

Minimizing Actuator-Induced Residual Error in Active Space Telescope Primary Mirrors

Aero/Astro Research Evaluation

Matthew W. Smith

m_smith@mit.edu

Advisor:

Prof. David W. Miller



- **Background:** Large telescopes
- **Motivation and Research Question:** Lightweight error-free mirrors
- **Literature Review**
- **Approach:** Parametric FE modeling
- **Results:** Rib shaping
- **Conclusion**

- Trend toward large aperture space telescopes
 - Higher resolving power
 - Increased light-gathering ability
 - Launch constraints drive new technology for segmented mirrors and *low areal density*

Previous generation:
Hubble Space Telescope (HST)



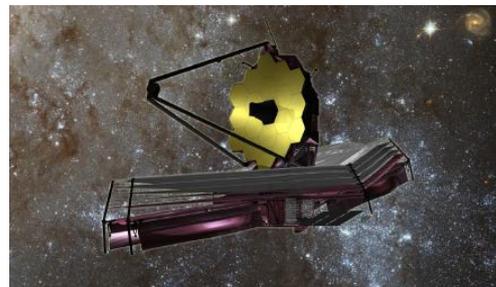
[1]



[1]

Material: Ultra low expansion (ULE) glass
 Areal density: 180 kg/m²
 Diameter: 2.4 m
 Shape control: passive
 Actuators: none

Current generation:
James Webb Space Telescope (JWST)



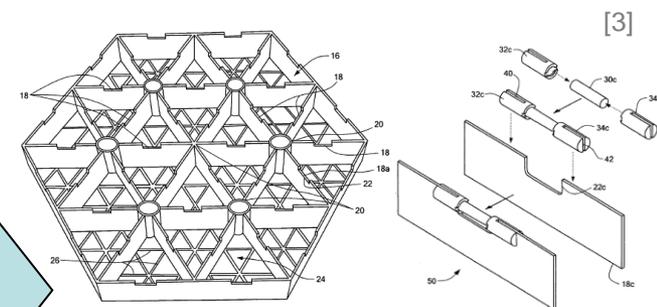
[1]



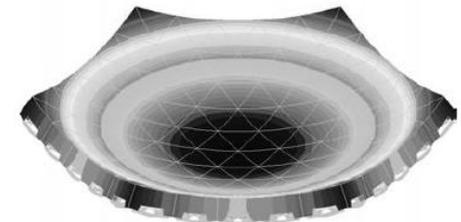
[2]

Material: Beryllium
 Areal density: ~30 kg/m²
 Diameter: 6.5 m (1.3 m segments)
 Shape control: active (7 DOF)
 Actuators: cryogenic stepper motors

Next generation:
“Highly Integrated” Mirror Design



[3]



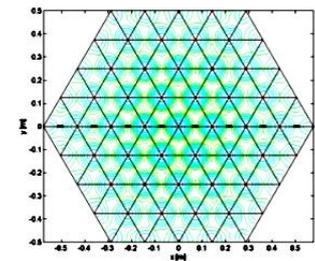
[4]

Material: Silicon Carbide (SiC)
 Areal density: <15 kg/m²
 Diameter: 1 m segments (baseline)
 Shape control: active (100+ DOF)
 Actuators: piezoelectric stacks

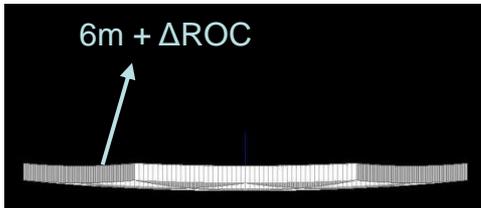


- Low areal density requires active shape control
 - Correct quasi-static disturbances (thermal)
 - Correct manufacturing errors (SiC casting process)
- Problem: actuator-induced residual error
 - Symptom of discrete actuators commanding low-order shapes
 - Degrades image sharpness

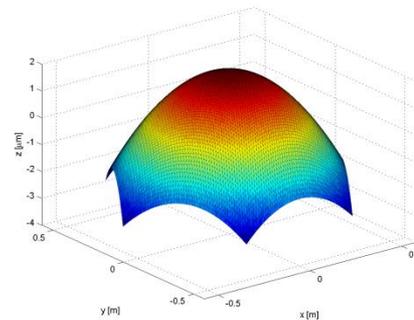
Actuator-induced residual error (actual – desired)



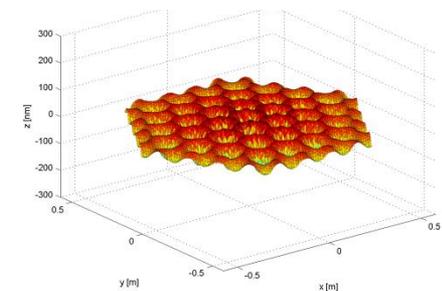
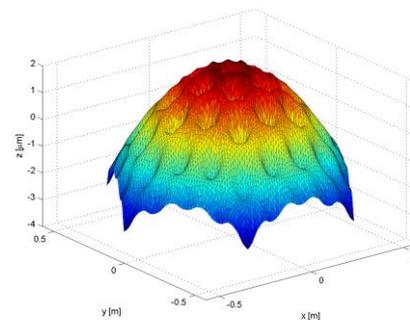
Un-actuated mirror



Desired surface change



Actual surface change

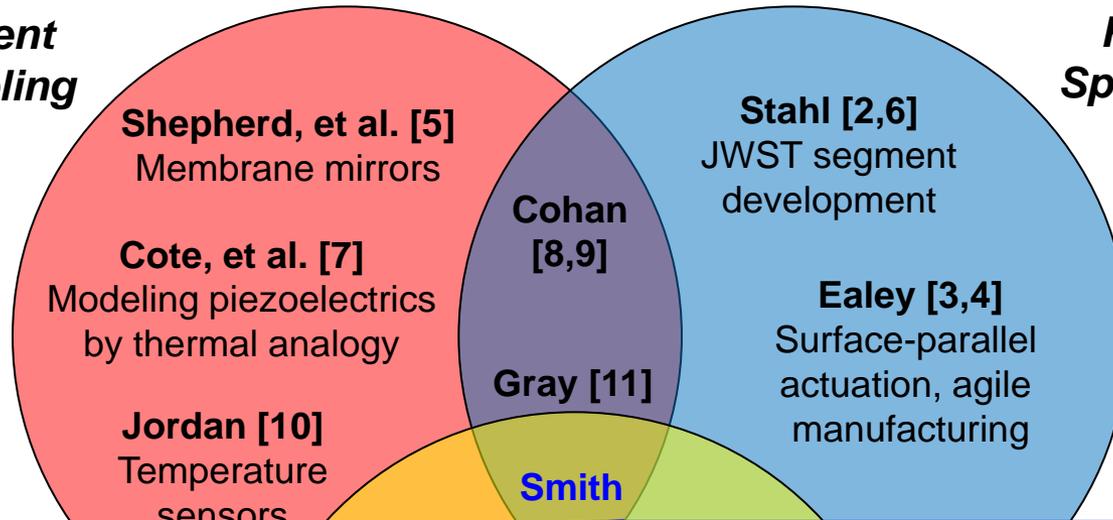


Research Objective

To reduce actuator-induced high spatial frequency residual error by taking advantage of changes in mirror geometry using a parametric finite element mirror model while keeping areal density (mass) and number of actuators (power, complexity) constant.

**Finite Element
Mirror Modeling**

**Rib-stiffened
Space Telescope
Mirrors**



Shepherd, et al. [5]
Membrane mirrors

Cote, et al. [7]
Modeling piezoelectrics
by thermal analogy

Jordan [10]
Temperature
sensors

**Budinoff
[12]**

Ealey [14]
Surface-
actuation

Bikkannavar, et al. [13]
Phase retrieval
deformable

Stahl [2,6]
JWST segment
development

**Cohan
[8,9]**

Gray [11]

Smith

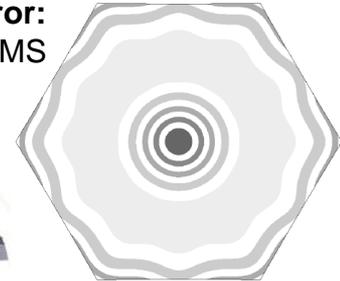
Ealey [3,4]
Surface-parallel
actuation, agile
manufacturing

