Status of the Forward Elastic Scattering Luminosity Monitor

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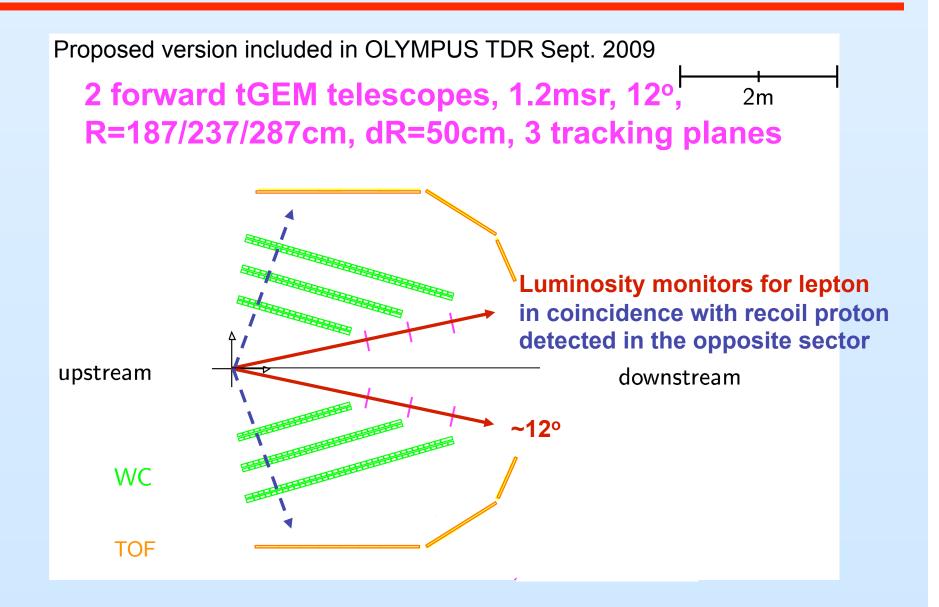
Proposed Experiment

- Electrons/positrons (100mA) in multi-GeV storage ring DORIS at DESY, Hamburg, Germany
- Unpolarized internal hydrogen target (buffer system)
 3x10¹⁵ at/cm² @ 100 mA → L = 2x10³³ / (cm²s)
- Large acceptance detector for e-p in coincidence BLAST detector from MIT-Bates available
- Measure ratio of positron-proton to electron-proton unpolarized elastic scattering to 1% stat.+sys.
- Redundant monitoring of (relative) luminosity (ratios) to ~1% per hour

Monitoring the Luminosities

- Pressure, temperature, flow, current measurements
 - limited in precision
- Moller/Bhabha scattering
 - 1.3 degrees for symmetric Moller setup
 - requires knowledge of Moller/Bhabha cross sections
- Small-angle elastic scattering
 - high count rate, no TPE at high epsilon / low Q²
 - single-arm and in coincidence with recoil proton
 - event-by-event with full track reconstruction
 - different acceptances for e⁺ and e⁻
 - requires knowledge of angular xsec. dependence

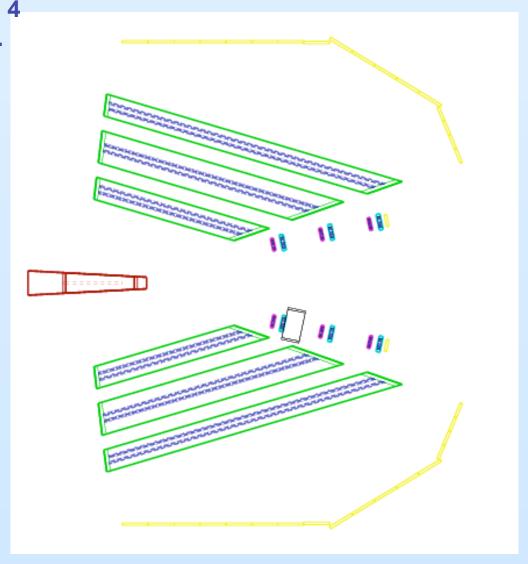
Luminosity Monitors: GEM Telescopes



Luminosity Monitors: GEM+MWPC

 Concept paper for TDR, Sec. 4 sent out to OLYMPUS collab.

