

Parity Violating Polarized eC Scattering Asymmetry $A_{RL}(eC)$ & New Physics

1) eC($\pm 0.3\%$) vs $\pm 25\%$ Bates

2) New Physics Sensitivity

$Q_W(C)$ vs $Q_W(Cs)$, $Q_W(e)$, $Q_W(p)$

i) Short-Distance Phys.

ii) Dark Parity Violation

**William J. Marciano
March 15, 2013
M.I.T.**



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For a recent review containing much of this material see:

“Low-Energy Measurements of the Weak Mixing Angle”

Annual Review of Nuclear and Particle Science (2013)

K. Kumar, S. Mantry, W. Marciano & P. Souder

Elastic Polarized PV $eC \rightarrow eC$

$$A_{RL} \equiv \sigma_R - \sigma_L / \sigma_R + \sigma_L,$$

- **G. Feinberg(1975)**: Simple $0^+ \rightarrow 0^+$ Amplitude

$$A_{RL}(eC) = -2^{1/2} G_F Q^2 / 48 \pi \alpha \times Q_W(^{12}C)$$

$$Q_W(^{12}C) = -24 \sin^2 \theta_W$$

Hadronic Effects Cancel at low Q^2 (see Donnelly Talk)

Robust Carbon Target can take high current

No Solenoid Spectrometer Polarized Target Issue for 0^+

$A_{\text{RL}}(\text{eC})$ (historical comparison)

BATES EXP $A_{\text{RL}}(\text{eC})$ (1978-1990)

P. Souder et al. PRL₆₅, 694(1990) (Pioneering Effort)

($0^+ \rightarrow 0^+$ Elastic Scattering)

Very Modest Effort by today's standards

$E_e = 0.25 \text{ GeV}$, $P_e = 0.37 \pm 0.02$, $I = 30\text{-}60 \mu\text{A}$, $\langle Q \rangle = 150 \text{ MeV}$, $T = 150 \text{ hrs}$

$A_{\text{RL}}(\text{eC})^{\text{SM}} = G_\mu Q^2 \sin^2 \theta_W / \sqrt{2} \pi \alpha$ $P_e A_{\text{RL}}(\text{eC})^{\text{exp}} = 0.60 \pm 0.14 \pm 0.02 \times 10^{-6}$

Measured $\sin^2 \theta_W = 0.20 \pm 0.05$

Current $\pm 25\%$ can be improved (statistically) to $\pm 0.2\%$ or better!

$P_e = 0.85$, $I = 150\text{-}500 \mu\text{A}$, $T = 1500\text{-}5000 \text{ hrs}$, $20 \times \text{Acceptance} \rightarrow \pm 0.2\%$!

Roughly Equivalent to 3xAPV(Cs) Sensitivity but no Atomic Theory

Main Issue: Polarization $\pm 0.2\%$, (Very Challenging!)

Note, if only $\pm 0.4\%$ Pol. Attainable, that will be \approx uncertainty

Elastic eC vs ep F.O.M

Statistical Figure of Merit $\approx A^2 \times \# \text{ scattering events}$
 $\pm 0.2\%$ eC compared to $\pm 1\%$ ep (requires **25x** larger F.O.M.)

All else equal (Pol., I, Target, Acceptance, Q^2 (small)...)

F.O.M.(eC) = $[(4s^2/(1-4s^2))^2 \times 36] \approx \mathbf{5500!} \rightarrow \text{eC takes } \approx \mathbf{1/200 \text{ time!}}$
F.O.M.(ep)

Alternative lower $\langle Q \rangle \approx 40 \text{ MeV}$ in eC statistically **easy!**

A_{RL} (eC) good exp at Mainz or JLAB (limited by Pol. Unc.)

Polarization: Needs R&D Effort $\pm 0.2\%$ - $\pm 0.4\%$

Standard Model Predictions

- $m_{\text{Higgs}}=125\text{-}126\text{GeV}$ SM Now Complete!
Combined with other fundamental parameters
- $\alpha^{-1} = 137.035999173(35)$
- $G_F = 1.1663787(6) \times 10^{-5} \text{ GeV}^{-2}$
- $m_Z = 91.1876(21) \text{ GeV}$
- $m_t=173.3(8) \text{ GeV}$

And Electroweak Radiative Corrections

Precise Predictions Possible – Test **“New Physics”**

Eg. S & T Loop Effects: 4th Generation, Technicolor..

$Z\chi$ from SO(10), Contact Interactions...

m_W & $\sin^2\theta_W(m_Z)_{MS}$ Predictions

$$m_W = 80.362(6)\text{GeV}[1 - 0.0036S + 0.0056T]$$

$$m_W = 80.385(15)\text{GeV} \text{ (experiment) } \textit{somewhat high}$$

$$\sin^2\theta_W(m_Z)_{MS} = 0.23124(6)[1 + 0.0157S - 0.0112T]$$

$$\sin^2\theta_W(m_Z)_{MS} = 0.23125(16) \text{ (Z pole experimental ave.)}$$

$$S = 0.07 \pm 0.09 \text{ and } T = 0.10 \pm 0.09$$

Very dependent on $\sin^2\theta_W(m_Z)_{MS}$ value!

Significantly Constrains: 4th Generation, Technicolor, SUSY...

