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# SAMI : Sydney-AAO Multi-object Integral-field spectrograph

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**University of Sydney  
AAO  
CAASTRO**



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# Outline

## The SAMI instrument:

- **Hexabundles:** The key to SAMI.
- Lessons learnt about fibres.
- SAMI survey specifications and performance.



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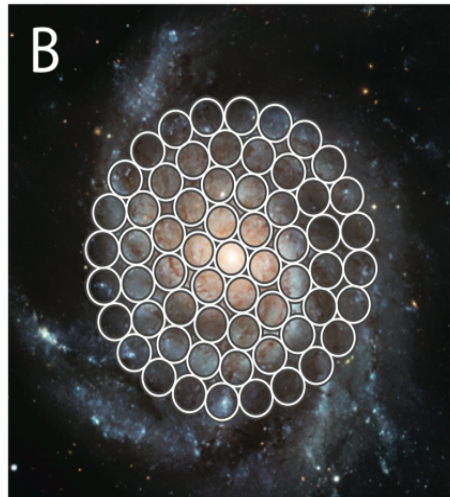
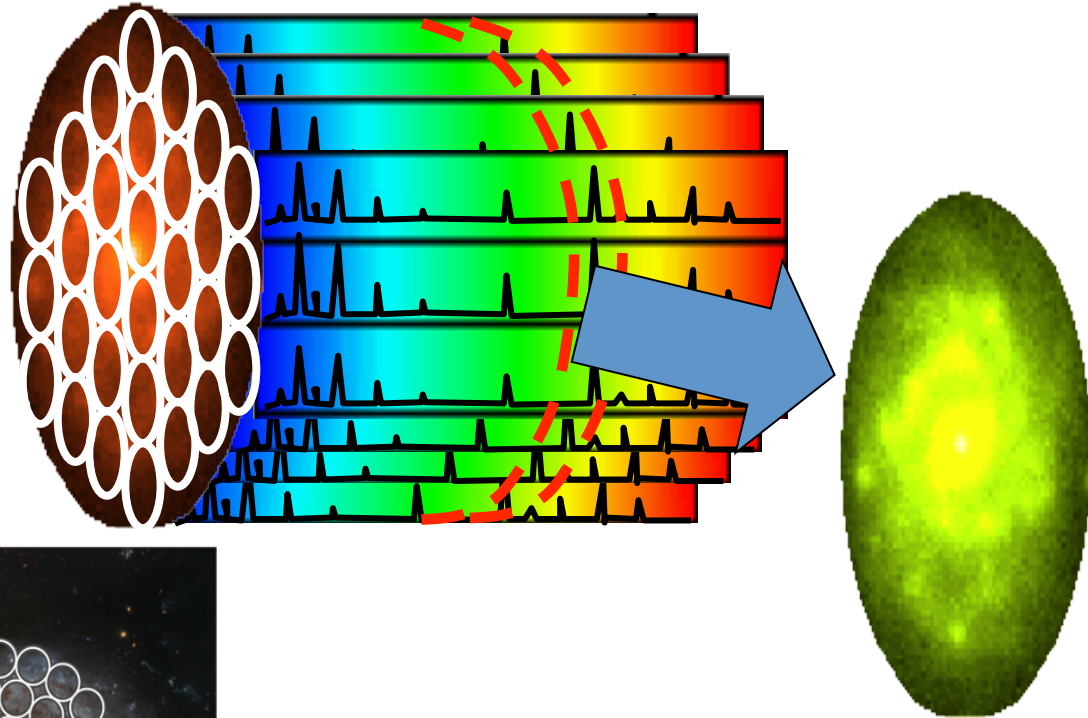


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# Hexabundles- optical fibre IFUs

Require multiple positionable IFUs in place of single fibres.

**Hexabundles** –optical  
fibre IFU, giving a  
spectrum from each  
optical fibre core.





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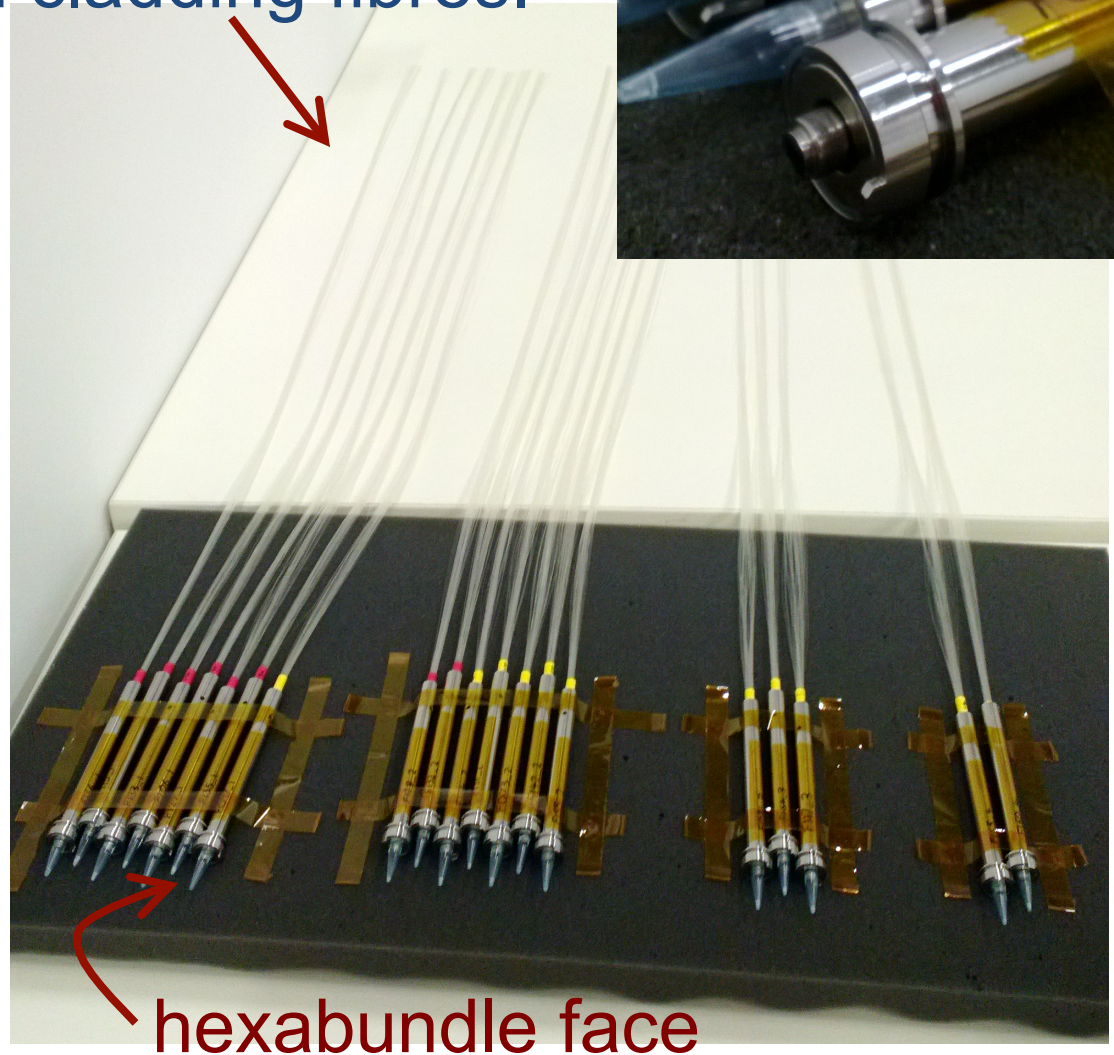


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# Hexabundles for SAMI

105 $\mu$ m core/125  $\mu$ m cladding fibres.

