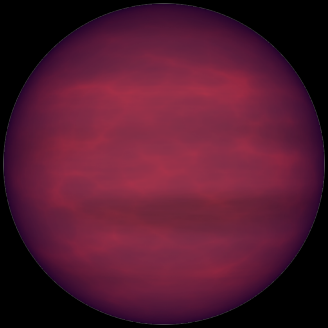
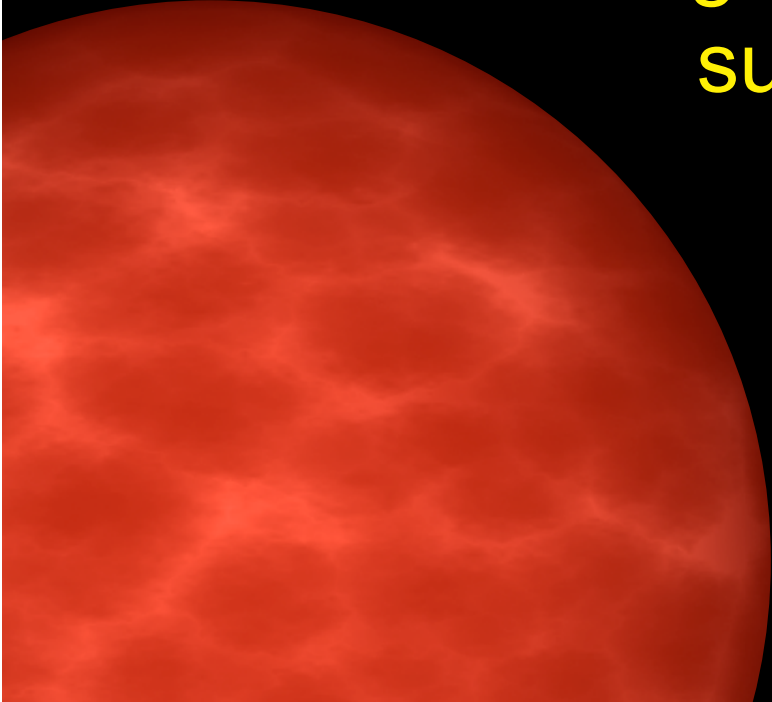
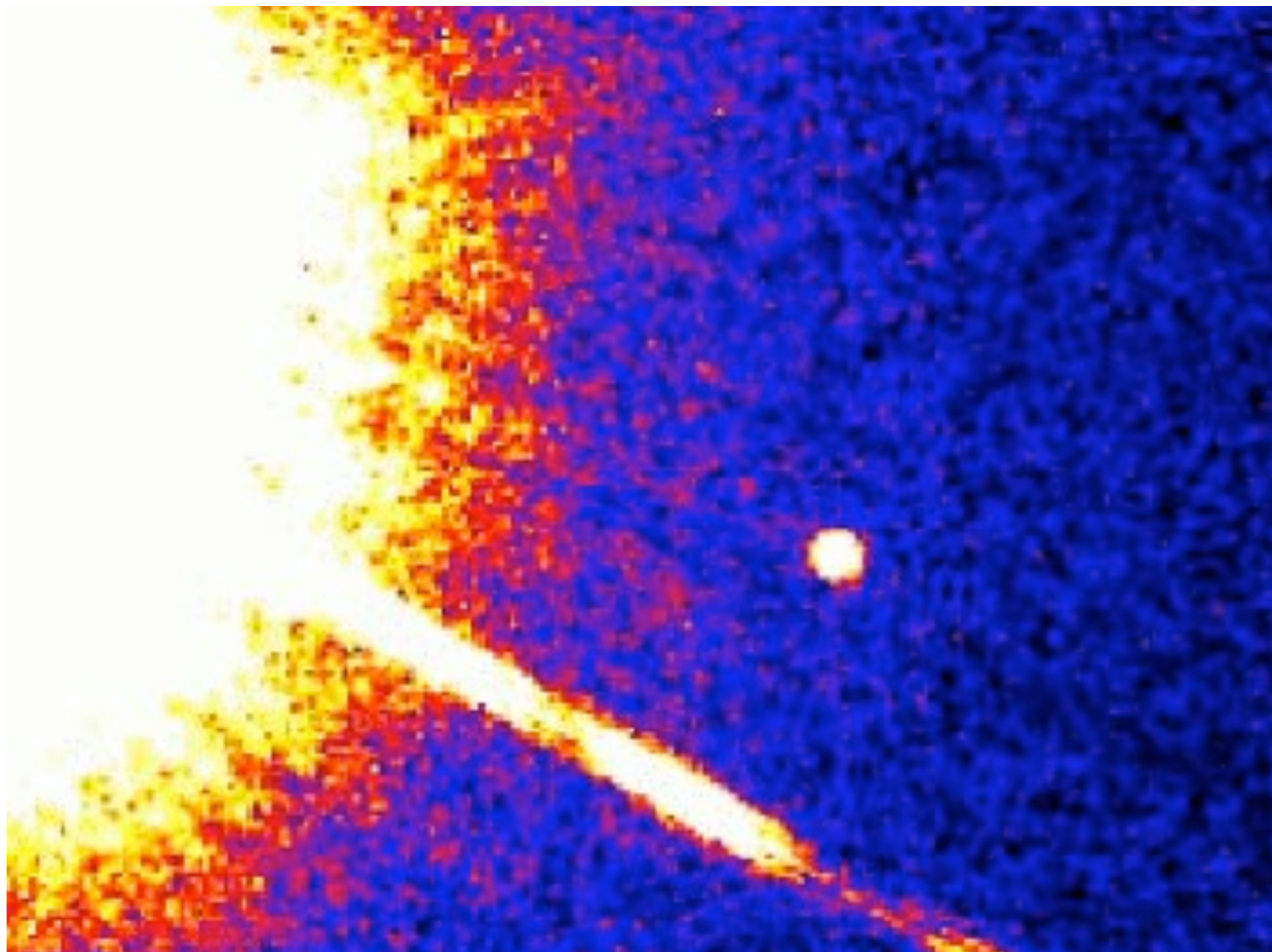




deconstructing brown:

unraveling the physical properties of
substellar objects

Adam J. Burgasser (MIT)



“The discovery of a new dish does more for human happiness than the discovery of a new star.”

Anthelme Brillat-Savarin (1825)

“The most exciting phrase to hear in science, the one that heralds new discoveries, is not 'Eureka!' (I found it!) but 'That's funny ...' ”

