MASSACHUSETTS INSTITUTE OF TECHNOLOGY Physics Department

Physics 8.286: The Early Universe Prof. Alan Guth October 2, 2007

QUIZ 1 FORMULA SHEET

DOPPLER SHIFT (For motion along a line):

z = v/u (nonrelativistic, source moving)

$$z = \frac{v/u}{1 - v/u} \quad \text{(nonrelativistic, observer moving)}$$
$$z = \sqrt{\frac{1 + \beta}{1 - \beta}} - 1 \quad \text{(special relativity, with } \beta = v/c\text{)}$$

COSMOLOGICAL REDSHIFT:

$$1 + z \equiv \frac{\lambda_{\text{observed}}}{\lambda_{\text{emitted}}} = \frac{R(t_{\text{observed}})}{R(t_{\text{emitted}})}$$

SPECIAL RELATIVITY:

Time Dilation Factor:

$$\gamma \equiv \frac{1}{\sqrt{1-\beta^2}} , \qquad \beta \equiv v/c$$

Lorentz-Fitzgerald Contraction Factor: γ

Relativity of Simultaneity:

Trailing clock reads later by an amount $\beta \ell_0/c$.

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EVOLUTION OF A MATTER-DOMINATED UNIVERSE:

$$\begin{split} H^2 &= \left(\frac{\dot{R}}{R}\right)^2 = \frac{8\pi}{3}G\rho - \frac{kc^2}{R^2} , \quad \ddot{R} = -\frac{4\pi}{3}G\rho R ,\\ \rho(t) &= \frac{R^3(t_i)}{R^3(t)} \,\rho(t_i) ,\\ \Omega &\equiv \rho/\rho_c , \text{ where } \rho_c = \frac{3H^2}{8\pi G} . \end{split}$$

Flat
$$(k = 0)$$
: $R(t) \propto t^{2/3}$,
 $\Omega = 1$.

Closed
$$(k > 0)$$
: $ct = \alpha(\theta - \sin \theta)$, $\frac{R}{\sqrt{k}} = \alpha(1 - \cos \theta)$,
 $\Omega = \frac{2}{1 + \cos \theta} > 1$,
where $\alpha \equiv \frac{4\pi}{3} \frac{G\rho}{c^2} \left(\frac{R}{\sqrt{k}}\right)^3$.
Open $(k < 0)$: $ct = \alpha \left(\sinh \theta - \theta\right)$, $\frac{R}{\sqrt{\kappa}} = \alpha \left(\cosh \theta - 1\right)$,
 $\Omega = \frac{2}{1 + \cosh \theta} < 1$,

$$\Omega = \frac{1}{1 + \cosh \theta} < 1 ,$$

where $\alpha \equiv \frac{4\pi}{3} \frac{G\rho}{c^2} \left(\frac{R}{\sqrt{\kappa}}\right)^3 ,$

$$\kappa \equiv -k > 0$$
 .