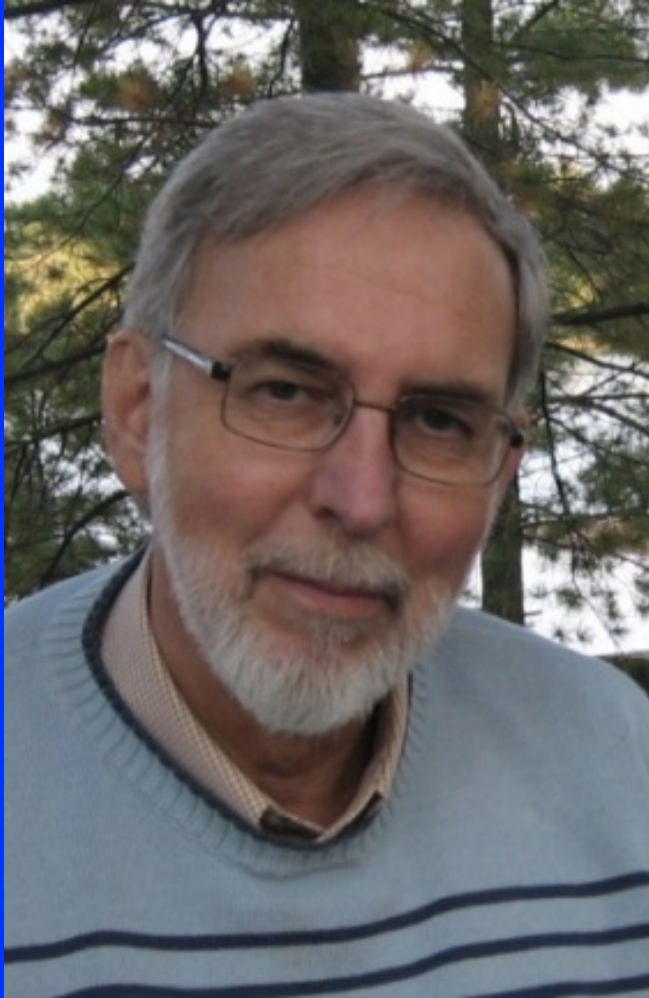


# The superfluid mass density



**Gordon Baym**  
**University of Illinois, Urbana**

**Finite Temperature Non-Equilibrium  
Superfluid Systems  
Heidelberg, 19 September 2011**

In fondest memory,  
Allan Griffin, d. May 19. 2011



# Landau Two-Fluid Model

Can picture superfluid  $^4\text{He}$  as two interpenetrating fluids:

Normal: density  $\rho_n(T)$ , velocity  $v_n$

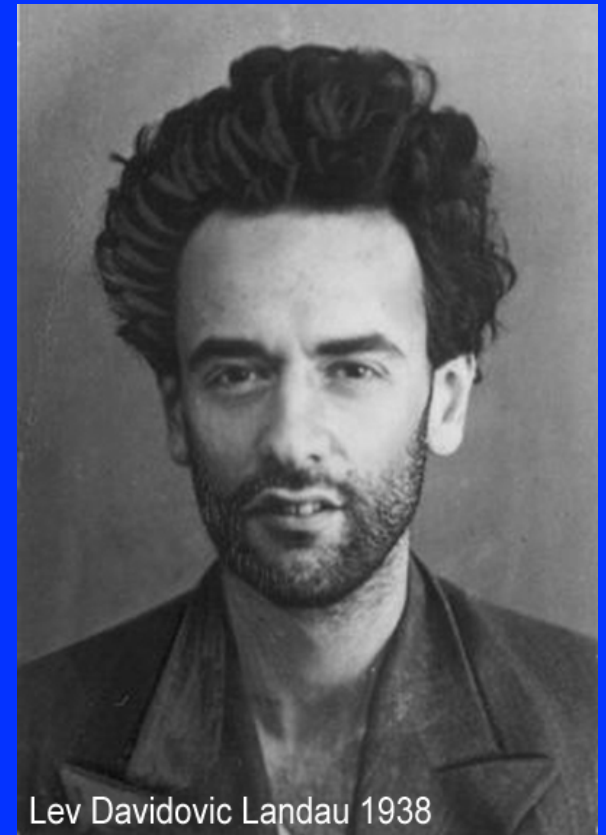
Superfluid: density  $\rho_s(T)$ , velocity  $v_s$

$$\rho = \rho_n(T) + \rho_s(T)$$

Mass current =  $\rho_s v_s + \rho_n v_n$

Entropy current =  $sv_n$

:carried by normal fluid only



Lev Davidovic Landau 1938

Second sound (collective mode) =

counter-oscillating normal and superfluids

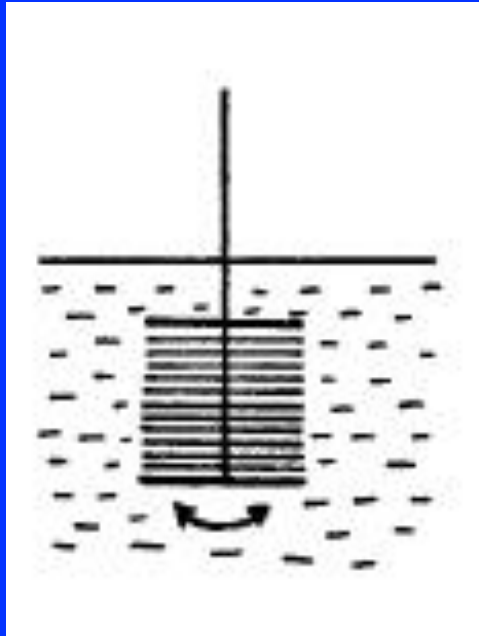
# **The discovery of superfluidity: a chronology of events in 1935-1938**

**Allan Griffin**  
**Department of Physics,**  
**University of Toronto**  
**Toronto, Ontario,**  
**Canada M5S 1A7**  
**griffin@physics.utoronto.ca**

**(September 21, 2006)**

The dramatic announcement of superfluidity of liquid  $^4\text{He}$  in 1938 is one of the defining moments in modern physics. The two short notes which were published back to back in the Jan. 8 issue of *Nature* (by Kapitza [1] working in Moscow and Allen and Misener [2] working in Cambridge University) immediately caught the attention of the physics community. This stimulated feverish activity in the period leading up to World War II, and in the 1950s developed into a major research area called “quantum fluids.”