105  125  
 $\mu$ m  $\mu$ m







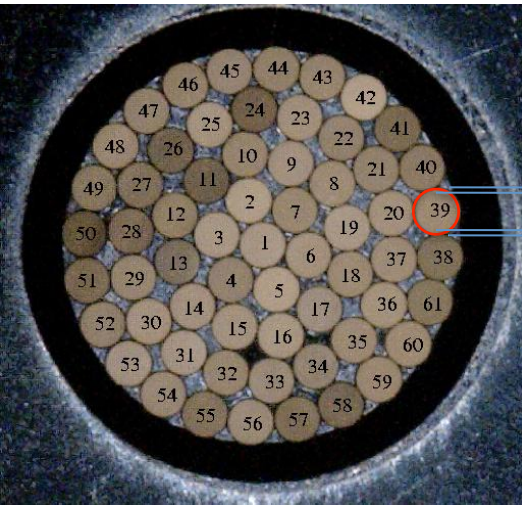
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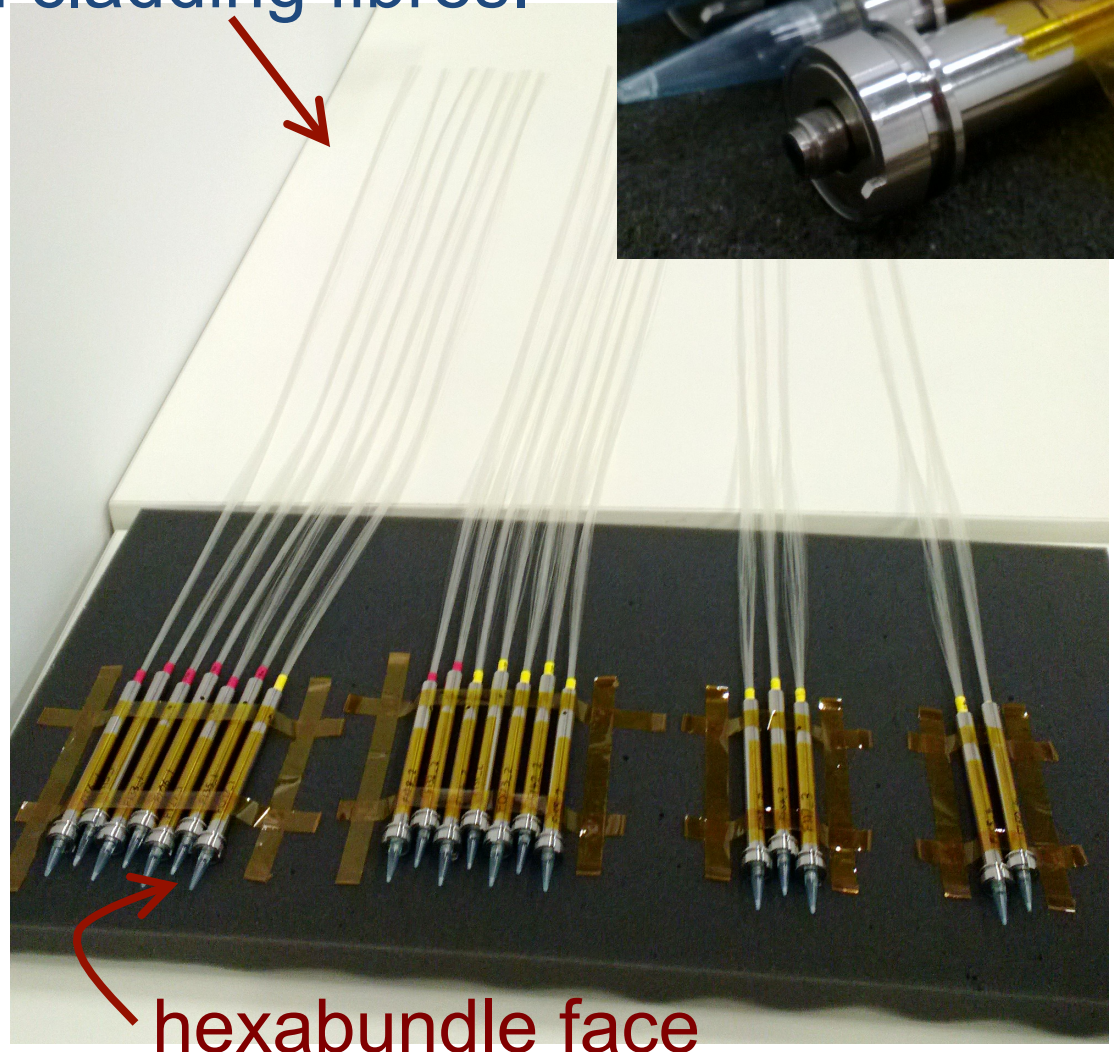
# Hexabundles for SAMI

105 $\mu$ m core/125  $\mu$ m cladding fibres.



105 $\mu$ m		125 $\mu$ m
105 $\mu$ m		115 $\mu$ m

61 optical fibres fused together using our glass fibre processing unit at Sydney Uni.





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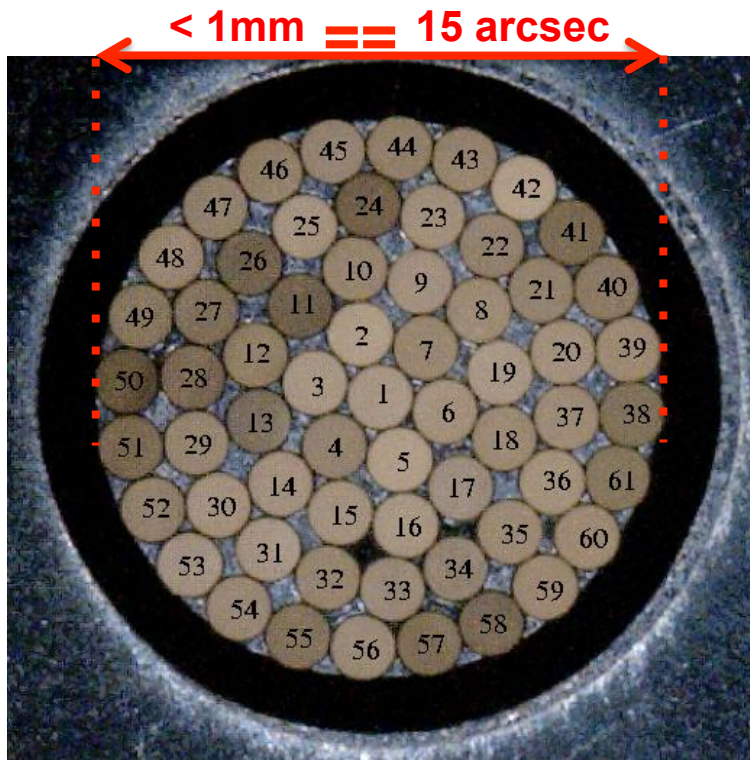


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# Hexabundles for SAMI

## Advantages of this design:

- 1mm diameter



**Bryant et al. 2014, 2012a, 2012b, 2011, 2010.**





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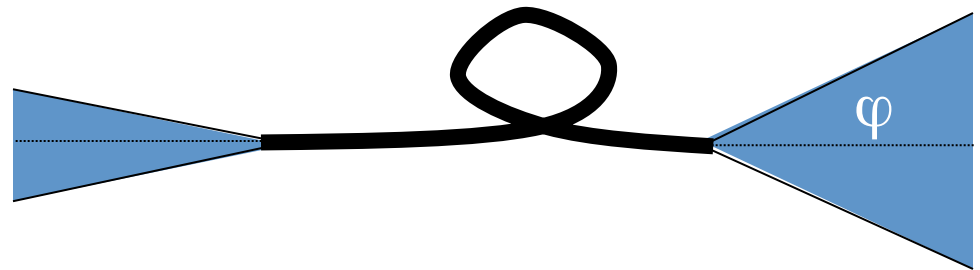
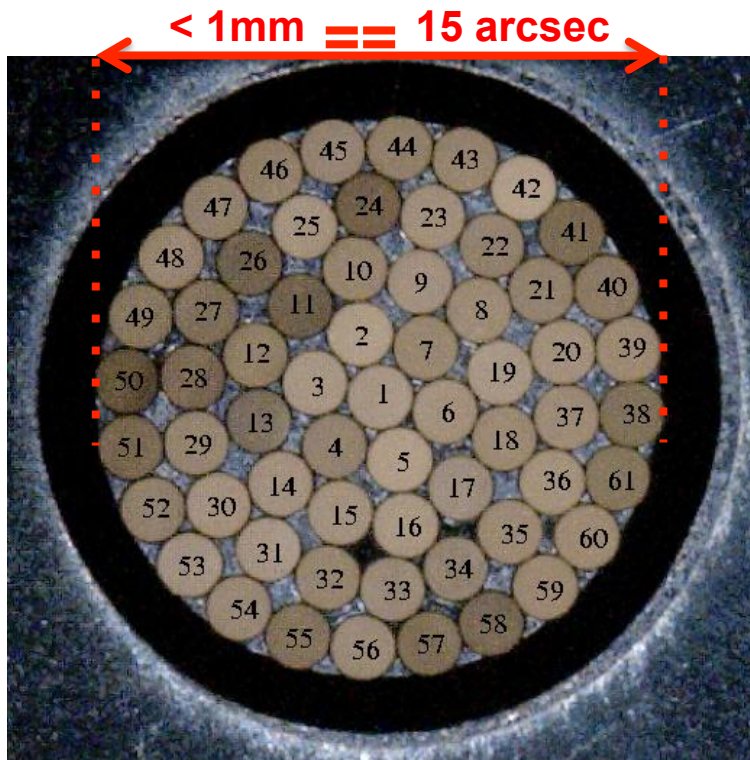


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# Hexabundles for SAMI

## Advantages of this design:

- 1mm diameter
- Low Focal Ratio Degradation (FRD) ....as good as for single fibres.



f/3.4 input hexabundles

f/3.15 acceptance into AAOmega spectrograph

**Bryant et al. 2014, 2012a, 2012b, 2011, 2010.**



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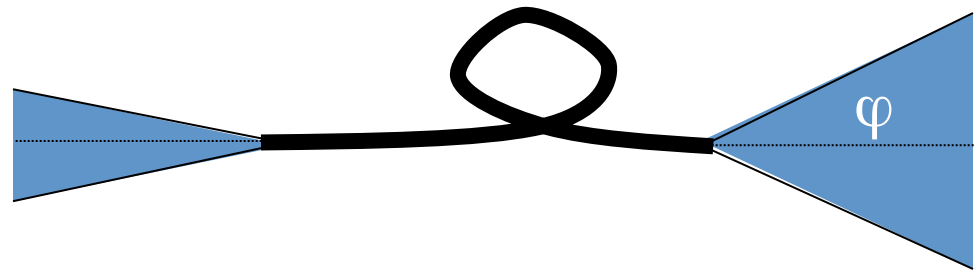
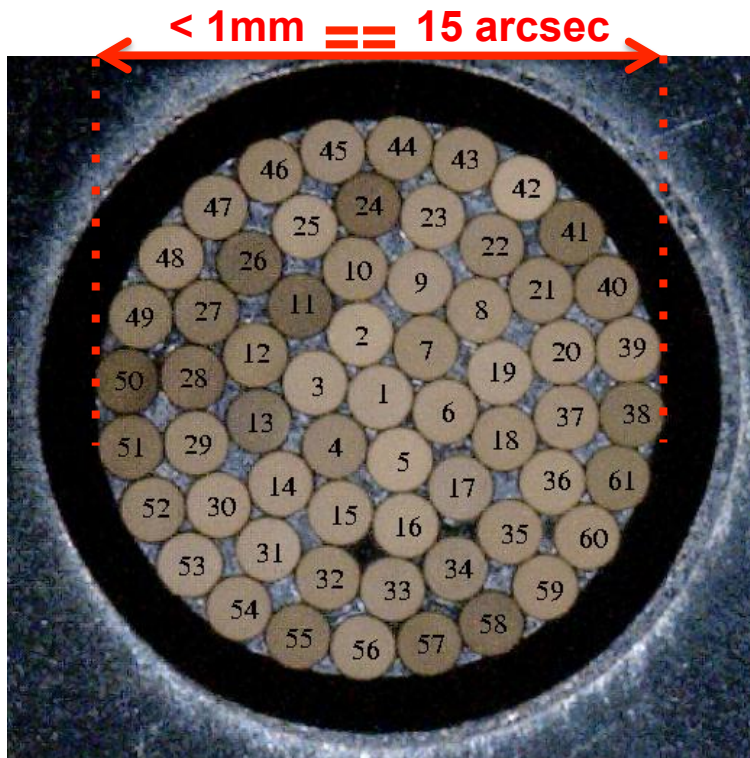