Isaac Asimov

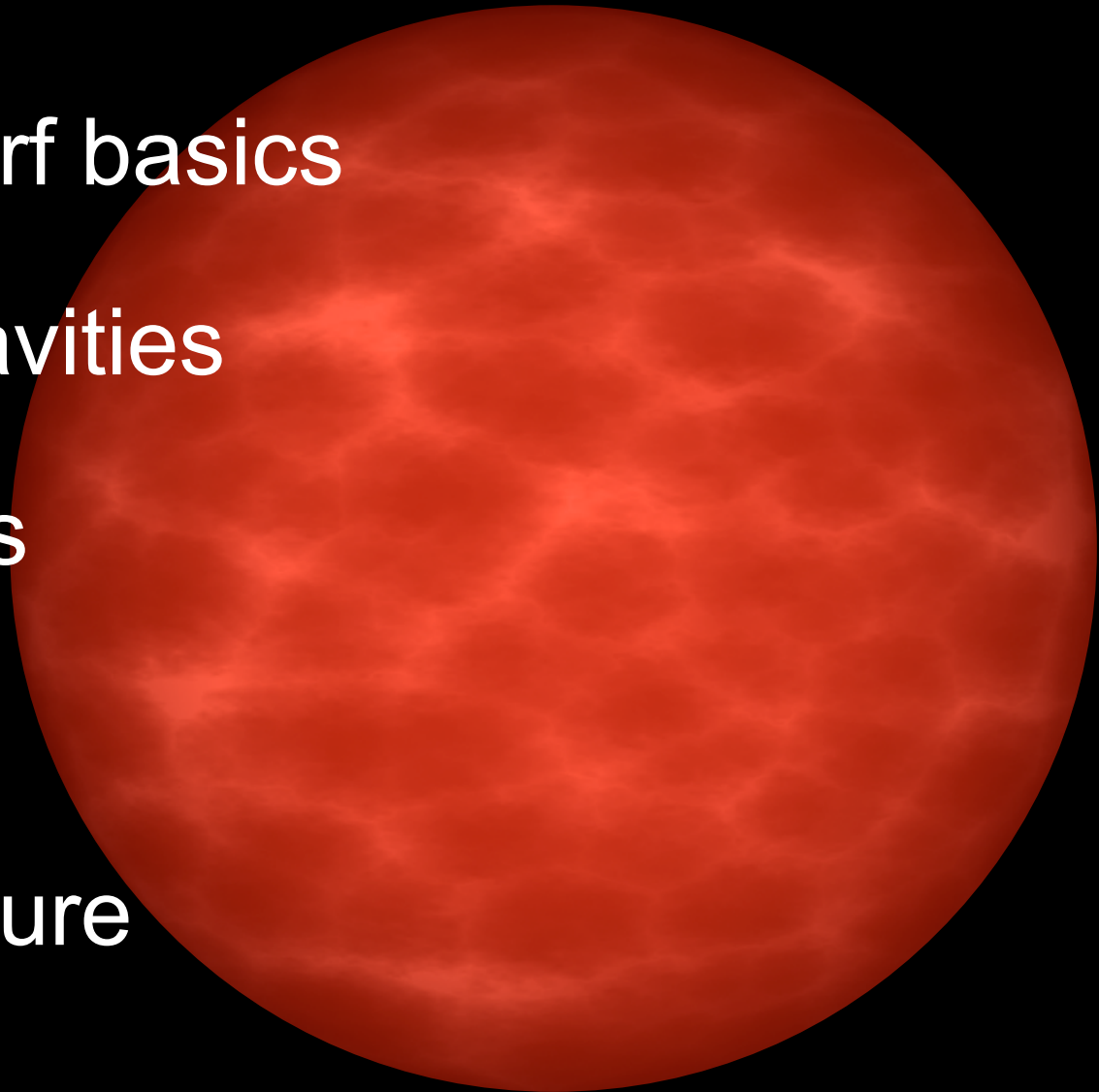
brown dwarf basics

surface gravities

metallicities

clouds

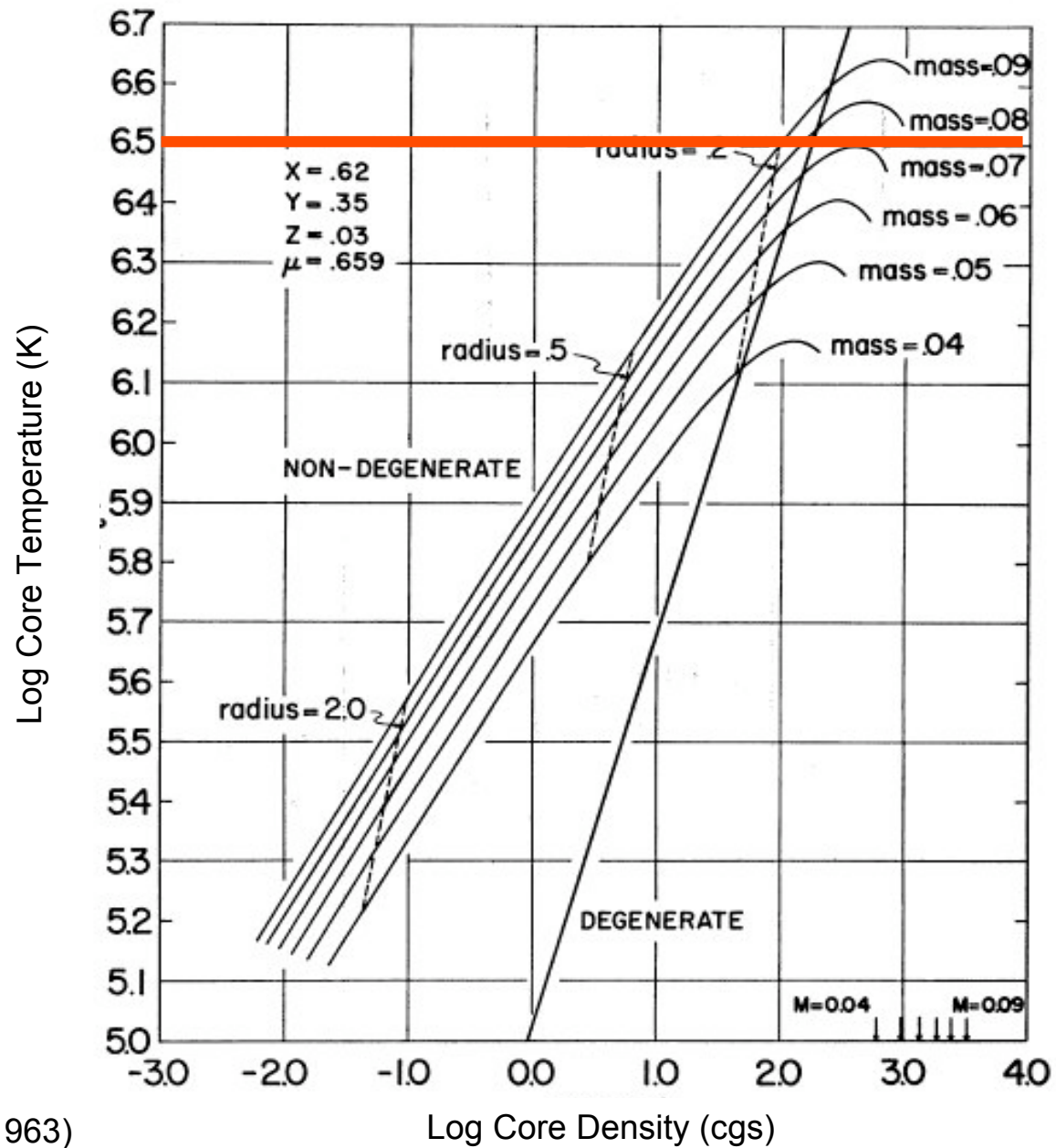
the big picture





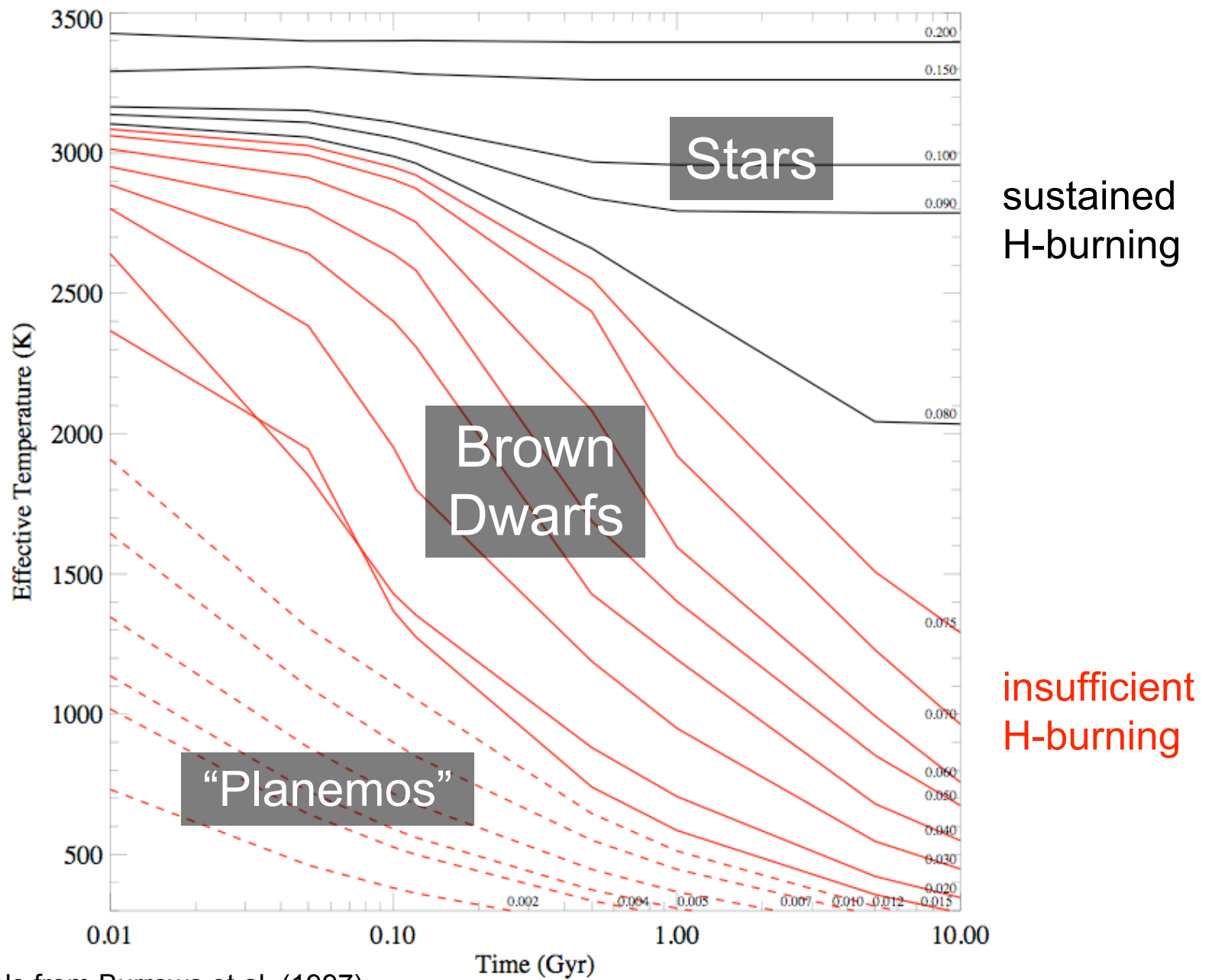
“meet your neighbors”