Low Q^2 Measurements of $\sin^2\theta_W(m_Z)_{MS}$

$\sin^2\theta_W(m_Z)_{MS} = 0.2283(20)$ Cs APV at $\langle Q \rangle \approx 2.4$ MeV

Dzuba, Berengut, Flambaum, Roberts 2012 PRL

**$Q_W(\text{Cs}) \rightarrow -72.58(29)(32)_{\text{th}}$ Update vs SM $-73.24(5)$
(1.5 sigma APV deviation)**

$\sin^2\theta_W(m_Z)_{MS} = 0.2329(13)$ Møller A_{PV} at $\langle Q \rangle \approx 160$ MeV

$A_{RL}(ee) = -131(14)(10) \times 10^{-9} \propto (1 - 4\sin^2\theta_W)$

Measured to $\pm 12\%$ $\rightarrow \sin^2\theta_W$ to $\pm 0.6\%$

$\sin^2\theta_W(m_Z)_{MS} = 0.2356(16)$

$\nu_\mu N$ at $\langle Q \rangle \approx 5$ GeV

Needs Reanalysis

Future Expectations

QWEAK exp at JLAB (Under Analysis)

$A_{RL}(ep \rightarrow ep) \approx -3 \times 10^{-7}$ $E = 1.165 \text{ GeV}$, $\langle Q \rangle \approx 160 \text{ MeV}$, $\text{Pol} \approx 89 \pm 1\%$

$\Delta \sin^2 \theta_W(m_Z)_{MS} = 0.0008$ via $\pm 4\%$ measurement of A_{RL}

Polarized Moller at JLAB $E = 11 \text{ GeV}$ $\langle Q \rangle = 75 \text{ MeV}$

$A_{RL}(ee \rightarrow ee)$ to $\pm 2.5\%$ $\Delta \sin^2 \theta_W(m_Z)_{MS} = \pm 0.00025!$

P2 MESA Flagship Experiment $A_{RL}(ep \rightarrow ep)$

Like **JLAB QWEAK** but Lower Energy & More Running Time

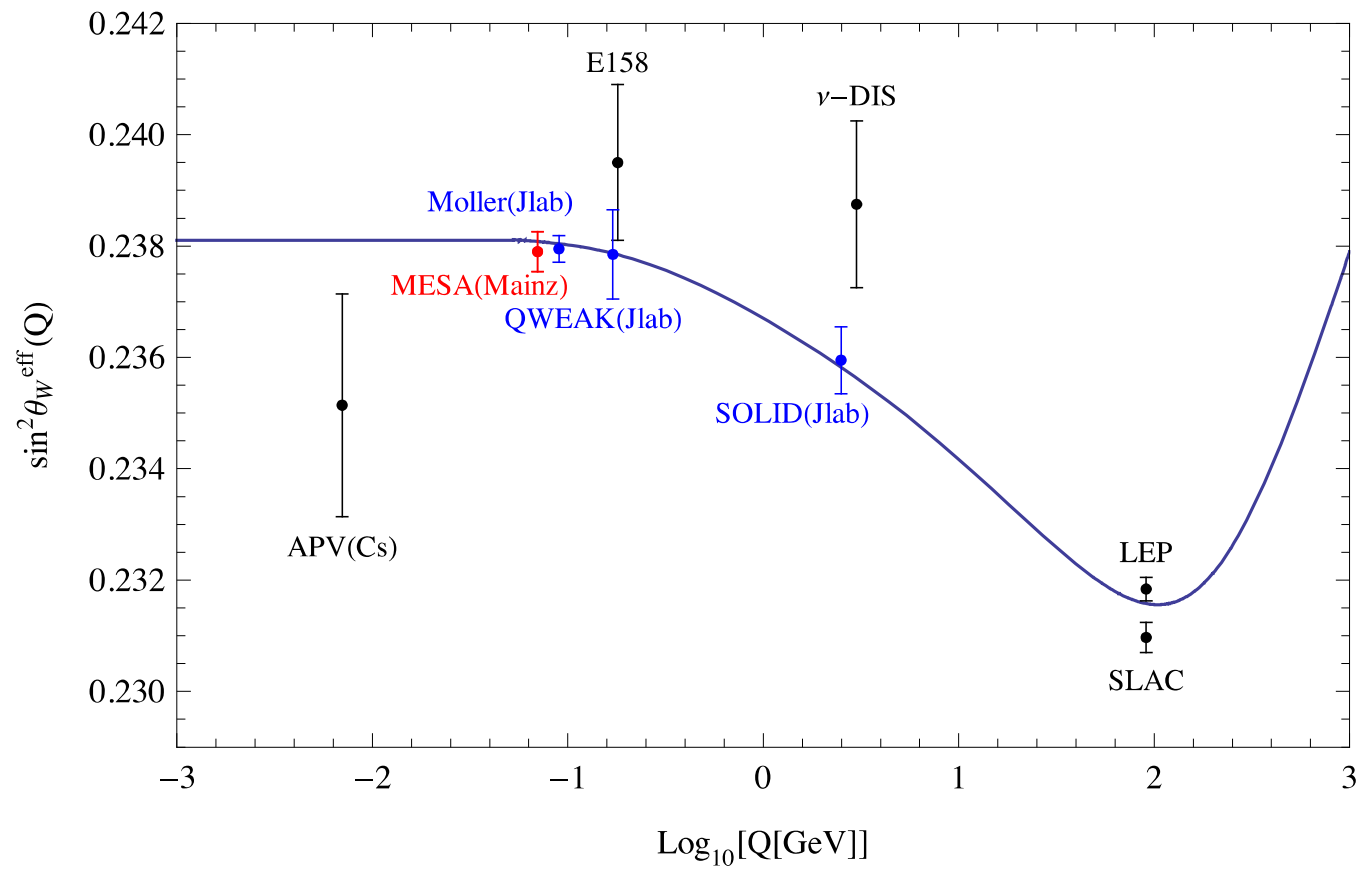
$E_e = 1.1 \text{ GeV} \rightarrow 0.14 \text{ GeV} (?)$ (Reduces Theory Unc.) $\langle Q \rangle \sim 100 \text{ MeV}?$