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# Hexabundles for SAMI

## Advantages of this design:

- 1mm diameter
- Low Focal Ratio Degradation (FRD) ....as good as for single fibres.
- Fill fraction (Area of fibre cores/total bundle area)  $\sim 75\%$ .
- Low cross-talk  $< 0.5\%$



f/3.4 input hexabundles

f/3.15 acceptance into AAOmega spectrograph

**Bryant et al. 2014, 2012a, 2012b, 2011, 2010.**



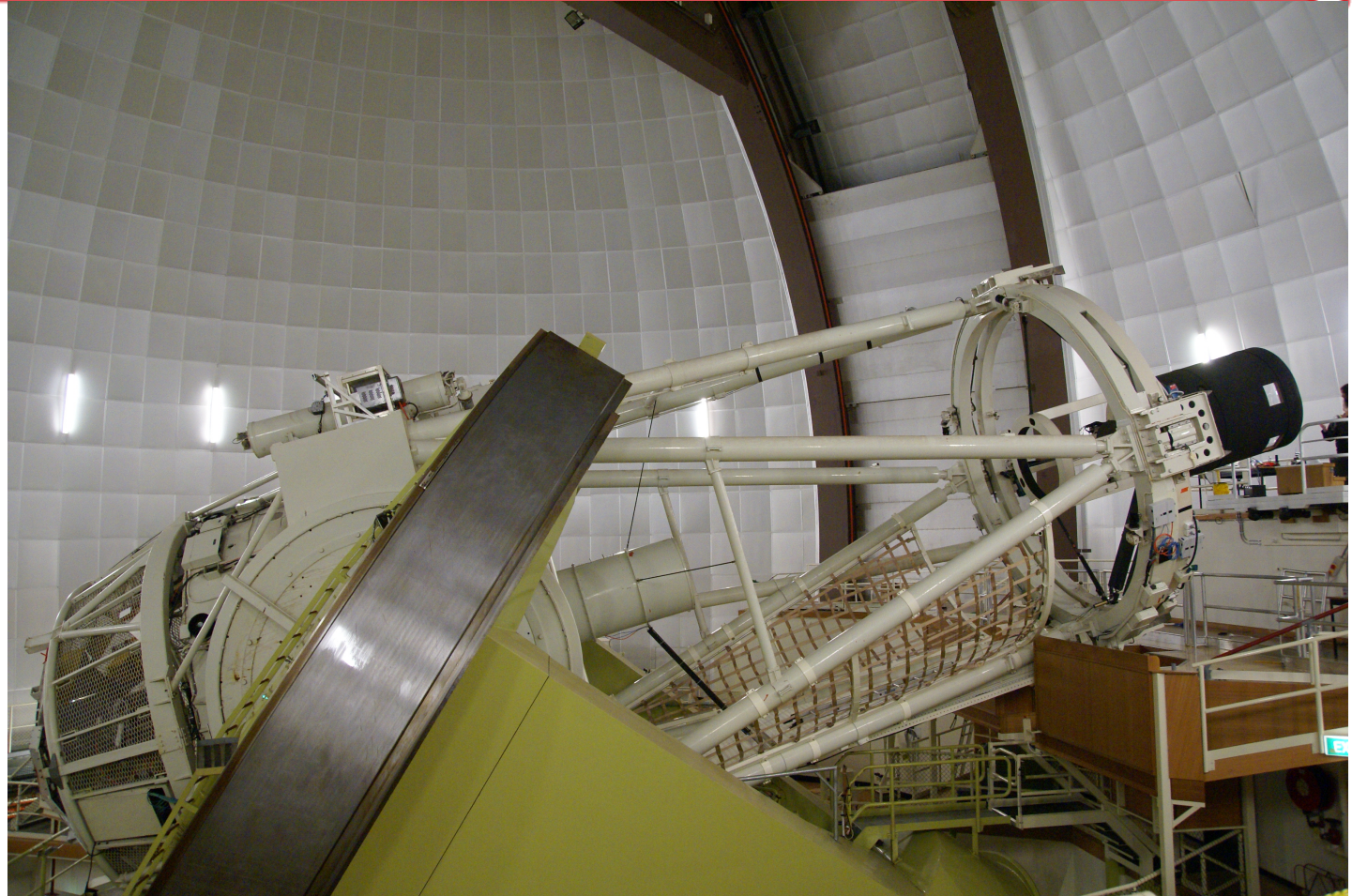


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# SAMI Sydney-AAO Multi-object Integral-field spectrograph





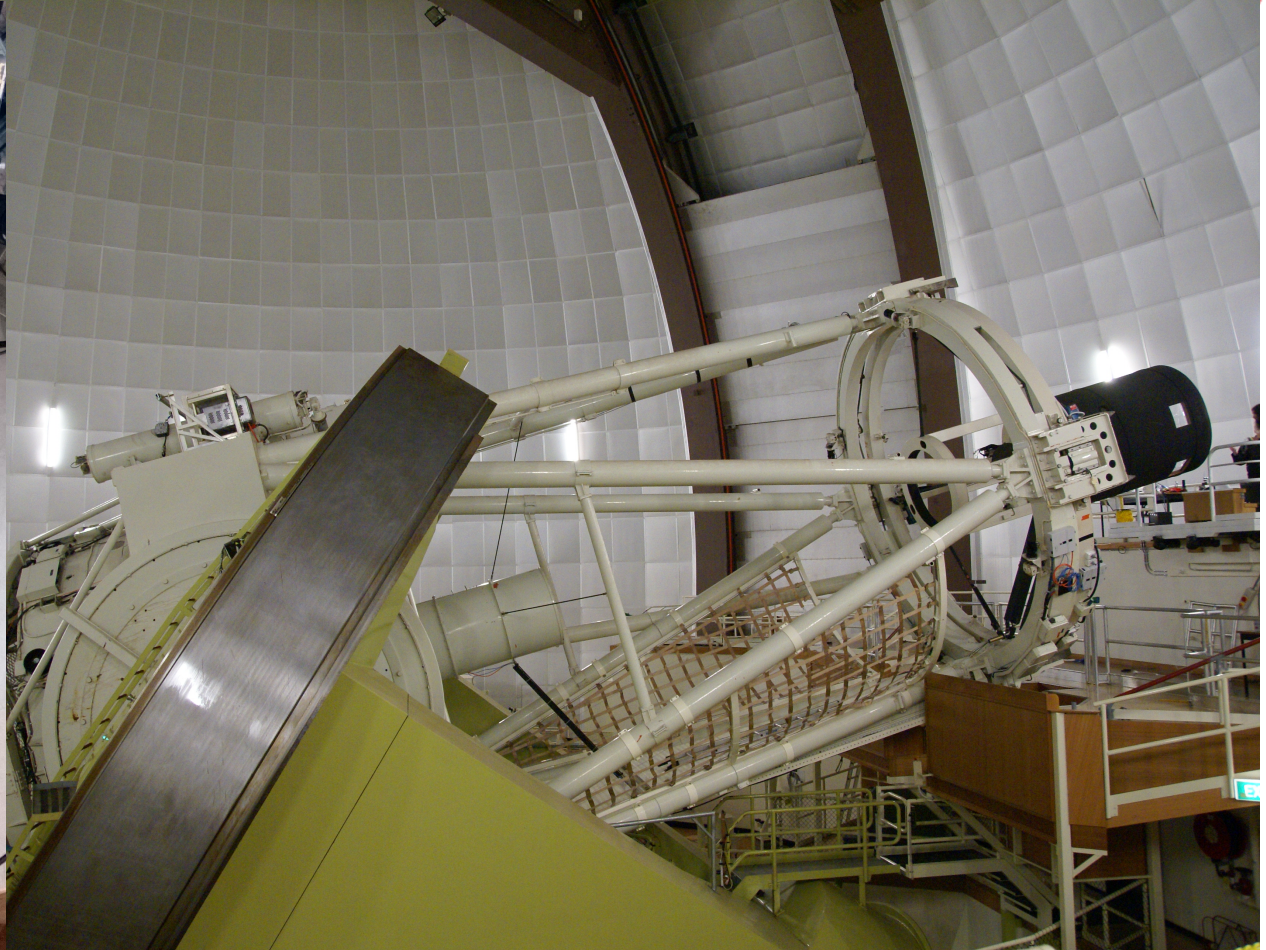
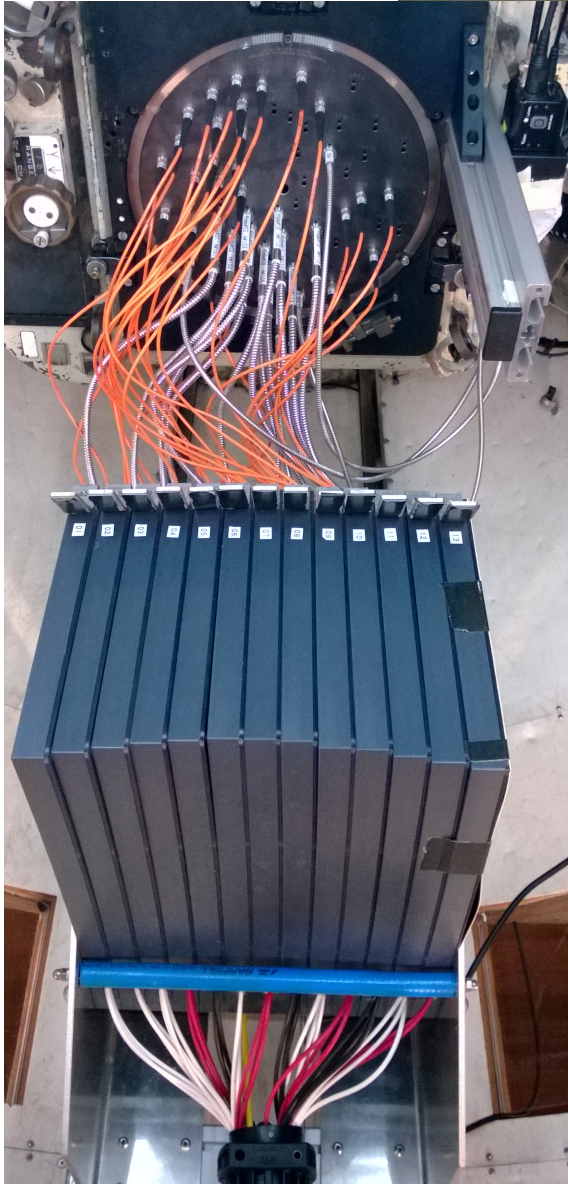