- Two MWPC telescopes with 3 x/u/v elements after each GEM element
- Scintillator for triggering and timing
- High redundancy No interference Two independent groups
- Well suited for efficiency and alignment calibration
- Both GEM and MWPC telescopes can operate independently



Luminosity Monitors: GEM+MWPC

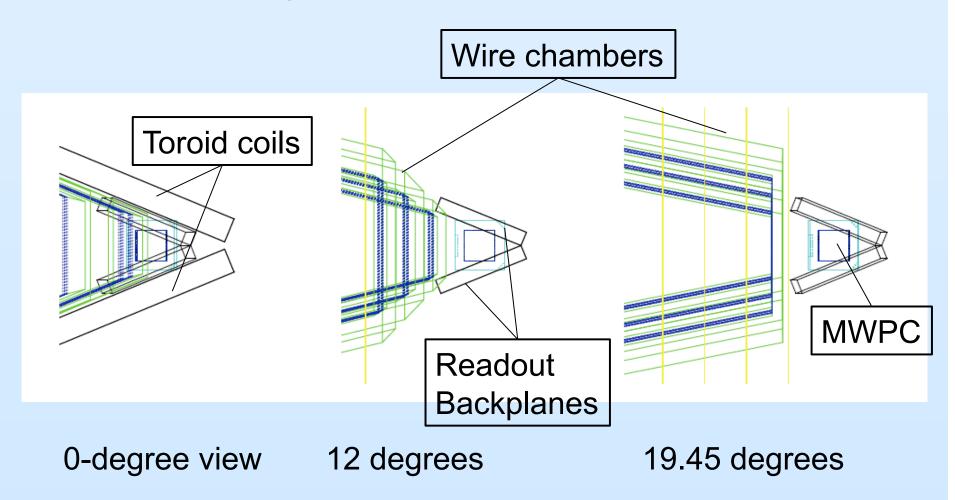
	$\delta z [\mathrm{cm}]$	$\delta \theta$	$\delta\phi$	$\delta p/p \ [\%]$
GEMs alone	1.67	0.137°	0.301°	4.57
GEMs alone + $MWPC$ material	1.74	0.142°	0.309°	4.86
MWPCs alone + GEM material	3.94	0.294°	0.407°	9.58
GEMs + MWPCs	1.60	0.132°	0.308°	4.43

Table 1: Obtained track resolutions for various configurations of the luminosity monitor, based on a GEANT3 simulation with Kalman filter technique for reconstruction.

- Concept paper for TDR, Sec. 4 sent out to OLYMPUS collaboration
- Each system alone is not deteriorated too much by the other
- Both systems combined give superior resolution

MWPC: Confined space ... but they fit

MWPC closest to target at 200 cm



Luminosity Monitors – Basic Properties

Proposed version included in OLYMPUS TDR Sept. 2009

E_0	Q^2	$p_{e'}$	ϵ	θ_p	p_p	Rate
[GeV]	$[(\mathrm{GeV/c})^2]$	$[\mathrm{GeV/c}]$			$[\mathrm{MeV/c}]$	$[h^{-1}]$
4.5	0.801	4.073	0.9736	58.7°	992	1846
2.0	0.167	1.911	0.9774	71.8°	418	49792

Table 4.1: Kinematics and count rates of the luminosity control measurement for beam energies of 2.0 and 4.5 GeV at $\theta_e = 12^{\circ}$. The assumed solid angle is 1.2 msr determined by the area of rearmost tracking plane farthest from the target.

