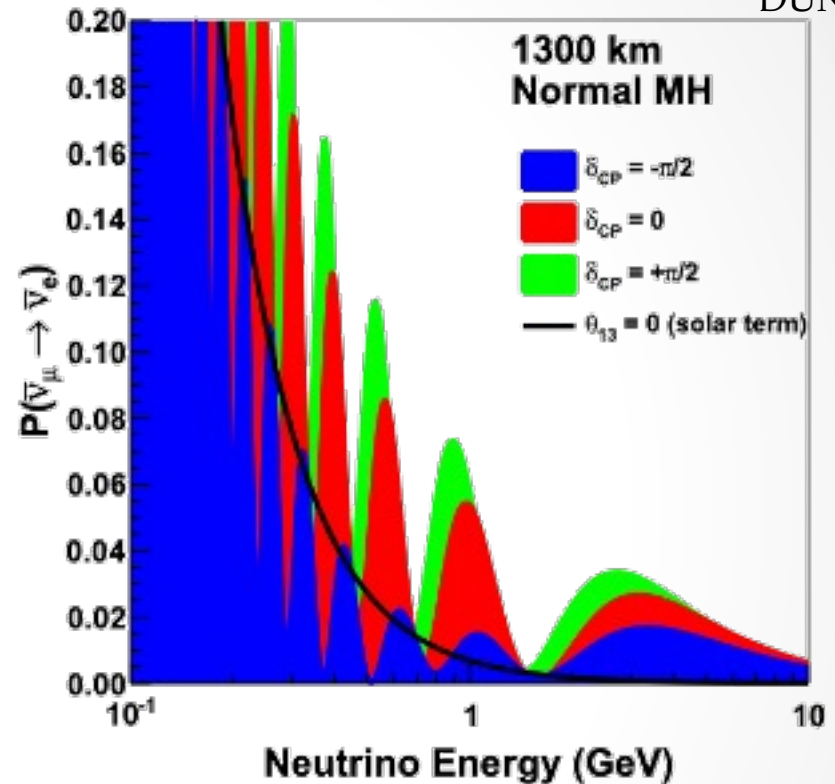
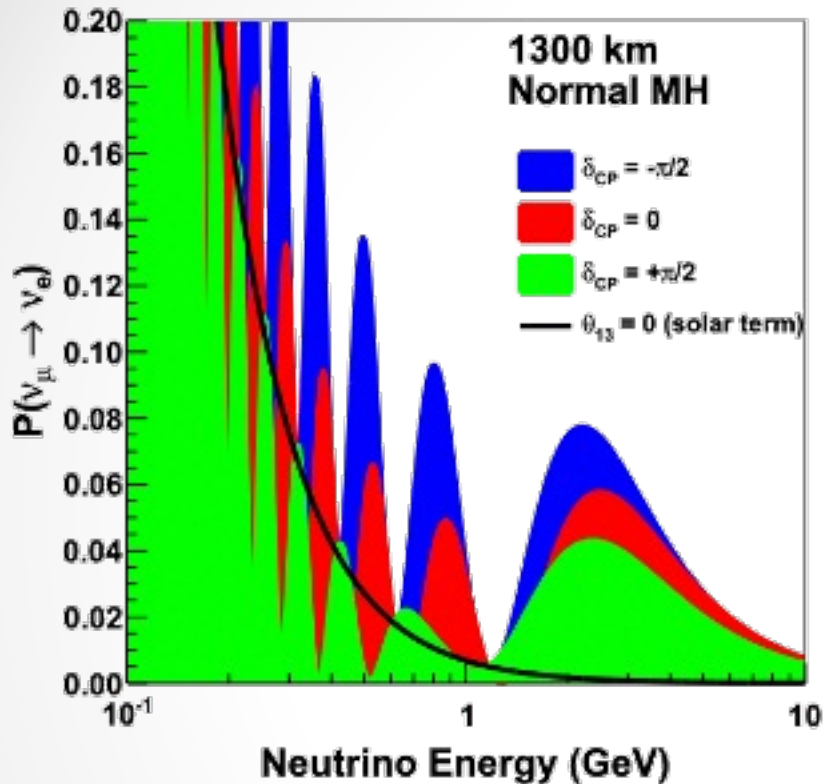
$$0.5 - 5 \text{ GeV } \nu_{\mu} \xrightarrow{\sim 1000 \text{ km}} \nu_e$$

DUNE 15



The Measurements

DUNE 2015



$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) \simeq & \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2(\Delta_{31} - aL)}{(\Delta_{31} - aL)^2} \Delta_{31}^2 \\
 & + \sin 2\theta_{23} \sin 2\theta_{13} \sin 2\theta_{12} \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \Delta_{31} \frac{\sin(aL)}{(aL)} \Delta_{21} \cos(\Delta_{31} + \delta_{CP}) \\
 & + \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(aL)}{(aL)^2} \Delta_{21}^2,
 \end{aligned}$$

Why oscillation?