the more brown dwarfs we know, the
better we understand them

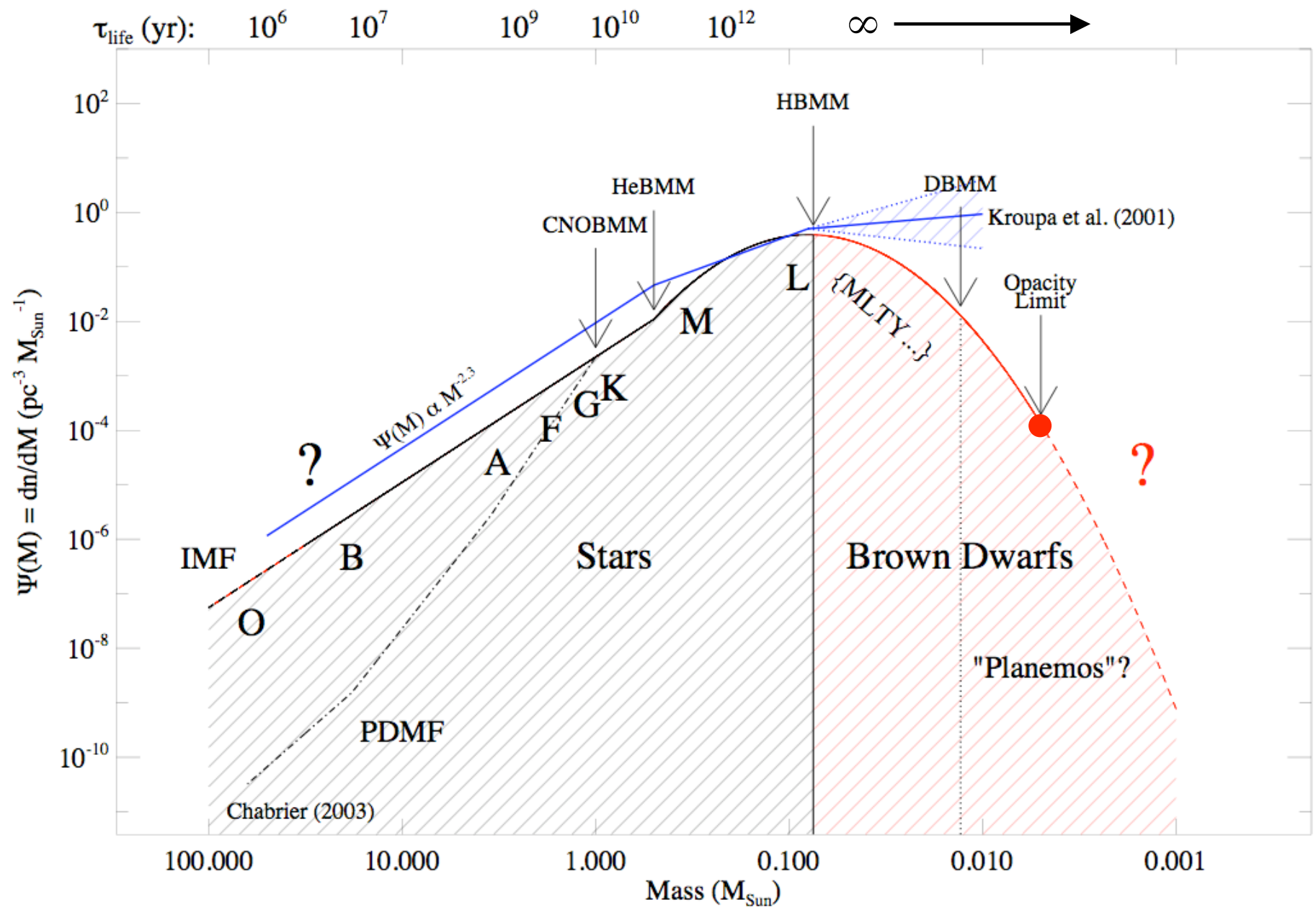


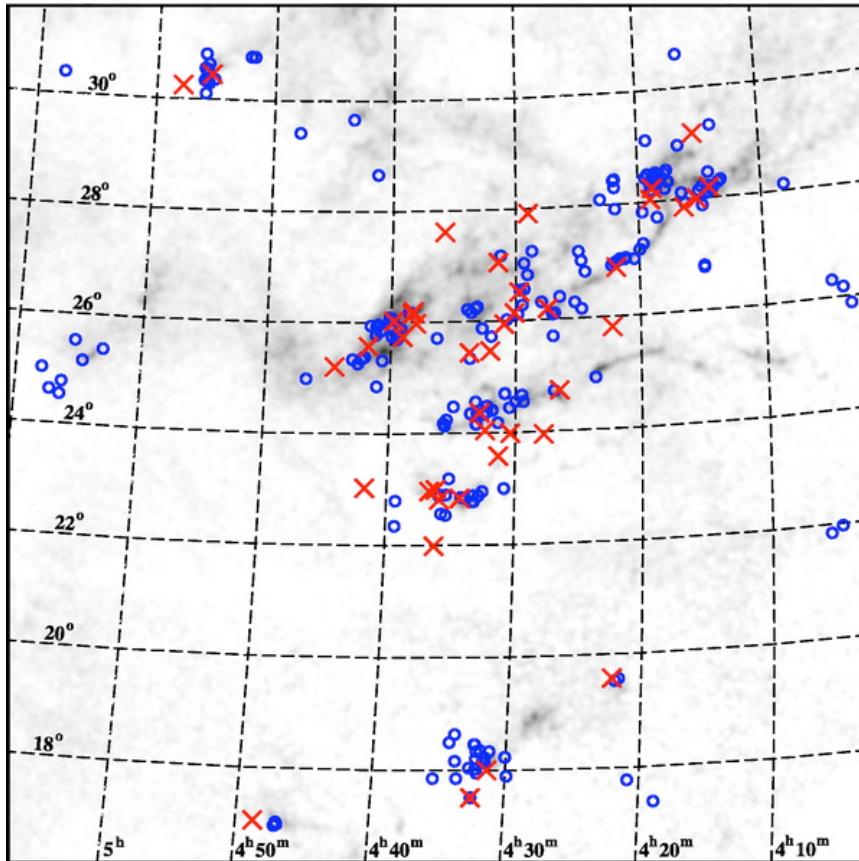
H-burning
threshold

Kumar (1962, 1963)
 see also Hayashi & Nakano (1963)



models from Burrows et al. (1997)





stellar and substellar distribution in Taurus

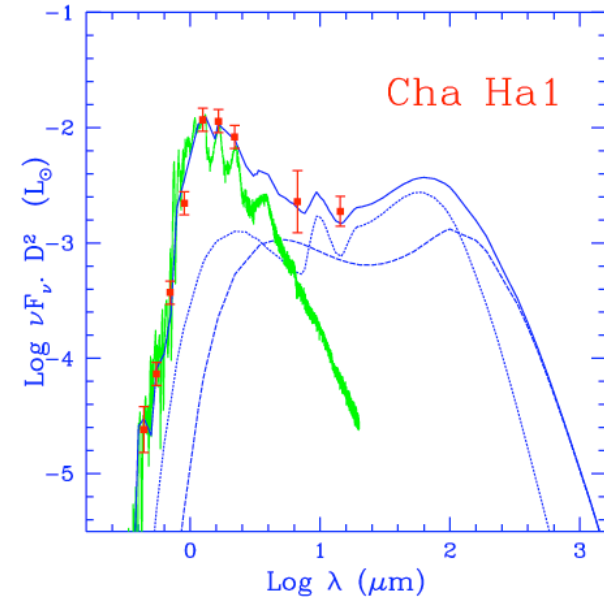
Luhman (2006)

see also Briceno et al. (2002); Guieu et al. 2006

disk accretion in a young brown dwarf

Jayawardhana et al. (2003)

