# 6.946 Assignment $4^{1}$ 

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## Exercise 2.2: Steiner's Theorem

The moment of inertia with respect to the first line is given by $I=\sum_{\alpha} m_{\alpha}\left(\xi_{\alpha}^{\perp}\right)^{2}$, where $m_{\alpha}$ is the mass of the individual particle, and $\xi_{\alpha}$ is the vector from the reference point to the individual point. Since the axis of rotation goes through the center of mass, the first line goes through the reference point.

Let $\vec{R}$ be a perpendicular vector from the first line to the second, and let $\xi_{\alpha}^{\perp \prime}=\xi_{\alpha}^{\perp}+\vec{R}$ be the reference vector from the second line to the individual particle. The new moment of inertia $I^{\prime}$ is given by:

$$
\begin{aligned}
I^{\prime} & =\sum_{\alpha} m_{\alpha}\left(\xi_{\alpha}^{\perp}\right)^{2}=\sum_{\alpha} m_{\alpha}\left(\xi_{\alpha}^{\perp}+\vec{R}\right)^{2}=\sum_{\alpha} m_{\alpha}\left(\left(\xi_{\alpha}^{\perp}\right)^{2}+2 \xi_{\alpha}^{\perp} \vec{R}+R^{2}\right)= \\
& =\sum_{\alpha} m_{\alpha}\left(\xi_{\alpha}^{\perp}\right)^{2}+\sum_{\alpha} m_{\alpha} 2 \xi_{\alpha}^{\perp} \vec{R}+R^{2} \sum_{\alpha} m_{\alpha}=I+\sum_{\alpha} m_{\alpha} 2 \xi_{\alpha}^{\perp} \vec{R}+M R^{2}
\end{aligned}
$$

By the definition of the center of mass, $\sum_{\alpha} m_{\alpha} \xi_{\alpha}^{\perp}=0$, and the middle term drops out. So

$$
I^{\prime}=I+M R^{2}
$$

## Exercise 2.3: Some useful moments of inertia

In the continuous case, we can replace the equation for the moment of inertia with

$$
I=\int_{M}\left(r_{\perp}\right)^{2} d m
$$

where $r_{\perp}$ is the perpendicular distance from the axis of rotation to the position of $d m$, and we integrate over the entire mass.
(a) We introduce a density $\delta=\frac{d m}{d V}=\frac{M}{\frac{4}{3} \pi R^{3}}$. Consider a spherical coordinate system, and let the axis of rotation be the z -axis.

$$
I=\int_{M} r_{\perp}^{2} d m=\int_{V} r_{\perp}^{2} \frac{d m}{d V} d V=\delta \int_{\rho=0}^{\rho=R} \int_{\theta=0}^{\theta=2 \pi} \int_{\phi=0}^{\phi=\pi} r_{\perp}^{2}\left(\rho^{2} \sin \phi d \phi d \theta d \rho\right)
$$

Since the axis of rotation is the z -axis, the perpendicular distance is just $r_{\perp}=\rho \sin \phi$, and we continue the calculation ${ }^{2}$ :

$$
I=\delta \int_{\rho=0}^{\rho=R} \int_{\theta=0}^{\theta=2 \pi} \int_{\phi=0}^{\phi=\pi}(\rho \sin \phi)^{2} \rho^{2} \sin \phi d \phi d \theta d \rho=\delta 2 \pi \int_{\rho=0}^{\rho=R} \int_{\phi=0}^{\phi=\pi} \rho^{4} \sin ^{3} \phi d \phi d \rho
$$

[^0]\[

$$
\begin{gathered}
=\delta 2 \pi \frac{R^{5}}{5} \int_{\phi=0}^{\phi=\pi} \sin ^{3} \phi d \phi=\delta 2 \pi \frac{R^{5}}{5}\left[\frac{1}{3} \sin ^{2} \phi \cos \phi+\frac{2}{3} \cos \phi\right]_{\phi=0}^{\phi=\pi} \\
=\delta 2 \pi \frac{R^{5}}{5} \frac{4}{3}=\frac{M}{\frac{4}{3} \pi R^{3}} 2 \pi \frac{R^{5}}{5} \frac{4}{3}=\frac{24 \pi M R^{5}}{60 \pi R^{3}}=\frac{2}{5} M R^{2}
\end{gathered}
$$
\]