- Two symmetric GEM telescopes at 12°
- Two-photon effect negligible at high-ε / low-Q²
- Sub-percent (relative) luminosity measurement per hour at 2.0 GeV, per day at 4.5 GeV
- 1.2 msr = $10 \times 10 \text{ cm}^2$ at ~290 cm distance (rearmost plane)
- Three GEM layers with ~0.1 mm resolution with ~50 cm gaps
 → Match vertex resolution (z) of ~ 0.1 1 cm at 12° with proton in BLAST
- Issue of acceptance loss and different average angle due to track bending can be alleviated with reduced toroid field

Luminosity Monitor based on GEMs

- Forward-angle electron/positron telescopes with good angular and vertex resolution
- Two telescopes with 3 triple-GEM detectors, left-right symmetric
- Coincidence with proton in opposite sector of main detector
- Single-arm tracks
- High rate capability
- Readout based on APV chip

MIT protoype:

Telescope of 3 Triple GEM prototypes (10 x 10 cm²) using TechEtch foils

F. Simon et al., NIM A598 (2009) 432



Providing GEM detectors

Collaboration HU-MIT-Rome

- TechEtch/MIT to provide GEM foils
- Assembly of detectors at MIT and MIT-Bates
- Rome: Readout system developed for Hall A / SBS
- Testing at HU / HUPTI / Jlab / J-PARC or DESY

Funding

- Secured NSF Nuclear Physics/ARRA basic research grant (\$405k) postdoc + travel; 1 graduate student supported by HU NSF group 2 undergraduate stipends for summer 2010
- Secured \$216k within NSF MRI-R2 for luminosity monitors (\$125k),
 1 graduate stipend + travel for commissioning

Tasks & Timeline for LuMo Construction

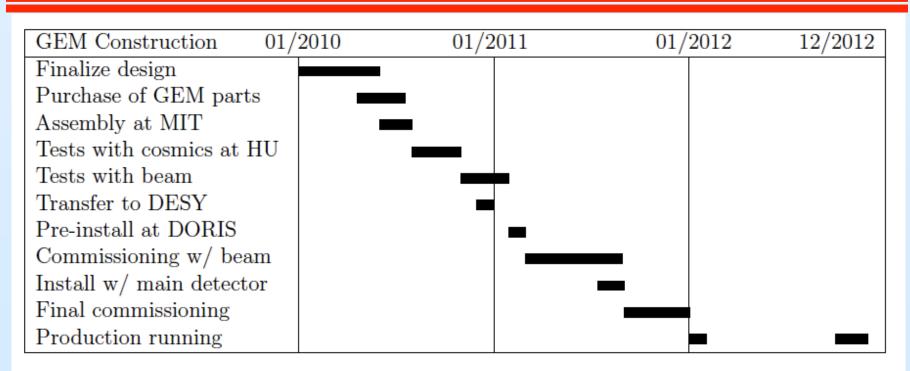


Table 4: Work plan for GEM luminosity monitor construction activities.

- Assembly of GEM detectors at MIT in summer 2010
- New research building at HU with lab space available by summer 2010
 - Central supply of (non-flammable) gases
 - Clean room on HU campus from ATLAS group available if necessary

Realization of Detectors

- Construction project fully funded (NSF, MRI-R2), grants are active
- Postdoc position possibly filled by J. Diefenbach eff. July 1, 2010
- Purchase of GEM and HV foils initiated, designed by D. Hasell,
 5-6 weeks delivery (early June)
- HU-MIT Research Agreement in preparation to enable Bates engineering services:
 - Drawings, direct ordering of parts, technical support for manufacturing
 - Parts to be available by June/July
- HU group to visit MIT and Bates from June 7 August 8, 2010
 - Two HU undergraduate students (Miles Campbell, Joshua McMahon)
 - One HU graduate student (Ozgur Ates); one postdoc (J.D., t.b.c.); and M.K.
 - Gluing and testing of GEM foils (Optical checks; HV conditioning)
 - Assembly of GEM detectors at MIT Bates
 - Some initial testing
- Have detectors constructed by August, transfer to HU

Realization of Detectors (cont'd)

- Prepare cosmic ray test stand at HU
 - Collaboration with M.E. Christy (HU)
 - Scintillators, eventually with SOS wire chambers for good track definition
- Implement test readout in August 2010 (Rome group)
 - New APV based system if ready
 - "Gassiplex" system (Jlab) as fallback for cosmic ray and source tests
- Testing with cosmic rays / sources from August October 2010
 - Signal to noise studies, cluster performance
 - Study gains and charge sharing
 - Operation stability, discharges
 - Efficiency versus voltage
 - Efficiency distributions versus x and y
 - Spatial resolution