**Try to rotate helium slowly. Normal fluid component rotates, but superfluid component stays put.**



Andronikashvili experiment – with stack of closely spaced disks oscillating back and forth – measure how much fluid rotates

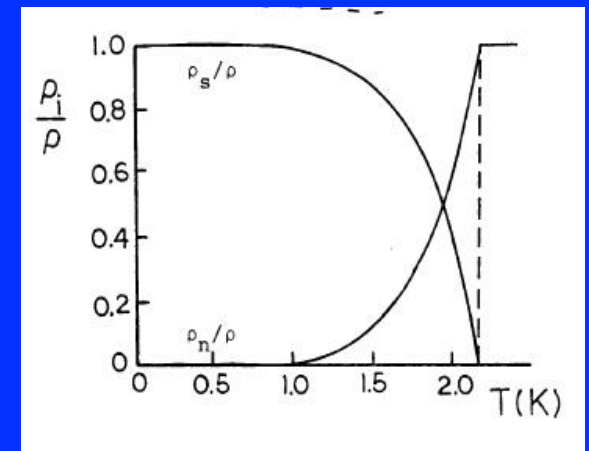
Moment of inertia

$$I = I_{\text{disk}} + I_{\text{fluid}}$$

*E.L. Andronikashvili,  
J. Physics, USSR, 1946*

Measure resonant frequency,  
and deduce  $I_{\text{fluid}}$  from

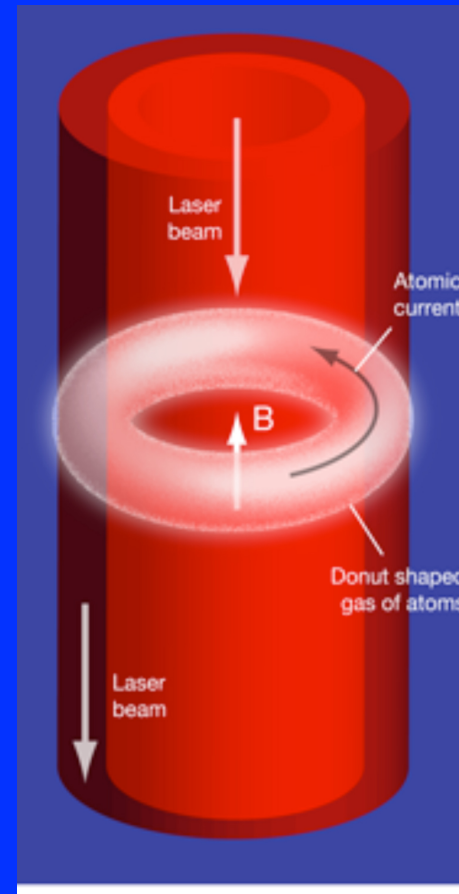
$$I \frac{d^2\theta}{dt^2} = -k\theta$$



# Andronikashvili experiment in cold atoms to measure superfluid mass density

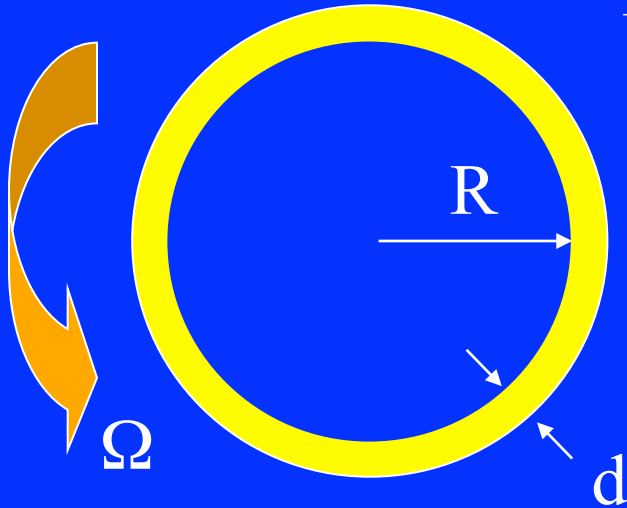
*N.R. Cooper and Z. Hadzibabic, PRL 104, 030401 (2010)*

Two-photon Raman coupling via beams with different orbital angular momentum difference, to simulate uniform rotation



# Hess-Fairbank experiment (Phys. Rev. Lett. 19, 216 (1967))

Rotate thin ( $d \ll R$ ) annulus of liquid  ${}^4\text{He}$  at  $\Omega$



1) Rotate slowly at  $T > T_\lambda$ :  $\Omega < \Omega_c \sim 1/mR^2$   
liquid rotates classically with angular momentum  $L = I_{\text{classical}} \Omega$ .

$$I_{\text{classical}} = NmR^2$$

2) **Cool to  $T < T_\lambda$** : liquid rotates with reduced moment of inertia  $I(T) < I_{\text{classical}}$ .  $I(T=0) = 0$ .

Only the normal fluid rotates.  $I(T) = (\rho_n/\rho)I_{\text{classical}}$   
The superfluid component remains stationary in the lab.

Reduction of moment of inertia is an equilibrium phenomenon.

# Moment of inertia of superfluid

Reduction of moment of inertia due condensation

= analog of Meissner effect.

$$I = \frac{\rho_n}{\rho} I_{classical}$$

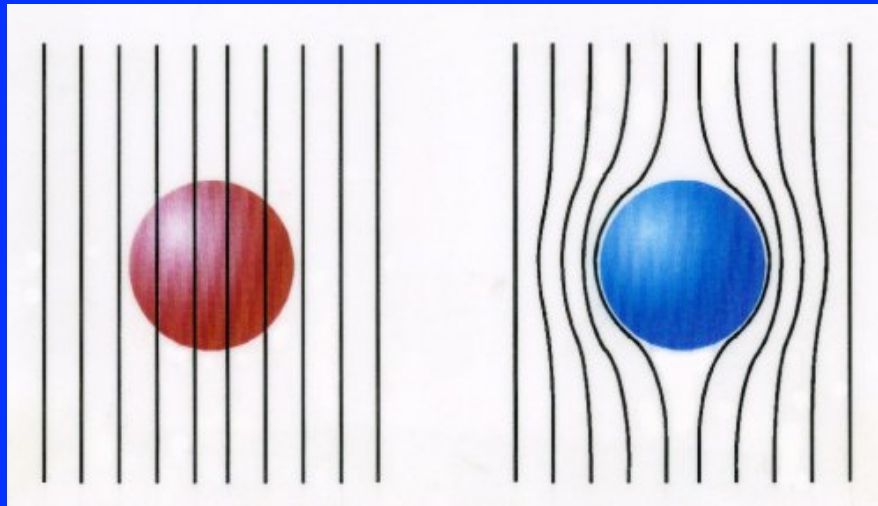
Rotational spectra of nuclei:  
 $E = J^2/2I$ , indicate moment of inertia,  $I$ , reduced from rigid body value,  $I_{cl}$ . Migdal (1959). BCS pairing.