$\Delta \sin^2 \theta_W = \pm 0.00018$ (stat) \pm (syst. 2 loop, pol.,)

Pol. $\pm 0.5\%$ + 2loop unc. $\rightarrow \Delta \sin^2 \theta_W = \pm 0.00020_{\text{syst}}$

Overall $\pm 0.0003 - 0.0004$

APV Update & Possible A_{RL} Measurements



Comparison of (static) weak charges

$$Q_W(e) = -0.0435(9)[1 + 0.25T - 0.34S + 0.7X(Q^2) + 7m_Z^2/m_{Z_X}^2]$$

$$Q_W(p) = 0.0707(9)[1 + 0.15T - 0.21S + 0.43X(Q^2) + 4.3m_Z^2/m_{Z_X}^2]$$

$$Q_W(^{133}\text{Cs}) = -73.24(5)[1 + 0.011S - 0.023X(Q^2) - 0.9m_Z^2/m_{Z_X}^2]$$

$$Q_W(^{12}\text{C}) = -5.510(5)[1 - 0.003T + 0.016S - 0.033X(Q^2) - m_Z^2/m_{Z_X}^2]$$

Exp. Goals: $Q_W(e) \pm 2.5\%$, $Q_W(p) \pm 1.5\%$, $Q_W(^{12}\text{C}) \pm 0.3\%$

Observations

$A_{RL}(eC)$ at $\pm 0.3\%$ about 3x more sensitive to new physics than current $Q_W(Cs)$ (now 1.5 sigma deviation from SM)

$A_{RL}(eC)$ at $\pm 0.3\%$ about the same or better than $A_{RL}(ee)$ at $\pm 2.5\%$ and $A_{RL}(ep)$ at $\pm 1.5\%$ for $m_{Z\chi}$ (2TeV) and some contact int.

$A_{RL}(eC)$ at $\pm 0.3\%$ about 2-2.5x worse than $A_{RL}(ep)$ & $A_{RL}(ee)$ for $\sin^2\theta_W$, S (lower $\langle Q \rangle \rightarrow$ better $X(Q^2)$ sensitivity).

$A_{RL}(eC)$ negligible 2 loop uncertainty

$A_{RL}(ep)$ & $A_{RL}(ee)$ currently 2 loop unc. $\pm 1.3\%$ & $\pm 2\%$

Polarized ee, ep, eC Asymmetries

- $A_{RL} = \sigma_R - \sigma_L / \sigma_R + \sigma_L$ Parity Violating $\propto Q^2$ very small

Experiment	$\langle Q \rangle$ MeV	$\Delta \sin^2 \theta_W$	Measurement
Cs APV	2.4	± 0.0020	Atomic Theory
E158 SLAC	160	± 0.0013	ee Completed
Q_{weak} JLAB	160	± 0.0008	ep in analysis
Moller JLAB	75	± 0.00029	ee approved
MESA (ep) P2	100?	± 0.00037	ep Low Energy
eC ($\pm 0.3\%$)	40-120?	$\pm 0.0007?$	Pol. & Syst.

Do Both P2 and A_{RL} (eC) at MESA (Complementary)

Dark Parity Violation

- $U(1)_d$ gauge symmetry from the Dark Sector
Dark Photon, U Boson, Secluded ... Dark Z (Z_d)

Interaction with our world:

Induced by heavy fermion loops & extended Higgs

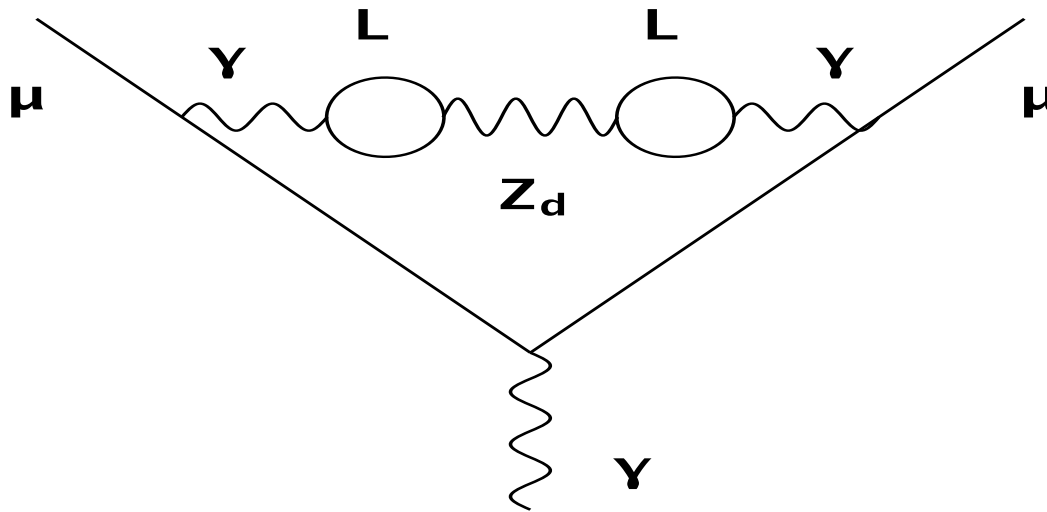
1) Kinetic Mixing $U(1)_Y \times U(1)_d$ $\epsilon e Z_d^\mu J_\mu^{\text{em}}$ $\epsilon \approx \alpha/\pi \approx 2 \times 10^{-3}$