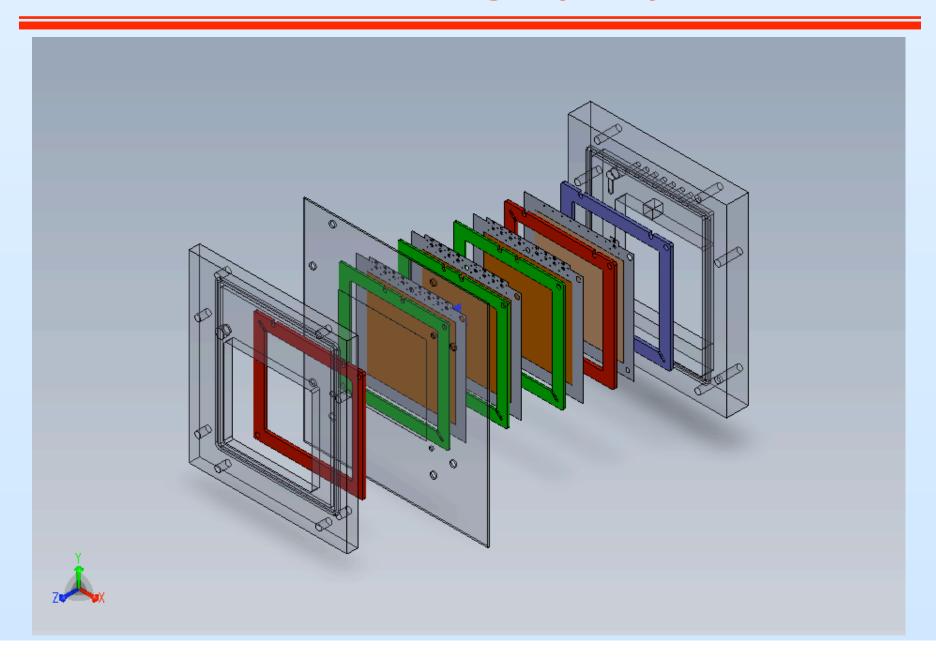
Realization of Detectors (cont'd)

- Testing with beam between October 2010 January 2011
 - HUPTI / 100 MeV protons with low intensity (initial proton track may be known)
 - Jlab / Qweak parasitically (scatt. electron tracks known if comb. w. Qweak)
 - Kaon 1.1BR commissioning at J-PARC (Oct-Nov. 2010)
 - DESY testbeam (tracks well known)
 - Or no separate beam testing at all? Decide by August
- Transfer detectors to DESY between Oct. 2010 and Jan. 2011
- Mount GEMs in DORIS in Feb. 2011 for spring test run
- Final installation in OLYMPUS in summer 2011

GEM Design Criteria / Issues

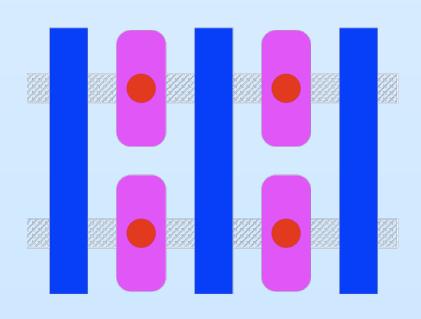
- Want 12.5x12.5 cm² total area gluing instead of bolting w/ O-rings
- Identical design for all chambers preferred easy replacements
- Can use inner frame design of new MIT prototype (D. Hasell) with minor modifications (like glue troughs), for everything but readout
- Use top and bottom spacer frames for gas pressure foils
- Access all infrastructure items voltage supply, gas connectors, readout connectors through top frame margin

GEM Detector Design (MIT)