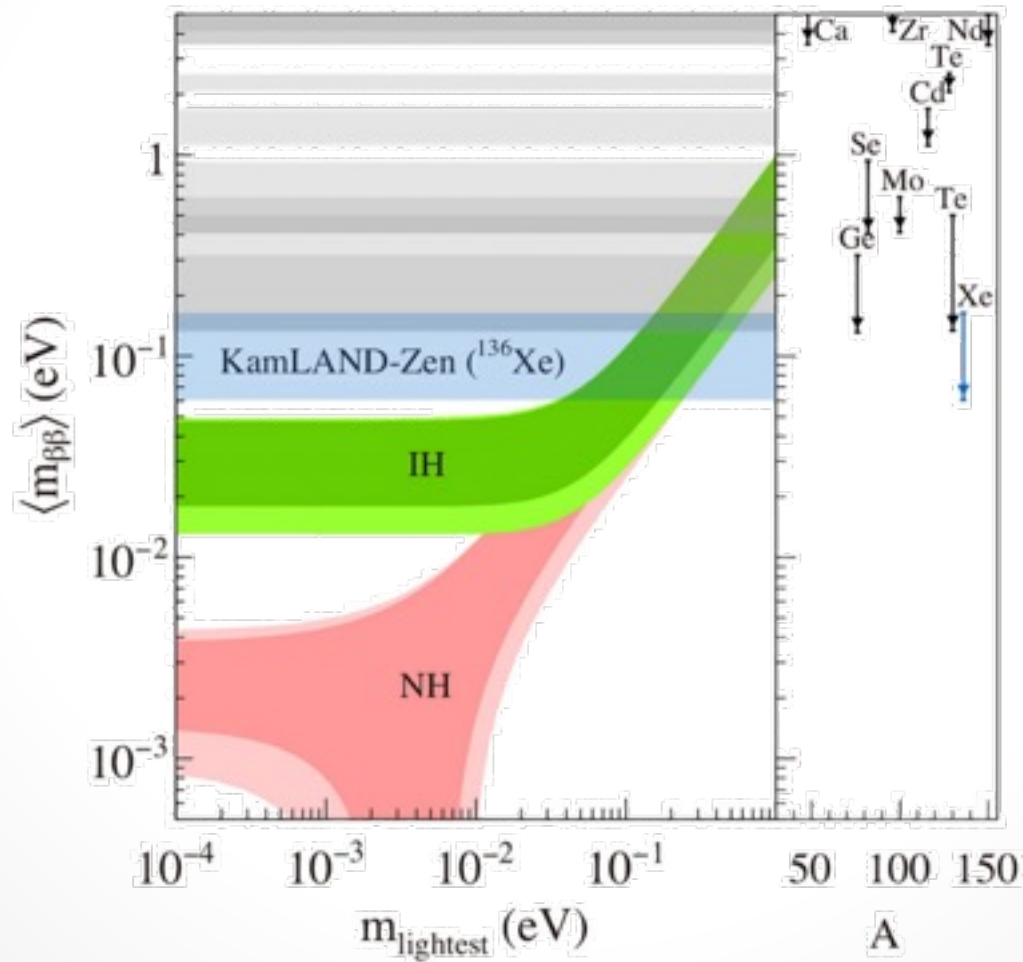
...