(b) We introduce a mass-per-surface-area constant $\delta=\frac{d m}{d S}=\frac{M}{4 \pi R^{2}}$. Consider a spherical coordinate system, and let the axis of rotation be the z -axis.

$$
I=\int_{M} r_{\perp}^{2} d m=\int_{S} r_{\perp}^{2} \frac{d m}{d S} d S=\delta \int_{\theta=0}^{\theta=2 \pi} \int_{\phi=0}^{\phi=\pi} r_{\perp}^{2}\left(R^{2} \sin \phi d \phi d \theta\right)
$$

Since the axis of rotation is the z-axis, the perpendicular distance is just $r_{\perp}=R \sin \phi$, and we continue the calculation:

$$
\begin{gathered}
I=\delta R^{2} \int_{\theta=0}^{\theta=2 \pi} \int_{\phi=0}^{\phi=\pi}(R \sin \phi)^{2} \sin \phi d \phi d \theta=\delta R^{4} 2 \pi \int_{\phi=0}^{\phi=\pi} \sin ^{3} \phi d \phi \\
=\delta R^{4} 2 \pi\left[\frac{1}{3} \sin ^{2} \phi \cos \phi+\frac{2}{3} \cos \phi\right]_{\phi=0}^{\phi=\pi}=\delta R^{4} 2 \pi \frac{4}{3}=\frac{M}{4 \pi R^{2}} R^{4} 2 \pi \frac{4}{3}=\frac{8 \pi M R^{4}}{12 \pi R^{2}}=\frac{2}{3} M R^{2}
\end{gathered}
$$

## Exercise 2.4: Jupiter

(a) We can write the rotational inertia of both a sphere with equal density and a planet as $\int_{V} \frac{d m}{d V} r^{2} d V$. In the case of the planet, $\frac{d m}{d V}$ is a constant density, but in the case of the planet, this term is large when $r$ is small and small when $r$ is large. Thus, those mass infinitesimals contribute less to the total value of the integral, since they are integrated at a lower value of $r$ than in the case of the sphere. Thus, the moment of inertia is smaller for a planet than for a sphere of uniform density of the same mass and radius.
(b) We use a spherical coordinate system, and align the z-axis with the axis of rotation. The variable density is given by $\frac{d m}{d V}=\frac{M}{R^{3}} \frac{\sin (\pi r / R)}{4 r / R}$.

$$
\begin{gathered}
I=\int_{M} r_{\perp}^{2} d m=\int_{V} r_{\perp}^{2} \frac{d m}{d V} d V=\int_{r=0}^{r=R} \int_{\theta=0}^{\theta=2 \pi} \int_{\phi=0}^{\phi=\pi}\left(r_{\perp}^{2}\right)\left(\frac{M}{R^{3}} \frac{\sin (\pi r / R)}{4 r / R}\right)\left(r^{2} \sin \phi d \phi d \theta d r\right) \\
=\frac{M}{4 R^{2}} \int_{r=0}^{r=R} \int_{\theta=0}^{\theta=2 \pi} \int_{\phi=0}^{\phi=\pi}\left(r_{\perp}^{2}\right) r \sin (\pi r / R) \sin \phi d \phi d \theta d r \\
=2 \pi \frac{M}{4 R^{2}} \int_{r=0}^{r=R} \int_{\phi=0}^{\phi=\pi}\left(r_{\perp}^{2}\right) r \sin (\pi r / R) \sin \phi d \phi d r
\end{gathered}
$$

Since the axis of rotation is the z -axis, the perpendicular distance is just $r_{\perp}=r \sin \phi$, and we continue the calculation:

$$
\begin{aligned}
= & 2 \pi \frac{M}{4 R^{2}} \int_{r=0}^{r=R} \int_{\phi=0}^{\phi=\pi}\left(r^{2} \sin ^{2} \phi\right) r \sin (\pi r / R) \sin \phi d \phi d r \\
& =2 \pi \frac{M}{4 R^{2}} \int_{r=0}^{r=R} \int_{\phi=0}^{\phi=\pi} \sin ^{2} \phi r^{3} \sin (\pi r / R) \sin \phi d \phi d r
\end{aligned}
$$

$$
=2 \pi \frac{M}{4 R^{2}} \int_{r=0}^{r=R} r^{3} \sin (\pi r / R) d r \int_{\phi=0}^{\phi=\pi} \sin ^{3} \phi d \phi
$$