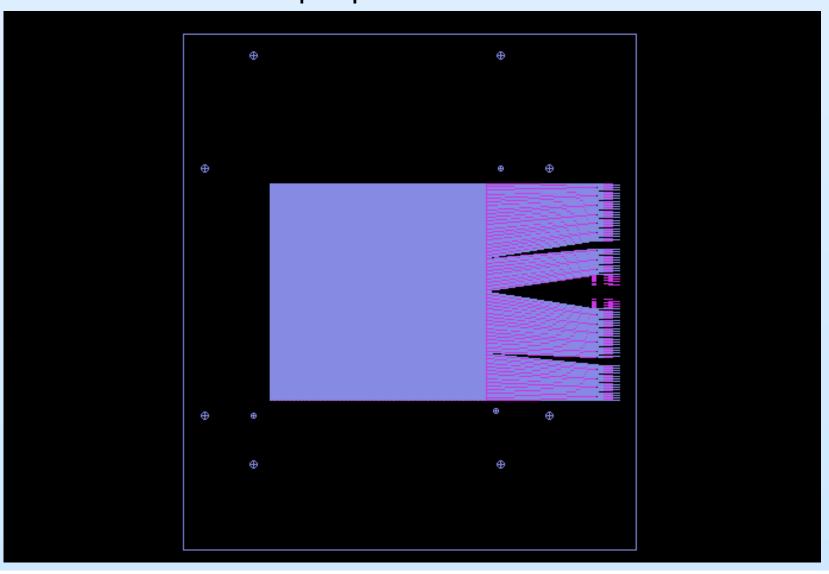


GEM Readout Board

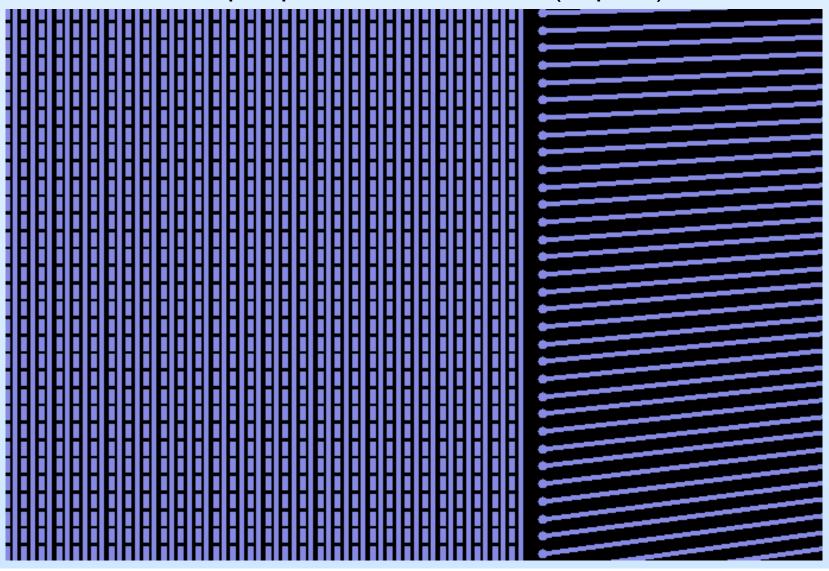
- Readout pitch of MIT prototype: 635 (800) μm, COMPASS: 400 μm
- Number of electronics channels per chamber: [100+100]/0.635(0.800,0.400) ~= 315 (250, 500), read out by 3(2, 4) APV chips
- Want 12.5x12.5 cm² total area
- Issue to be decided fast: Readout board design
 - CERN design (R. de Oliveira), based on laser ablation technique, "standard" size too large (>18x18). Is used by INFN&UVA at PREX
 - CERN can produce any other design based on Gerber file
 - Laser ablation technique @ MIT (Compunetics) discontinued
 - Readout with strips and pads vias-connected on rear side
 - Is currently still in R&D phase
 - Existing design for 10x10 cm²
 with 800 μm pitch