Element	$\beta$ [7]	$x_p$	$x_n$	$\left(\frac{J}{J_0}\right)_{rect.}$	$\left(\frac{J}{J_0}\right)_{osc.}$	$\left(\frac{J}{J_0}\right)_{exper.}$ [7]
Nd <sup>150</sup>	0.26	0.54	0.94	0.15	0.38	0.35
Sm <sup>152</sup>	0.24	0.65	1.02	0.17	0.43	0.38
Gd <sup>154</sup>	0.26	0.52	0.88	0.13	0.35	0.36
Gd <sup>156</sup>	0.33	0.87	1.37	0.22	0.57	0.48
Gd <sup>157</sup>	0.29	0.93	1.60	0.22	0.64	0.60
Dy <sup>162</sup>	0.30	0.84	1.43	0.23	0.57	0.50
Hf <sup>179</sup>	0.20	0.99	1.75	0.27	0.66	0.52
Os <sup>186</sup>	0.18	0.44	0.69	0.09	0.26	0.28
Th <sup>230</sup>	0.22	0.63	0.95	0.15	0.40	0.43
Th <sup>232</sup>	0.22	0.84	1.42	0.24	0.60	0.44
U <sup>238</sup>	0.24	0.83	1.29	0.22	0.54	0.43

# The Meissner effect

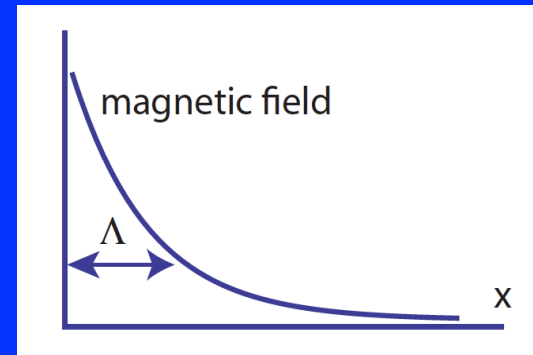
*W. Meissner and R. Ochsenfeld, Berlin 1933*

Superconductors expel magnetic fields  
(below critical field):



normal      superconducting

Screening of magnetic field  
within penetration depth  $\Lambda$

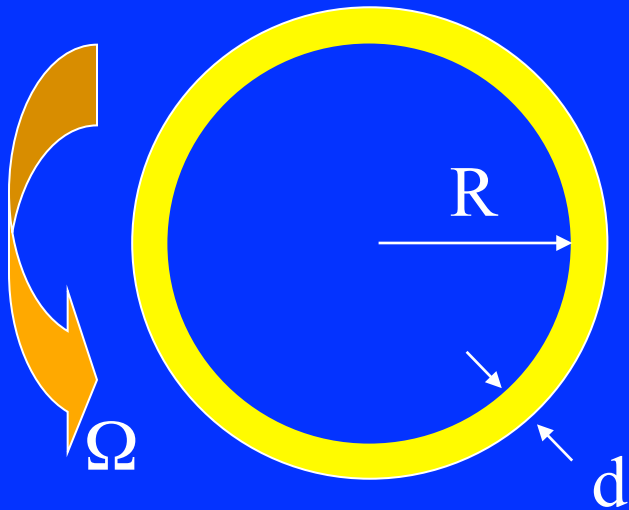


Fundamental property of superconductors:  
perfect diamagnets -- not perfect  
conductors! Equivalent to reduced  
moment of inertia in neutral superfluids

$$\frac{1}{\Lambda^2} = \frac{4\pi n e^2 \rho_s}{m c^2 \rho}$$



# Superfluid flow



1) Rotate rapidly at  $T > T_\lambda$ :  $\Omega > \Omega_c$   
liquid rotates classically with angular momentum  $L = I_{\text{classical}} \Omega$ .

2) Continue rotating, cool to  $T < T_\lambda$ :  
liquid rotates classically

3) Stop rotation of annulus. Liquid keeps rotating with  $L = I_s \Omega$ ,  
where  $I_s = (\rho_s / \rho) I_{\text{classical}}$

Only the superfluid rotates. The normal component is stationary.

Superfluid flow is metastable (albeit with huge lifetime in macroscopic system)

**Order parameter of condensate**  $\Psi(\vec{r}) = |\psi|e^{i\phi(\vec{r})}$

wave function of mode into which particles condense

**Defined more rigorously by eigenfunction of**

**largest eigenvalue of density matrix**

$$\langle \psi(\vec{r})\psi^\dagger(\vec{r}') \rangle \rightarrow \Psi(\vec{r})\Psi(\vec{r}')^*$$

Superfluid velocity:

$$\vec{v}_s(\vec{r}) = \frac{\hbar}{m} \nabla \phi$$

Chemical potential:

$$\mu = \partial \phi / \partial t$$

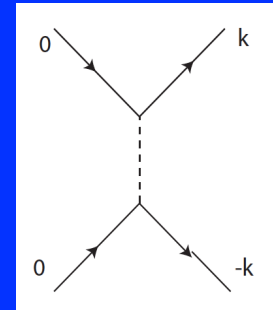
Superfluid acceleration eqn.:

$$\frac{\partial \vec{v}_s}{\partial t} + \nabla \mu = 0$$

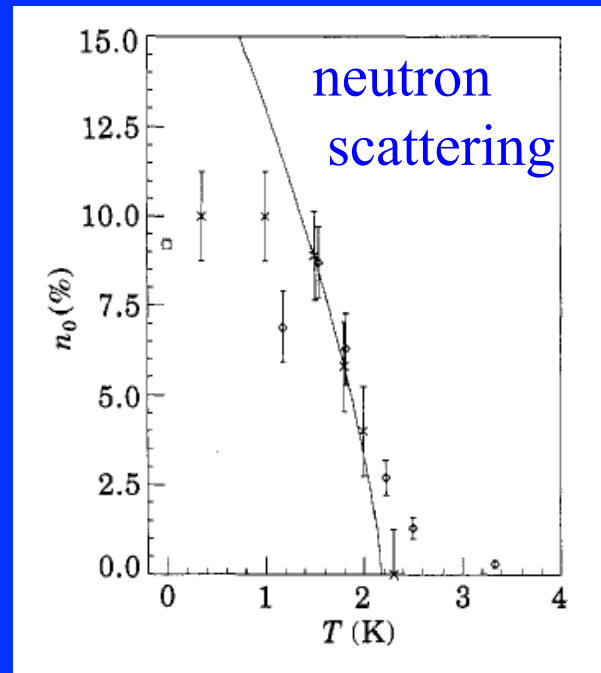
# Condensate density is NOT superfluid density

$$\rho_s \neq m|\psi|^2$$

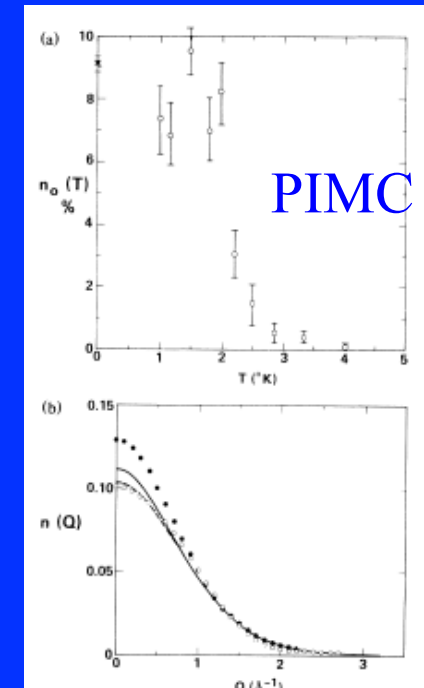
In ground state, interactions drive particles into non-zero momentum single particle states:



In  $^4\text{He}$  at  $T=0$ ,  
 $\rho_s/\rho = 1$ , while  
<10% of particles  
are in condensate



*Snow, Wang, & Sokol,  
Europhys. Lett. 19 (1992)*



*Ceperley & Pollock,  
PRL 56 (1986)*

# Order parameter of BCS paired fermions

Paired seen in amplitude to remove a pair of fermions ( $\uparrow\downarrow$ ) then add pair back, and come back to same state:

$$\langle \psi_{\uparrow}^{\dagger}(1)\psi_{\downarrow}^{\dagger}(2)\psi_{\downarrow}(3)\psi_{\uparrow}(4) \rangle \simeq \langle \psi_{\uparrow}^{\dagger}(1)\psi_{\downarrow}^{\dagger}(2) \rangle \langle \psi_{\downarrow}(3)\psi_{\uparrow}(4) \rangle$$

[Cf.,  $\langle \psi(\vec{r})\psi^{\dagger}(\vec{r}') \rangle \rightarrow \Psi(\vec{r})\Psi(\vec{r}')^*$  in Bose system]

Order parameter  $\langle \psi_{\downarrow}(r)\psi_{\uparrow}(r) \rangle \rightarrow \Psi(r)$ , as in Bose system

Similar physics as in Bose system

$$\Psi(\vec{r}) = |\psi| e^{i\phi(\vec{r})}$$

Supercurrent velocity:

$$\vec{v}_s(\vec{r}) = \frac{\hbar}{m} \nabla \phi$$

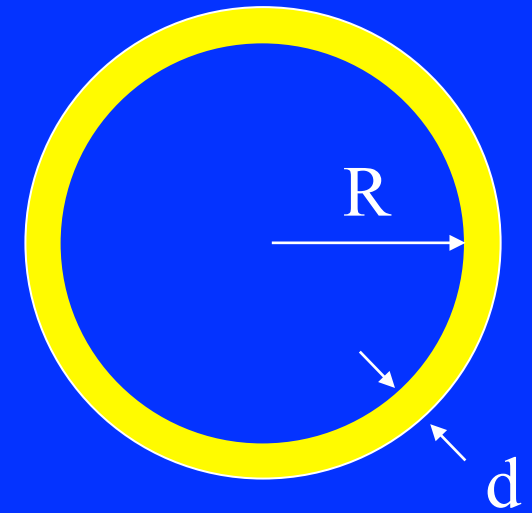
Chemical potential:

$$\mu = -\partial\phi/\partial t$$

# (Meta)stability of superfluid flow

Bosons of density  $n$  in annulus,  $T=0$

$$H = \frac{\hbar^2}{2m} \nabla \psi^* \cdot \nabla \psi + \frac{g}{2} \psi^* \psi^* \psi \psi$$



Condensates:  $\psi_0 = \sqrt{n}$  at rest

$\psi_1 = e^{i\phi} \sqrt{n}$  single vortex

Can one slip continuously from single vortex state to rest state, via condensate  $\psi = a\psi_0 + b\psi_1$  with  $|a|^2 + |b|^2 = 1$  ?

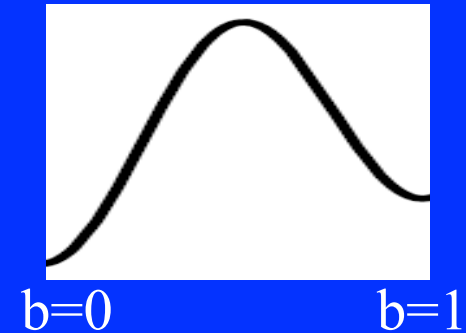
Energy density:  $E/V = \frac{\hbar^2}{2mR^2} n |b|^2 + \frac{g}{2} n^2 (|a|^4 + |b|^4 + 4|a|^2 |b|^2)$

$$= \frac{\hbar^2}{2mR^2} n |b|^2 + \frac{g}{2} n^2 + 2gn^2 |b|^2 (1 - |b|^2)$$

Energy density:

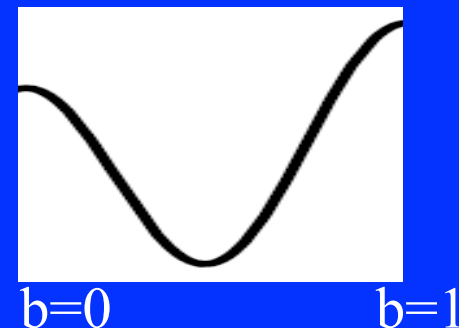
$$= \frac{\hbar^2}{2mR^2} n |b|^2 + \frac{g}{2} n^2 + 2gn^2 |b|^2 (1 - |b|^2)$$