SPOT

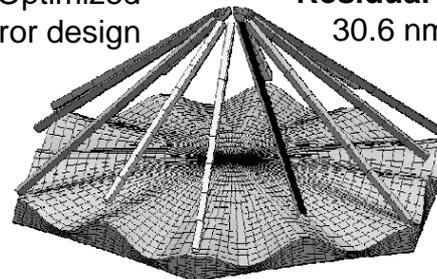
Non-optimized
mirror design



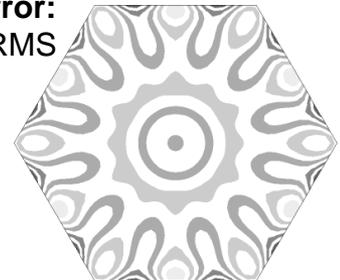
Residual Error:
225 nm RMS



Optimized
mirror design



Residual Error:
30.6 nm RMS

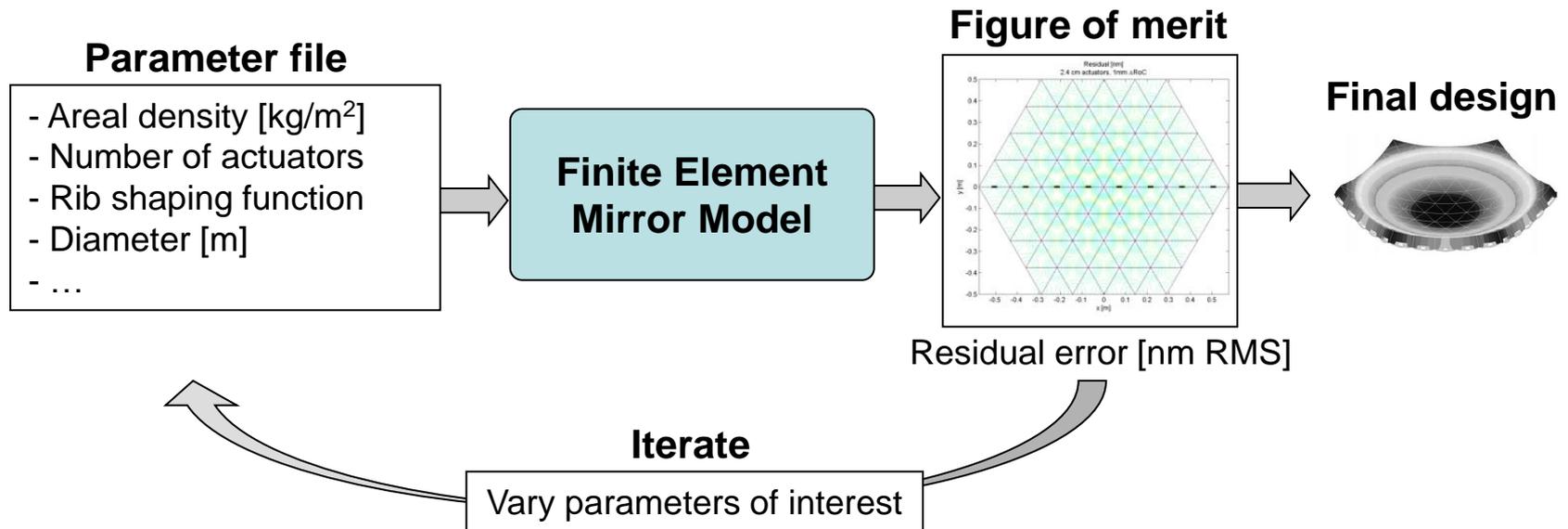


Shape Optimization to Decrease Residual Error

Research Objective

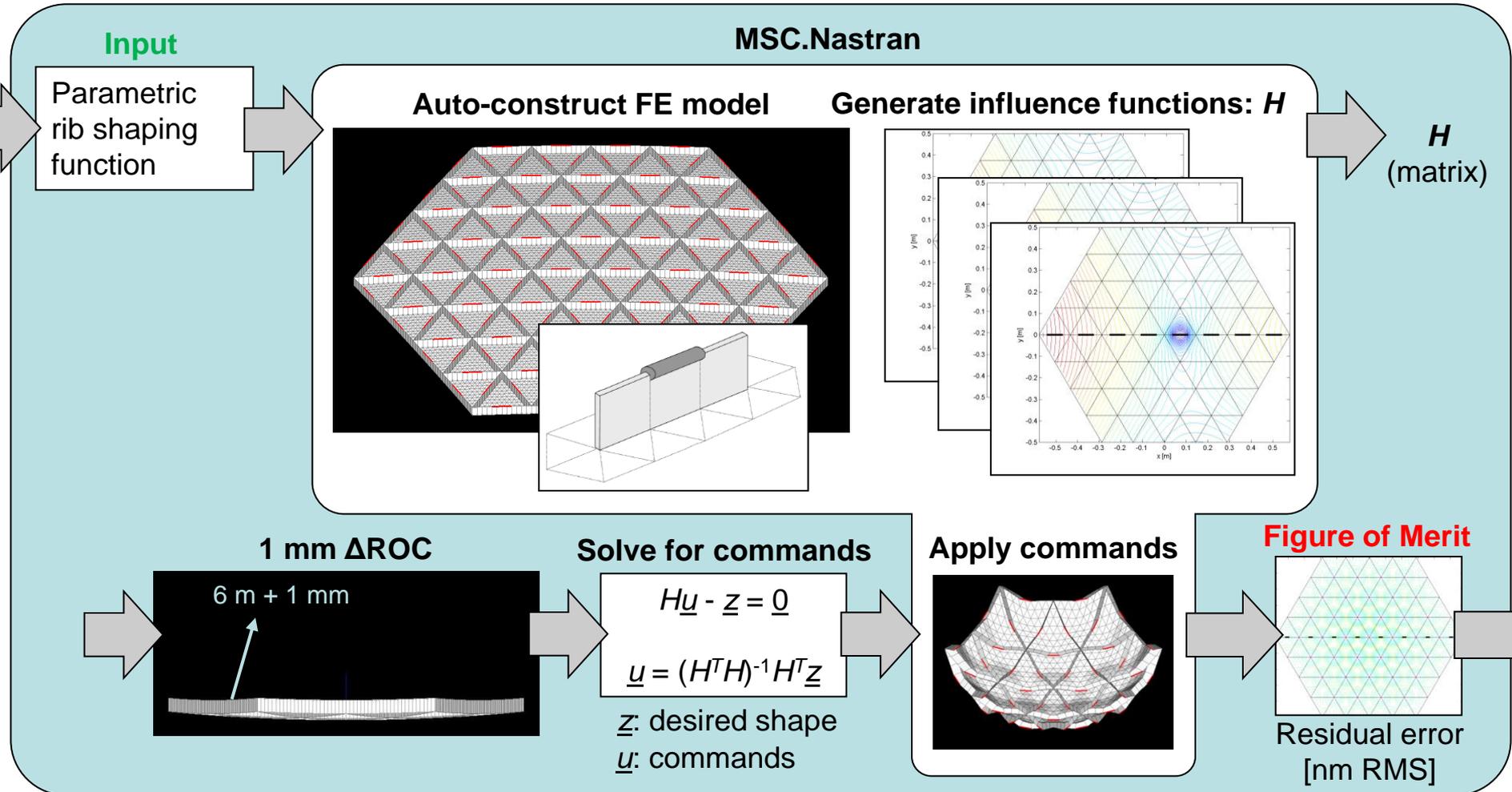
To reduce actuator-induced high spatial frequency residual error by taking advantage of changes in mirror geometry using a parametric finite element mirror model while keeping areal density (mass) and number of actuators (power, complexity) constant.

- Model of primary mirror
 - Parameterization allows for rapid iteration

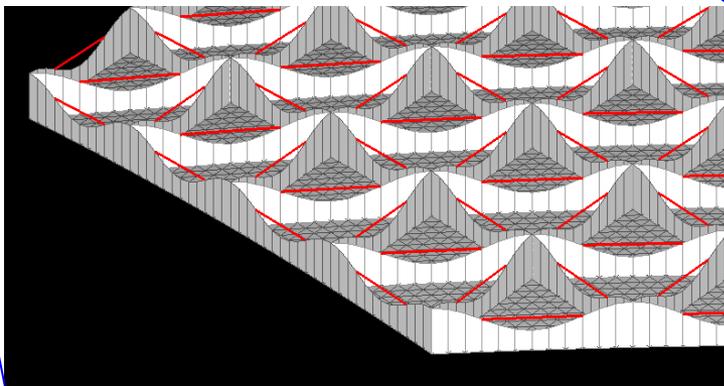
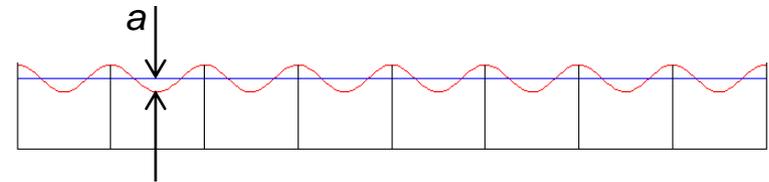
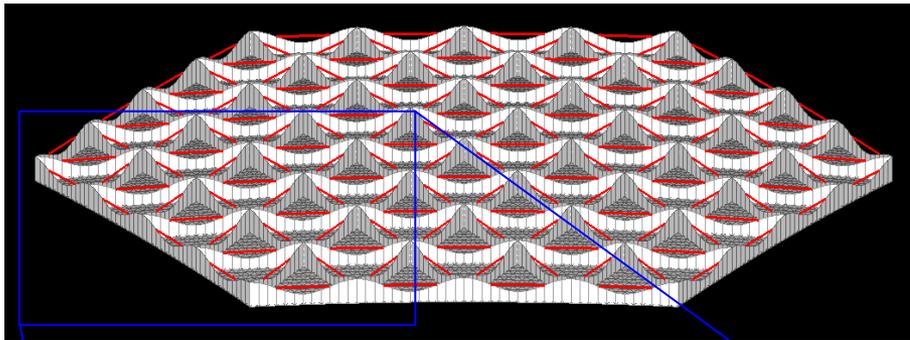