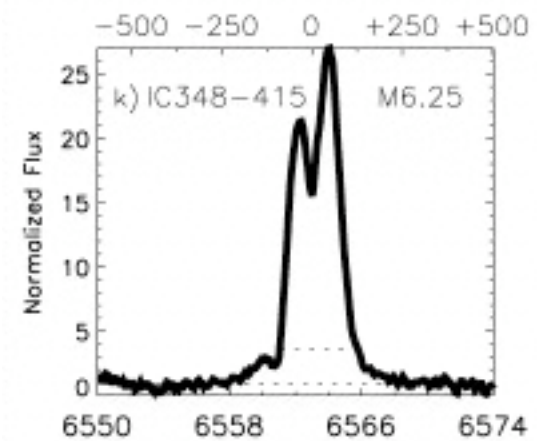
see also White & Basri (2003); Muzerolle et al. (2003,2005)

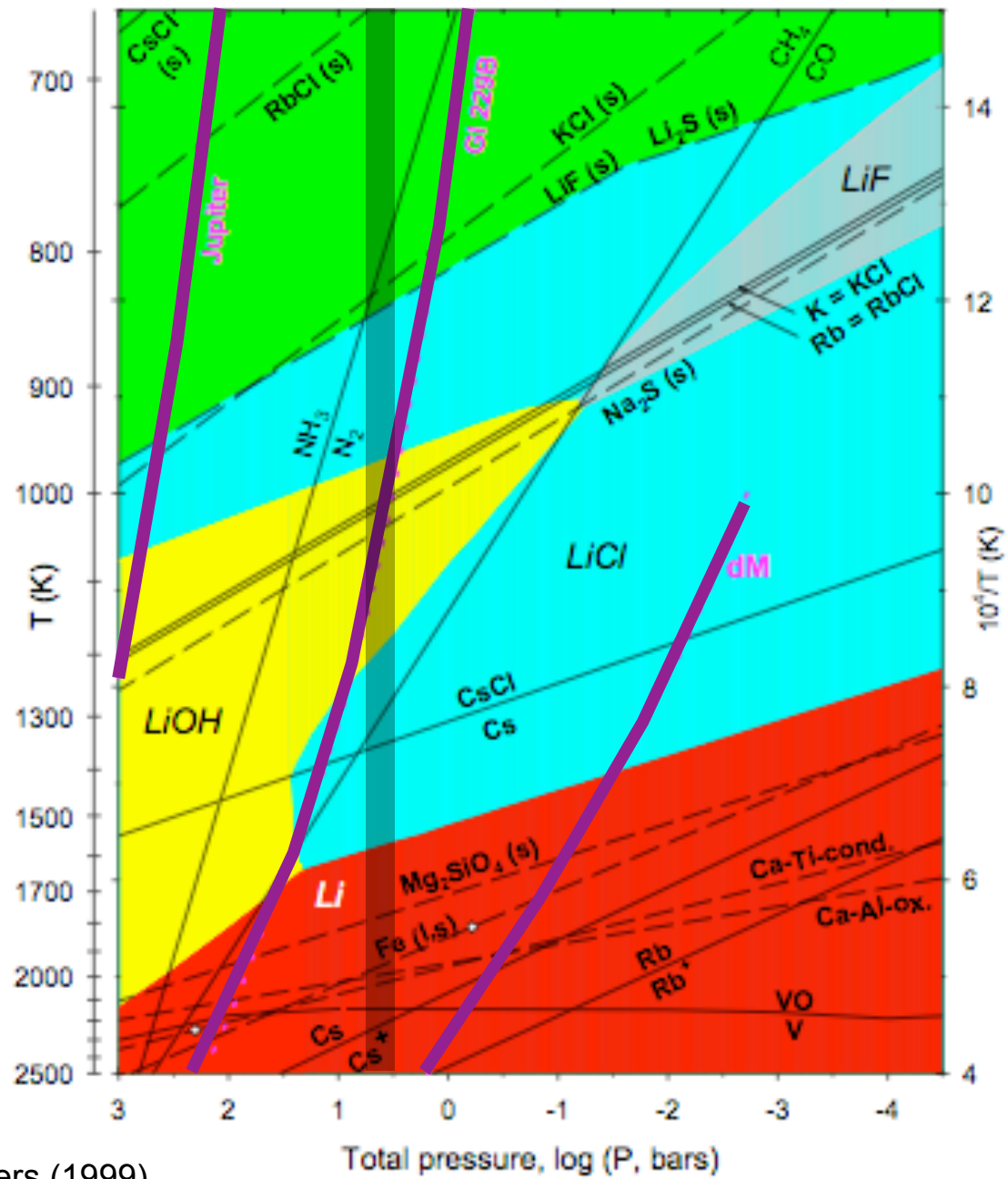


young brown dwarf with disk excess

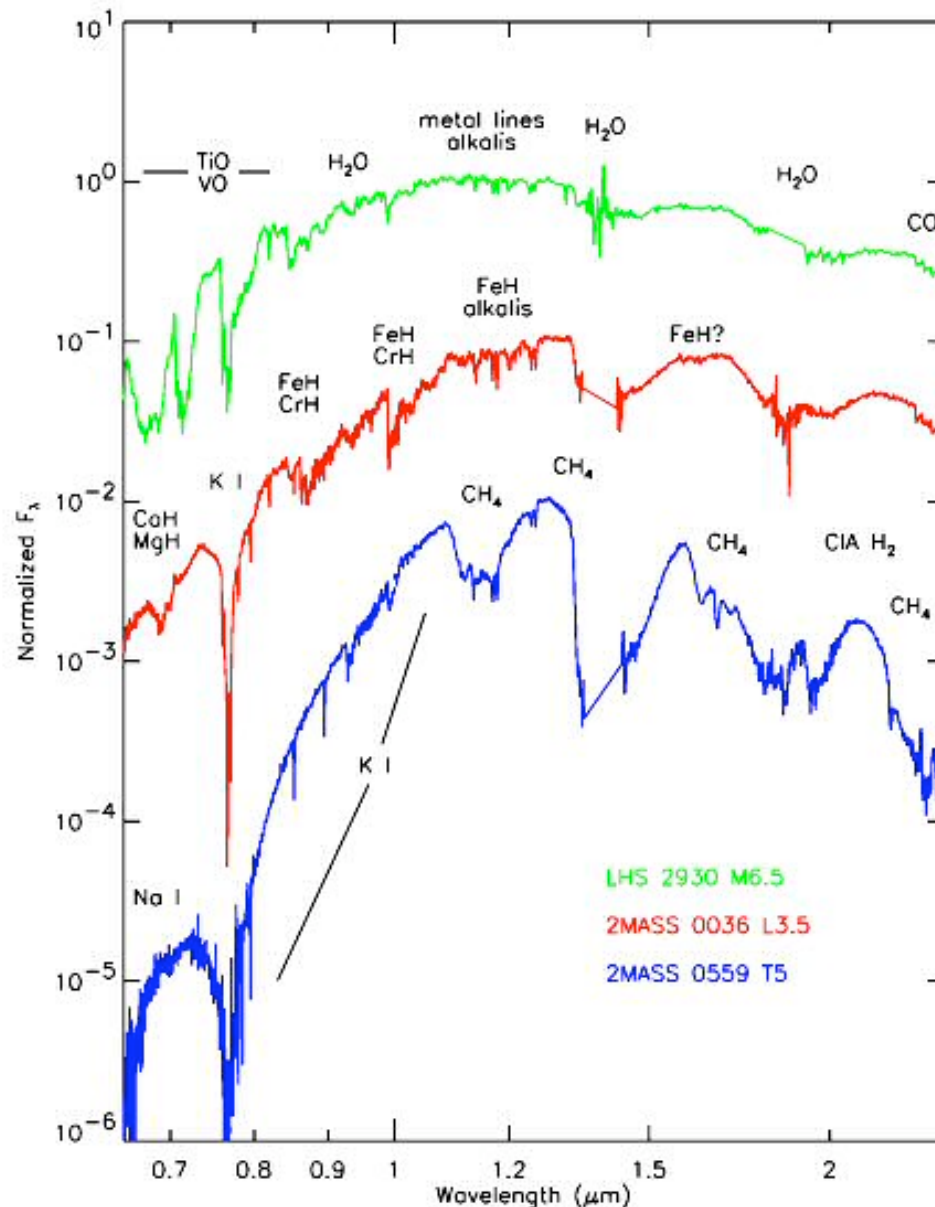
Natta & Testi (2001)

see also Comeron et al. (1998,1999)