$g > 0$ : have barrier of height  $\sim gnN$   
for  $0 < |b|^2 < 1$



Superfluid flow state (vortex) with  $b=1$  is metastable

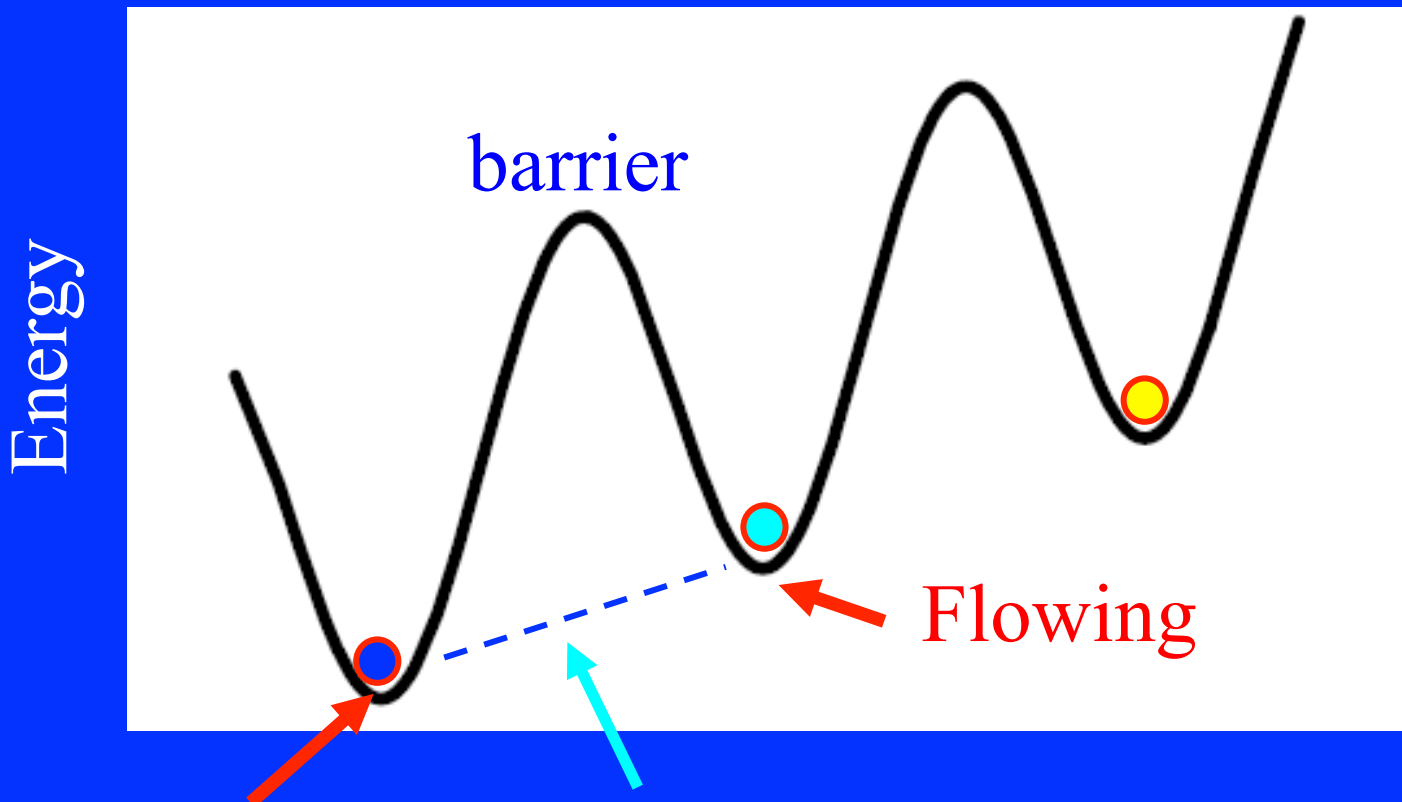
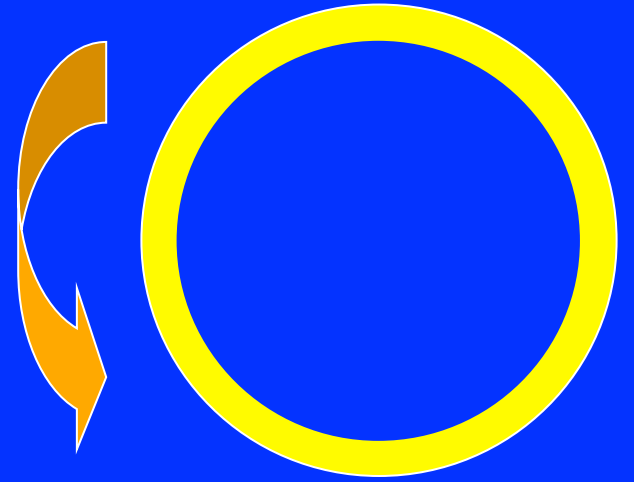
But for  $g < 0$  have minimum



Flow is unstable!

Cf. H atom decaying from 2p to 1s state, emitting energy

Superfluid flow difficult to stop  
because of enormous energy  
barrier, a hill:



No flow

Normally (not super), roll down with no barrier,  
from flowing state to resting

# Condensate density and superfluid mass density

*M. Holzmann & GB, Phys. Rev. B 76, 092502 (2007)*

For a superfluid flowing down a pipe (at rest) with superfluid velocity  $v_s$  in  $z$  direction

$$F(v_s, T, \mu) = F(0, T, \mu) + \frac{1}{2} \rho_s v_s^2$$

$$\partial F / \partial v_s = - \langle P_z - M v_s \rangle / V$$

where  $P_z$  is the total momentum.

$$\partial^2 F / \partial v_s^2 = \rho - \beta \langle P_z^2 \rangle / V,$$

so that

$$\rho_n = \beta \langle P_z^2 \rangle / V.$$

In normal state, total momentum is Gaussianly distributed:

$$\propto \exp(-\beta P^2 / 2M) \quad \text{so that} \quad \beta \langle P_z^2 \rangle / V = M / V \quad \text{and} \quad \rho_n = \rho$$

In superfluid phase, the total momentum and  $v_s$  are entangled and total momentum distribution is not classical.



# Exact relation between $\rho_s$ and the condensate density, via the single particle Green's function

*P.C. Hohenberg & P.C. Martin, PRL 22 (1963); B.D. B.D. Josephson, PL21 (1966); GB, St. Andrews lectures (1967), A. Griffin PR B30 (1984)*

$$G(k, z) = -i \langle T(\psi^\dagger \psi) \rangle(k, z) \quad z = \text{complex frequency}$$

$$\rho_s = - \lim_{k \rightarrow 0} \frac{n_0 m^2}{k^2 G(k, 0)}$$

Ex. in Bogoliubov mean field ( $n_0=n$ ),

$$G(k, z) = \frac{z + gn + k^2/2m}{z^2 - gnk^2/m + k^4/4m^2}$$

$$\Rightarrow \rho_s = nm$$

Valid in 2D as well as 3D:

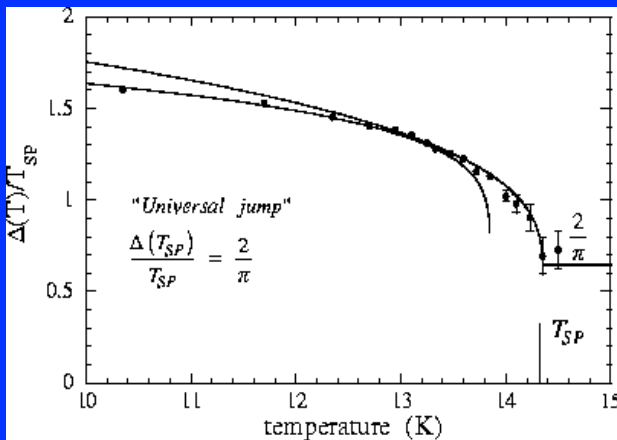
*M. Holzmann & GB, PR B 76 (2007);  
M. Holzmann, GB, J-P Blaizot,  
& F Laloë, PNAS 104 (2007)*