- Mirror model details
 - Varying the rib shaping function

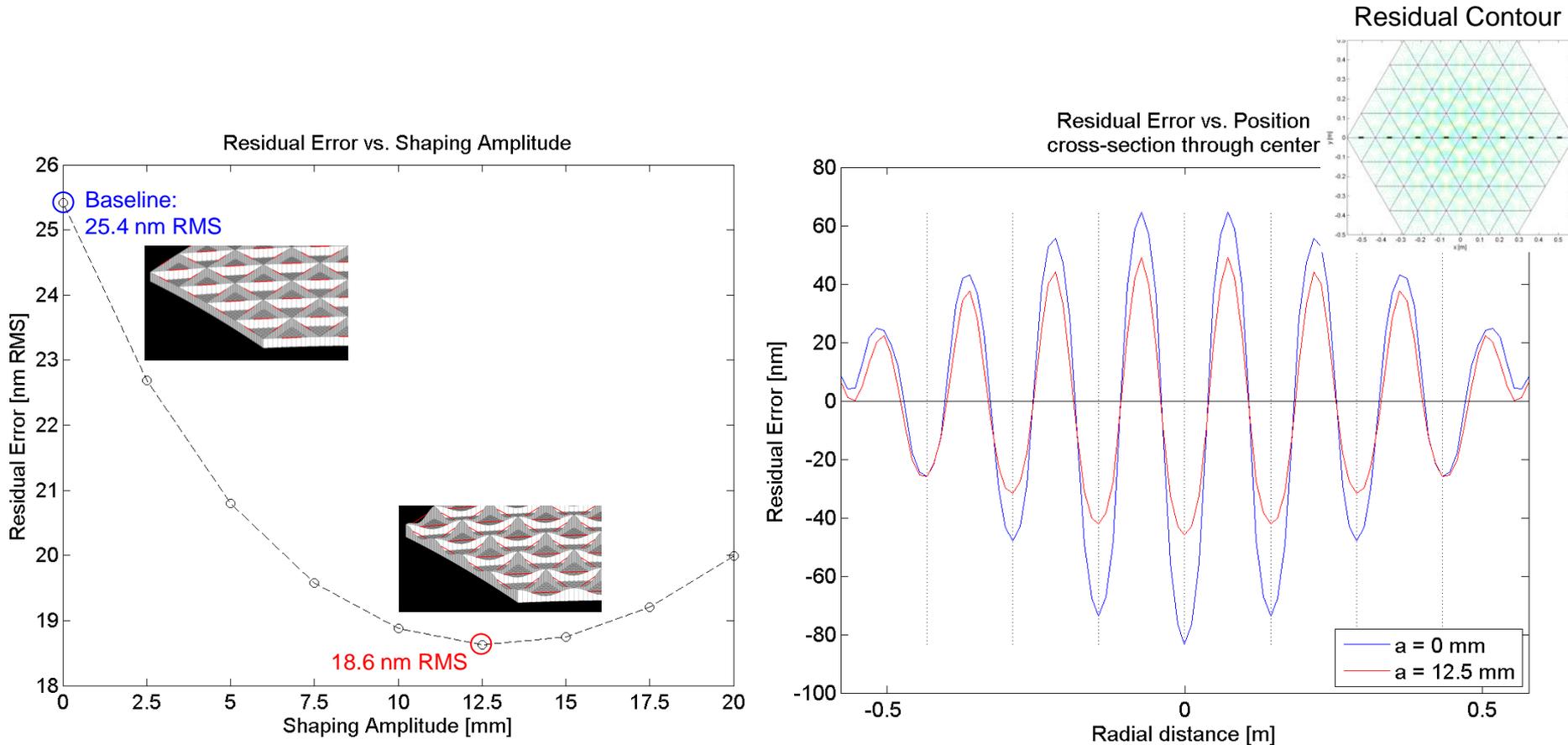
MATLAB



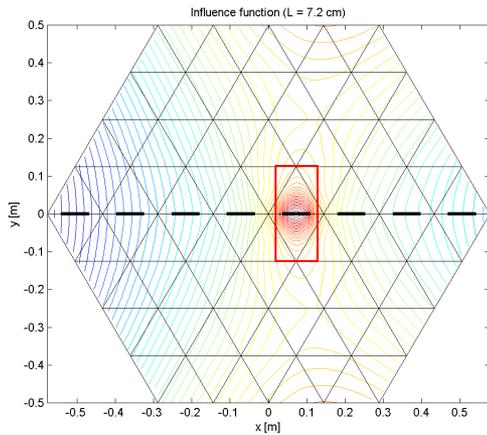
- Sinusoidal rib shaping
 - Idea: spread the load/influence of a localized actuator further along a rib
 - Parameterize shaping function by amplitude a ($0 \text{ mm} \leq a \leq 20 \text{ mm}$)
 - Areal density stays constant



- Sinusoidal shaping function
 - $a = 12.5$ mm minimizes residual error (27% reduction compared to the baseline)



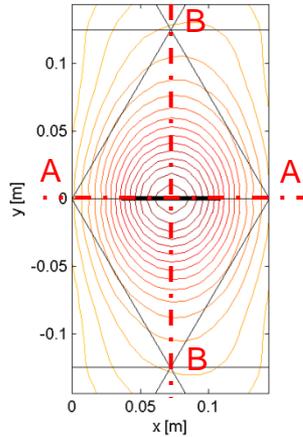
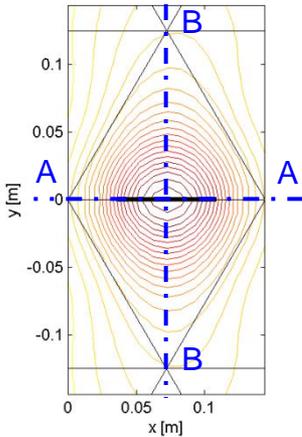
- Sinusoidal shaping function
 - Reason for improvement: “broader” influence function
 - Actuator becomes less “discrete”



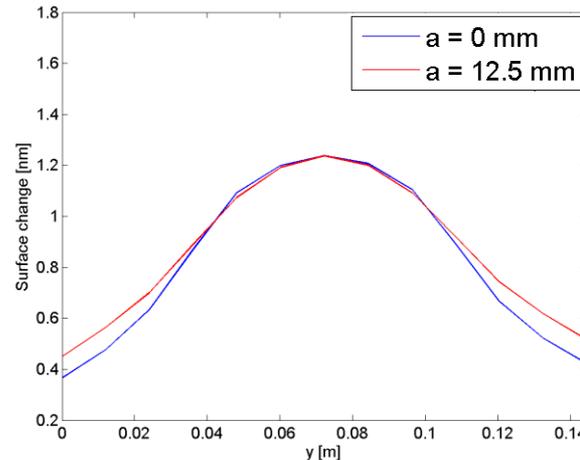
Influence Function

$a = 0$ mm

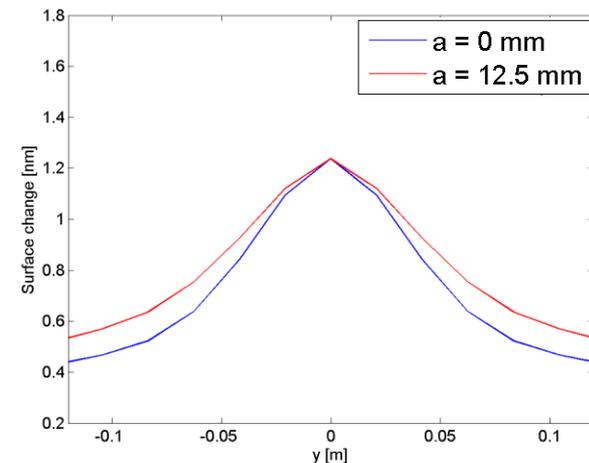
$a = 12.5$ mm



Slice A-A (horizontal)



Slice B-B (vertical)



Research Objective

To reduce actuator-induced high spatial frequency residual error by taking advantage of changes in mirror geometry using a parametric finite element mirror model while keeping areal density (mass) and number of actuators (power, complexity) constant.

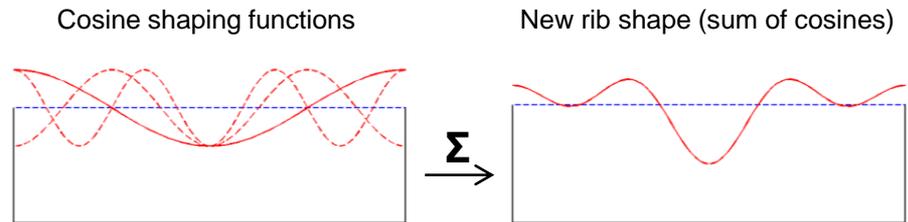
- Sinusoidal rib shaping
 - Rib shaping can reduce residual error
 - Mass is constant
 - Number of actuators is constant
 - $a = 12.5$ mm (best case) reduces residual by 27%
 - Broadens influence functions and “spreads” the effect of discrete actuators

Research Objective

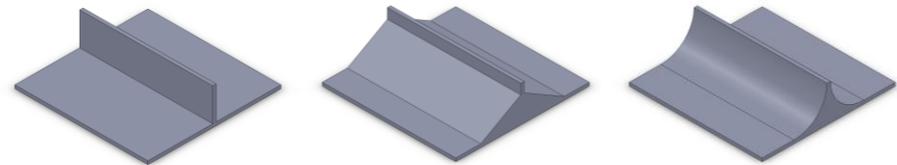
To reduce actuator-induced high spatial frequency residual error by taking advantage of changes in mirror geometry using a parametric finite element mirror model while keeping areal density (mass) and number of actuators (power, complexity) constant.

• Path ahead

- Rib shaping: Explore additional “basis functions” for ribs and optimize for best combination (e.g. series of cosines)



- Rib blending: Blend the ribs smoothly into the facesheet, instead of the current 90° junction

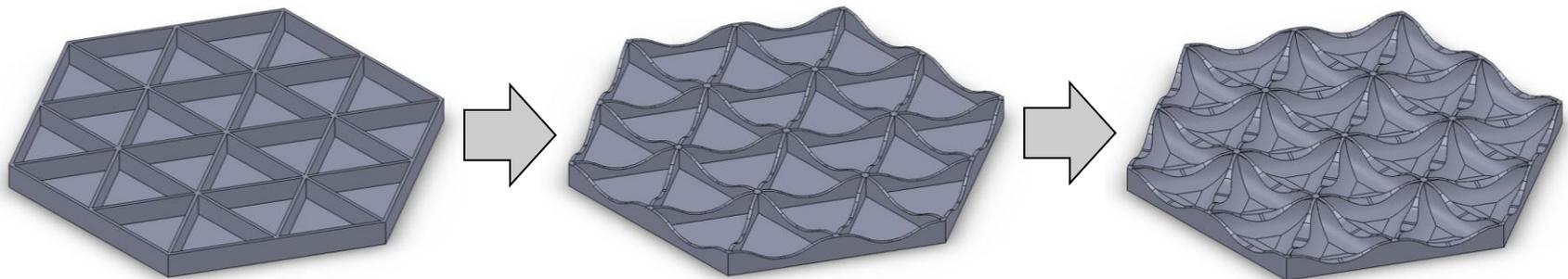


- General shape optimization: Using parameterized rib shaping and blending functions, optimize to find mirror shape that best reduces actuator-induced residual, while meeting stiffness requirements

Research Objective

To reduce actuator-induced high spatial frequency residual error by taking advantage of changes in mirror geometry using a parametric finite element mirror model while keeping areal density (mass) and number of actuators (power, complexity) constant.