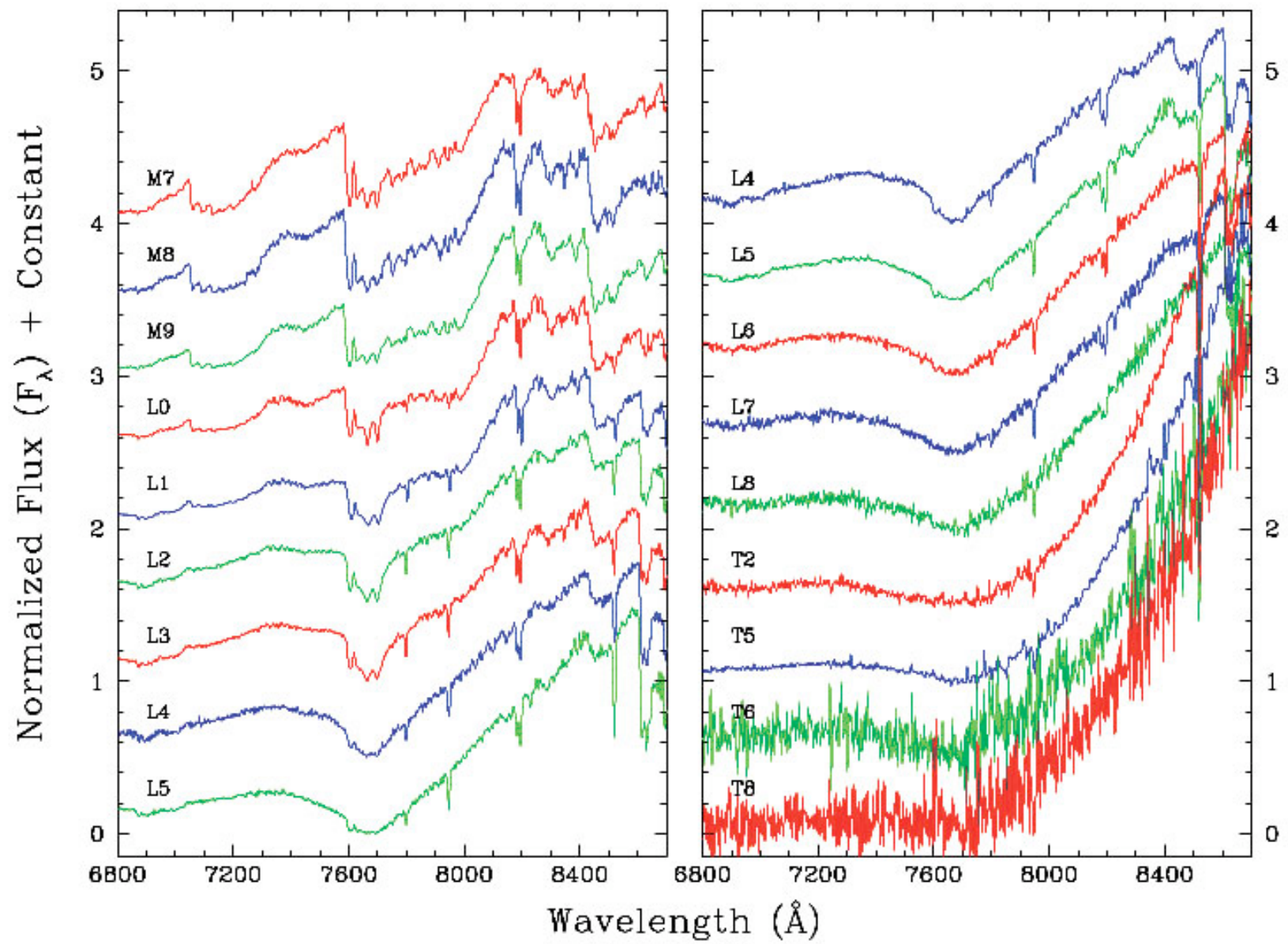
Lodders (1999)



M dwarfs are dominated by TiO, VO, H₂O, CO absorption plus metal/alkali lines.

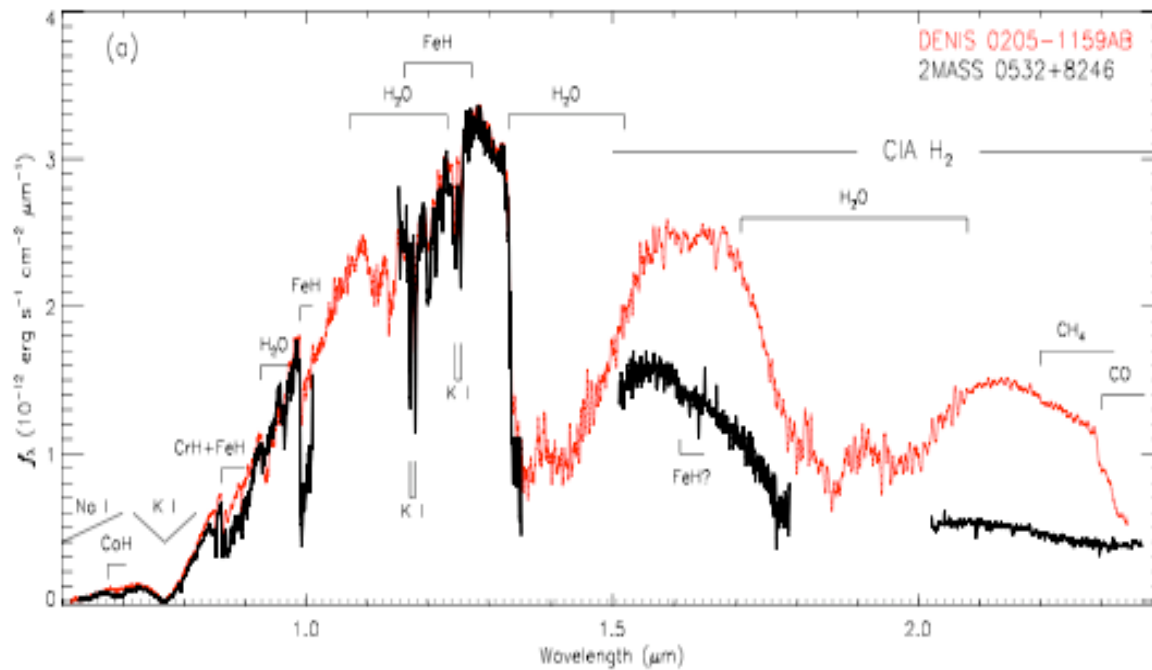
L dwarfs replace oxides with hydrides (FeH, CrH, MgH, CaH), alkalis are prominent, condensate clouds.

T dwarfs exhibit strong CH₄ and H₂O and extremely broadened Na I and K I.