What we do not know about neutrino
oscillation

Inputs for Other Physics

KamLAND 2016

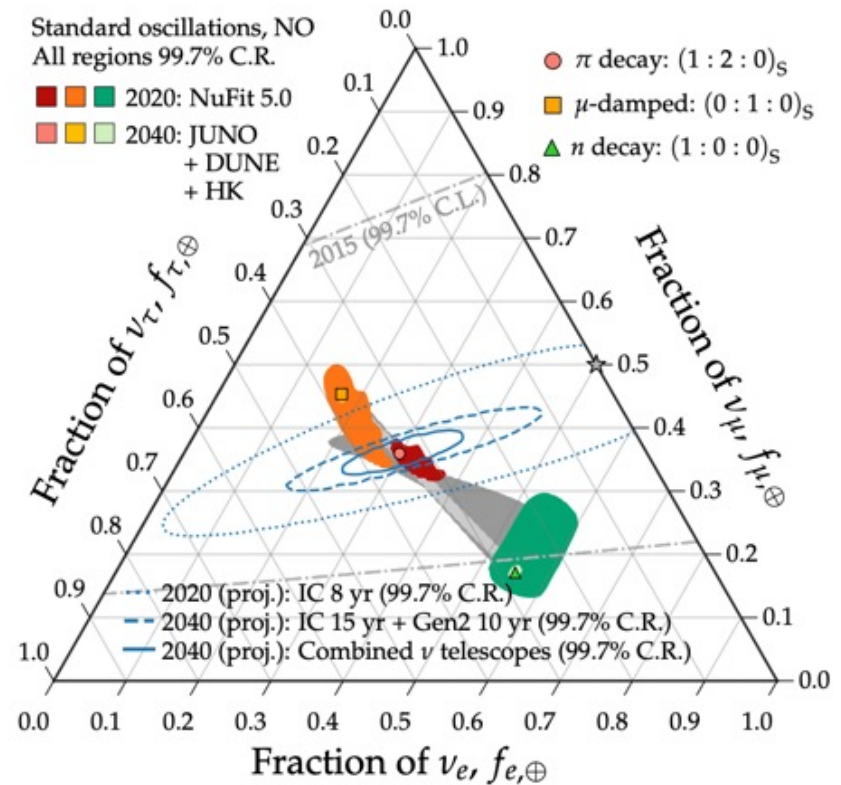
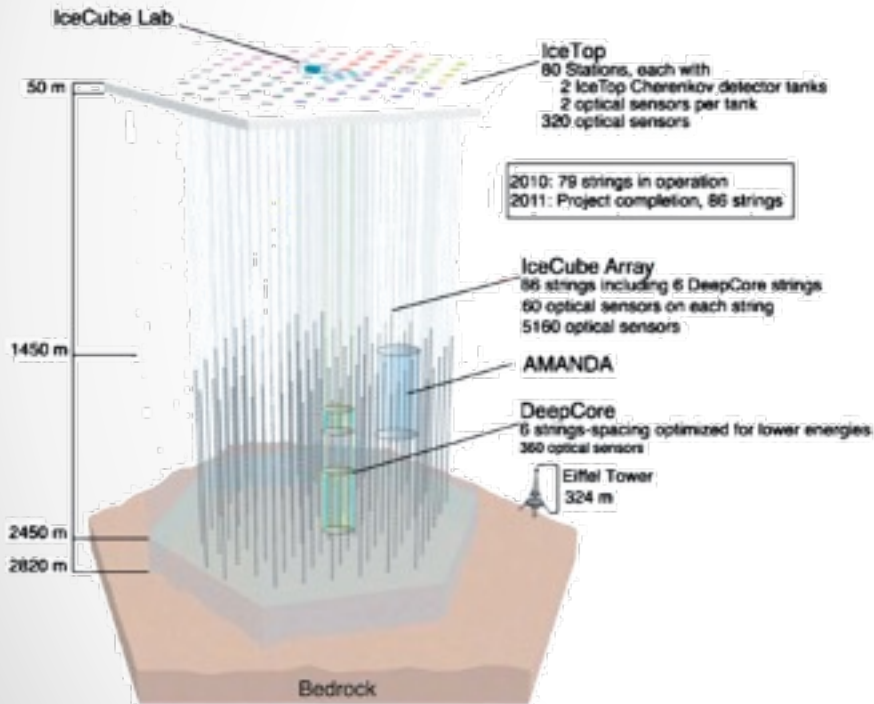
$0\nu\beta\beta$ decay



Inputs for Other Physics

Astrophysical neutrino sources

Song et al., 2020



Testing Oscillation Paradigm

A unitary 3*3 mixing matrix isn't the whole story...