- Expected contributions
 - Set of optimized mirror shapes that reduce actuator-induced residual error (constant mass and constant number of actuators)
 - New approach to mirror design: start with actuator specification, then design mirror
 - Compatible with existing manufacturing techniques

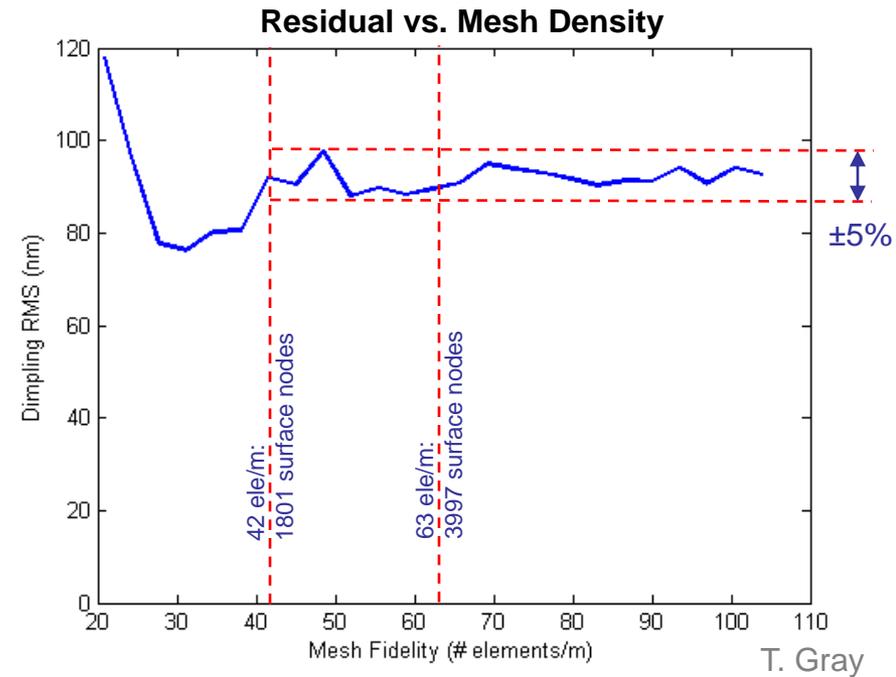
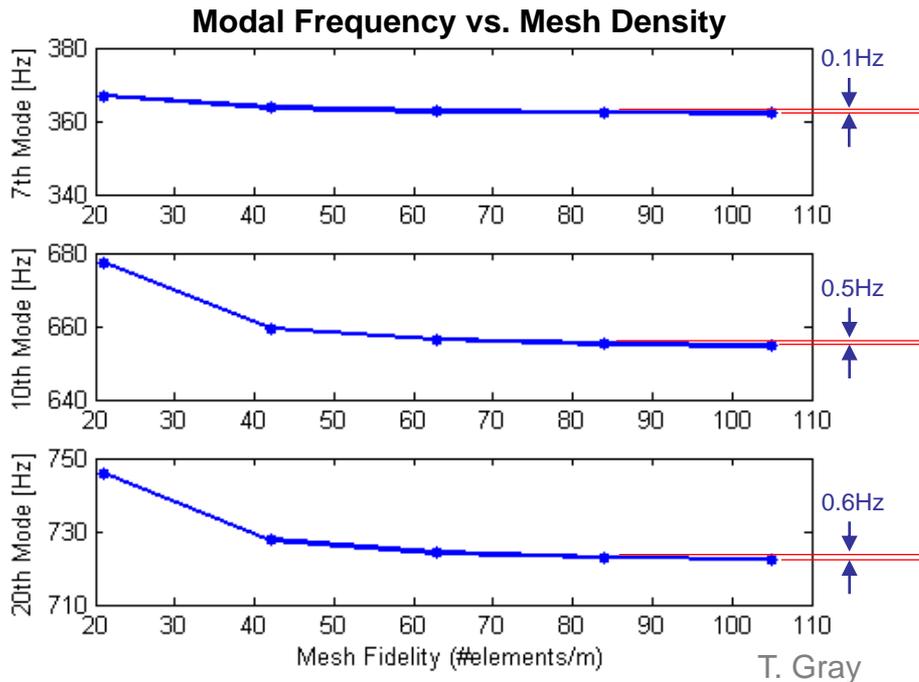


- [1] National Aeronautics and Space Administration, image archive.
- [2] H. Philip Stahl, “JWST Lightweight Mirror TRL-6 Results”. *Proc. IEEE Aerospace Conference* (2007).
- [3] Mark A. Ealey, “Agile Mandrel Apparatus and Method”. US Patent Application No. US 2006/0250672 A1 (2006).
- [4] Mark A. Ealey and John A. Wellman, “Highly adaptive integrated meniscus primary mirrors”. *Proc. SPIE* 5166 (2004).
- [5] Michael J. Shepherd, et al., “Modal Transformation Method for Deformable Membrane Mirrors,” *Journal of Guidance, Control, and Dynamics*, Vol. 32, No. 1 (2009).
- [6] H. Philip Stahl, “JWST mirror technology development results”. *Proc. SPIE* 6671 (2007).
- [7] F. Cote, et al., “Dynamic and Static Modelling of Piezoelectric Composite Structures Using a Thermal Analogy with MSC/Nastran”, *Composite Structures*, Vol. 65, No. 3 (2004).
- [8] Lucy E. Cohan, “Integrated Modeling to Facilitate Control Architecture Design for Lightweight Space Telescopes”, *MS Thesis*, MIT (2007).
- [9] Lucy E. Cohan, “Integrated Modeling and Design of Lightweight, Active Mirrors for Launch Survival and On-Orbit Performance”, *PhD Thesis*, MIT (2010).
- [10] Elizabeth Jordan, “Design and Shape Control of Lightweight Mirrors for Dynamic Performance and Athermalization”, *MS Thesis*, MIT (2007).
- [11] Thomas L. Gray, “Minimizing High Spatial Frequency Residual in Active Space Telescope Mirrors”, *MS Thesis*, MIT (2008).
- [12] Jason G. Budinoff & Gregory J. Michels, “Design & optimization of the spherical primary optical telescope (SPOT) primary mirror segment”, *Proc. SPIE* 5877 (2005).
- [13] Bruce H. Dean, et al., “Phase retrieval algorithm for JWST Flight and Testbed Telescope”, *Proc. SPIE* 6265 (2006).
- [14] Mark A. Ealey and John A. Wellman, “High Authority Deformable Optical System”. US Patent Application No. US 2006/0262434 A1 (2006).
- [15] Maureen L. Mulvihill and Mark A. Ealey, “Large Mirror Actuators in the 21st Century”, *Proc SPIE* 5166 (2004).
- [16] Siddarayappa Bikkannavar, et al., “Autonomous phase retrieval control for calibration of the Palomar Adaptive Optics system”, *Proc. SPIE* 7015 (2008).

Backup

- Convergence

- Sensitivity of model outputs to FE mesh density
- Does the model go to one value as mesh density changes?

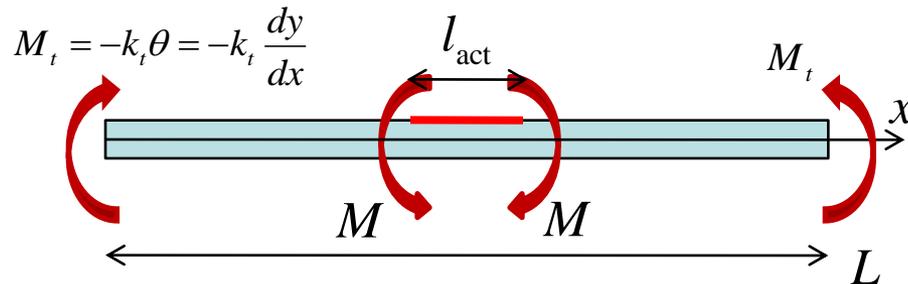


- Benchmarking against empirical data
 - Does that one value match real life?

Model output	Error compared to empirical data
Stiffness (fundamental frequency)	2%
Mirror stress due to launch vibration	8%
Residual error for 1 mm Δ ROC	7%

- Feature-level validation

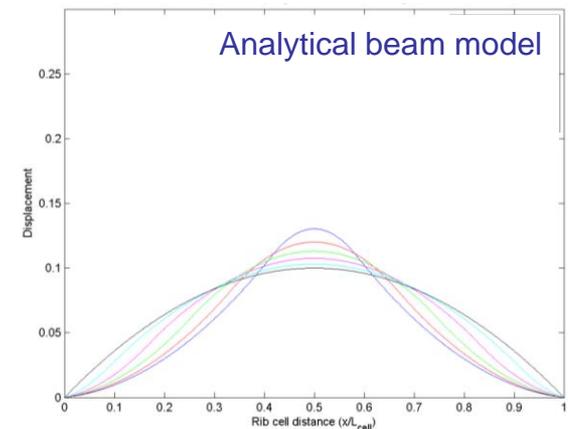
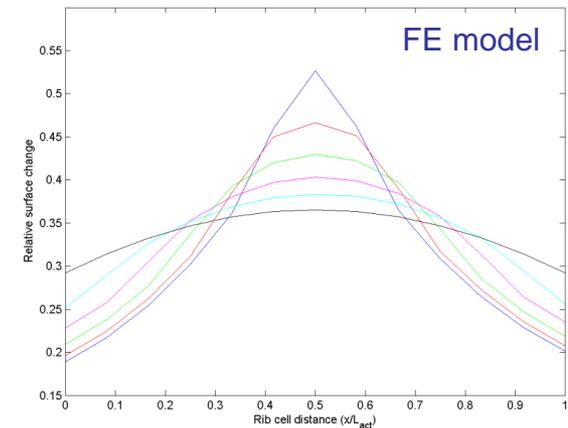
- Validate individual features of the model for which there is no “global” test data
- Compare with first-principles models, published test results, published analysis
- E.g.: modeling and validation of piezoelectric actuators in [7,9]
- E.g.: actuator length validation via analytical beam model (below)