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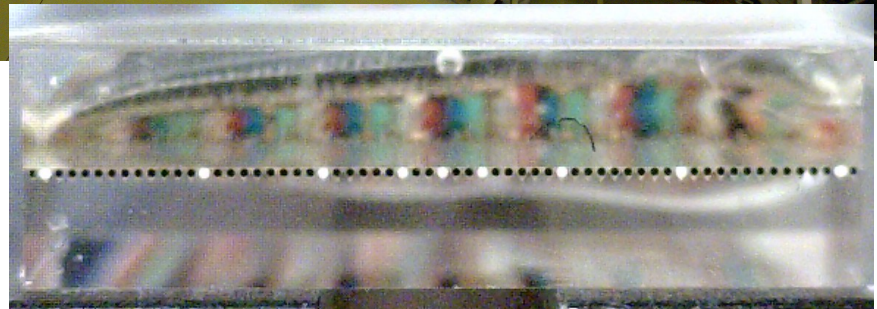
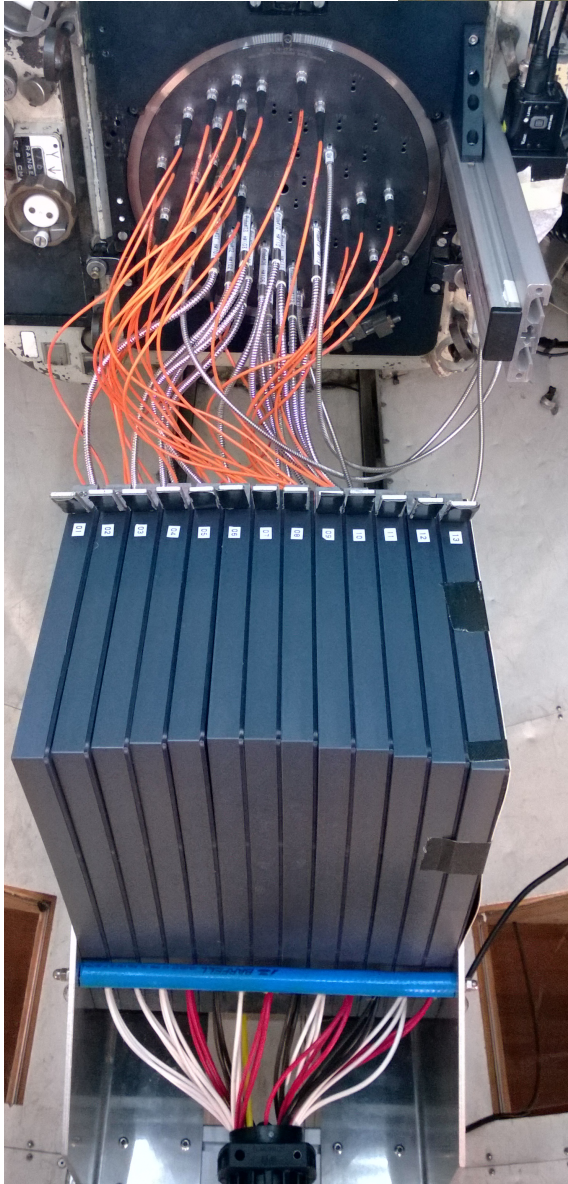
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# SAMI Sydney-AAO Multi-object Integral-field spectrograph

- AAOmega slit allows 819 fibres.  
13 bundles + 26 sky fibres in  $1^\circ$  field.







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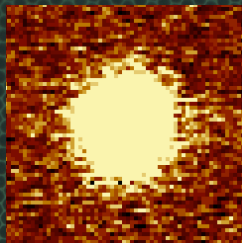
# SAMI guide bundles

- Coherent Polymer imaging bundles



- 7095 cores, 16 microns each.
- f.o.v 22.8" (0.081"/pix)
- 70-80% throughput
- $V < 14$  mag stars
- Fitted into same connector as hexabundle to plug into plate.
- Guide stars can be anywhere in the 1 degree field.
- Allows fast telescope focus.
- Eliminates flexure issues.

200  $\mu\text{m}$





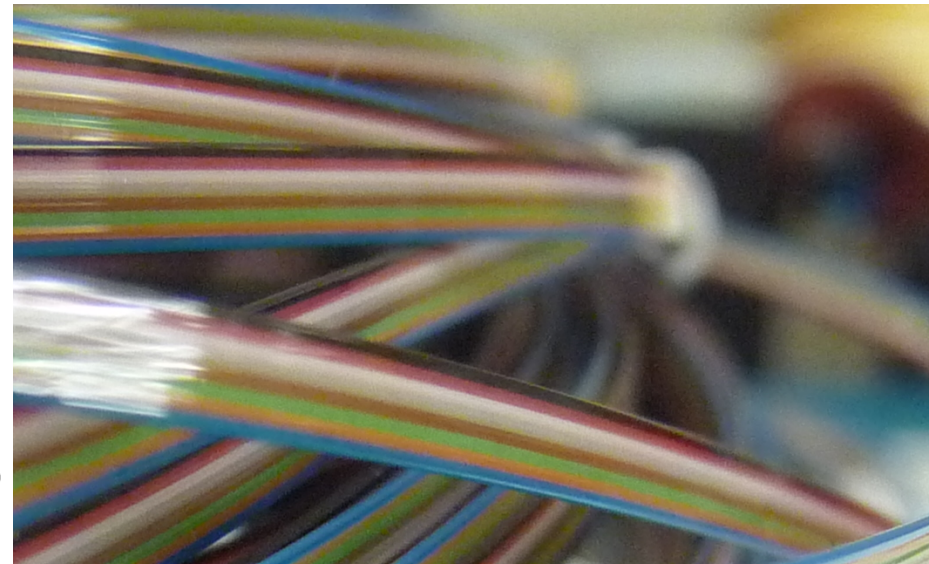
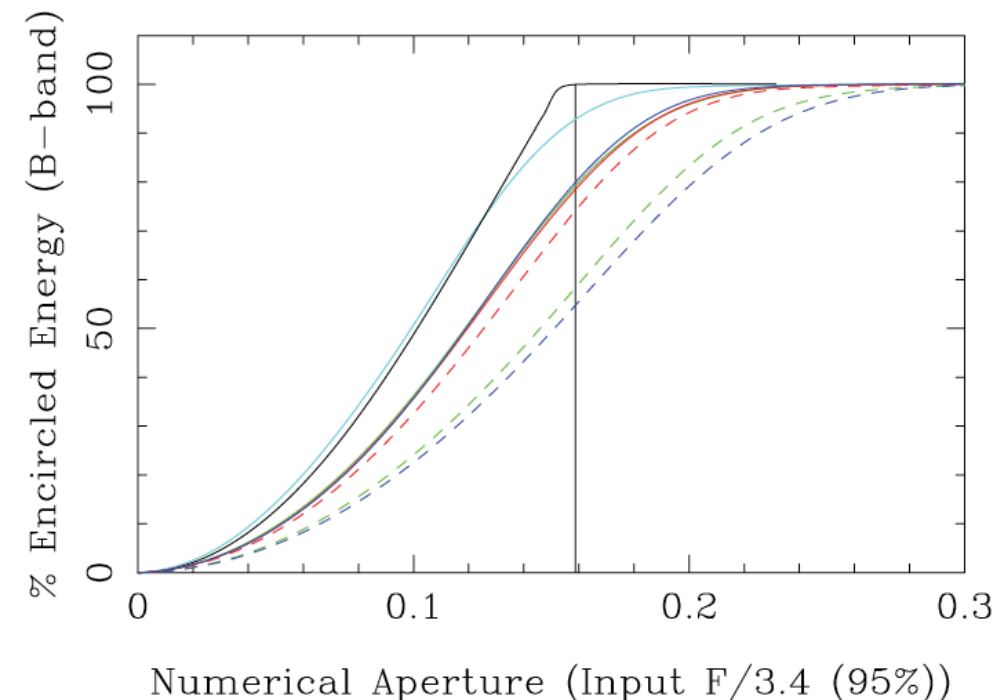
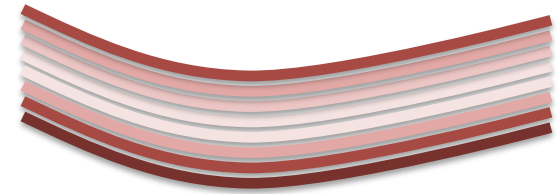
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# What did we learn?