2) Z- Z_d Mass Mixing $\epsilon_Z g/2 \cos \theta_W Z_d^\mu J_\mu^{\text{NC}}$
 $\epsilon_Z = m_{Z_d}/m_Z \delta = O(m_{Z_d}/m_Z)^2 \approx \underline{10^{-6}}$

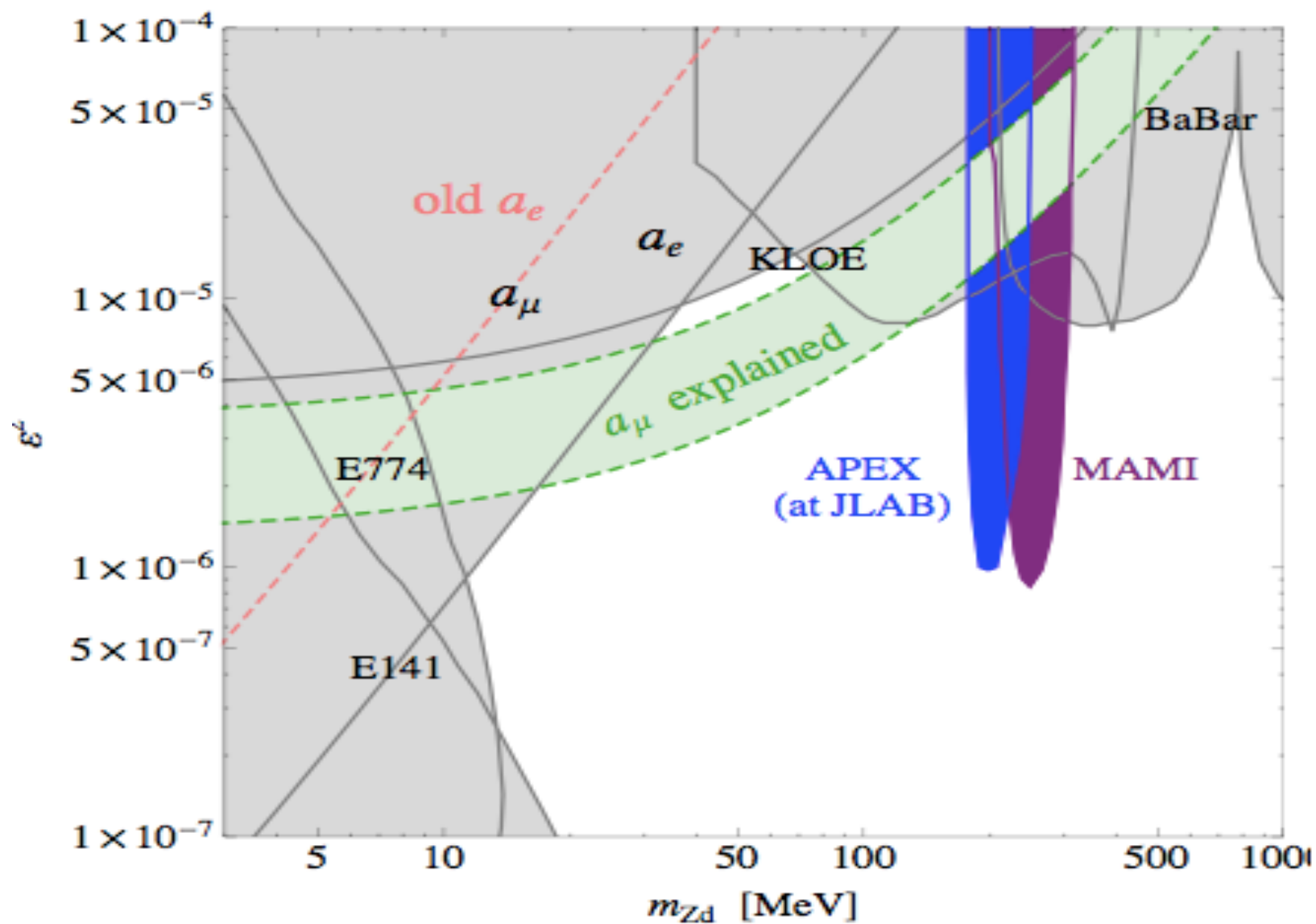
$$\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = 286(80) \times 10^{-11} \text{ (3.6}\sigma \text{ discrepancy!})}$$
$$\approx \frac{1}{4}(\alpha/\pi)^3 \text{ (Effective 3 loop physics?)}$$

