

Problem Set #1

Due: Thursday, February 16, 2006, in class

Boldface changes added on February 10, 2006

1. Geodesic in Newtonian Gravitational Potential

Dodelson, Chapter 2, exercise 3

In part b, the sum $p^0 + m\phi$ is conserved under some condition. What is the condition? Note: more components of connection may be needed in addition to those in part (a).

In part c, Drop the m in the expression for Newton's 2nd Law.

2. Robertson-Walker Metric (from Fall 2001)

Consider the general Robertson-Walker metric, written as the invariant line element:

$$ds^2 = -dt^2 + a^2(t) \left[\frac{dr^2}{1 - kr^2} + r^2(d\theta^2 + \sin^2\theta d\phi^2) \right]$$

Find coordinate transformations that will put the line element in the following forms:

$$ds^2 = a^2(\tau) \left[-d\tau^2 + d\chi^2 + r^2(\chi)(d\theta^2 + \sin^2\theta d\phi^2) \right]$$

$$ds^2 = a^2(\tau) \left[-d\tau^2 + \frac{d\bar{r}^2 + \bar{r}^2(d\theta^2 + \sin^2\theta d\phi^2)}{(1 + \frac{1}{4}k\bar{r}^2)^2} \right]$$

$$ds^2 = -dt^2 + a^2(t) \frac{(dx^2 + dy^2 + dz^2)}{\left[1 + \frac{1}{4}k(x^2 + y^2 + z^2)\right]^2}$$

For each case, indicate the full range of the variables. Give explicit formulae for $r(\chi)$ and $\bar{r}(r)$. (Hint: Different forms may be required for $k > 0$, $k < 0$, and $k = 0$. Note also that $a(\tau)$ is not the same function of its argument as $a(t)$.)

3. Isotropy implies Homogeneity

- a. Use elementary thought experiments to convince yourself (as well as me), that a 3-dimensional hypersurface of constant time which is isotropic at all locations must be totally homogeneous.
- b. Use the space-like components of the contracted Bianchi identities, $G^{ij}{}_{;j} = 0$, as a statement of isotropy to more formally show that the **spatial components of $R_{\mu}{}^{\nu}$** must be uniform on the hypersurface (the gradient of curvature is zero everywhere).

4. Rotating Flat space time

Consider the Minkowski line element in cylindrical coordinates:

$$ds^2 = -dt^2 + dr^2 + dz^2 + r^2 d\phi^2$$

and introduce a change in variable ϕ to account for uniform rotation: $\phi = \phi' - \omega t$.

- a. Write down the metric in the new coordinate system, both $g_{\mu\nu}$ and its inverse $g^{\mu\nu}$.
- b. Find the non-zero Christoffel symbols corresponding to this metric.
- c. Calculate the corresponding Ricci Tensor, $R_{\mu\nu}$ and Scalar, R .
- d. What do Einstein's field equations imply for the energy momentum tensor, $T_{\mu\nu}$?