Neutrinos have masses

Weinberg operator:

$$L \supset c_\nu \frac{LHLH}{\Lambda} + \text{h. c.}$$

$$U_{\text{PMNS}}^{\text{Extended}} = \begin{pmatrix} \overbrace{\begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix}}^{U_{\text{PMNS}}^{3 \times 3}} & \cdots & U_{en} \\ \vdots & \ddots & \vdots \\ U_{s_n1} & U_{s_n2} & U_{s_n3} & \cdots & U_{s_nn} \end{pmatrix}$$

New Physics at Scale $\Lambda \Rightarrow$

New States \Rightarrow

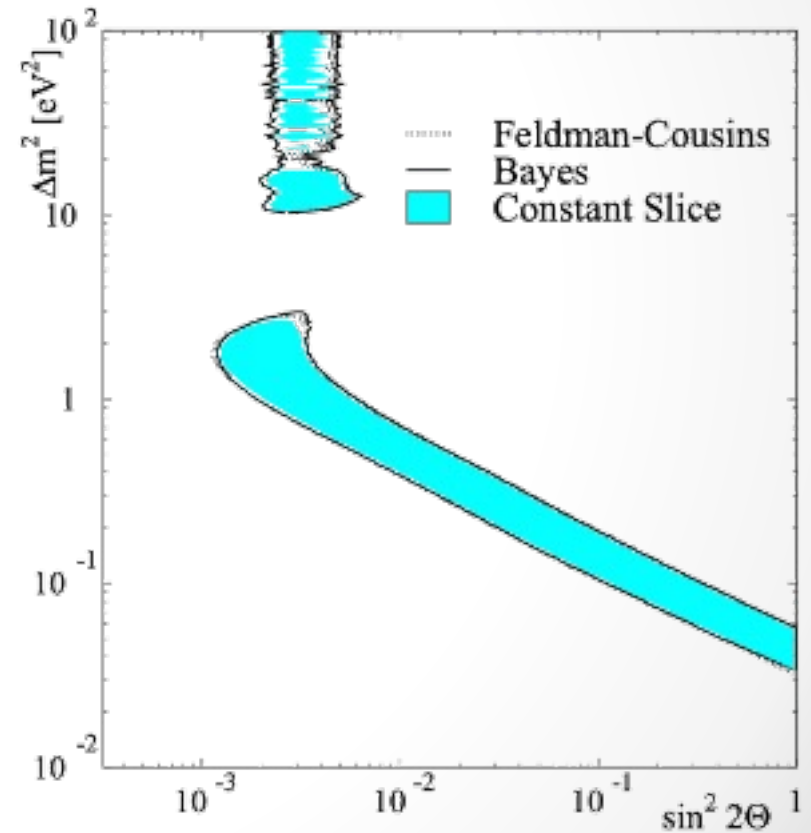
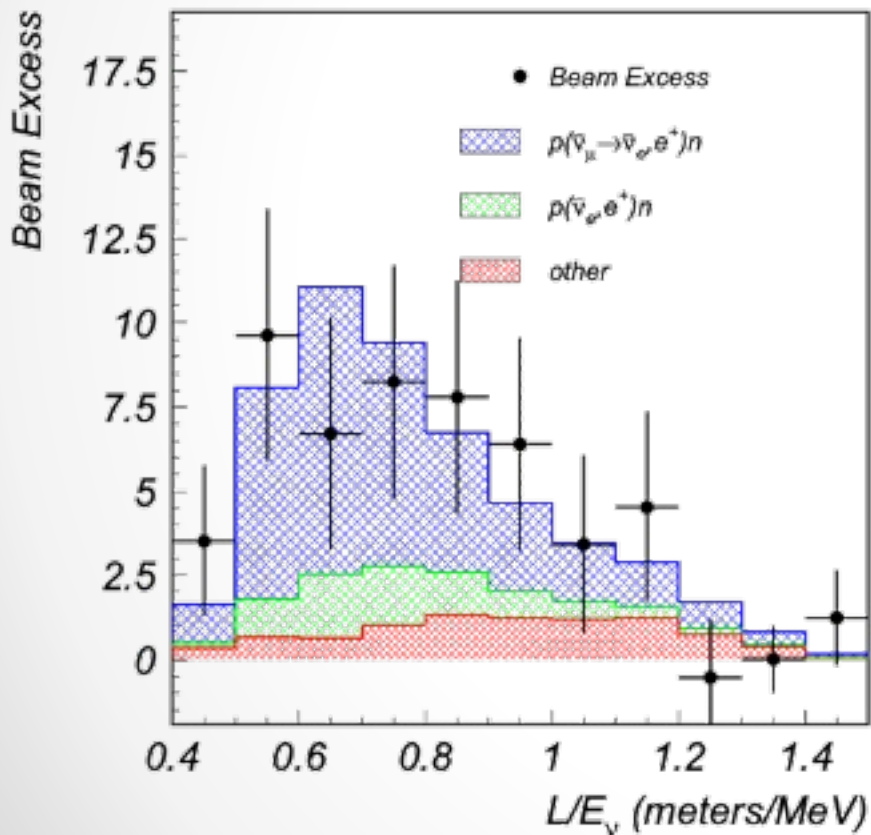
$$(3 \times 3) U_{\alpha k} \rightarrow (n \times n) \mathcal{U}_{\alpha k}$$

Sterile neutrinos?