In 2D finite size Berestetskii-Kosterlitz-Thouless system,

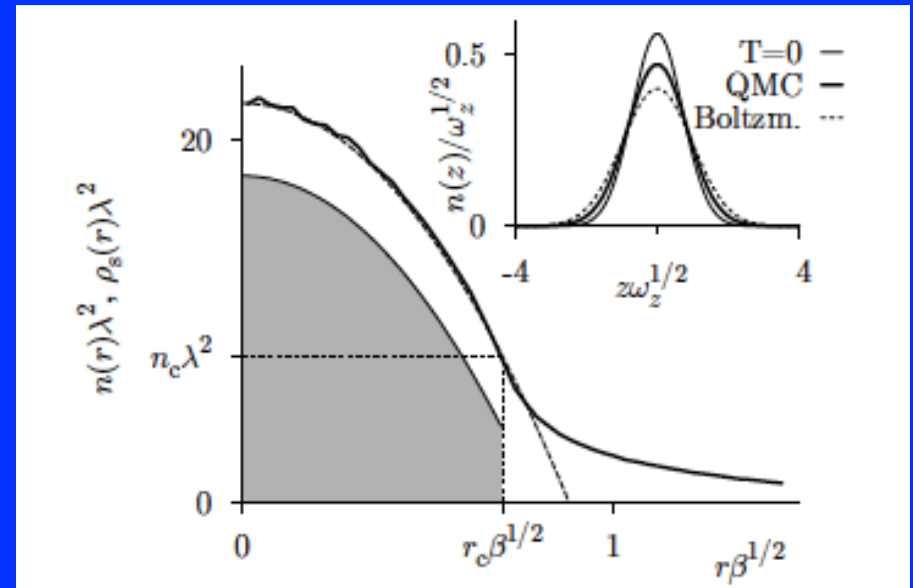
$$n_0 \sim 1/(\text{size})^{2-\eta}$$

$$G(k, 0) \sim 1/k^{2-\eta}$$

At  $T_c$   $\rho_s = \frac{m^2 T}{2\pi\eta}$   $\eta = 1/4$



Order parameter in CuGeO<sub>3</sub>  
*Lorenzo et al., EPL. 45 (1999)*



Density profile in 2D trap.  
Shaded region  $\Leftrightarrow \rho_s$

*Holtzmann & Krauth, EPL 82 (2008)*

# Exact (and equivalent) definition of $\rho_s$ in terms of current-current correlation functions

$$Y_{ij}(\mathbf{r}, \mathbf{r}', \omega) = \int dt e^{i\omega(t-t')} \langle [j_i(\mathbf{r}, t), j_j(\mathbf{r}', t')] \rangle$$

$$\mathbf{j}(\mathbf{r}) = \frac{1}{2im} [\psi^\dagger(\mathbf{r}) \nabla \psi(\mathbf{r}) - \nabla \psi^\dagger(\mathbf{r}) \psi(\mathbf{r})]$$

Decompose into longitudinal and transverse components:

$$Y_{ij}(\mathbf{k}, \omega) = \frac{k_i k_j}{k^2} Y_L(k, \omega) + \left( \delta_{ij} - \frac{k_i k_j}{k^2} \right) Y_T(k, \omega).$$

f-sum rule  $\Rightarrow$

$$\rho = \lim_{k \rightarrow 0} m^2 \int \frac{d\omega}{2\pi} \frac{Y_L(k, \omega)}{\omega}$$

motion in tube with  
closed ends

Define normal mass density

$$\rho_n = \lim_{k \rightarrow 0} m^2 \int \frac{d\omega}{2\pi} \frac{Y_T(k, \omega)}{\omega}$$

motion in tube with open ends



$$\rho_s = \rho - \rho_n$$

## How does it work?

$$\text{At } T=0, \quad \int \frac{d\omega}{2\pi} \frac{\Upsilon(k, \omega)}{\omega} = 2 \sum_{a \neq 0} \frac{|\langle a | j_q | 0 \rangle|^2}{E_a - E_0}$$

In general, for low-lying states,

$$\langle a | j_k | 0 \rangle_L \sim k^{1/2} \quad E_a - E_0 \sim k$$

and  $\rho_n = \rho$ . Same for transverse in normal .

In BEC,  $\langle \text{phonon} | \vec{j}_k | 0 \rangle \sim k^{1/2} \hat{k}$  .

Thus  $\langle \text{phonon} | \vec{j}_k | 0 \rangle_T = 0$  and  $\rho_n = 0$ .

In superconductor with gap, matrix elements vanish in long wavelength limit, while denominators in T integral remain finite and  $\rho_n = 0$ .

# Meissner effect

The  $\langle jj \rangle_T$  xx correlation gives fundamental description of the Meissner effect as well:

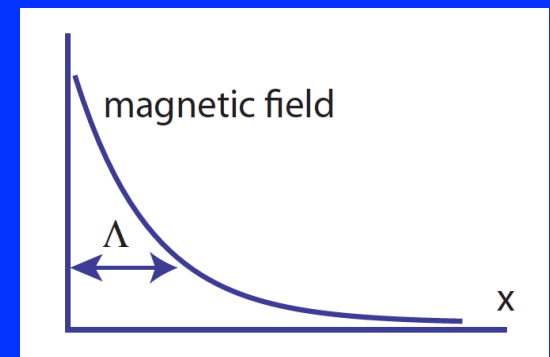
Penetration depth  $\Lambda$  in superconductor

$\Leftrightarrow$  screening of magnetic field

$$\frac{1}{\Lambda^2} = \frac{4\pi n e^2 \rho_s}{m c^2 \rho}$$

$$A_{\text{tot}}(\mathbf{k}) = \frac{k^2}{k^2 + (1/\Lambda^2)} A(\mathbf{k})$$

$\Rightarrow$



# Moment of inertia of superfluid

$$I = \frac{\rho_n}{\rho} I_{\text{classical}}$$

In terms of  $\langle jj \rangle$ :

$$\mathcal{I}_{ij} = \int \frac{d\mathbf{k}}{(2\pi)^3} \int d\mathbf{r} \int d\mathbf{r}' e^{i\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')} \varepsilon_{isl} \varepsilon_{jmn} r'_s r'_m \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{\Upsilon_{ln}(\mathbf{k}\omega)}{\omega}$$

# Landau calculation of $\rho_s$ for system w. quasiparticles

$$\langle \mathbf{P} \rangle = \sum_{\mathbf{p}} \mathbf{p} \langle N_{\mathbf{p}} \rangle$$

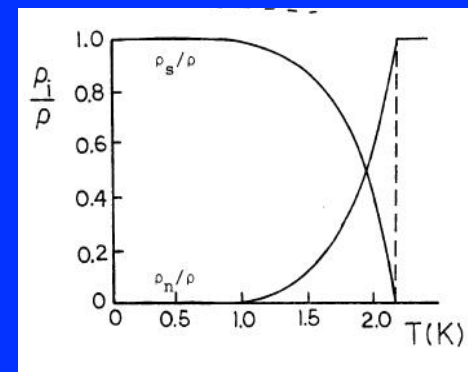
$$\langle N_{\mathbf{p}} \rangle = [e^{\beta(\epsilon_{\mathbf{p}} + \mathbf{p} \cdot (\mathbf{v}_s - \mathbf{v}_n))} - 1]^{-1}$$

$$\langle \mathbf{P} \rangle = - \sum_{\mathbf{p}} \mathbf{p} (\mathbf{p} \cdot \mathbf{v}_n) \frac{\partial}{\partial \epsilon_{\mathbf{p}}} \frac{1}{e^{\beta \epsilon_{\mathbf{p}}} - 1}$$

$$\rho_n = - \int \frac{d\mathbf{p}}{(2\pi)^3} \frac{p^2}{3} \frac{\partial}{\partial \epsilon_{\mathbf{p}}} \frac{1}{e^{\beta \epsilon_{\mathbf{p}}} - 1}$$