Effective 3 loop g_μ -2 Diagram

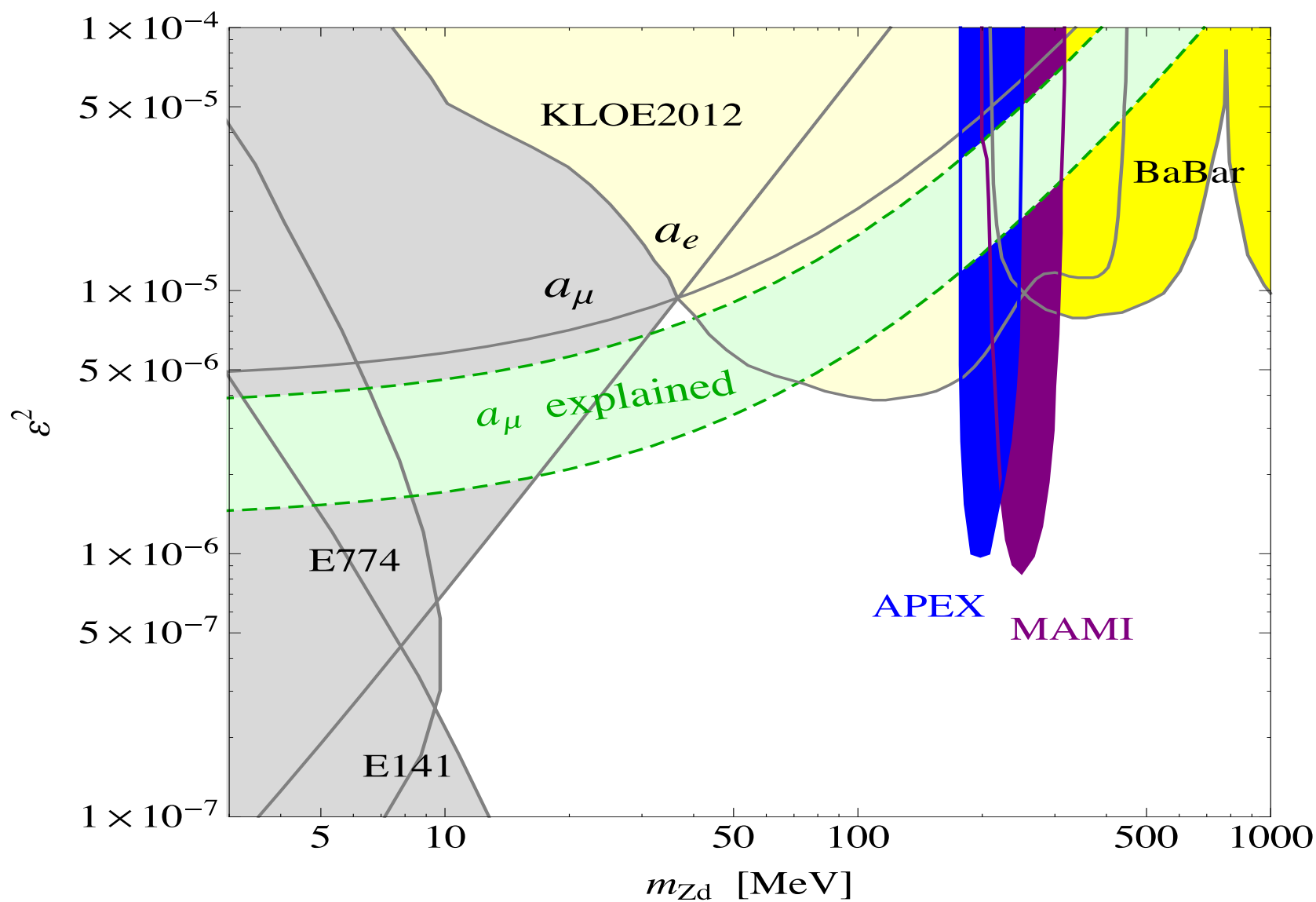
$a_\mu^{Z_d} = \alpha/2\pi\epsilon^2 F(m_{Z_d}/m_\mu)$, $F(0)=1$ solves g_μ -2 discrepancy
for $\epsilon^2 \approx 3-5 \times 10^{-6}$ & $m_{Z_d} \approx 20-60 \text{ MeV}$ (see figure)



***Old a_e vs Recent a_e (3 sigma bound)
(Davoudiasl, Lee & WJM)***



+2012 KLOE Update PL



Dark Parity Violation

Effect of ε & ε_z together: (at low $Q^2 \ll m_Z^2$)

$$\Delta \sin^2 \theta_W(Q^2) = -0.42 \varepsilon \delta m_Z m_{Z_d} / (Q^2 + m_{Z_d}^2)$$

For $\delta \approx m_{Z_d} / m_Z$, $\Delta \sin^2 \theta_W(Q^2) = \pm 0.42 \varepsilon m_{Z_d}^2 / (Q^2 + m_{Z_d}^2)$

Shift largest at small $Q^2 \ll m_{Z_d}^2$ ($\approx O(1\%)$! Eg APV)

(1.5 sigma APV deviation) fit $\rightarrow \varepsilon \delta = 4 \times 10^{-6}$