Short-Baseline Anomaly

LSND (Liquid Scintillator Neutrino Detector)

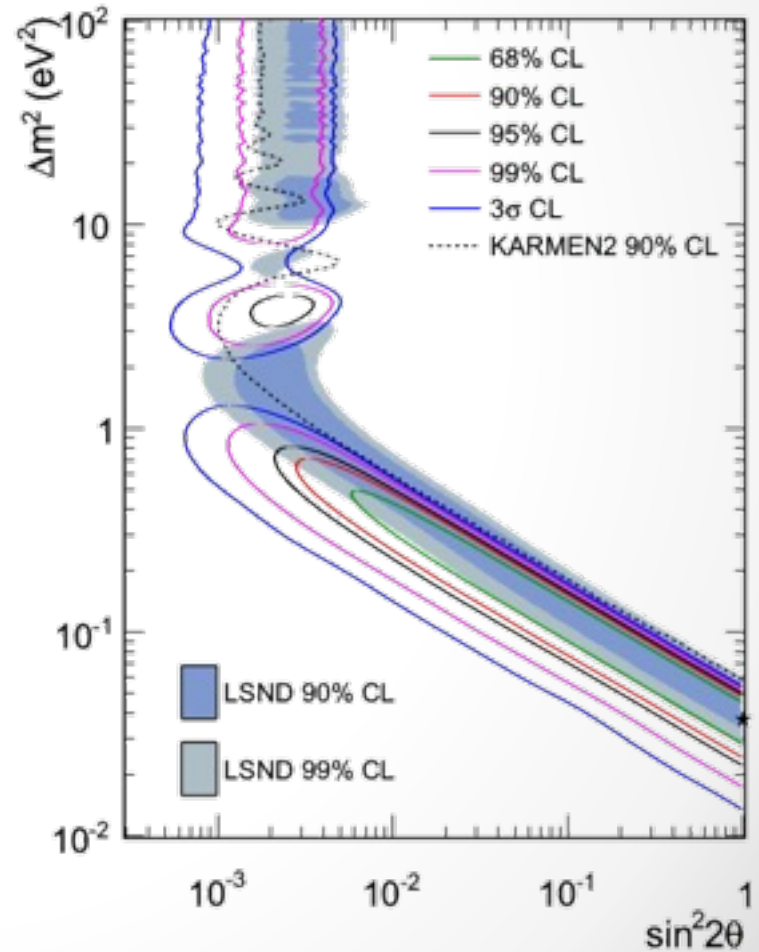
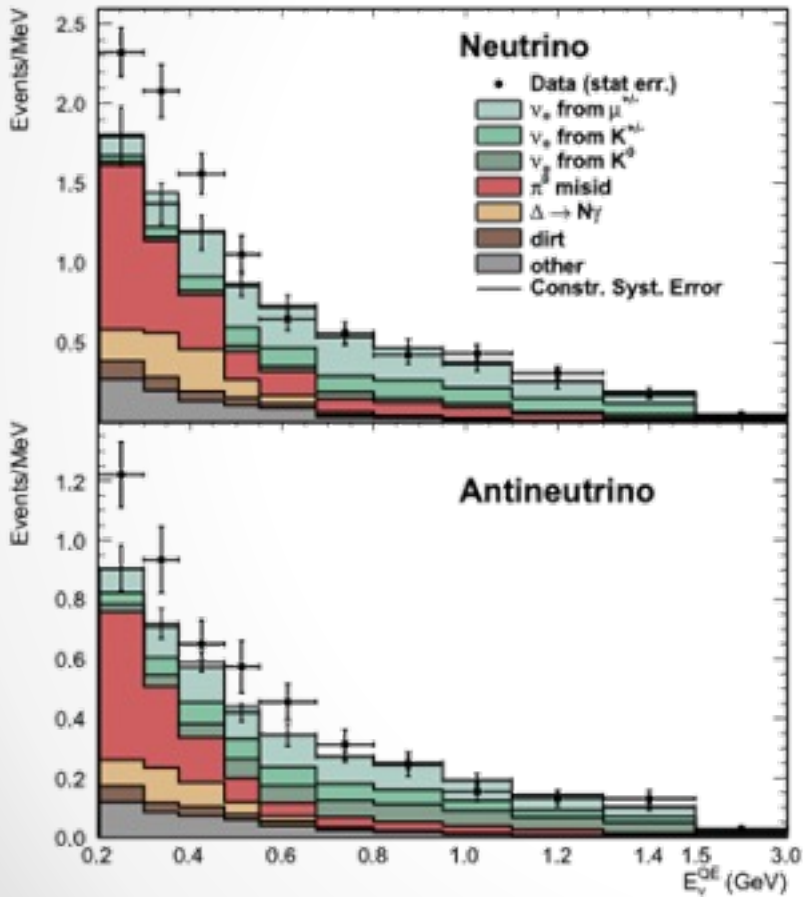
LSND 2001