- Ribbonising fibres is a bad idea
  - Ribbonising causes FRD loss due to differential stress.
  - >50% loss in throughput due to FRD





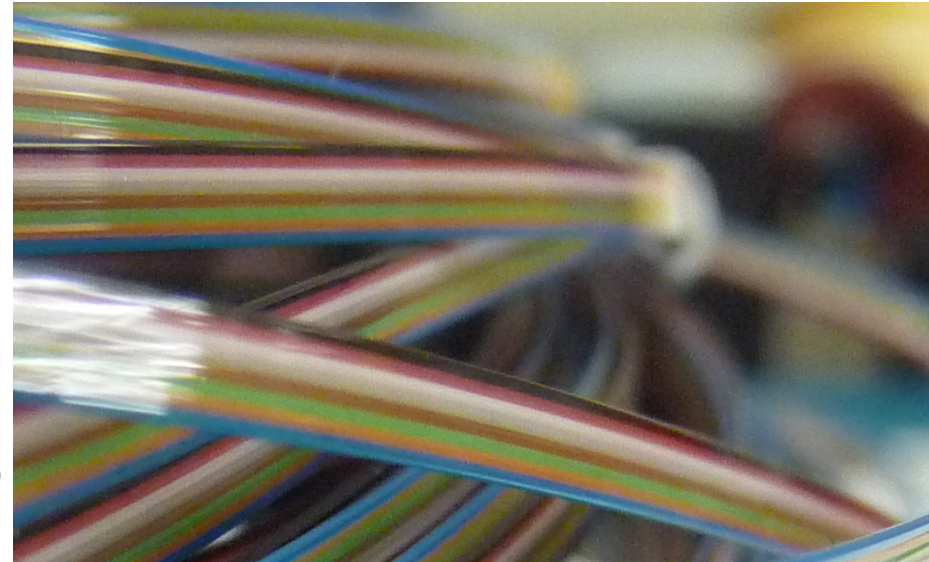
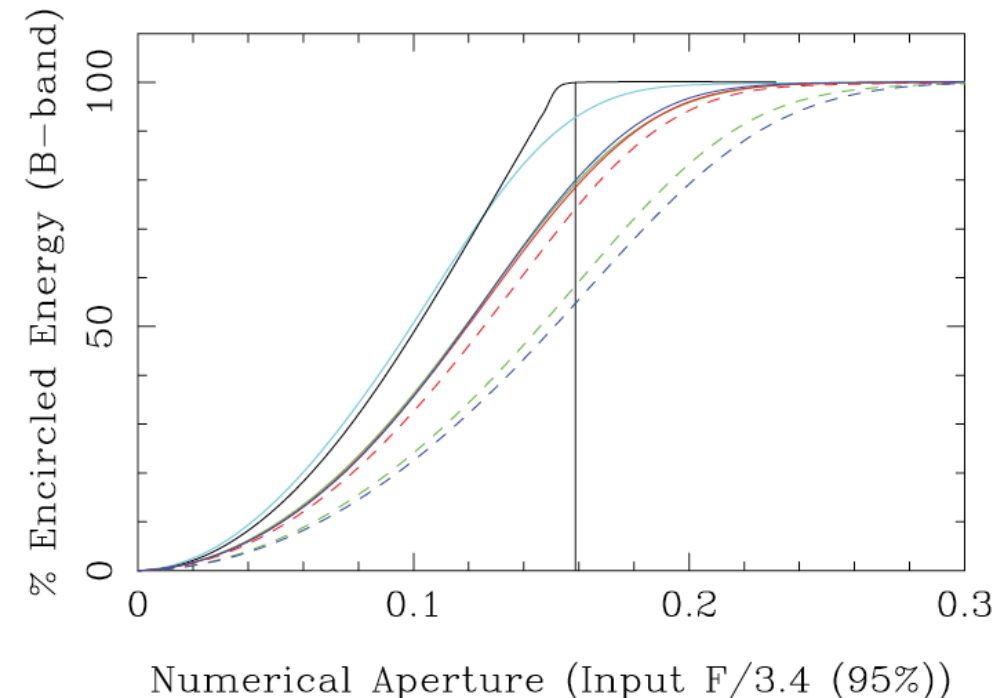
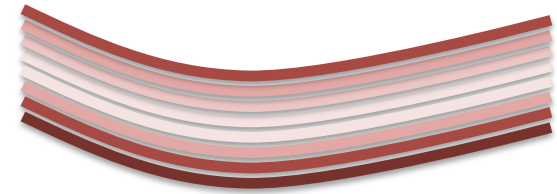
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# What did we learn?

- Ribbonising fibres is a bad idea
  - Ribbonising causes FRD loss due to differential stress.
  - >50% loss in throughput due to FRD
- Hytrel coating increases FRD





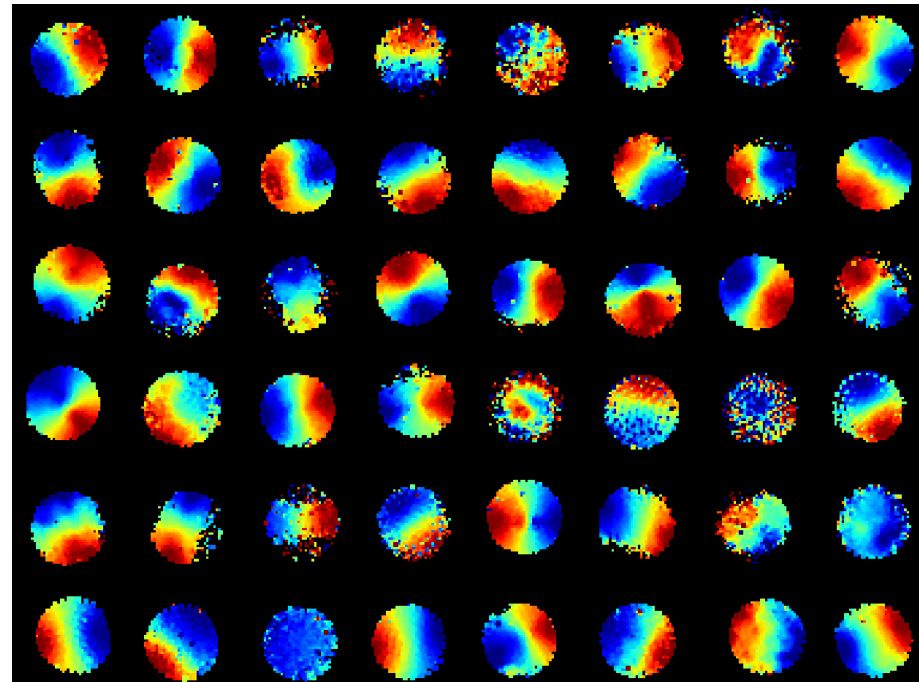
<http://sami-survey.org/>

- SAMI Galaxy Survey team of >60 astronomers from Australia and internationally.
- Key science:
  - physical processes driving galaxy transformations,
  - build up of mass and angular momentum in galaxies,
  - how does gas get into/out of galaxies (feedback and feeding)
- The galaxy survey began in March 2013.
- 151-181 nights awarded – from 2013B-16A.
- 3400 galaxies = 2800 field/group galaxies + 600 cluster targets