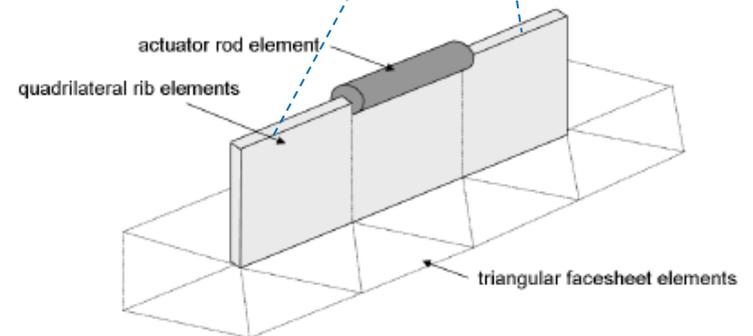
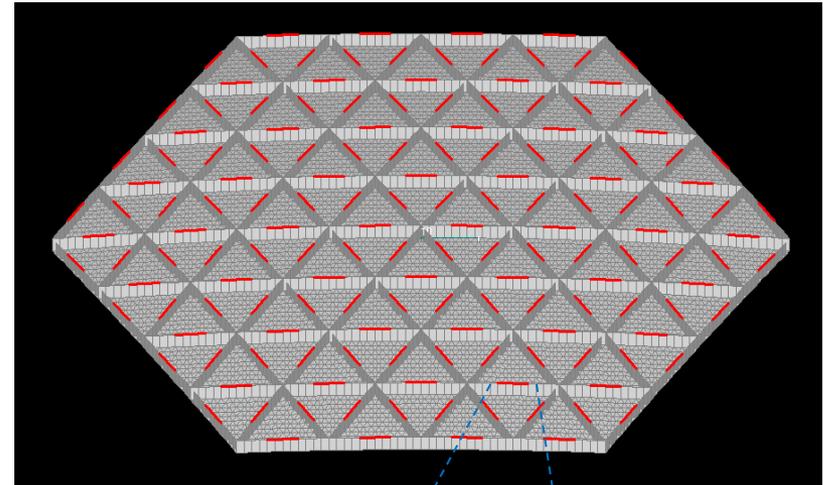
$$y(x) = \frac{M_t}{2EI} x^2 - \frac{M}{2EI} \left\langle x - \frac{L-l_{act}}{2} \right\rangle^2 + \frac{M}{2EI} \left\langle x - \frac{L+l_{act}}{2} \right\rangle^2 + \frac{Ml_{act} - M_t L}{2EI} x$$

$$M_t = \frac{Ml_{act}}{L + EI/k_t}$$

$$\langle x - x_0 \rangle^2 = \begin{cases} 0 & x < x_0 \\ (x - x_0)^2 & x \geq x_0 \end{cases}$$

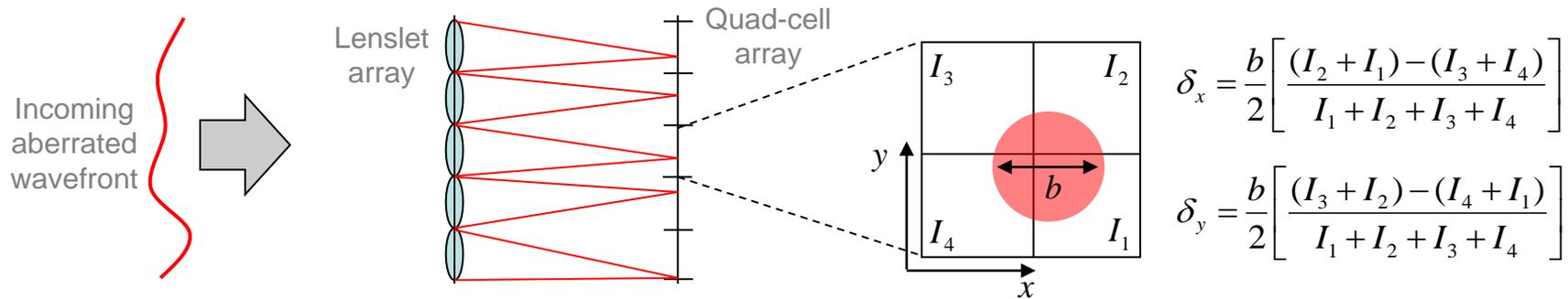


Parameter	Value
Diameter (flat-flat)	1.0 m
Areal density (SiC only)	8 kg/m ²
Rib rings	4
Primary rib height	25.4 mm
Primary rib thickness	1 mm
Face sheet thickness	1.8 mm
Actuator length	7.2 cm
Actuator length/cell	50%
Number of actuators	156
Radius of curvature	6 m



- Shack-Hartmann Wavefront Sensor

- Detects the slope of the wavefront by measuring spot displacements
- Measures the wavefront at discrete locations on the pupil
- Ignores non-common path errors

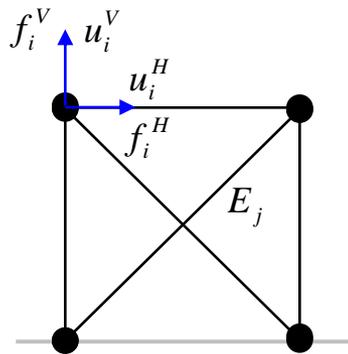


- Modified Gerchberg-Saxton (MGS)

- Computes wavefront by inverse transforming a defocused point image
- Uses the focal plane array sampling (more of a continuous map)
- FT and IFT between image and pupil, enforcing pupil plane image constraints
- Known defocus removes ambiguity in phase map
- Captures non-common path errors
- Baseline method for fine phasing in JWST

- Static Analysis

- Goal: Compute displacements from known forces and constraints
- Relate forces to displacements via a global stiffness matrix K
- Build the stiffness matrix from kinematic relationships, Hooke's Law, and force balance



u_i Horizontal and vertical displacements of node i

f_i Horizontal and vertical forces acting at node i

E_j Elastic modulus of element j

Elongations (strain)
from displacements

$$\underline{\varepsilon} = A\underline{u}$$



Internal force (stress)
from elongation (strain)

$$\underline{\sigma} = E\underline{\varepsilon}$$



Force balance
(internal and external)

$$\underline{f} = A^T \underline{\sigma}$$

Force from displacements (combining above equations)

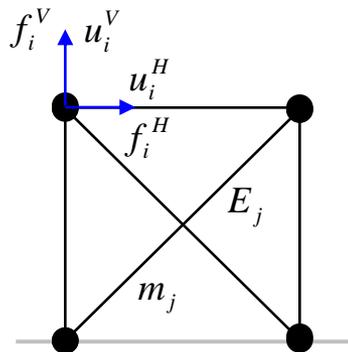
$$\underline{f} = A^T E A \underline{u} \quad (1)$$

$$\underline{f} = K \underline{u}$$

- Solve (1), typically by elimination

- Normal Modes Analysis

- Goal: Compute the normal frequencies and normal modes of a system
- Transform into an eigenvalue problem



u_i Horizontal and vertical displacements of node i

f_i Horizontal and vertical forces acting at node i

m_j Mass of element j

M Global mass matrix

K Global stiffness matrix (see previous slide)

$$M\ddot{\underline{u}} + K\underline{u} = 0$$

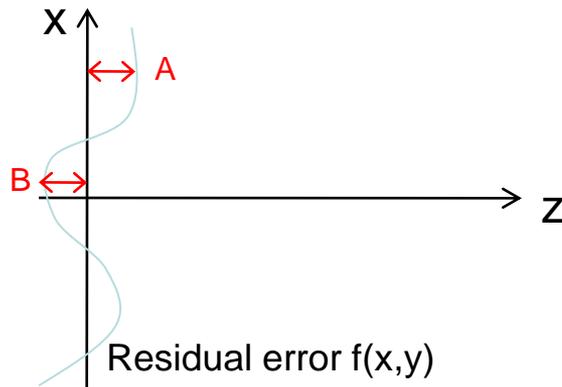
$$\underline{u} = \underline{\phi} \sin \omega t$$

$$\boxed{(K - \omega^2 M)\underline{\phi} = 0} \quad (2)$$

Last line is an eigenvalue equation of the form $A\underline{x} = \lambda\underline{x}$

- Solve for normal frequencies (eigenvalues) and normal modes (eigenvectors)

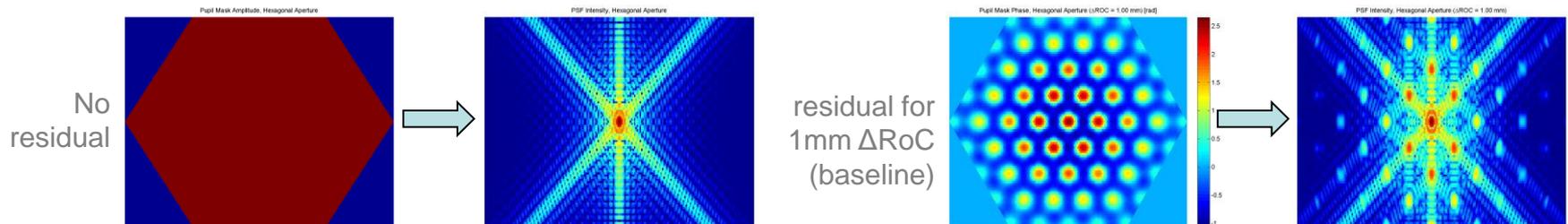
- Impact of residual error on imaging
 - Residual error causes wavefront error (WFE = 2*residual error)
 - This adds an additional phase term to the pupil function
 - Result: degraded image quality via reduced peak energy and image artifacts