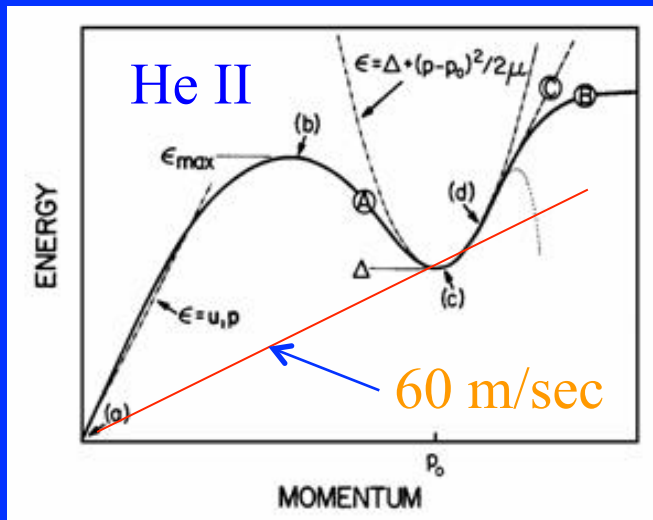
for phonons:

$$\rho_n(T) = \frac{2\pi^2}{45\hbar^3 s^3} T^4$$



# The Landau criterion for superfluidity

Superfluid with elementary excitation spectrum  $\epsilon(q)$



Fluid flowing in pipe in x direction, velocity  $v$  with respect to walls. In wall frame excitation energy is

$$\epsilon_v(q) = \epsilon(q) + vq_x$$

According to Landau:

For  $v < \epsilon(q)/q$  cannot make spontaneous excitations (which would decay superflow) and *flow is superfluid*.

For  $v$  opposite to  $q_x$  and  $v > \epsilon(q)/q$  have  $\epsilon_v(q) < 0$

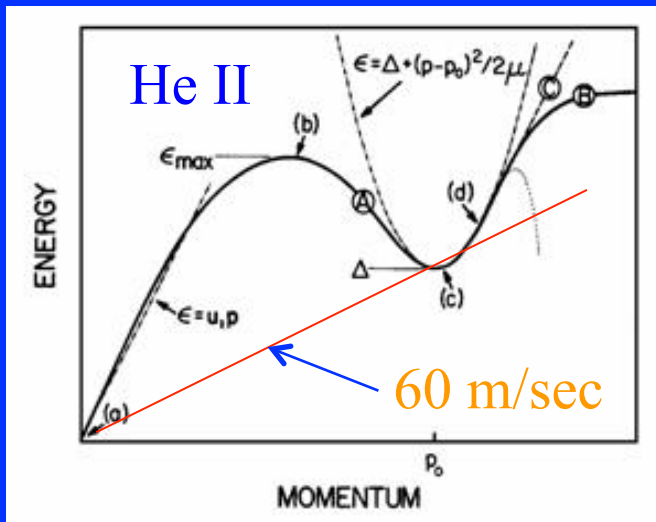
Can then make excitations spontaneously, and

*superfluidity ceases.*  $v_{crit} = 60$  m/sec in superfluid He.

In this way we see that neither phonons nor rotons can be excited if the velocity of flow in helium II is not too large. This means that the flow of the liquid does not slow down, i.e. helium II discloses the phenomenon of superfluidity †. **SUFFICIENT**

It must be remarked that already the reasons given above are enough to make the superfluidity vanish at sufficiently large velocities. We leave aside the question as to whether superfluidity disappears at smaller velocities for some other reason (the velocity limit obtained from (4.2) is large—the velocity of sound in helium equals 250 m/sec; (4.4) gives a value only several times lower). **NECESSARY**

*L.D. Landau, J. Phys. USSR 5, 71 (1941)*



At Landau critical velocity, group and phase velocity of excitations are equal:

$$\frac{\partial \epsilon}{\partial q} = \frac{\epsilon}{q}$$



# The Landau criterion is neither necessary nor sufficient

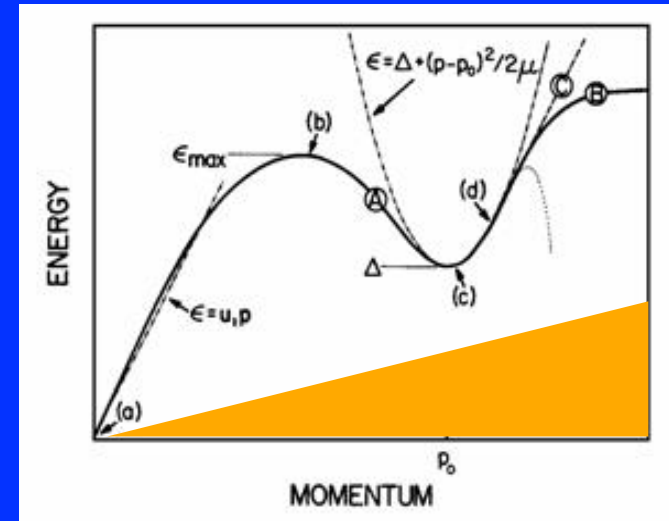
Superfluid systems with no “gap”:

1) Dilute solutions of degenerate  $^3\text{He}$  in superfluid  $^4\text{He}$ :

Particle-hole spectrum

$$\omega = (\vec{p} + \vec{q})^2 / 2m - \vec{p}^2 / 2m$$

reaches down to  $\omega = 0$  at  $q \neq 0$ .



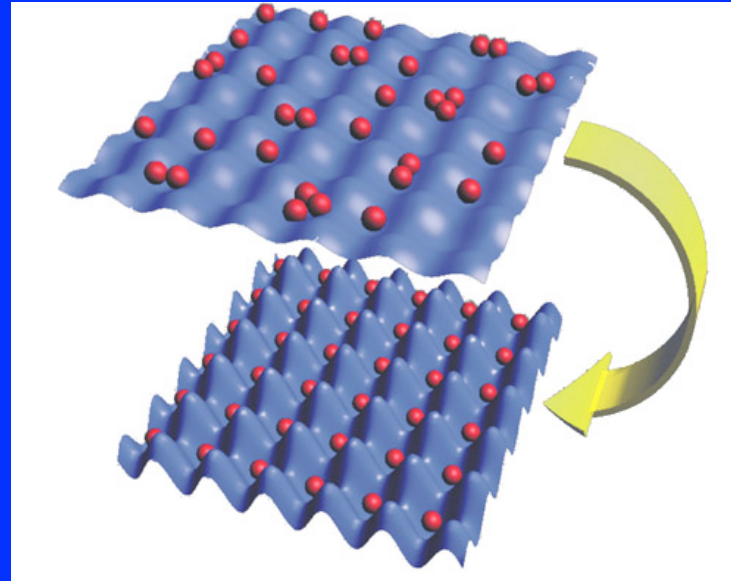
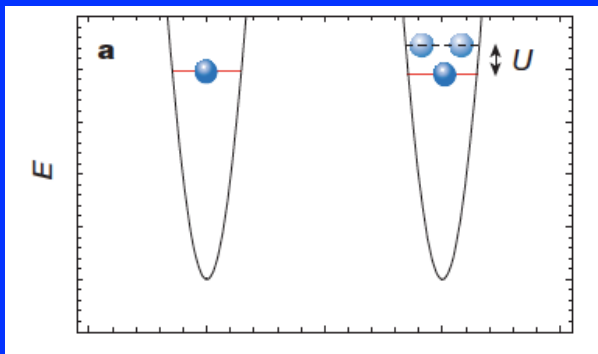
Landau critical velocity vanishes, but system is superfluid.