Short-Baseline Anomaly

MiniBooNE

MiniBooNE 2012

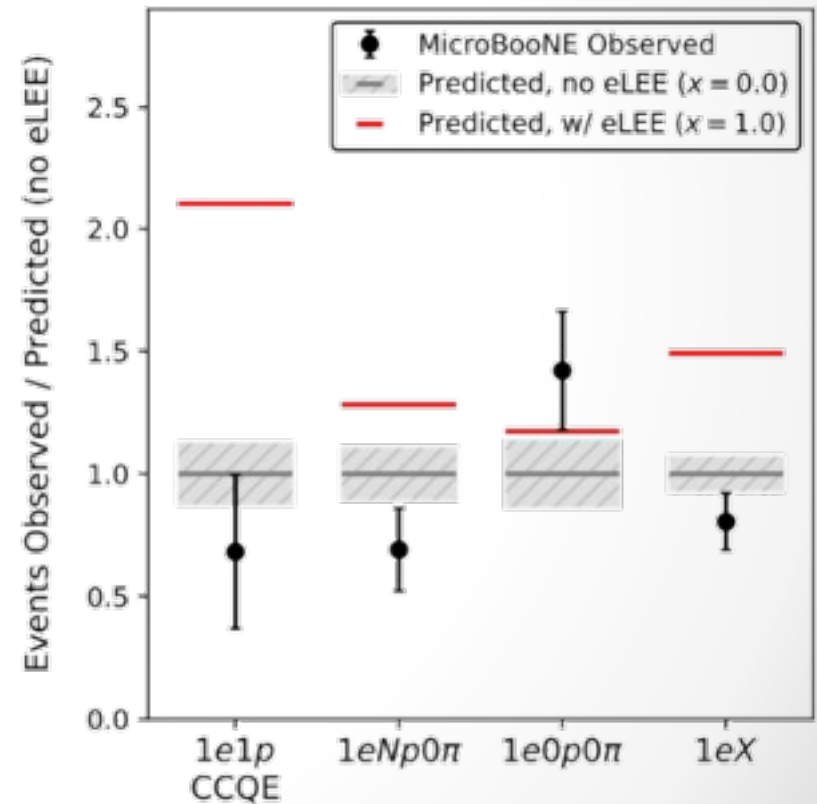
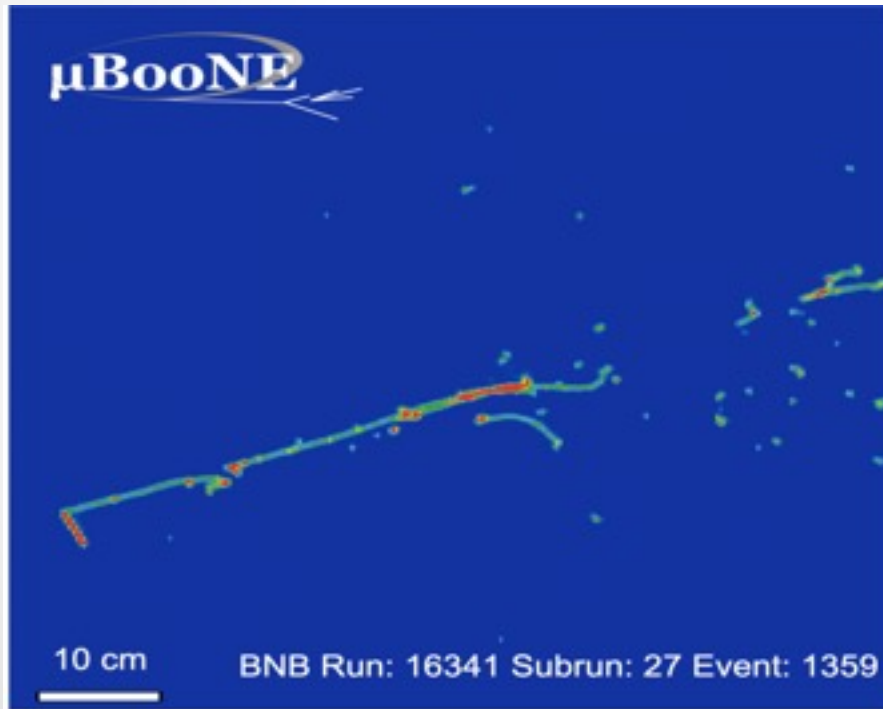