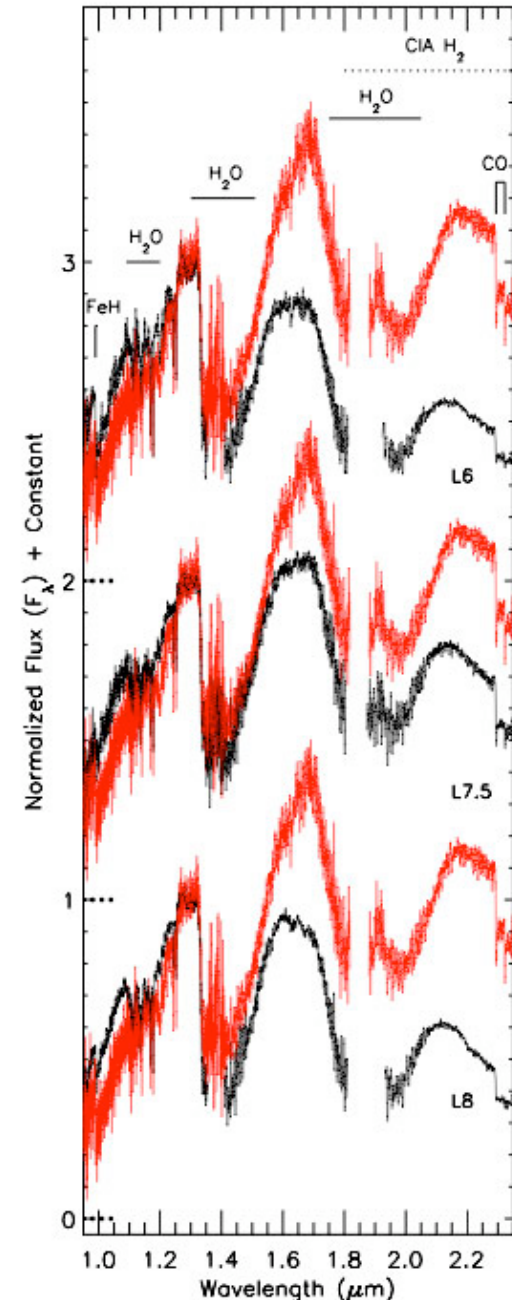
Kirkpatrick (2005)

peculiar sources

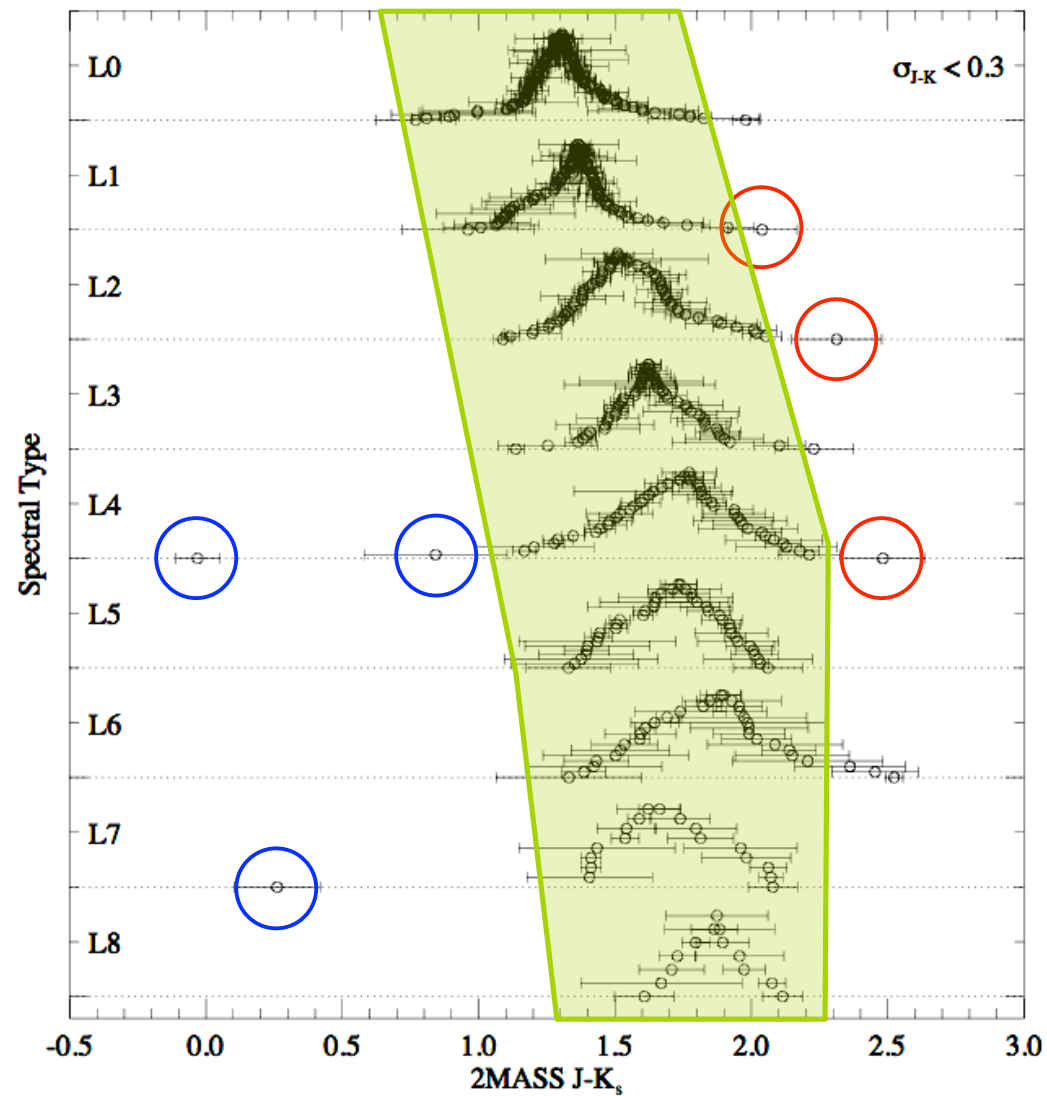


Burgasser et al. (2003)

Looper et al. (in prep.)



the “shrimp plot”

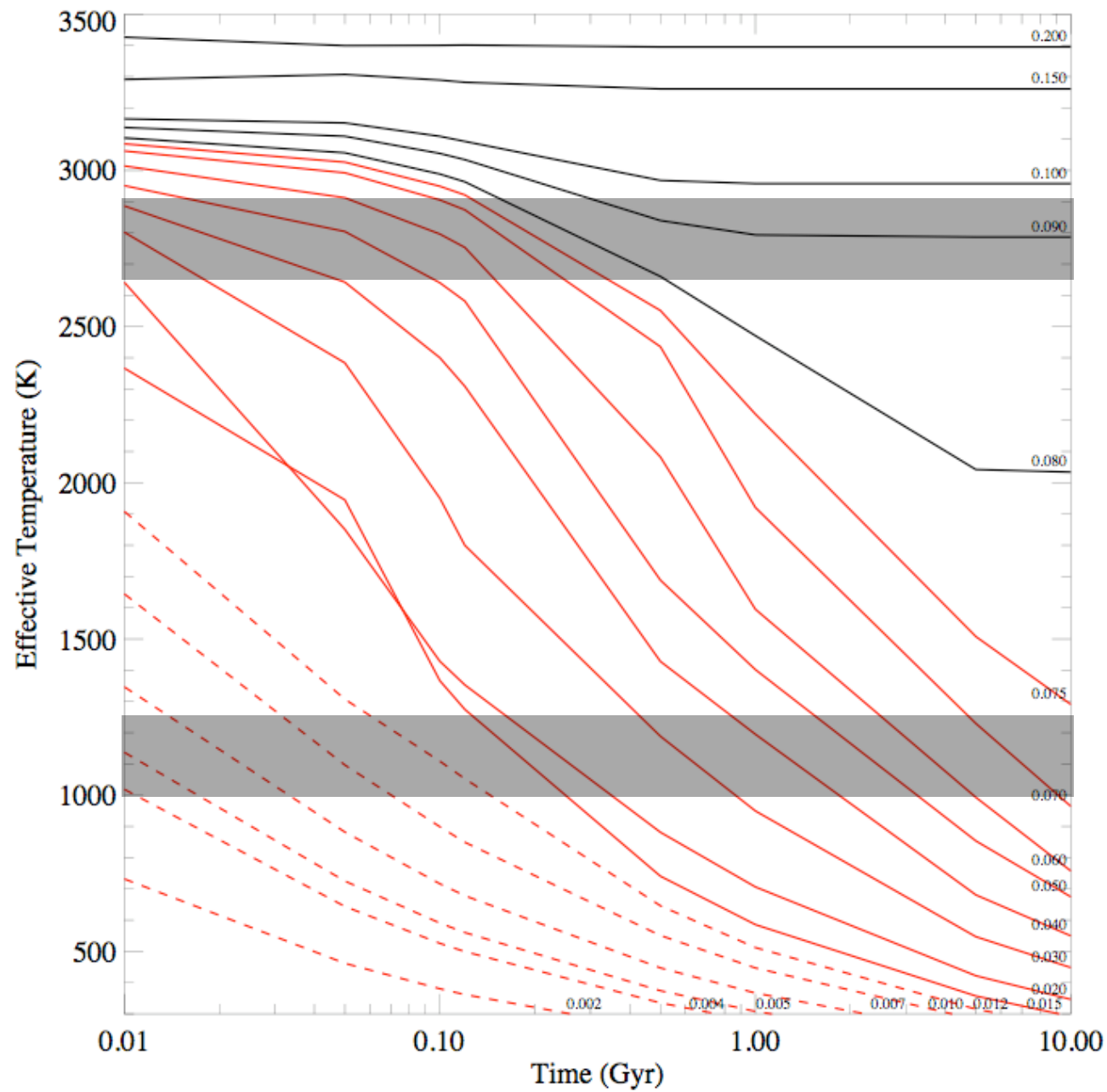


Kirkpatrick et al. (in prep.)