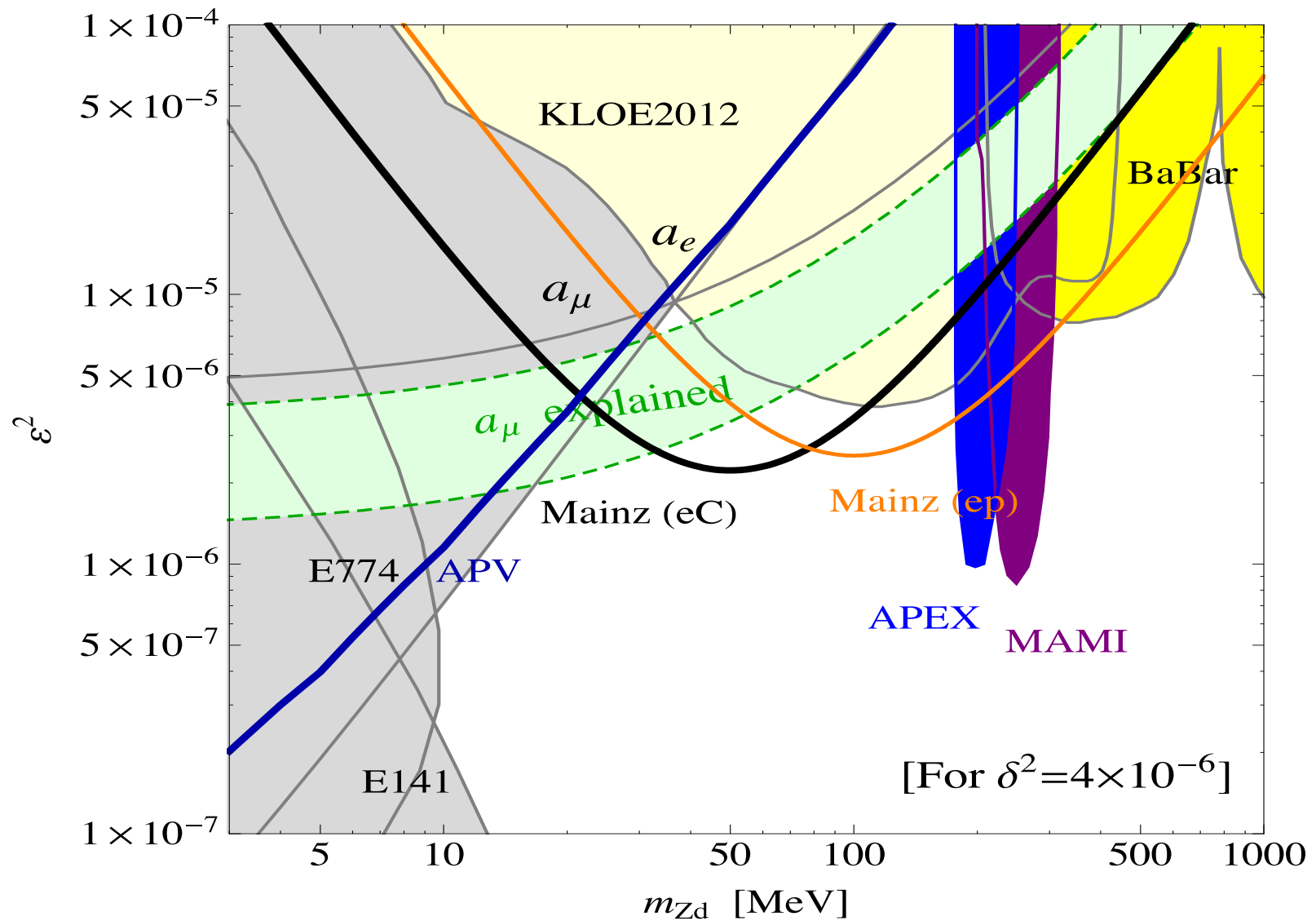
or $\varepsilon \approx \delta \approx 2 \times 10^{-3}$ for $(g_\mu - 2)$ & APV $\rightarrow m_{Z_d} \approx 50 \text{ MeV}$ region

$\sin^2 \theta_W(Q \approx 75 \text{ MeV})$ shift by $\pm O(0.5-1\%)$!!

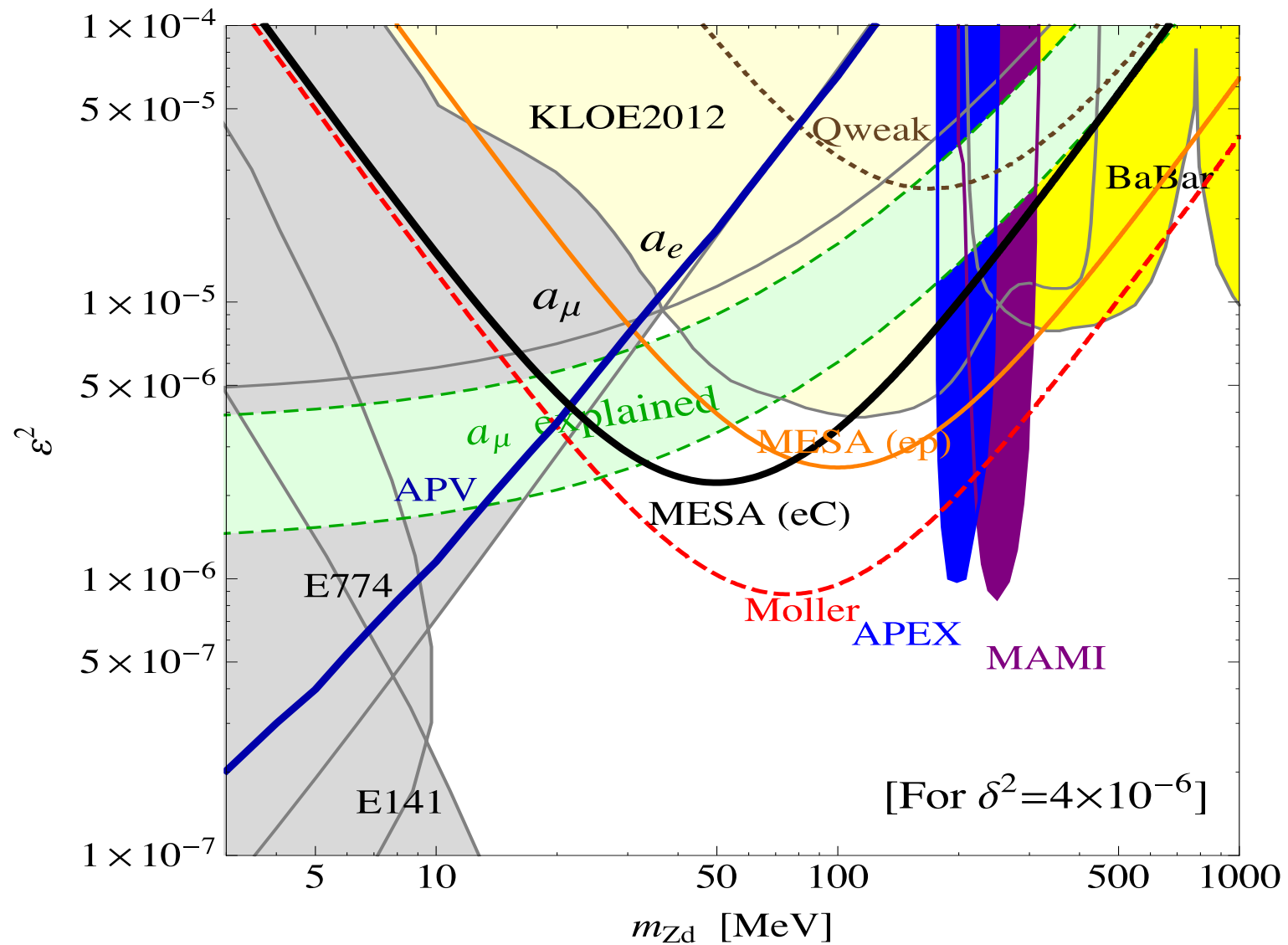
δ down to $\approx 10^{-3}$ Potentially Observable