Short-Baseline Anomaly

MicroBooNE

MicroBooNE 2021

Liquid argon TPC

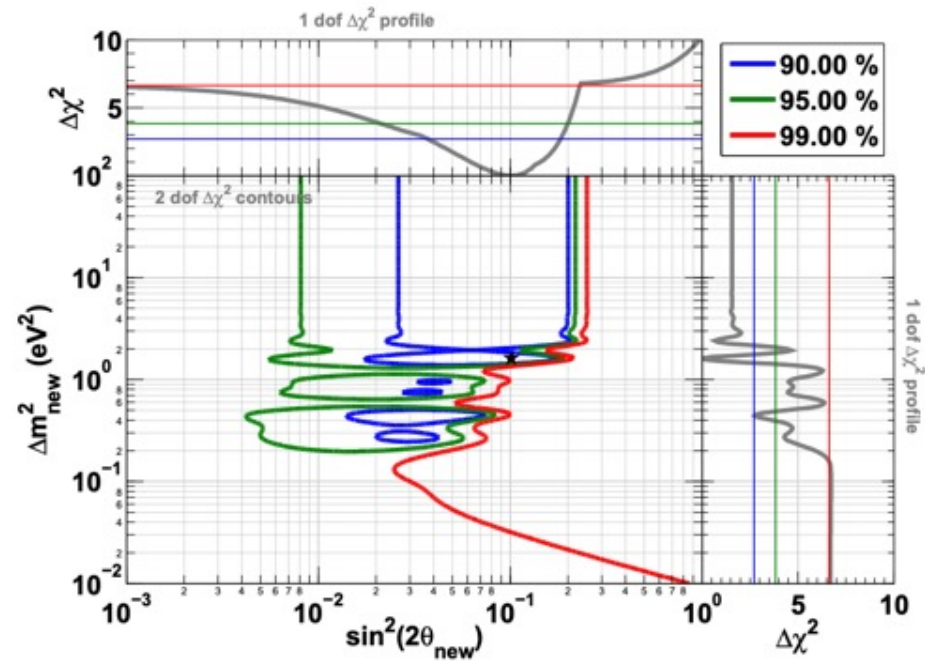
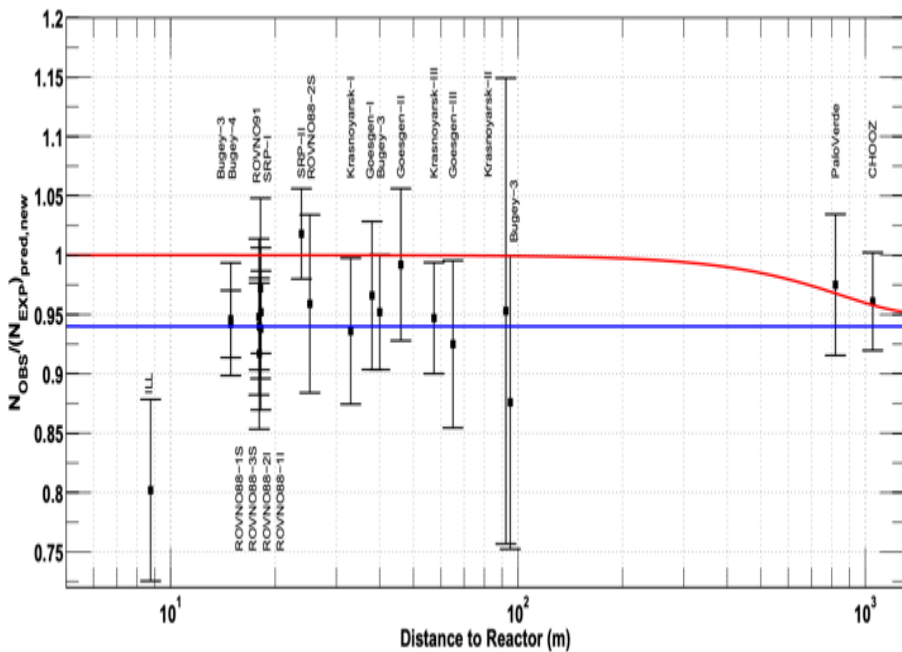