We know from 2.3 (a) that the integral on the right is $\frac{4}{3}$. We look up the integral on the left on the internet ${ }^{3}$.

$$
\begin{gathered}
=2 \pi \frac{M}{4 R^{2}}\left[\frac{3\left(\left(\frac{\pi}{R}\right)^{2} r^{2}-2\right) \sin (\pi r / R)}{\left(\frac{\pi}{R}\right)^{4}}-\frac{r\left(\left(\frac{\pi}{R}\right)^{2} r^{2}-6\right) \cos (\pi r / R)}{\left(\frac{\pi}{R}\right)^{3}}\right]_{r=0}^{r=R}\left[\frac{4}{3}\right] \\
=2 \pi \frac{M}{3 R^{2}}\left[\left(0-\frac{R\left(\left(\frac{\pi}{R}\right)^{2} R^{2}-6\right) \cos (\pi)}{\left(\frac{\pi}{R}\right)^{3}}\right)-(0-0)\right]=\frac{2 M \pi}{3 R^{2}}\left[R^{4} \frac{\pi^{2}-6}{\pi^{3}}\right] \\
=\frac{2 M R^{2}}{3} \frac{\pi^{2}-6}{\pi^{2}}=\frac{2}{3}\left(1-\frac{6}{\pi^{2}}\right) M R^{2}
\end{gathered}
$$

Which is less than the moment of inertia of a sphere with equal, albeit evenly distributed, mass $-\frac{2}{3} M R^{2}$.

## Exercise 2.5: A constraint on the moments of inertia

Let there be some arrangement of particles $\{\alpha\}$ with mass $\left\{m_{\alpha}\right\}$ and position $r_{\alpha}=r_{\alpha, 1} \hat{e}_{1}+$ $r_{\alpha, 2} \hat{e}_{2}+r_{\alpha, 3} \hat{e}_{3}$ with respect to some set of orthonormal axes $\left\{\hat{e}_{i}\right\}$. Consider the sum of any two moments of inertia, say $I_{11}$ and $I_{22}$, and compare it with $I_{33}$.

$$
\begin{gathered}
I_{11}+I_{22}=\sum_{\alpha} m_{\alpha}\left(\left(\hat{e}_{1} \times r_{\alpha}\right)^{2}+\left(\hat{e}_{2} \times r_{\alpha}\right)^{2}\right)=\sum_{\alpha} m_{\alpha}\left(\left(r_{\alpha, 2}^{2}+r_{\alpha, 3}^{2}\right)+\left(r_{\alpha, 1}^{2}+r_{\alpha, 3}^{2}\right)\right) \\
I_{33}=\sum_{\alpha} m_{\alpha}\left(\hat{e}_{3} \times r_{\alpha}\right)^{2}=\sum_{\alpha} m_{\alpha}\left(r_{\alpha, 1}^{2}+r_{\alpha, 2}^{2}\right)
\end{gathered}
$$

So $I_{11}+I_{22} \geq I_{33}$. (In particular, there is only equality if $\sum_{\alpha} r_{\alpha, 3}^{2}=0$.) By symmetry, this is true for any choice of orthonormal axes. The sum of any two moments of inertia are greater than or equal to the third.

## Exercise 2.6: Principal moments of inertia

(a) We choose a set of axes, and show that they are the principal axes. Let $(0,0,0)$ be the origin and center of mass, and let $r_{1}=\left(\frac{1}{\sqrt{3}}, 0,-\frac{1}{4} \sqrt{\frac{2}{3}}\right), r_{2}, r_{3}=\left(-\frac{1}{2 \sqrt{3}}, \pm \frac{1}{2},-\frac{1}{4} \sqrt{\frac{2}{3}}\right)$, and $p_{4}=\left(0,0, \frac{3}{4} \sqrt{\frac{2}{3}}\right)$ be the four points. Thus, the distance between any two points is 1 . The orthogonal axis we choose is $(\hat{x}, \hat{y}, \hat{z})$. We intend to calculate the components of the moments of inertia for every combination $(i, j)$ :

$$
I_{i j}=\sum_{\alpha} m_{\alpha}\left(e_{i} \times r_{\alpha}\right) \cdot\left(e_{j} \times r_{\alpha}\right)
$$

