

# Determining The Structure and Dynamics of the Milky Way Galaxy

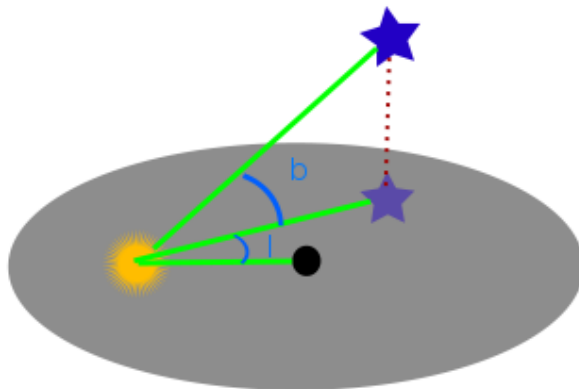
Shawn Westerdale

MIT - Department of Physics

# Experimental Goals

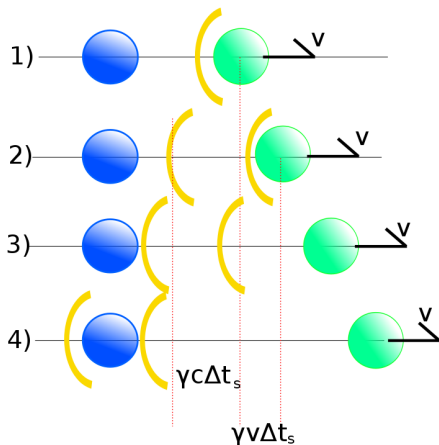
- **Overlying Question:** What is the mass distribution of the Milky Way?
- Determine how velocity varies with distance from the center of the galaxy
- Determine the density distribution of the galaxy

# Galactic Coordinates



**Figure:**  $b$  is the angle above the galactic plane,  $l$  is points angle in the galactic plane

# Relativistic Doppler Shift



- $\lambda_{obs} = (c\gamma + v\gamma)\Delta t_{src}\lambda_{src}$   
 $= c\sqrt{\frac{1+\beta}{1-\beta}}$
- $\nu = \frac{c}{\lambda}$

## Doppler Shift

$$\nu_{obs} = \nu_{src} \sqrt{\frac{1-\beta}{1+\beta}} \quad (1)$$

# Known Structure of Galaxy

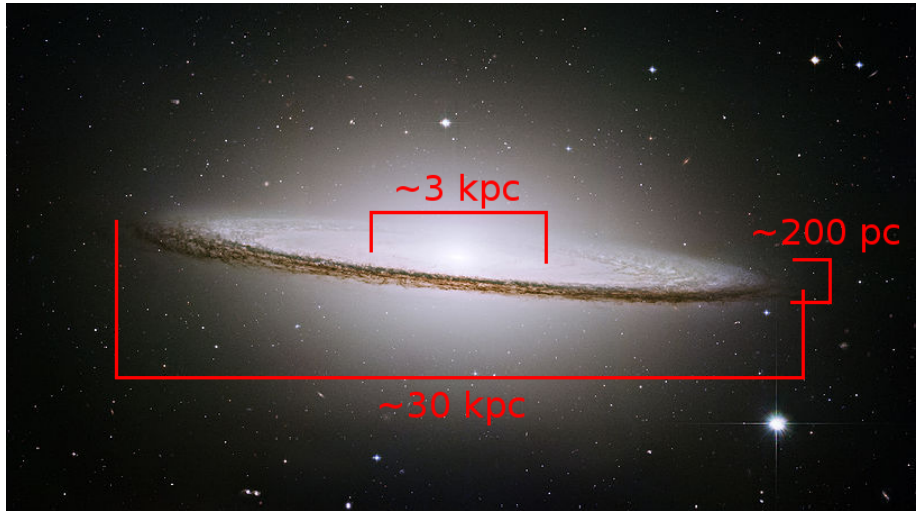
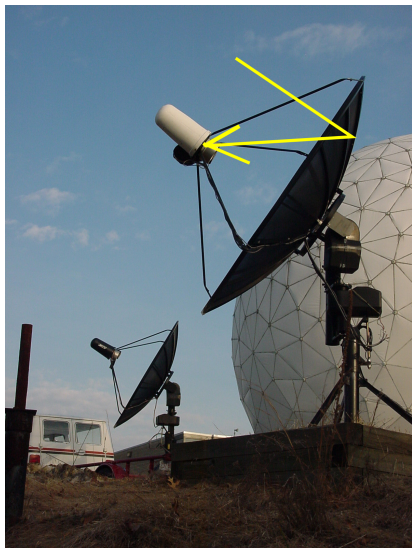


Image from: [http://en.wikipedia.org/wiki/File:M104\\_ngc4594\\_sombrero\\_galaxy\\_hi-res.jpg](http://en.wikipedia.org/wiki/File:M104_ngc4594_sombrero_galaxy_hi-res.jpg)