$A_{RL}(ee)$ & $A_{RL}(eC)$ at low Q^2 Potentially Important

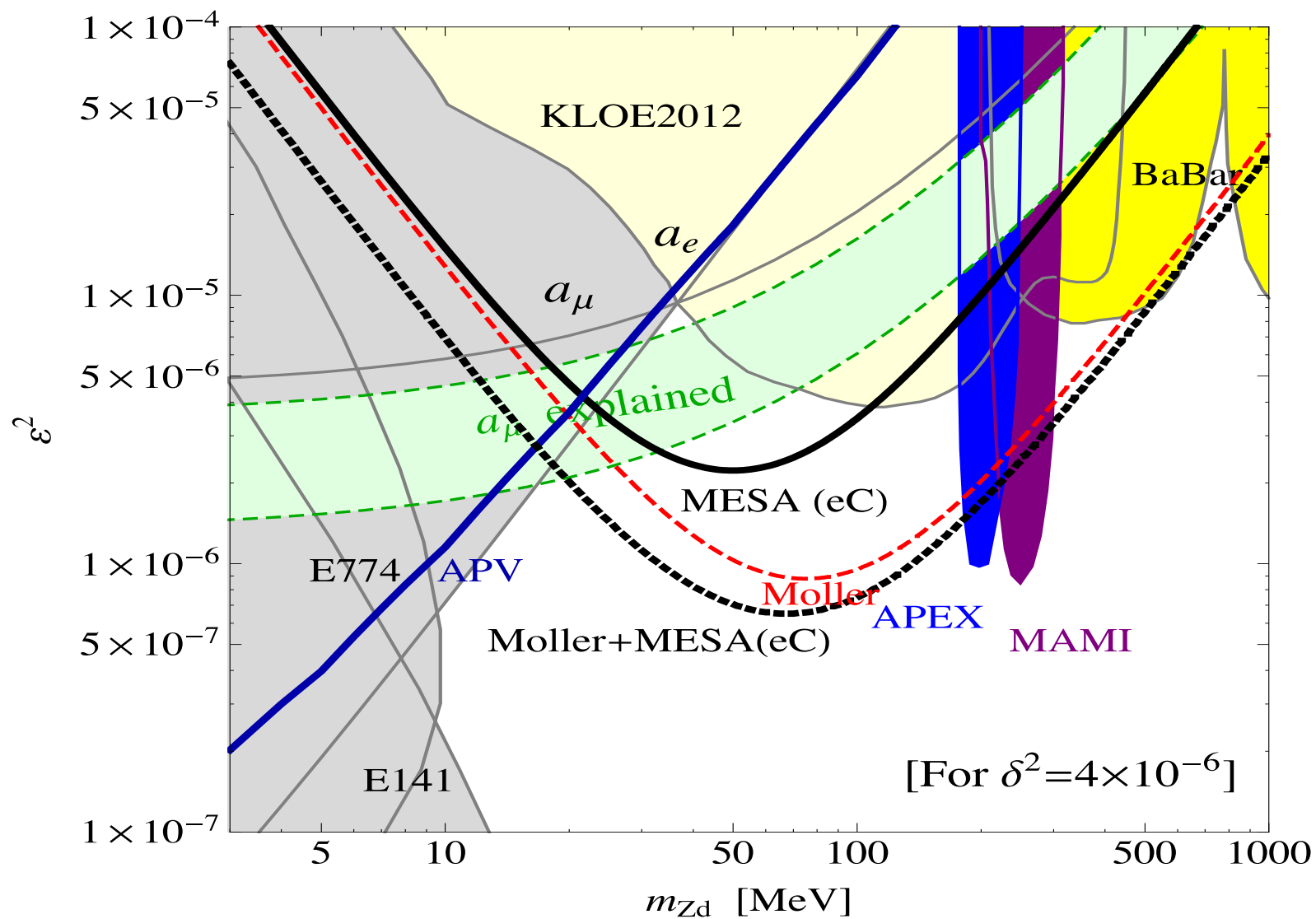
$A_{\text{RL}}(\text{eC})$ at $\langle Q \rangle = 50 \text{ MeV}$ vs $A_{\text{RL}}(\text{ep})$ at $\langle Q \rangle = 100 \text{ MeV}$