Croom et al. 2012

**>1000 observed already!**

**Public data release of 107 galaxies available**  
<http://sami-survey.org/edr>





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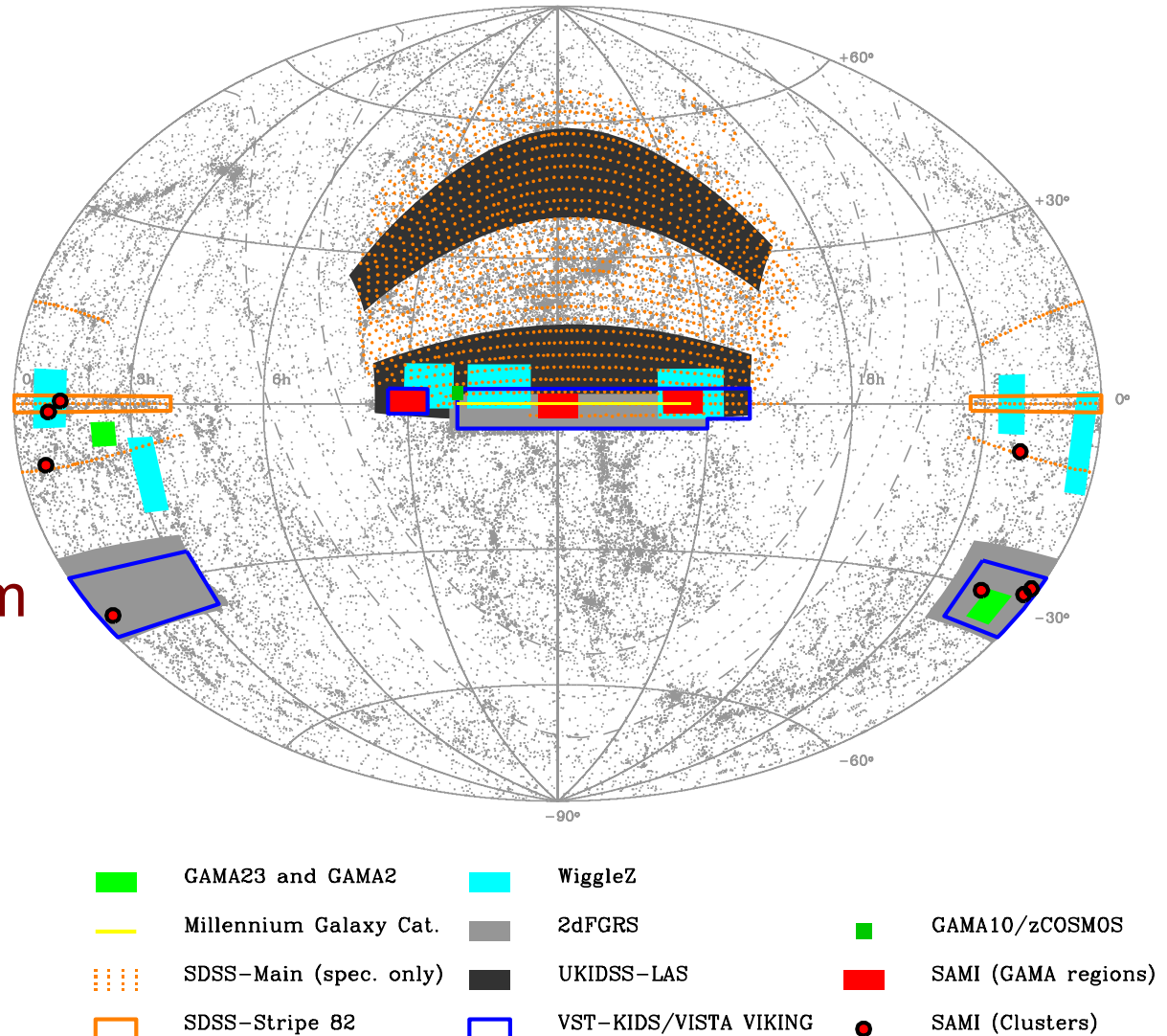
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# SAMI Galaxy Survey

## Target selection

**Aim: Broad range in environment and stellar mass**

- SAMI target density: 12 hexabundles; 1 degree field.
- Field/group galaxies from GAMA
- Cluster galaxies from 8 clusters.
- Ancillary data



SAMI Target selection paper: Bryant et al. 2014



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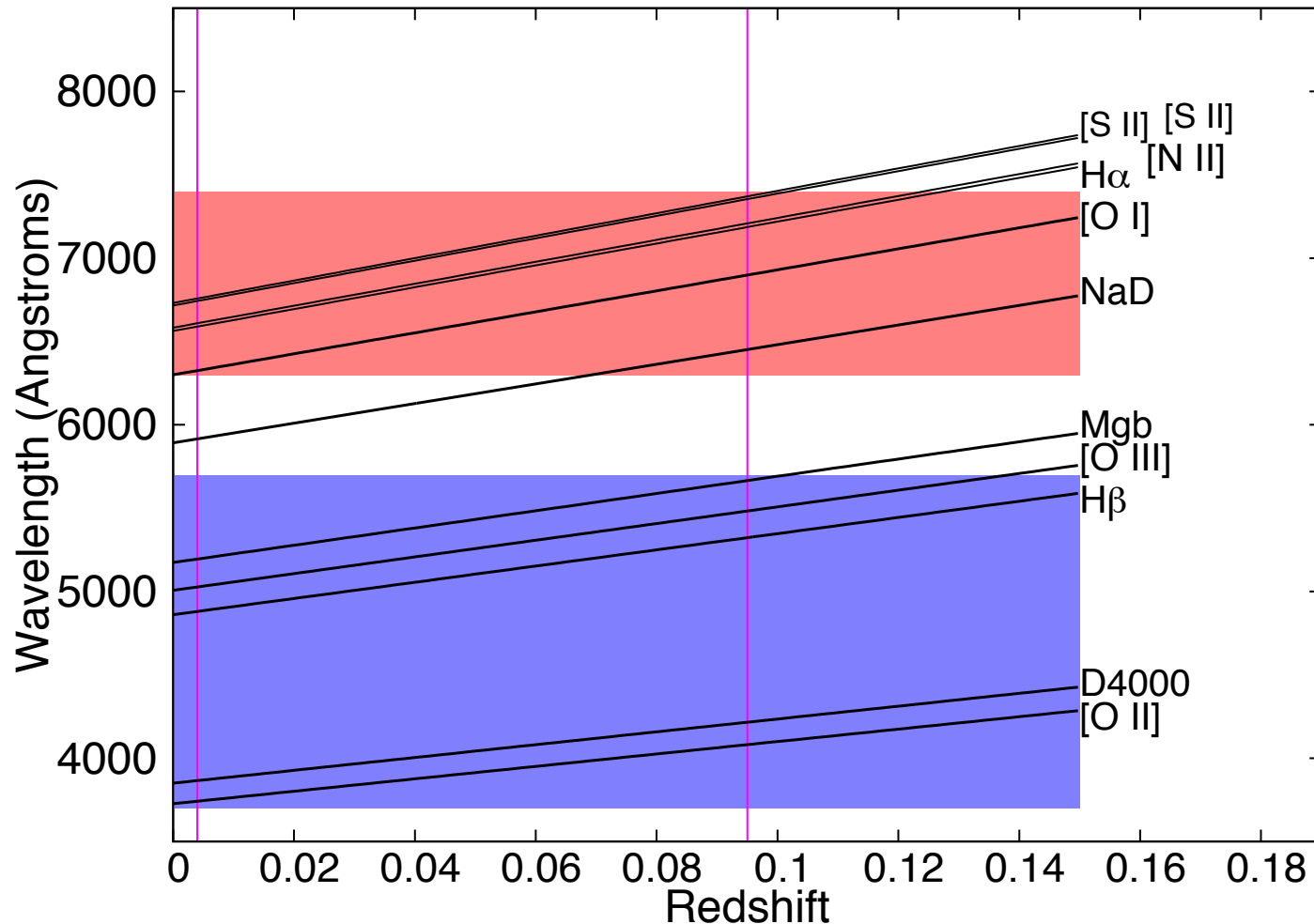


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# SAMI Galaxy Survey

## Instrument setup

- Galaxy features vs. redshift:



- 1000R grating,  
6300-7400Å,  
R=4500
- 580V grating,  
3700-5600Å,  
R=1700



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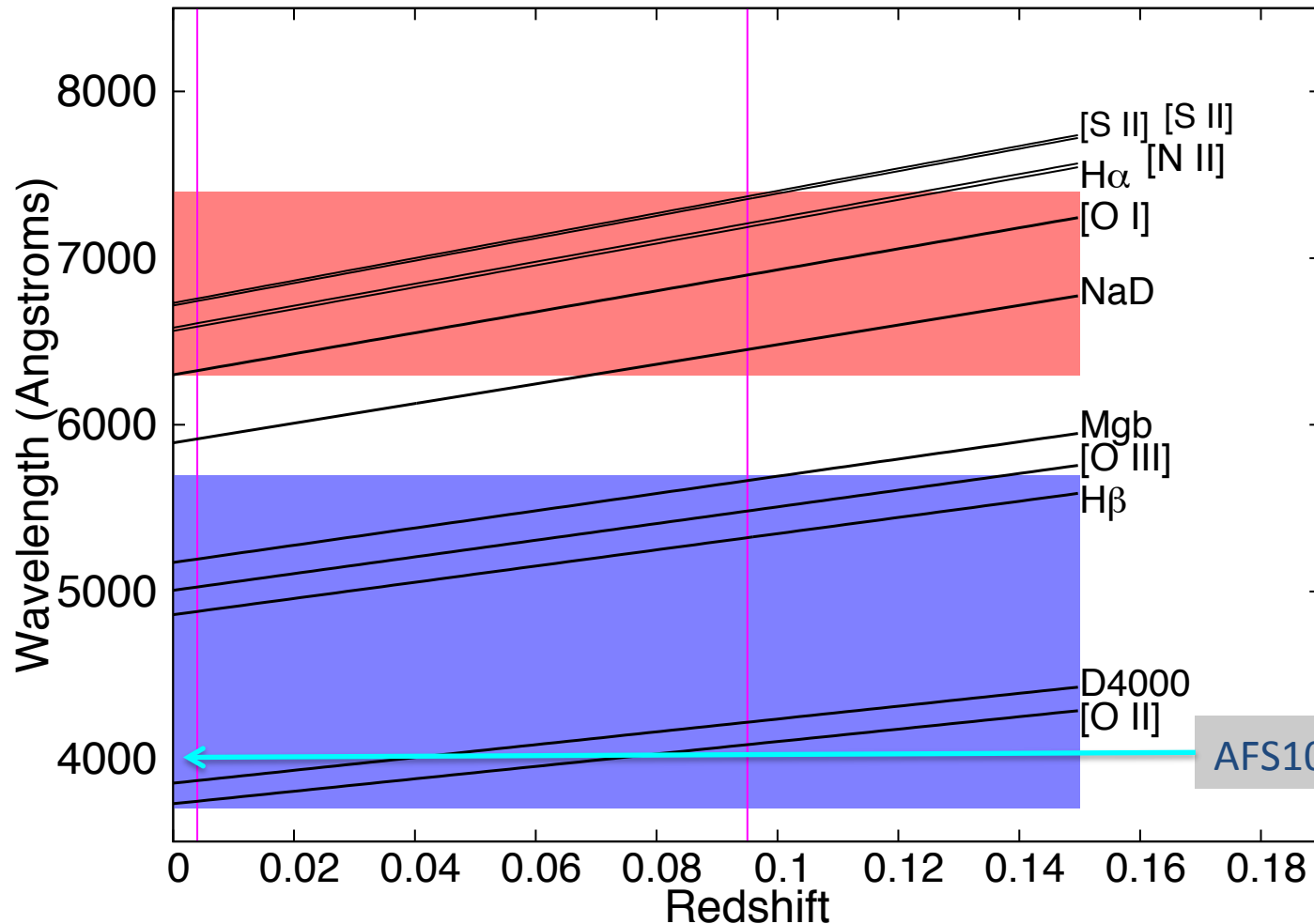