A: Optical “acceleration” → decrease phase

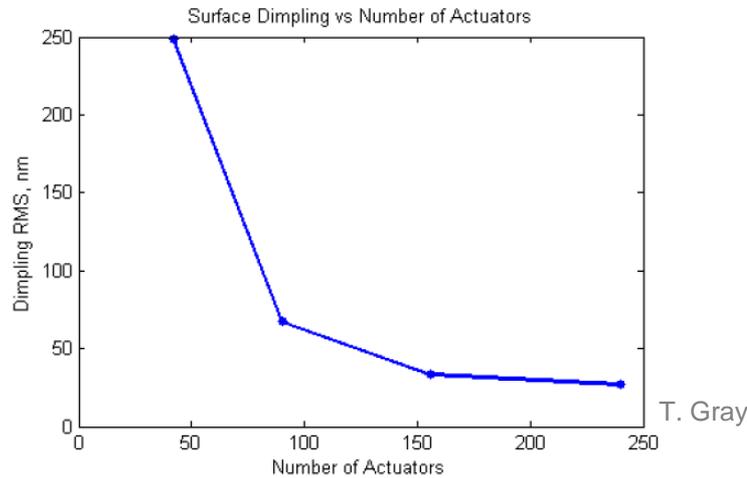
B: Optical delay → increase phase

$$\phi \text{ [rad]} = -2 * 2\pi/\lambda * f(x,y) \text{ [nm]}$$

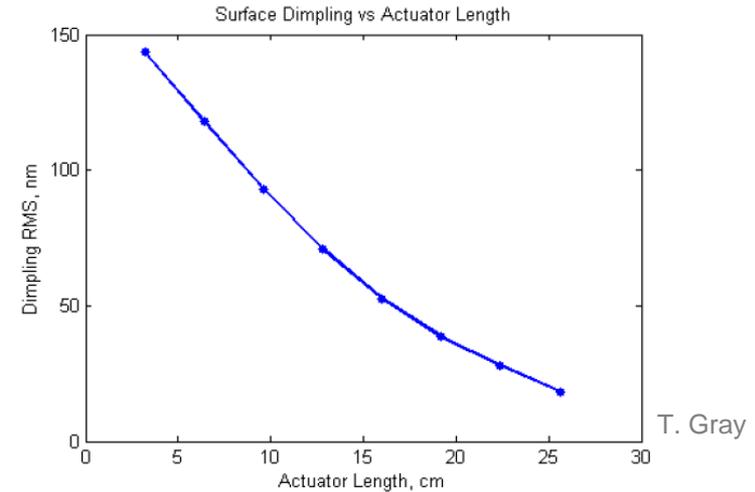


- Effect of geometric parameters on residual error

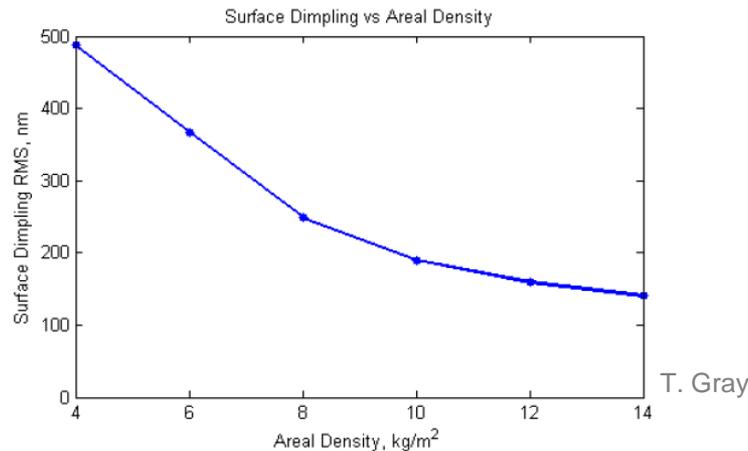
- Number of actuators (rib rings)



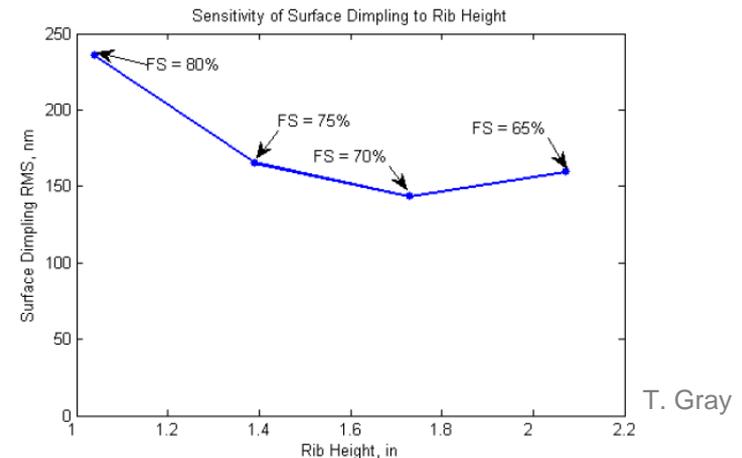
- Actuator length



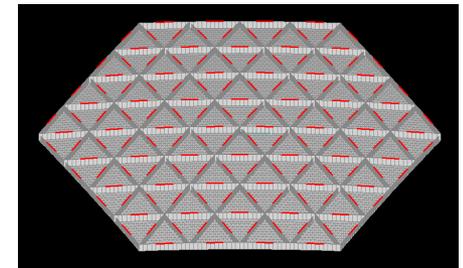
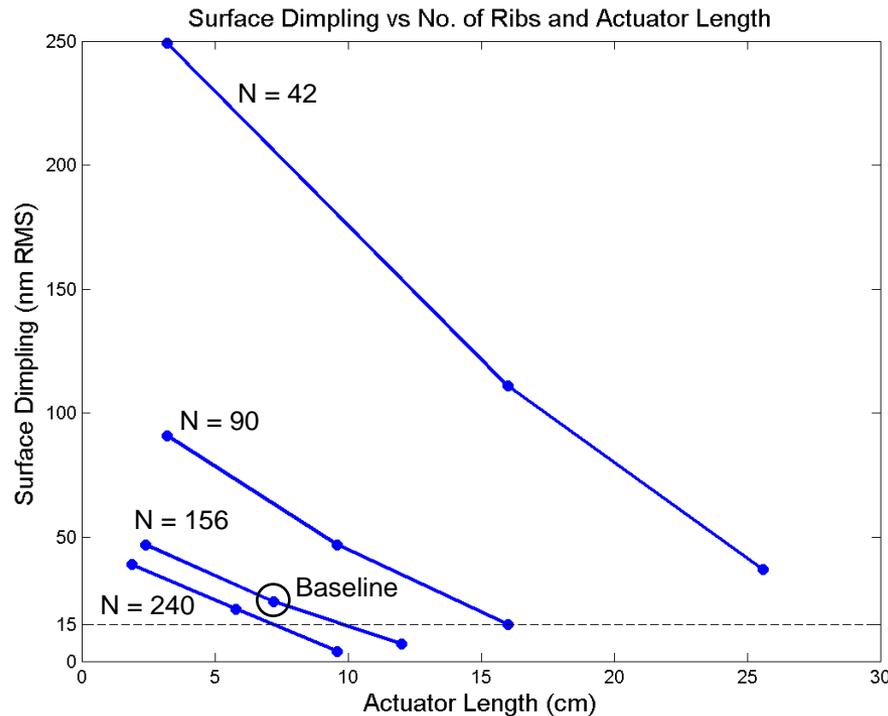
- Areal density



- Rib height

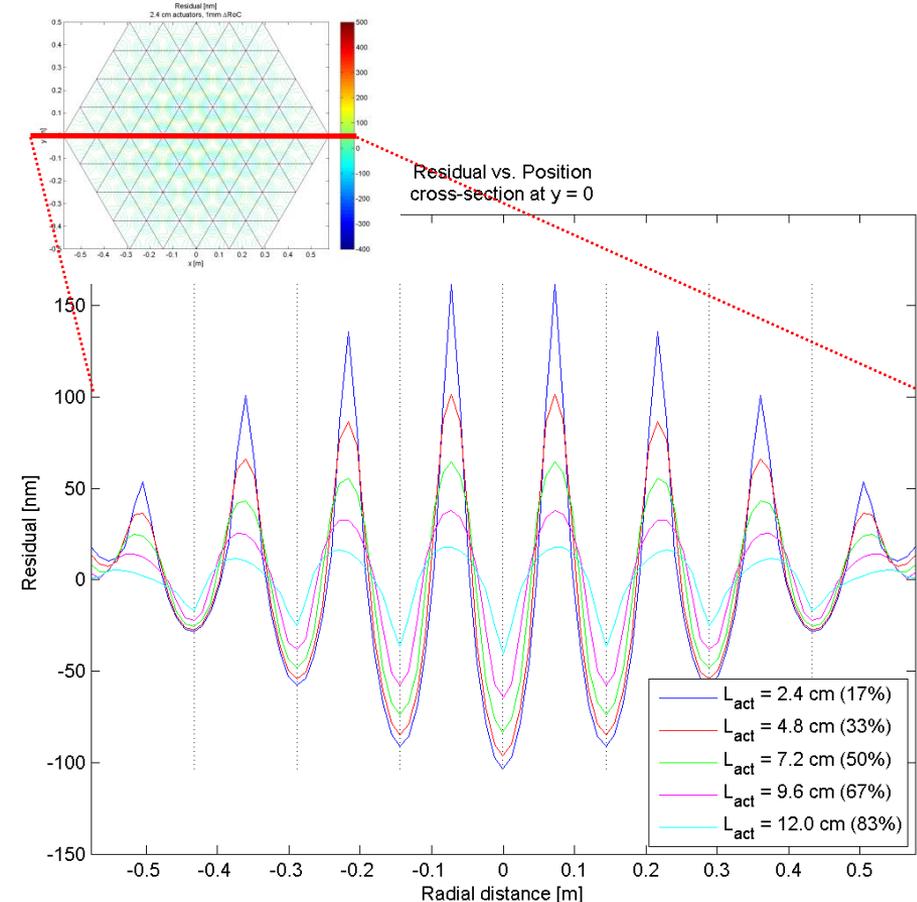
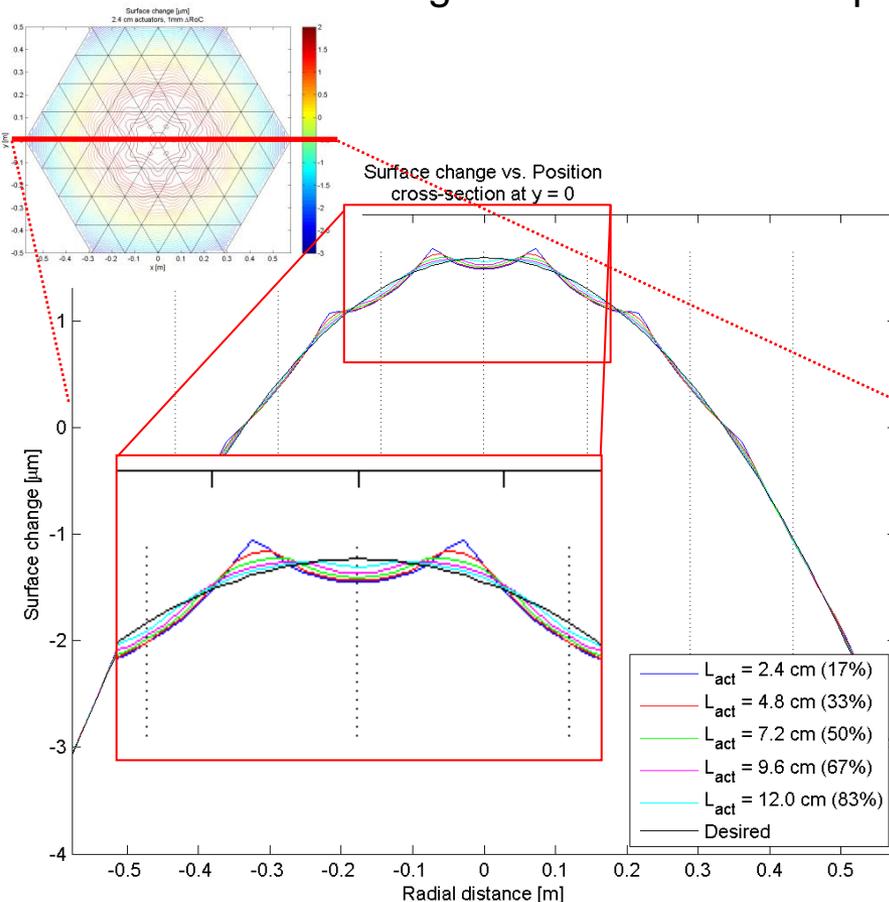