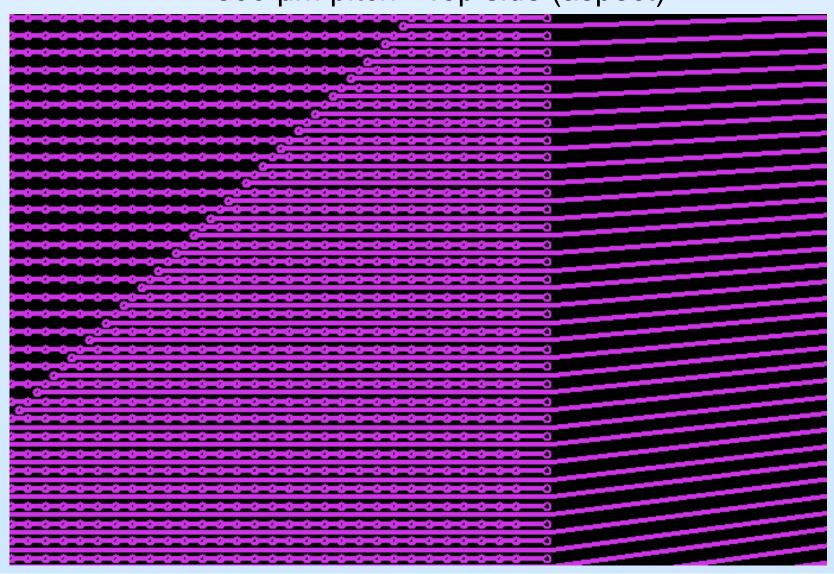
800 µm pitch - Total view



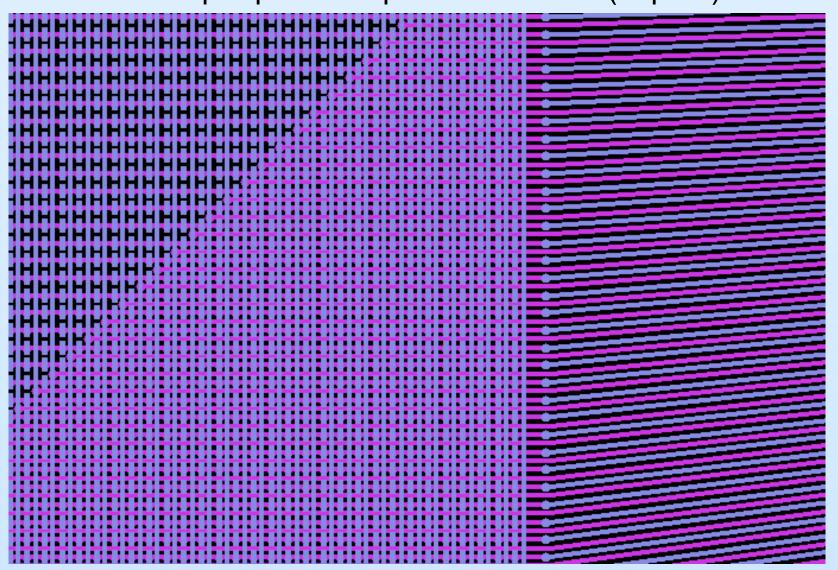
800 µm pitch - Bottom side (aspect)



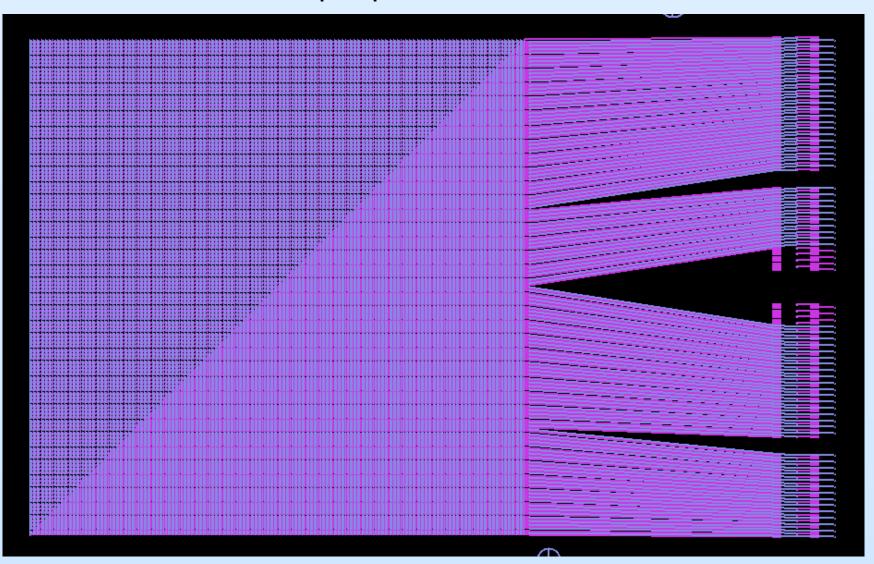
800 µm pitch - Top side (aspect)