“weighing the baby”

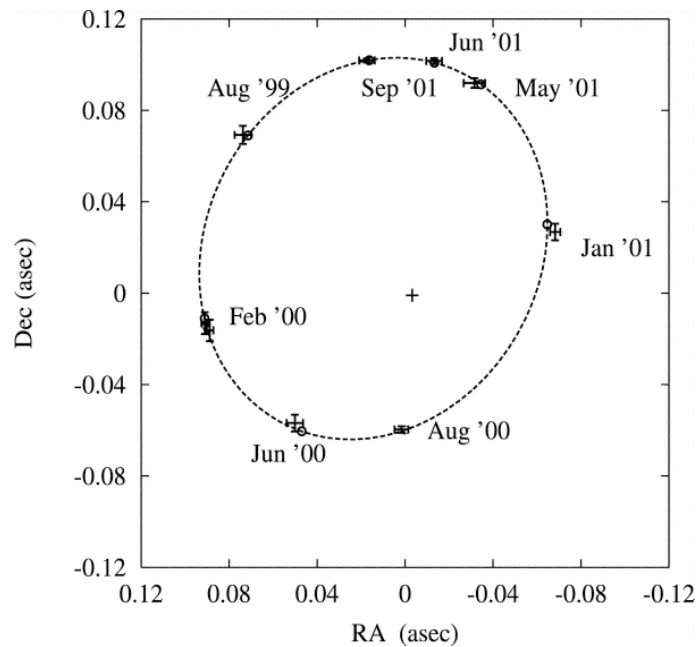
surface gravities provide mass and age estimates for individual brown dwarfs - and may uncover new populations



star or
brown dwarf?

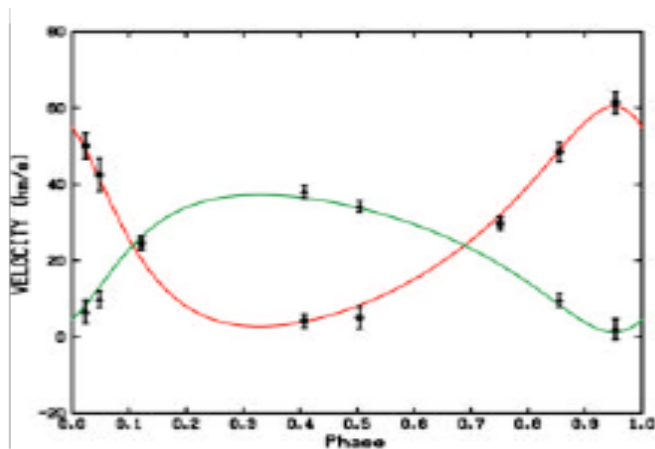
brown dwarf
or planemo?

Direct Mass Determinations



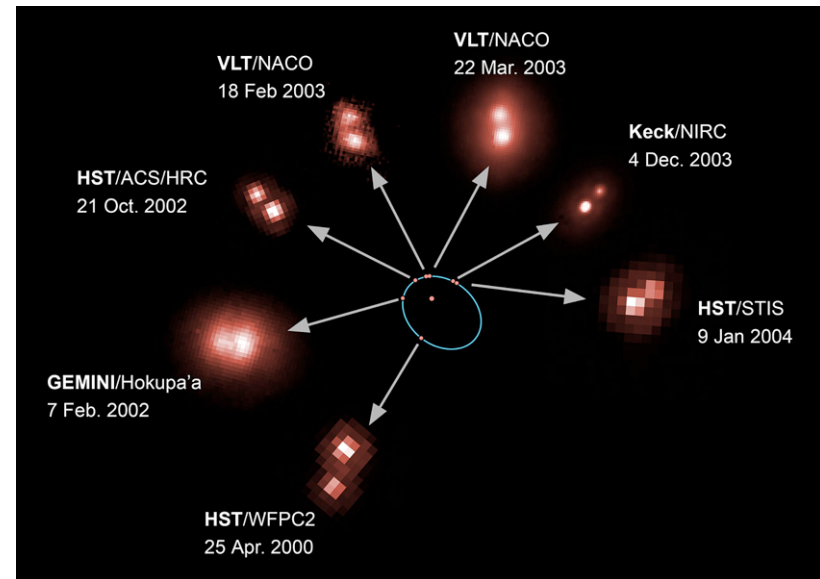
GJ 569Bab

Lane et al. (2001); Zapatero Osorio et al. (2004)



2MASS 0535-0546AB

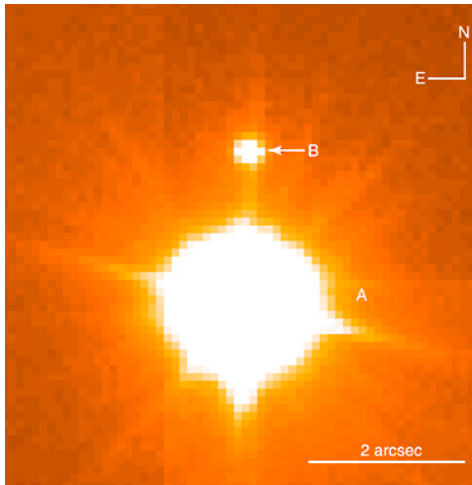
Stassun et al. (2006)



2MASS 0746+2000

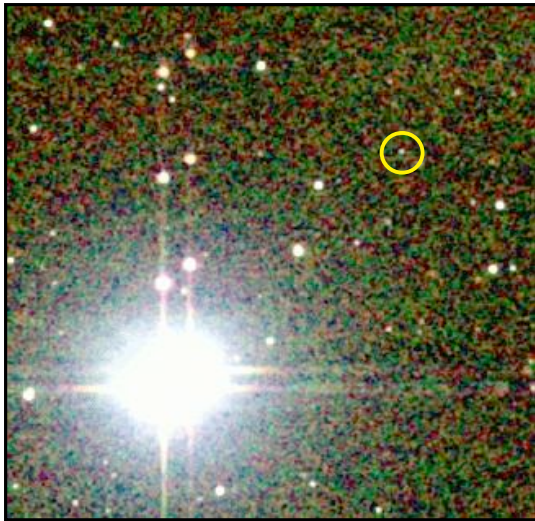
Bouy et al. (2004)

Direct Age Determinations



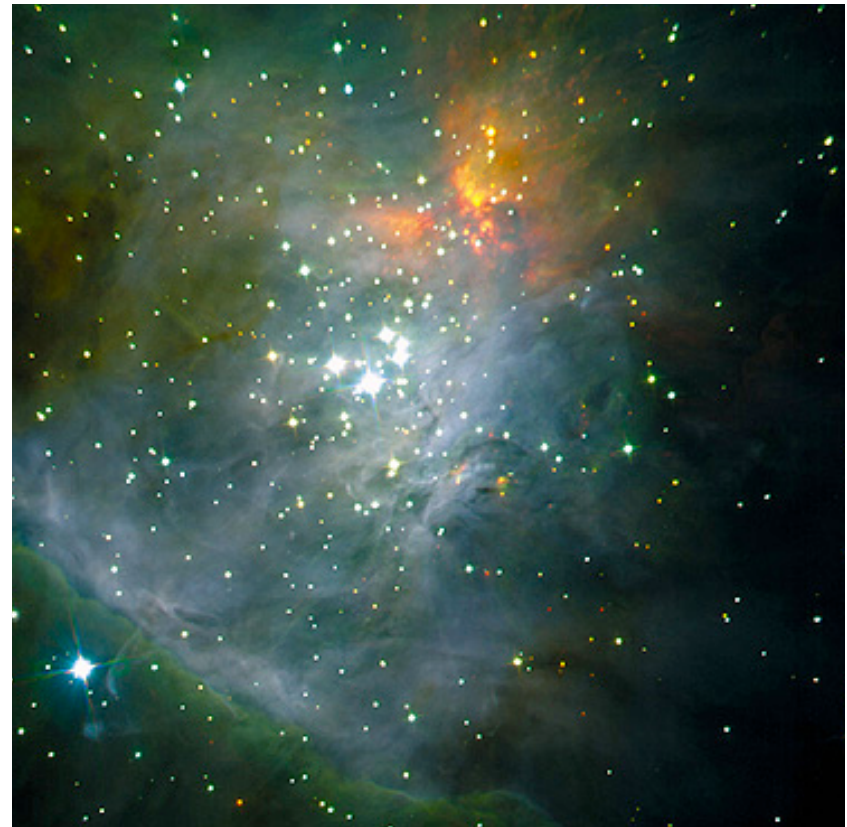
TWA 5B, ~12 Myr

Lowrance et al. (1999); Neuhäuser et al. (2000)