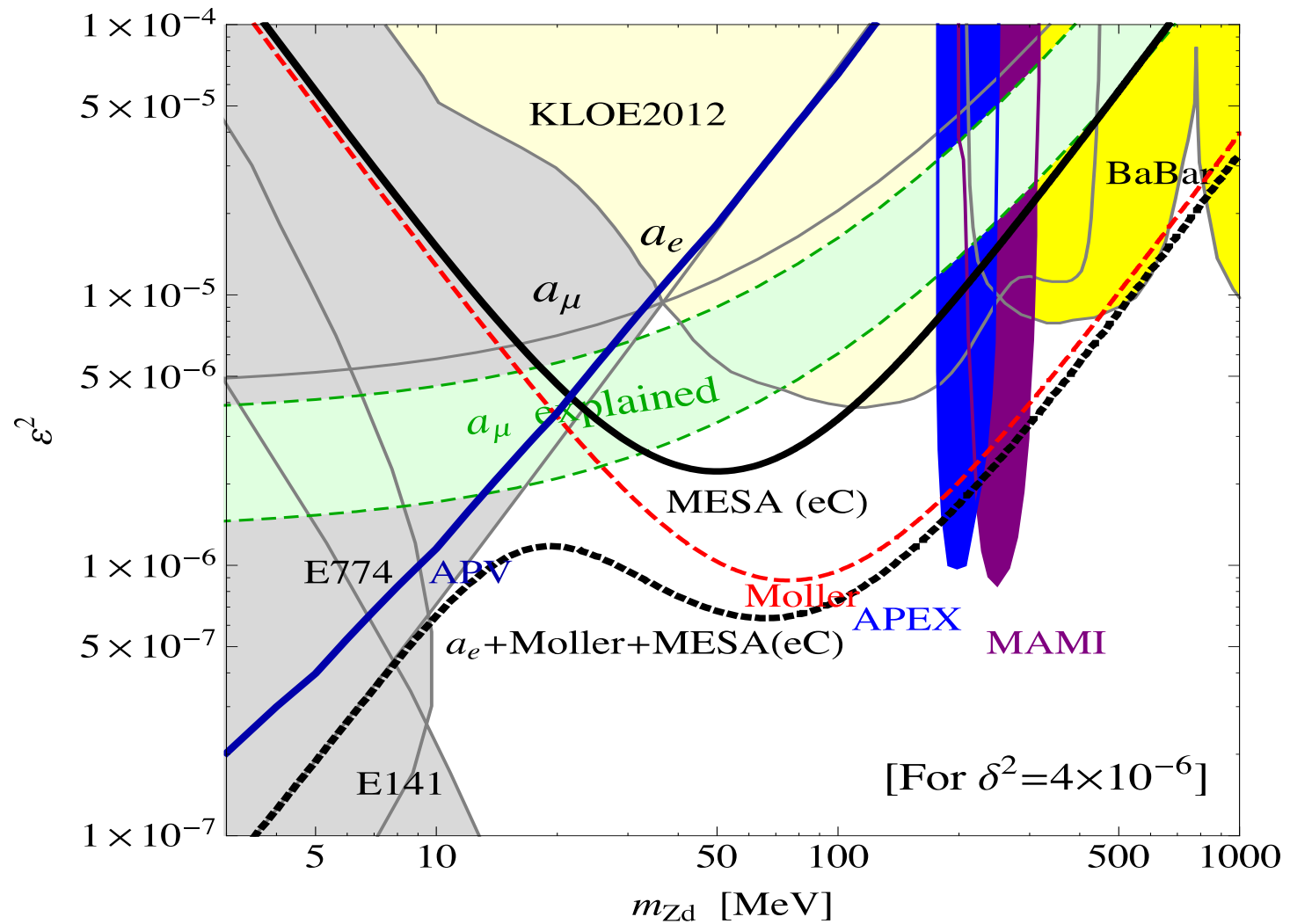
+Moller $A_{RL}(ee)$ at $\langle Q \rangle = 75 \text{ MeV}$



Moller $\langle Q \rangle = 75 \text{ MeV}$ + $A_{\text{RL}}(\text{eC}) \langle Q \rangle = 50 \text{ MeV}$



Combined Constraints



Conclusion

$A_{RL}(eC)$ at $\pm 0.3\%$ about 3x sensitivity of $Q_W(Cs)$ **1.5 σ deviation?**

No Atomic Theory Unc. But Polarization Unc. (Dominant?)

Statistically $A_{RL}(eC)$ much easier than $A_{RL}(ep)$ at low energy
Complementary. Motivates improved ($\pm 0.2-0.4\%$) polarimetry

If “dark photon” (20-60MeV) evidence emerges in direct
searches \rightarrow motivation for low Q^2 $A_{RL}(eC)$

Do $\langle Q \rangle \approx 40-120\text{MeV}$ look for **relative** $\sin^2\theta_W(Q^2)$ change

Overcomes pol. Uncertainty!

Deserves Further Study