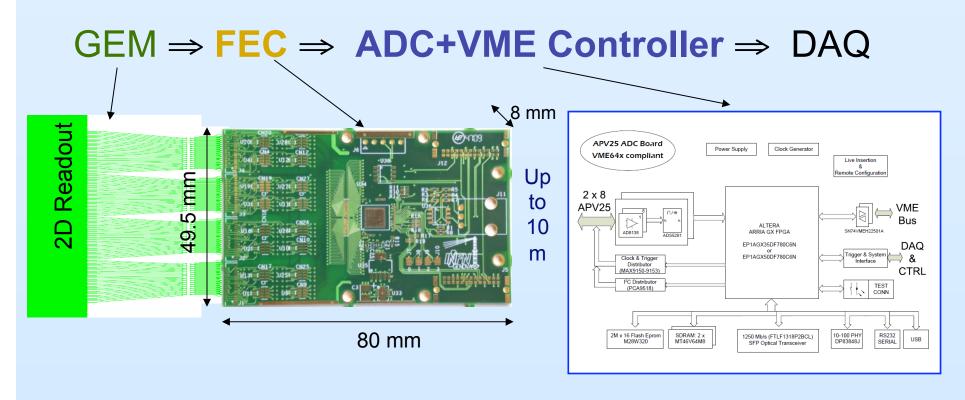
800 µm pitch - Top+bottom sides (aspect)



800 µm pitch - Total view



Readout Electronics



- Frontend card (APV + VME) by INFN Rome → see next talk, by S. Frullani
- First test of 10x10 GEMs from INFN and UVa at PREX/Hall A in March 2010
- Second test, with new APV readout toward end of PREX in May 2010
- System to become available for OLYMPUS GEMs in August 2010

Backup slides

Luminosity Monitor MC Tasks

- Continue to work on GEANT4 simulation (Ozgur Ates, HU graduate student of HU nuclear physics group)
 - use new, faster codes with better bookkeeping features
 - higher MC statistics
- Simulations of phase space integral(s) (acceptance) expected counts, acceptance-averaged cross section
- Study of systematic effects (beam offset, slope, width; etc.) on counts per bin
- Simulation of backgrounds (Moller/Bhabha; Inelastics)

Luminosity Monitors: Cost estimate

Proposed version included in OLYMPUS TDR Sept. 2009

Item	Amount	Cont.(%)	Total/k\$	Remarks
Support frame	3	20	1.8	\$500/frame (2+1 spare)
GEM chamber mechanics	9	20	10.8	1000/chamber (6+3)
GEM foils $10 \times 10 \text{ cm}^2$	40	20	9.6	\$200/GEM foil (27+13)
Readout layer	9	20	21.6	\$2000/board (6+3)
Hybrids	80	20	19.2	\$200/hybrid (54+26)
APV25 chips	80	20	2.4	\$25/chip (54+26)
Cables	18	20	2.2	Signal and HV (6+3)
FEE	2880	20	34.6	$10/\text{channel } (6+3) \times 320$
Readout system	1		5.0	
HV distribution	9	20	0.5	\$50/chamber
Power Supply		20	5.0	HV pods
Gas system	9	20	3.2	300/line (6+3)
Misc. items			9.1	
Total			125.0	

Table 4.3: Cost estimate for the OLYMPUS luminosity monitors based on two plus one spare forward-angle GEM telescopes, each based on three triple-GEM detectors.