2) Superfluid  $^4\text{He}$  at non-zero temperature: Can scatter a phonon of momentum  $k$  to  $-k$  with zero energy change.

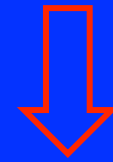
Again Landau critical velocity vanishes, but system is a perfectly good superfluid.

# Gap also not sufficient to guarantee superfluidity:

ex. bosons in optical lattice:

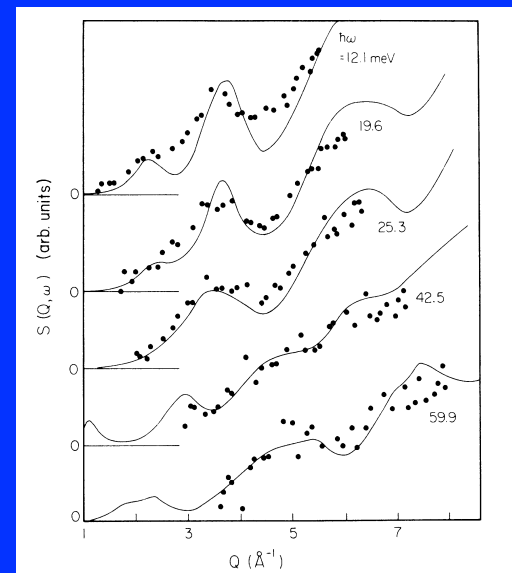


superfluid



Mott insulator

Amorphous solids, e.g., Si doped with H, not superfluid.



## What happens when the Landau criterion is violated?

The superfluid mass density becomes less than the total mass density. It does not necessarily vanish!

In dilute solutions of  $^3\text{He}$  in superfluid  $^4\text{He}$ ,

$$\rho_s = \rho - (m^* - m_3)n_3$$

$m^*$  =  $^3\text{He}$  effective mass,  $m_3$  = bare mass.

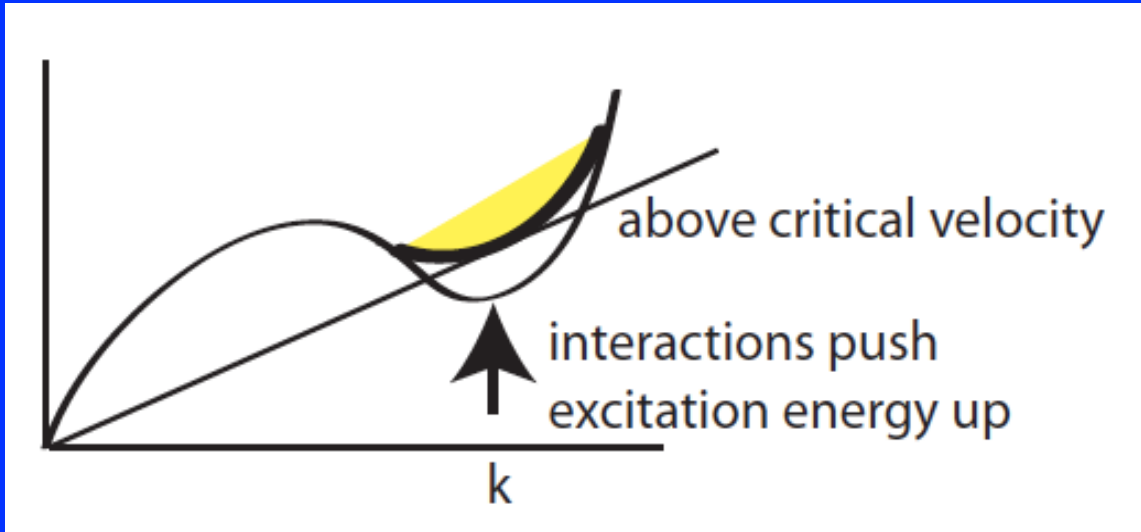
In superfluid  $^4\text{He}$  at nonzero temperature,

$$\rho_s = \rho - aT^4 - \dots$$

Formation of non-uniform states

# Formation of non-uniform states

*L. Pitaevskii, JLTP 87, 127 (1992), GB & CJ Pethick*



Beyond the critical velocity  
spontaneously form  
excitations of finite  
momentum  $k$ .

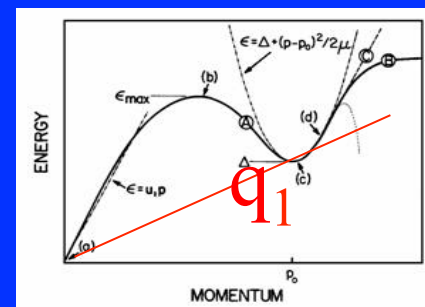
Interactions of these excitations, when repulsive, raises energy of unstable mode  $k$ , to make velocity just critical.

Mixing in of modes of momentum  $k$  causes condensate to become non-uniform.

# Simple model when Landau critical velocity is exceeded

(GB and CJ Pethick)

Weakly interacting Bose gas with finite range interaction  $g(r)$ , and thus  $g(q) (> 0)$ , to produce  $v_{\text{crit}}$  at non-zero  $q = q_1$



In Bogoliubov approx:

$$\varepsilon(q) = \left[ \frac{ng(q)q^2}{m} + \left( \frac{q^2}{2m} \right)^2 \right]^{1/2}$$

Critical point: where group velocity = phase velocity,

$$\frac{d(gn)}{d(q^2/2m)} = -1,$$

Study stability of uniform condensate for  $q > q_1$

$$\psi(x) = \sqrt{n} e^{iqx}$$

Since excitations at  $q_1$  can form spontaneously, generate new condensate of form

$$\psi(x) = e^{iqx} [\sqrt{n_0} + ue^{-iq_1x} + ve^{iq_1x}]$$

Let  $u = \zeta \cosh(\phi/2), \quad v = \zeta \sinh(\phi/2)$

Stable solution  $\zeta^2 = (v - v_{crit}) \frac{q_1}{2G}$  above critical velocity

$$G = \frac{g_1}{4\epsilon_1^2} \left[ g_1 g_2 n^2 - \frac{q_1^2}{m} \left( \frac{q_1^2}{m} + 2g_1 n \right) \right] = \text{effective repulsion of excitations near } q_1$$

$$g_\kappa \equiv g(\kappa q_1)$$

Non uniform density:  $n(x) = n + 2\sqrt{n_0}\zeta \cosh \phi \cos(q_1x)$

Reduction of superfluid mass density

$$\rho_s = mn - mq_1 \frac{\partial \zeta^2}{\partial q} = mn - \frac{q_1^2}{2G}$$

THANK YOU