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# SAMI Galaxy Survey

## Instrument setup

- Galaxy features vs. redshift:



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AFS105/125 60% throughput



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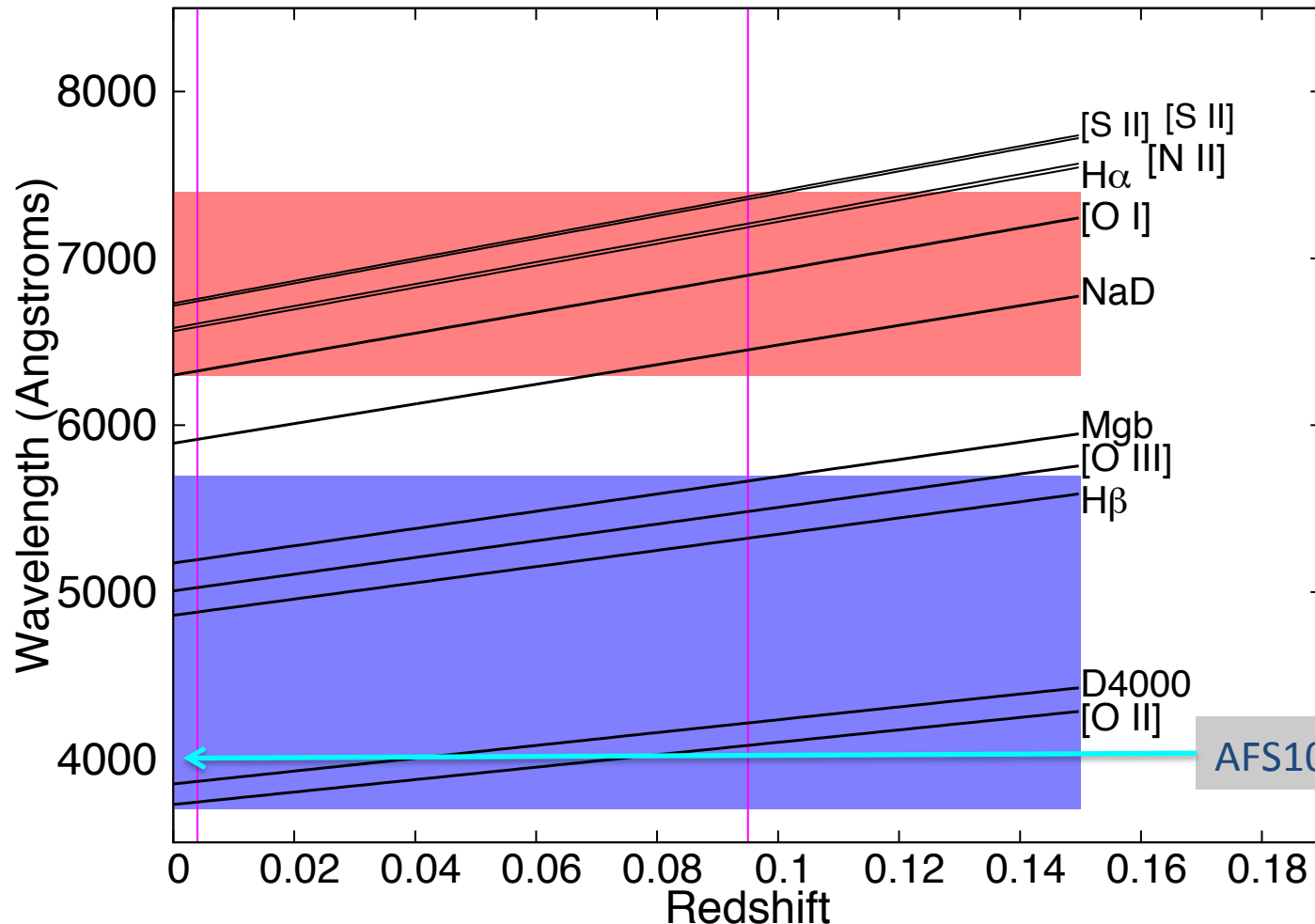


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# SAMI Galaxy Survey

## Instrument setup

- Galaxy features vs. redshift:



- 1000R grating,  
6300-7400A,  
R=4500
- 580V grating,  
3700-5600A,  
R=1700

AFS105/125 60% throughput

WF105/125Y >80%





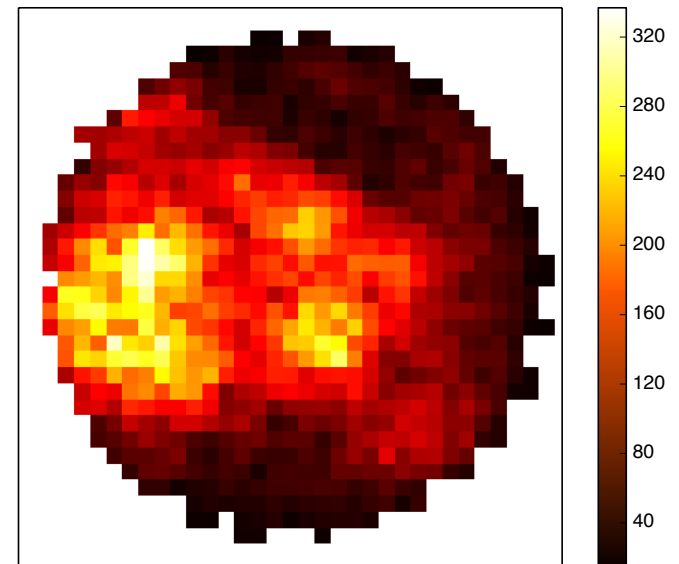
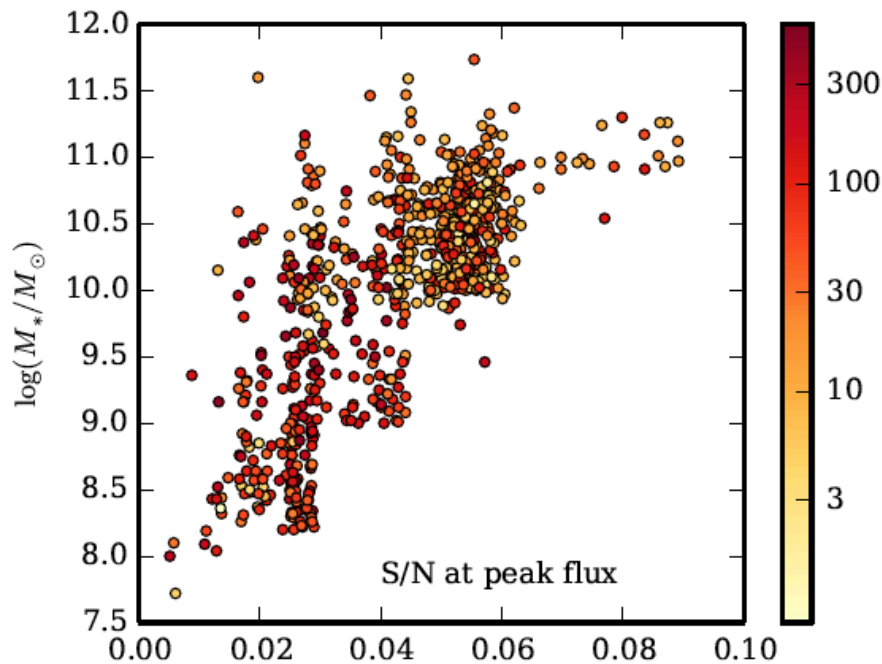
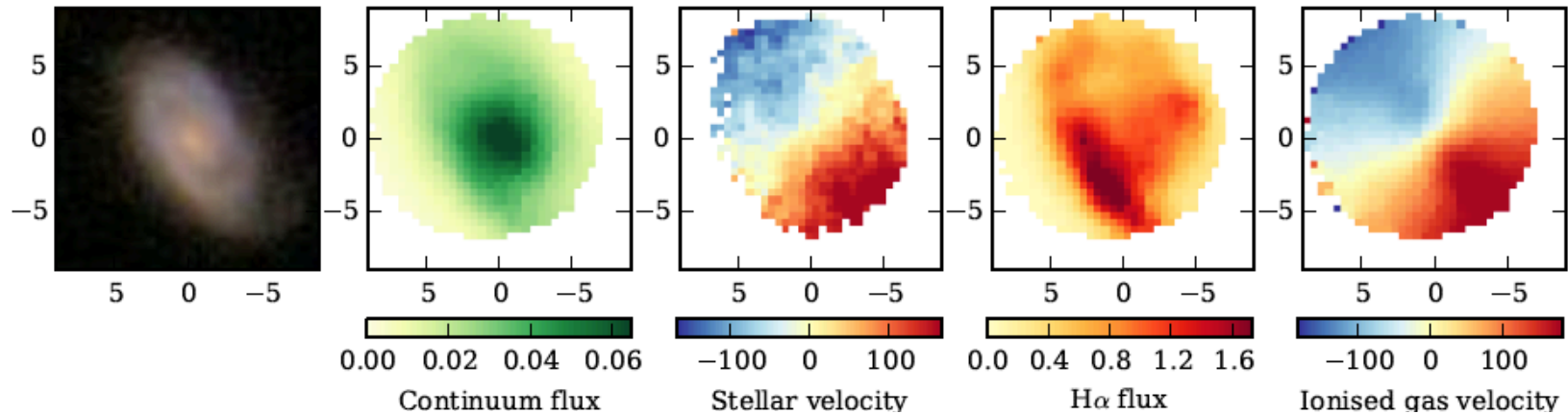
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# SAMI Galaxy Survey