- Changing actuator length
 - Longer actuators give better performance
 - Increasing actuator length increases the distance between the applied moments, smoothing the commanded shape



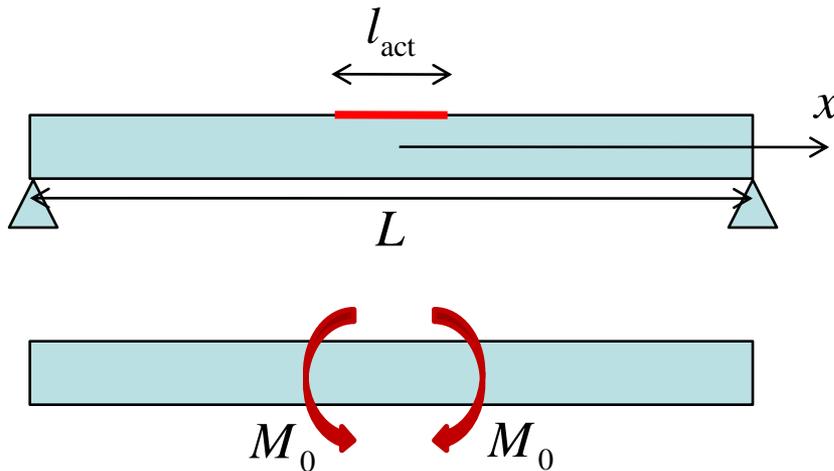
Rib "rings"	Actuators
2	42
3	90
4	156
5	240

- Changing actuator length
 - Longer actuators give better performance (3.6x improvement over the baseline)
 - Increasing actuator length increases the distance between the applied moments, smoothing the commanded shape



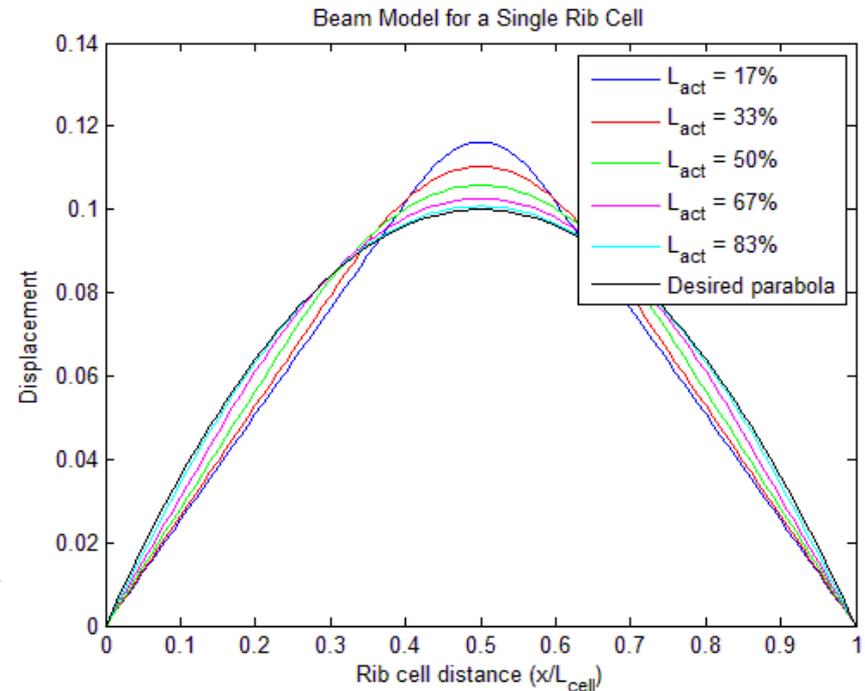
- Beam model

- Single rib cell modeled; actuator modeled as a moment couple
- As space between couple increases, more of the beam experiences parabolic deflection
- Influence function becomes less localized

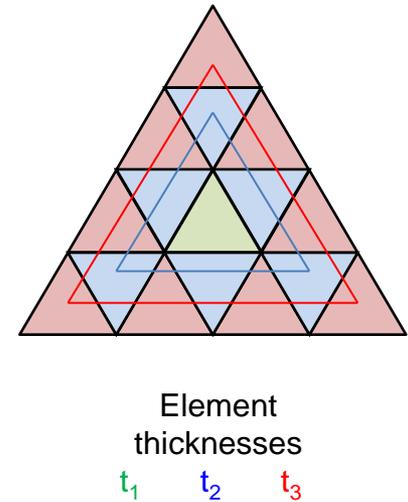
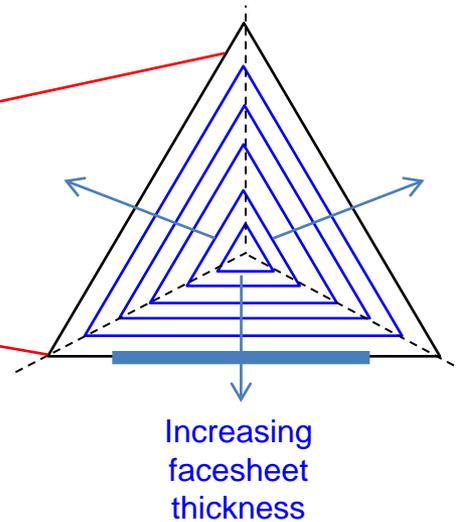
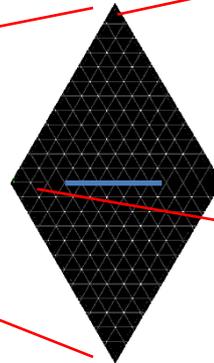
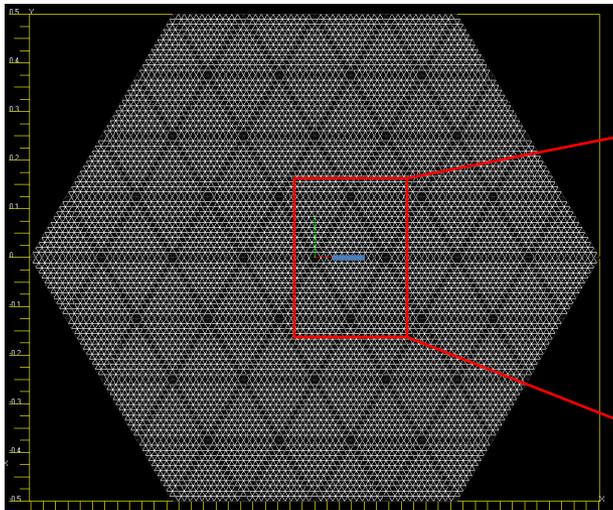
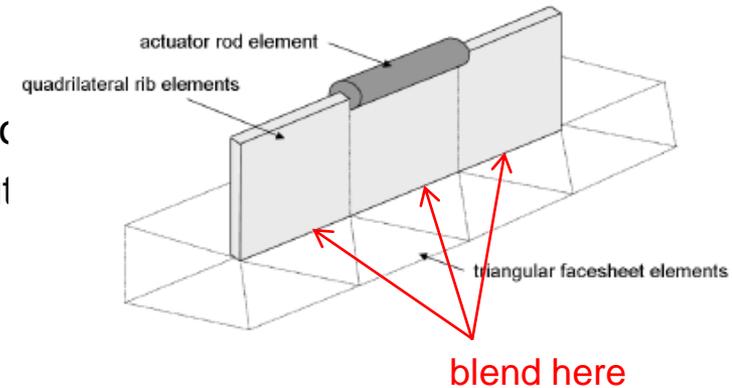


$$y(x) = \frac{M_0}{2EI} \left\langle x - \frac{a+l_{act}}{2} \right\rangle^2 - \frac{M_0}{2EI} \left\langle x - \frac{a-l_{act}}{2} \right\rangle^2 + \frac{M_0 l_{act}}{EI} \left(1 - \frac{a}{2L} \right) x$$