Before we do this, however, we determine the cross product of the axes with the different $\left\{r_{i}\right\}$ :

[^1]\[

$$
\begin{gathered}
\hat{x} \times r_{1}=\frac{1}{4} \sqrt{\frac{2}{3}} \hat{y}, \hat{y} \times r_{1}=-\frac{1}{4} \sqrt{\frac{2}{3}} \hat{x}-\frac{1}{\sqrt{3}} \hat{z}, \hat{z} \times r_{1}=\frac{1}{\sqrt{3}} \hat{y} \\
\hat{x} \times r_{2}=\frac{1}{4} \sqrt{\frac{2}{3}} \hat{y}+\frac{1}{2} \hat{z}, \hat{y} \times r_{2}=-\frac{1}{4} \sqrt{\frac{2}{3}} \hat{x}+\frac{1}{2 \sqrt{3}} \hat{z}, \hat{z} \times r_{2}=-\frac{1}{2} \hat{x}-\frac{1}{2 \sqrt{3}} \hat{y} \\
\hat{x} \times r_{3}=\frac{1}{4} \sqrt{\frac{2}{3}} \hat{y}-\frac{1}{2} \hat{z}, \hat{y} \times r_{3}=-\frac{1}{4} \sqrt{\frac{2}{3}} \hat{x}+\frac{1}{2 \sqrt{3}} \hat{z}, \hat{z} \times r_{3}=\frac{1}{2} \hat{x}-\frac{1}{2 \sqrt{3}} \hat{y} \\
\hat{x} \times r_{4}=-\frac{3}{4} \sqrt{\frac{2}{3}} \hat{y}, \hat{y} \times r_{4}=\frac{3}{4} \sqrt{\frac{2}{3}} \hat{x}, \hat{z} \times r_{4}=0
\end{gathered}
$$
\]

We calculate the off-diagonal components $I_{x y}, I_{y z}$ and $I_{x z}$, and let the mass term be a constant.

$$
\begin{aligned}
& I_{x y}=m \sum_{\alpha}\left(\hat{x} \times r_{\alpha}\right) \cdot\left(\hat{y} \times r_{\alpha}\right)=m\left(0+\frac{1}{4 \sqrt{3}}-\frac{1}{4 \sqrt{3}}+0\right)=0 \\
& I_{y z}=m \sum_{\alpha}\left(\hat{y} \times r_{\alpha}\right) \cdot\left(\hat{z} \times r_{\alpha}\right)=m\left(0+\frac{1}{8} \sqrt{\frac{2}{3}}-\frac{1}{8} \sqrt{\frac{2}{3}}+0\right)=0 \\
& I_{x z}=m \sum_{\alpha}\left(\hat{x} \times r_{\alpha}\right) \cdot\left(\hat{z} \times r_{\alpha}\right)=m\left(\frac{\sqrt{2}}{12}-\frac{\sqrt{2}}{24}-\frac{\sqrt{2}}{24}+0\right)=0
\end{aligned}
$$

By symmetry, $I_{y x}=I_{z y}=I_{z x}=0$, and the off-diagonal elements are zero. Thus, our chosen axes are the principal axes. We now find the principal moments of inertia.

$$
\begin{gathered}
I_{x x}=m \sum_{\alpha}\left(\hat{x} \times r_{\alpha}\right) \cdot\left(\hat{x} \times r_{\alpha}\right)=m \frac{1}{16} \frac{2}{3}+2 m\left(\frac{1}{16} \frac{2}{3}+\frac{1}{4}\right)+m \frac{9}{16} \frac{2}{3}=m \\
I_{y y}=m \sum_{\alpha}\left(\hat{y} \times r_{\alpha}\right) \cdot\left(\hat{y} \times r_{\alpha}\right)=m\left(\frac{1}{16} \frac{2}{3}+\frac{1}{3}\right)+2 m\left(\frac{1}{16} \frac{2}{3}+\frac{1}{12}\right)+m \frac{9}{16} \frac{2}{3}=m \\
I_{z z}=m \sum_{\alpha}\left(\hat{z} \times r_{\alpha}\right) \cdot\left(\hat{z} \times r_{\alpha}\right)=m \frac{1}{3}+2 m\left(\frac{1}{4}+\frac{1}{12}\right)+0=m
\end{gathered}
$$