The jury is still out

Reactor Anomaly

From a combination of experiments

Mention et al., 2011

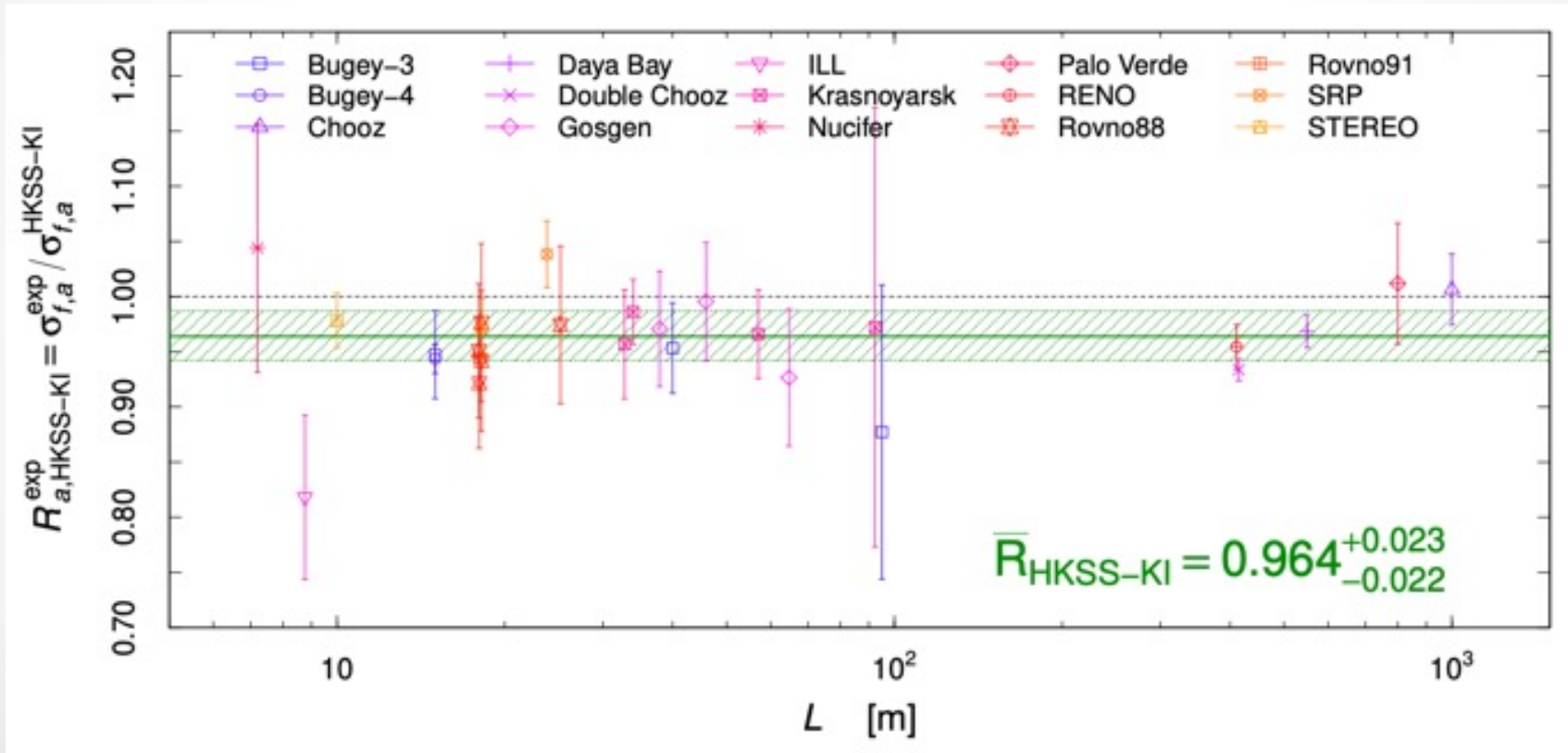


Similar region as the short-baseline anomaly hints at

Reactor Anomaly

The key is evaluation of reactor neutrino fluxes

Giunti et al., 2022



The jury is still out

Conclusions

➤ Standard three-flavor neutrino oscillation

- Two-level system
- Matter effect
- Three-flavor oscillation

➤ Sterile neutrinos

- Neutrino-nucleus cross sections
- Reactor flux evaluations
- Experimental capabilities