## Galaxy data



Sharp et al. 2014, Allen et al. 2014

Bryant et al. 2014



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# Conclusion

## The SAMI instrument:

- **Hexabundles** have been successfully demonstrated on sky.
- The SAMI Galaxy Survey is nearly 1/3<sup>rd</sup> through with 1000 galaxies observed and 107 in a public data release.
- Minimizing FRD in the hexabundles and fibre cables was the biggest challenge.
- Science papers are now coming out (Fogarty+, Ho+, Sharp+, Allen+, Bryant+, Richards+, Cortese+.....)



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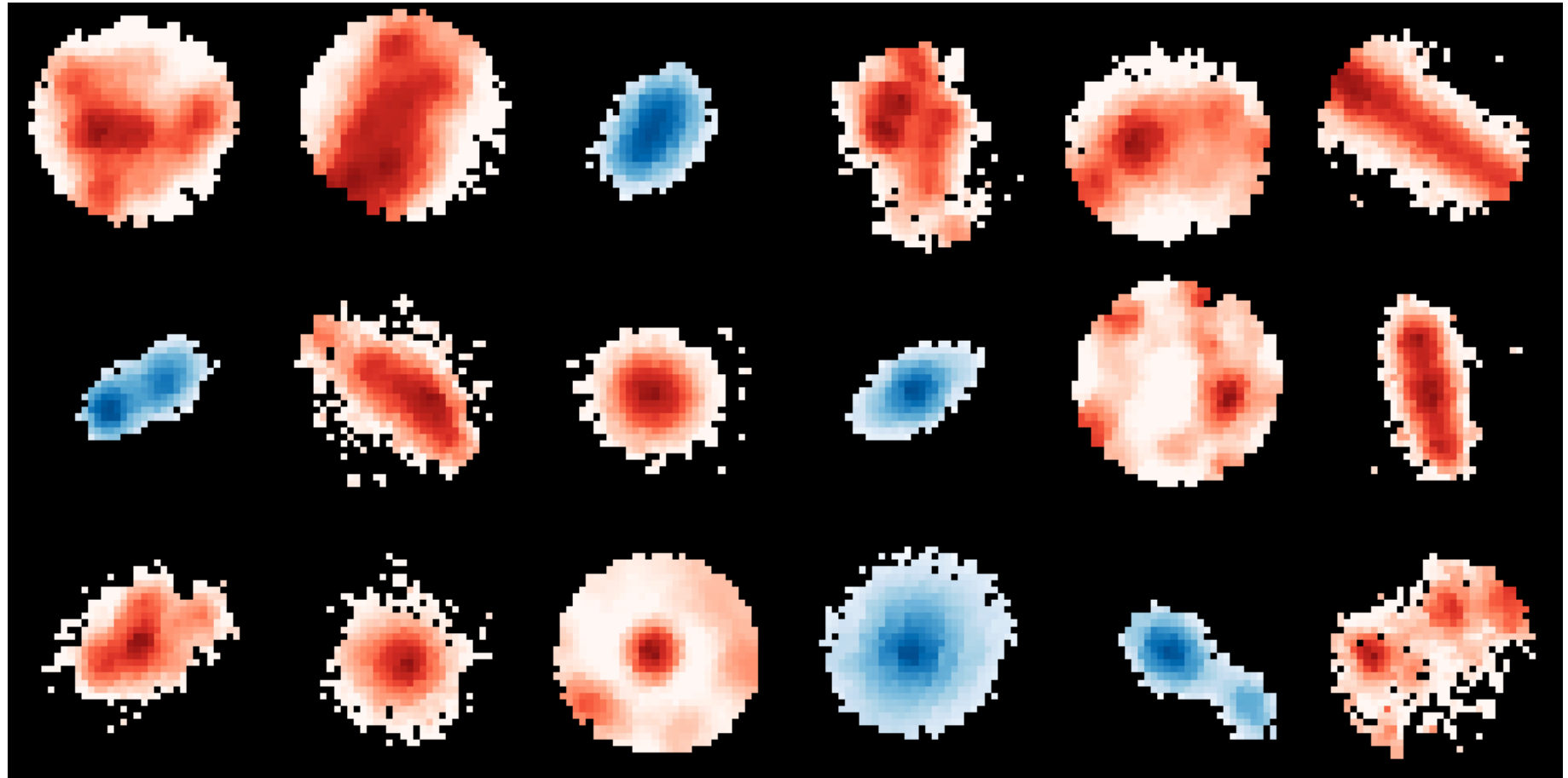




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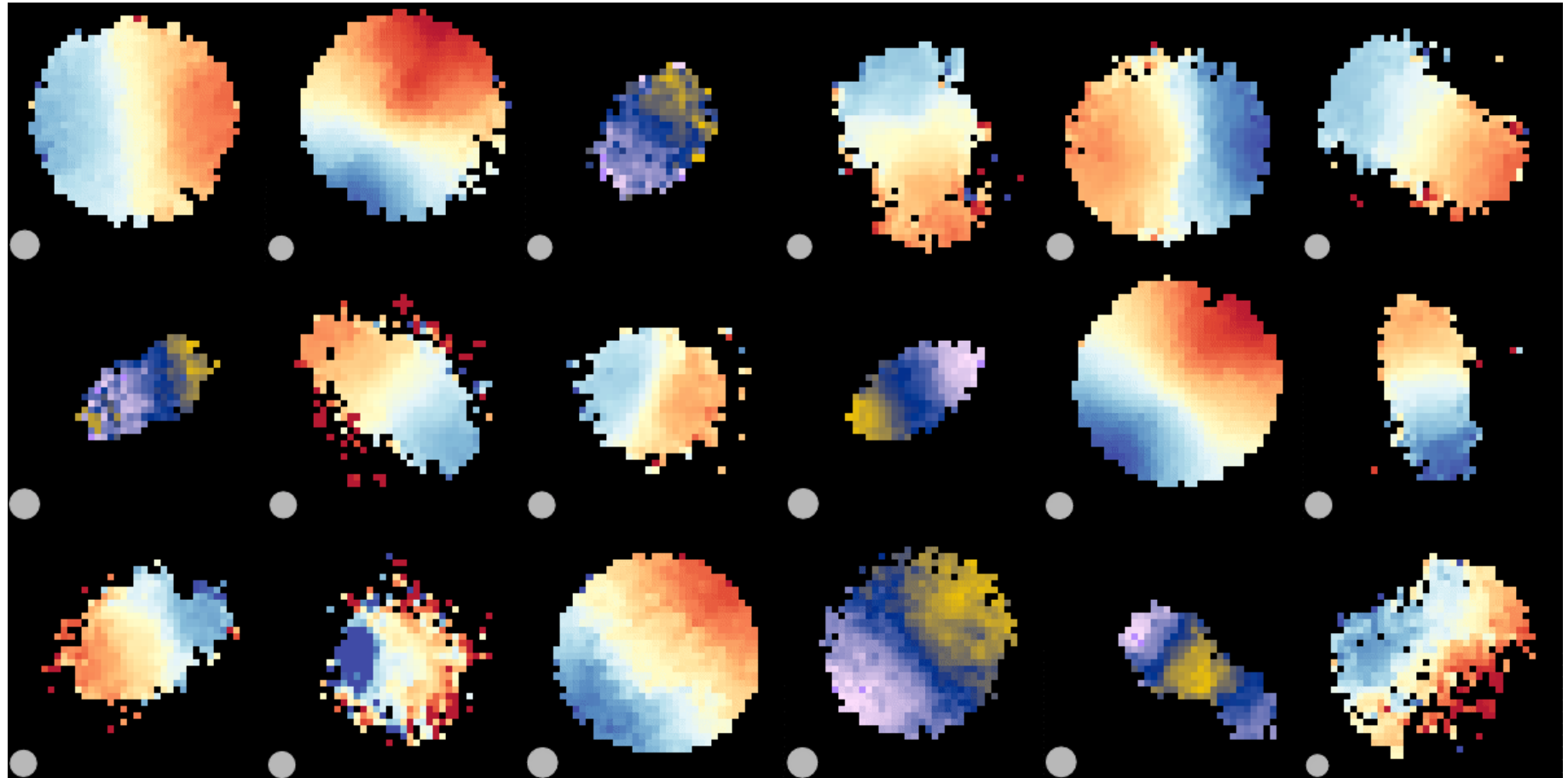




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# SAMI Galaxy Survey

## Target selection