Let $R$ be the distance between the points, or, the scale of the arrangement. We chose $R$ to be 1 earlier, but we can extend this calculation to a general tetrahedral arrangement by multiplying the principal moments of inertia by $R^{2}$. Thus, the principal axes are $\hat{x}, \hat{y}, \hat{z}$, and the corresponding principal moments of inertia are each $m R^{2}$, which is what we wanted to show.
(b) Let the cube with side length $R$ centered at $(0,0,0)$ be described by $-\frac{R}{2} \leq x, y, z \leq \frac{R}{2}$. We will show that $(\hat{x}, \hat{y}, \hat{z})$ are the principal axes. Let the cube have a uniform density $\delta=\frac{M}{R^{3}}$.

First, we compute the off-diagonal inertial moments. Let $r(x, y, z)=x \hat{x}+y \hat{y}+z \hat{z}$ be the vector from the origin to the point at $(x, y, z)$.

$$
I_{x y}=\int_{V} \delta(\hat{x} \times r) \cdot(\hat{y} \times r) d V=\delta \int_{V}(-z \hat{y}+y \hat{z}) \cdot(z \hat{x}-x \hat{z}) d V=\delta \int_{V}-x y d V
$$

But by symmetry, this is zero, and so are the rest of the off-diagonal components. Thus, the axes we have chosen are the principal axes, and we go to calculate the principal moments.

$$
\begin{gathered}
I_{x x}=\int_{V} \delta(\hat{x} \times r) \cdot(\hat{x} \times r) d V=\delta \int_{V}(-z \hat{y}+y \hat{z}) \cdot(-z \hat{y}+y \hat{z}) d V=\delta \int_{V} z^{2}+y^{2} d V \\
=\delta \int_{V} z^{2}+y^{2} d V=\delta R \int_{-R / 2}^{R / 2} \int_{-R / 2}^{R / 2} z^{2}+y^{2} d z d y=2 \delta R^{2} \int_{-R / 2}^{R / 2} z^{2} d z=2 \delta R^{2} \frac{1}{12} R^{3}=\frac{1}{6} M R^{2}
\end{gathered}
$$

Similarly, $I_{y y}=I_{z z}=\frac{1}{6} M R^{2}$.
(c) The center of mass of this arrangement is $\left(\frac{-1+1+1+0+0}{5}, \frac{0+0+1+0+0}{5}, \frac{0+0+0+0+1}{5}\right)=\left(\frac{1}{5}, \frac{1}{5}, \frac{1}{5}\right)$. Our plan of attack is to choose any set of orthogonal axes, and to calculate the tensor matrix $I$. The eigenvalues of the matrix are the principal moments of inertia, and the eigenvectors are the principal axes. Since it contributes only a multiplicative term, we let the mass $m_{\alpha}=1$ for all $\alpha$.

We choose $(\hat{x}, \hat{y}, \hat{z})$ as our axes, and calculate the cross product of each of these with the relative displacement vectors for the points. First we calculate these new vectors:

$$
\begin{gathered}
r_{1}=\left(\frac{4}{5},-\frac{1}{5},-\frac{1}{5}\right), r_{2}=\left(-\frac{6}{5},-\frac{1}{5},-\frac{1}{5}\right), r_{3}=\left(\frac{4}{5}, \frac{4}{5},-\frac{1}{5}\right) \\
r_{4}=\left(-\frac{1}{5},-\frac{1}{5},-\frac{1}{5}\right), r_{5}=\left(-\frac{1}{5},-\frac{1}{5}, \frac{4}{5}\right)
\end{gathered}
$$

Then the crossproducts.