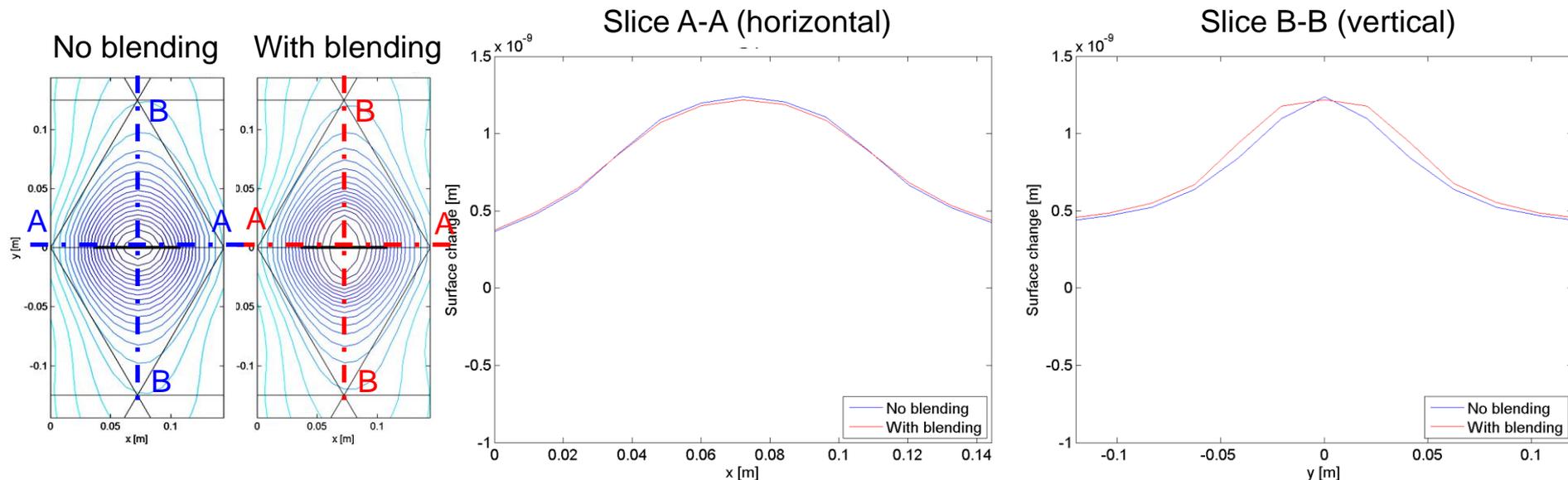
$$\langle x - x_0 \rangle^2 = \begin{cases} 0 & x < x_0 \\ (x - x_0)^2 & x \geq x_0 \end{cases}$$



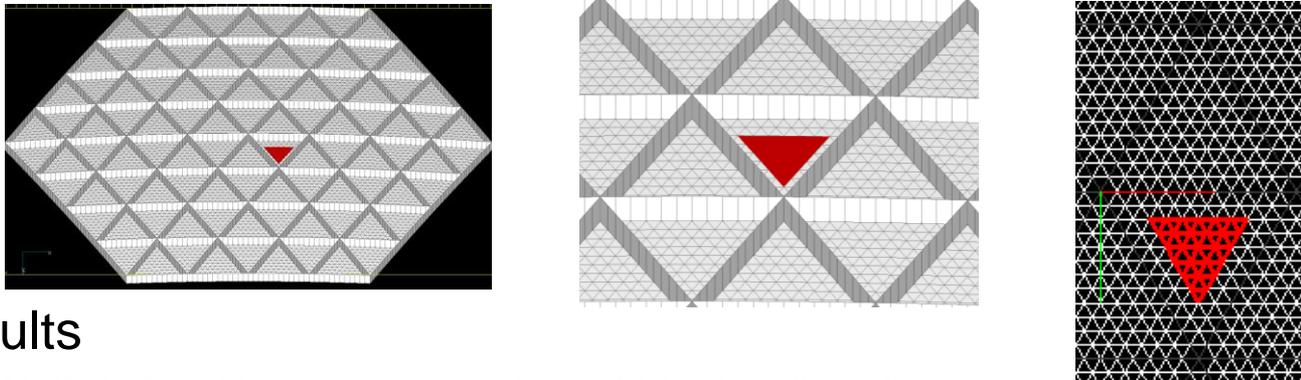
- Rib-to-facesheet blending
 - Idea: smooth the transition from rib to facesheet spread the influence of a given actuator
 - Initial implementation: Use 2D shell elements but vary the thickness to approximate 3D blending
- Implementation details:



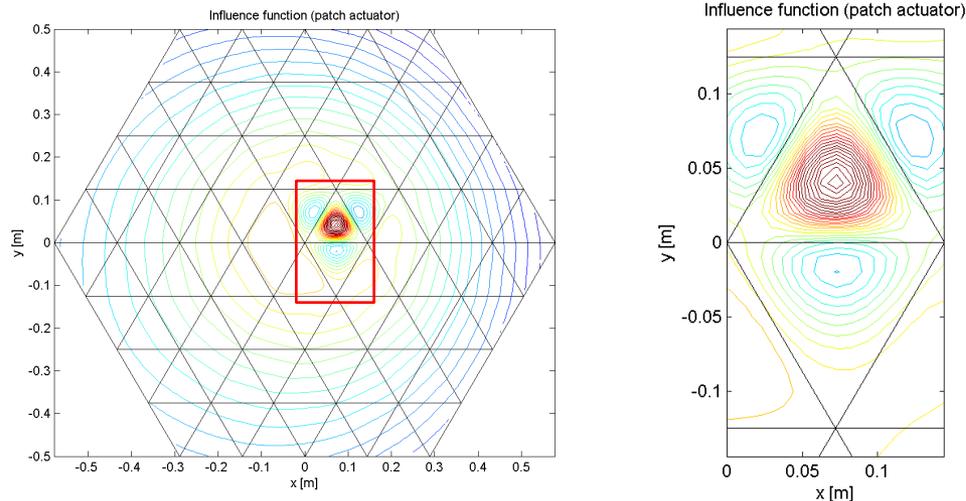
- Rib-to-facesheet blending
 - Idea: smooth the transition from rib to facesheet spread the influence of a given actuator
 - Initial implementation: Use 2D shell elements but vary the thickness to approximate 3D blending
- Results:
 - Increases the “vertical” extent of the actuator’s influence function



- 2D patch actuator
 - Idea: use a thin piezoelectric patch and take advantage of the d_{31} behavior
 - Place within the cells to add another set of influence functions
 - Initial implementation: Use 2D shell elements attached at nodes



- Results
 - Well-defined influence function within the rib cell



- Spherical Primary Optical Telescope

- NASA Goddard
- Single segment test bed

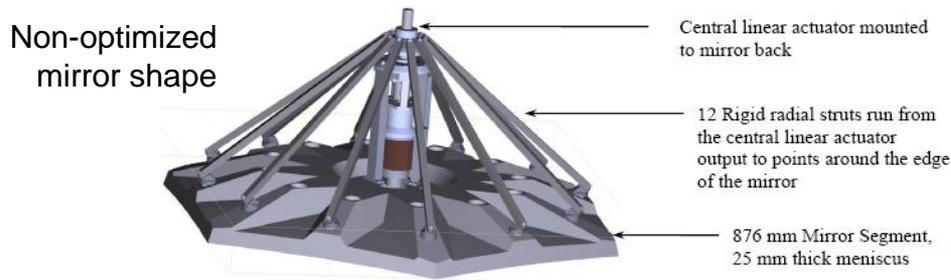
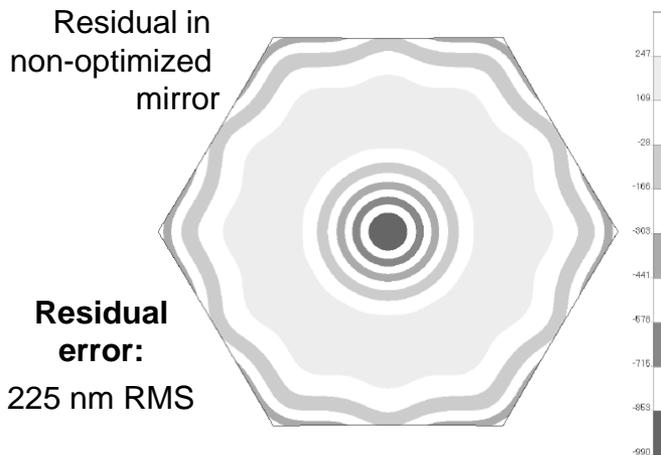
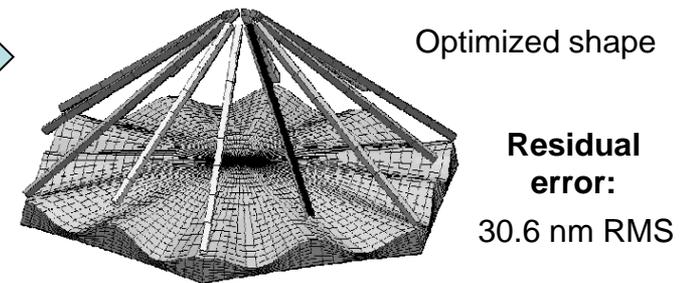


Figure shape	sphere
Diameter	0.876 m
RoC	5 m (nominal)
Max surface error for 400 μm ΔRoC	15 nm RMS
Material	Pyrex
Minimum areal density (10 mm thickness throughout)	22.3 kg/m ² (Pyrex only)



Select basis functions
 $f(r) = a_1 r + a_2 r^2 + a_3 r^3 + \dots$
 $g(\theta) = b_1 \sin(6\theta) + b_2 \sin(12\theta)$
 Optimize to find coefficients



Jason G. Budinoff, Gregory J. Michels, "Design & Optimization of the Spherical Primary Optical Telescope (SPOT) Primary Mirror Segment", *Proc. SPIE*, Vol. 5877 (2005).

- Past MOST results
 - Parametric space telescope model
 - Primary mirror and kinematic mounts
 - Secondary mirror and support tower
 - Solar arrays and bus
 - Allows rapid tradespace exploration
 - Compare designs and identify driving parameters