# Experimental Setup and Data Collection



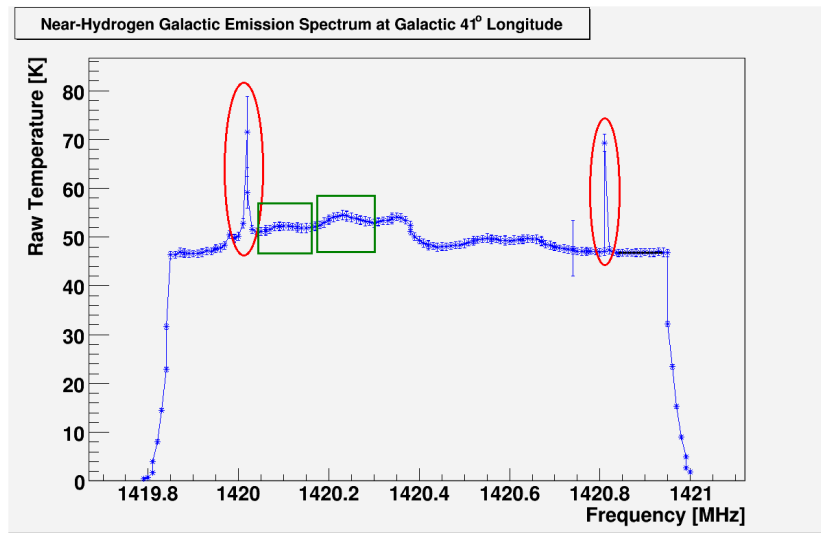
- MIT's small radio telescope
- Measures antenna temperature  $\propto$  power
- Scan 35 points for 10 minutes each from  $l = 25^\circ$  to  $l = 60^\circ$ 
  - $b = 0^\circ$
- Beamwidth =  $7.4 \pm .5^\circ$

# Keplerian and Newtonian Predictions

- Under the assumption that nearly all mass is located in the bulge. Given this, Newton and Kepler predict:
- Inside the bulge
  - Assuming constant density
  - $\frac{4G\rho\pi r^3}{r^2} = \frac{v^2}{r}$
  - $v = \sqrt{\frac{4}{3}G\rho\pi r}$
- Outside the bulge
  - $\frac{GM}{r^2} = \frac{v^2}{r}$
  - $v = \sqrt{GM} \frac{1}{\sqrt{r}}$

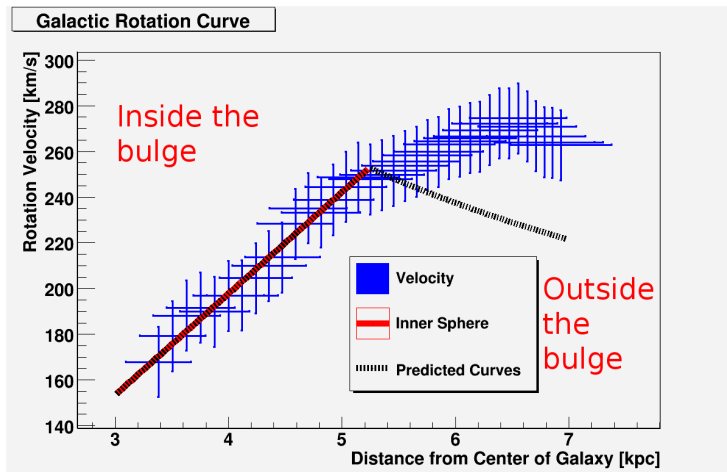
# Data

## Frequency Spectrum



# Analysis

## Galactic Rotation Curve



# Results

## Rotation Curve

Results do not match predictions!

- Keplerian predictions match data within the bulge, but remain roughly constant outside
- Fitting predictions to observations yields terrible fit
- Possible existence of dark matter may explain observations

# Analysis

## Density Distribution

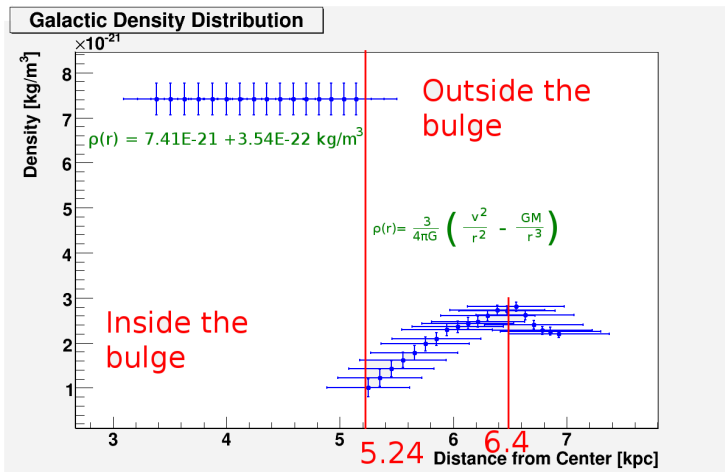
- Assuming an isotropic dark matter distribution
- $$\frac{G\left(M + \frac{4}{3}\pi r^3 \rho(r)\right)}{r^2} = \frac{v^2}{r^2}$$

### Isotropic Radial Density Distribution

$$\rho(r) = \frac{3}{4\pi G} \left( \frac{v^2}{r^2} - \frac{GM}{r^3} \right) \quad (2)$$

# Results

## Density Distribution



# Error Analysis

## Random Errors

- $\Delta T_{Raw} \approx 1\%$
- Many runs, low random error in measurements
- Fitting curve
  - $\Delta \nu_{min} \approx 3\%$
  - $\Delta \text{Baseline} \approx .4\%$

# Error Analysis

## Systematic Errors

- $\Delta l \approx .5^\circ$
- Frequency binning:  $\Delta \nu \approx .007\text{MHz}$
- Uncertainty in known values
  - $\Delta R_{sun} = 5\text{ kpc}$
  - $\Delta t_{sun} = 15359\text{ m/s (15\%)}$

# Conclusions

- The velocity of the galaxy within the central bulge behaves as expected
- Outside the central bulge, Kepler's and Newton's predictions, assuming all of the mass is contained in the center of the galaxy, fail
  - Velocity curve levels off around at around 5 kpc
  - Dark matter may be able to explain

# Acknowledgements

I would like to thank my partner Anna Waldman-Brown for helping me collect and analyze data as well as the Junior Lab staff for helping me work out problems I encountered along the way. I would also like to thank Sukrit Ranjan for providing advice on interpreting the output of the SRT.