$$
\begin{aligned}
& \hat{x} \times r_{1}=\frac{1}{5} \hat{y}-\frac{1}{5} \hat{z}, \hat{y} \times r_{1}=-\frac{1}{5} \hat{x}-\frac{4}{5} \hat{z}, \hat{z} \times r_{1}=\frac{1}{5} \hat{x}+\frac{4}{5} \hat{y} \\
& \hat{x} \times r_{2}=\frac{1}{5} \hat{y}-\frac{1}{5} \hat{z}, \hat{y} \times r_{2}=-\frac{1}{5} \hat{x}+\frac{6}{5} \hat{z}, \hat{z} \times r_{2}=\frac{1}{5} \hat{x}-\frac{6}{5} \hat{y} \\
& \hat{x} \times r_{3}=\frac{1}{5} \hat{y}+\frac{4}{5} \hat{z}, \hat{y} \times r_{3}=-\frac{1}{5} \hat{x}-\frac{4}{5} \hat{z}, \hat{z} \times r_{3}=-\frac{4}{5} \hat{x}+\frac{4}{5} \hat{y} \\
& \hat{x} \times r_{4}=\frac{1}{5} \hat{y}-\frac{1}{5} \hat{z}, \hat{y} \times r_{4}=-\frac{1}{5} \hat{x}+\frac{1}{5} \hat{z}, \hat{z} \times r_{4}=\frac{1}{5} \hat{x}-\frac{1}{5} \hat{y} \\
& \hat{x} \times r_{5}=-\frac{4}{5} \hat{y}-\frac{1}{5} \hat{z}, \hat{y} \times r_{5}=\frac{4}{5} \hat{x}+\frac{1}{5} \hat{z}, \hat{z} \times r_{5}=\frac{1}{5} \hat{x}-\frac{1}{5} \hat{y}
\end{aligned}
$$

Now we calculate entries in the matrix $I$, using $I_{i j}=\sum_{\alpha}\left(\hat{e}_{i} \times r_{\alpha}\right) \cdot\left(\hat{e}_{i} \times r_{\alpha}\right)$.

$$
\begin{aligned}
& I_{x x}=\frac{2}{25}+\frac{2}{25}+\frac{17}{25}+\frac{2}{25}+\frac{17}{25}=\frac{8}{5} \\
& I_{y y}=\frac{17}{25}+\frac{37}{25}+\frac{17}{25}+\frac{2}{25}+\frac{17}{25}=\frac{18}{5} \\
& I_{z z}=\frac{17}{25}+\frac{37}{25}+\frac{32}{25}+\frac{2}{25}+\frac{2}{25}=\frac{18}{5}
\end{aligned}
$$

And the off-diagonal moments.

$$
\begin{aligned}
& I_{x y}=\frac{4}{25}-\frac{6}{25}-\frac{16}{25}-\frac{1}{25}-\frac{1}{25}=\frac{4}{5} \\
& I_{y z}=-\frac{1}{25}-\frac{1}{25}+\frac{4}{25}-\frac{1}{25}+\frac{4}{25}=\frac{1}{5} \\
& I_{x z}=\frac{4}{25}-\frac{6}{25}+\frac{4}{25}-\frac{1}{25}+\frac{4}{25}=\frac{1}{5}
\end{aligned}
$$

Since the $I^{T}=I$, we have

$$
I=\frac{1}{5}\left(\begin{array}{ccc}
8 & 4 & 1 \\
4 & 18 & 1 \\
1 & 1 & 18
\end{array}\right)
$$

I used Scientific Python (SciPy) to solve for the eigenvalues and eigenvectors of this matrix.

```
>>> import scipy; import scipy.linalg
>>> scipy.linalg.eig(1/5. * scipy.array([[8,4,1],[4,18,1],[1,1,18]]))
(array([ 1.31273832+0.j, 4.03274534+0.j, 3.45451633+0.j]),
    array([[-0.94401077, 0.30599901, -0.12332182],
    [ 0.3254508 , 0.80245031, -0.50015525],
    [ 0.05408738, 0.51228712, 0.85710936]]))
```

Thus, the principal axes are, to three decimal places, $-.944 \hat{x}+.305 \hat{y}+-.123 \hat{z}, .325 \hat{x}+.802 \hat{y}+$ $-.500 \hat{z}$, and $.054 \hat{x}+.512 \hat{y}+.857 \hat{x}$, with associated principal moments of inertia $1.313 \mathrm{~m}, 4.033 \mathrm{~m}$ and 3.455 m .


[^0]:    ${ }^{1}$ This only contains the first four of six problems. I'm turning the other two problems in on Tuesday.
    ${ }^{2}$ We use the easily verifiable fact that $\int \sin ^{3}(x) d x=\frac{2}{3} \sin ^{2}(x) \cos (x)+\frac{1}{3} \cos (x)$.

[^1]:    ${ }^{3}$ http://integrals.wolfram.com/index.jsp