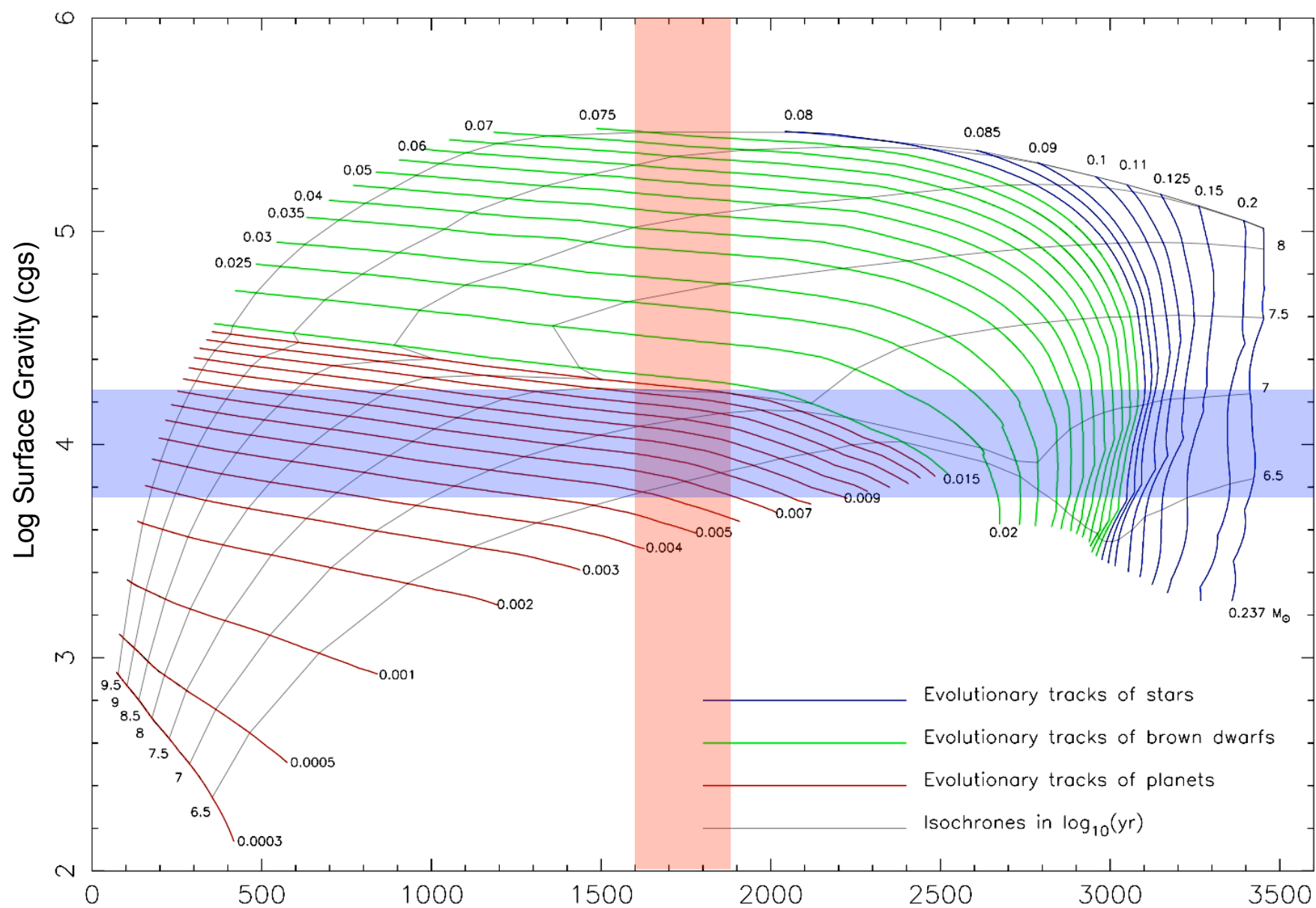


Gliese 570D, 2-5 Gyr

Burgasser et al. (2000)

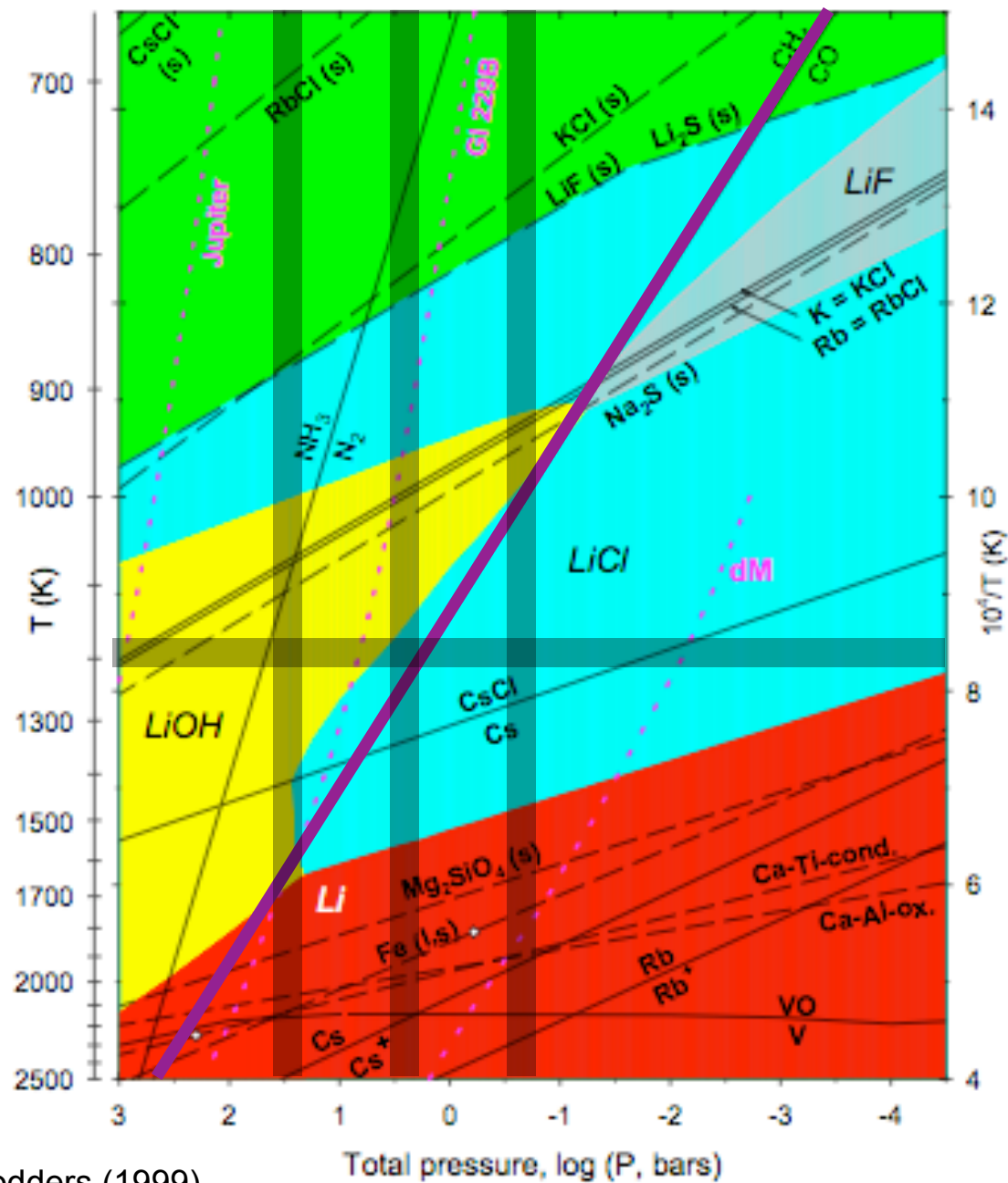


Orion Nebula Cluster ~ 1 Myr



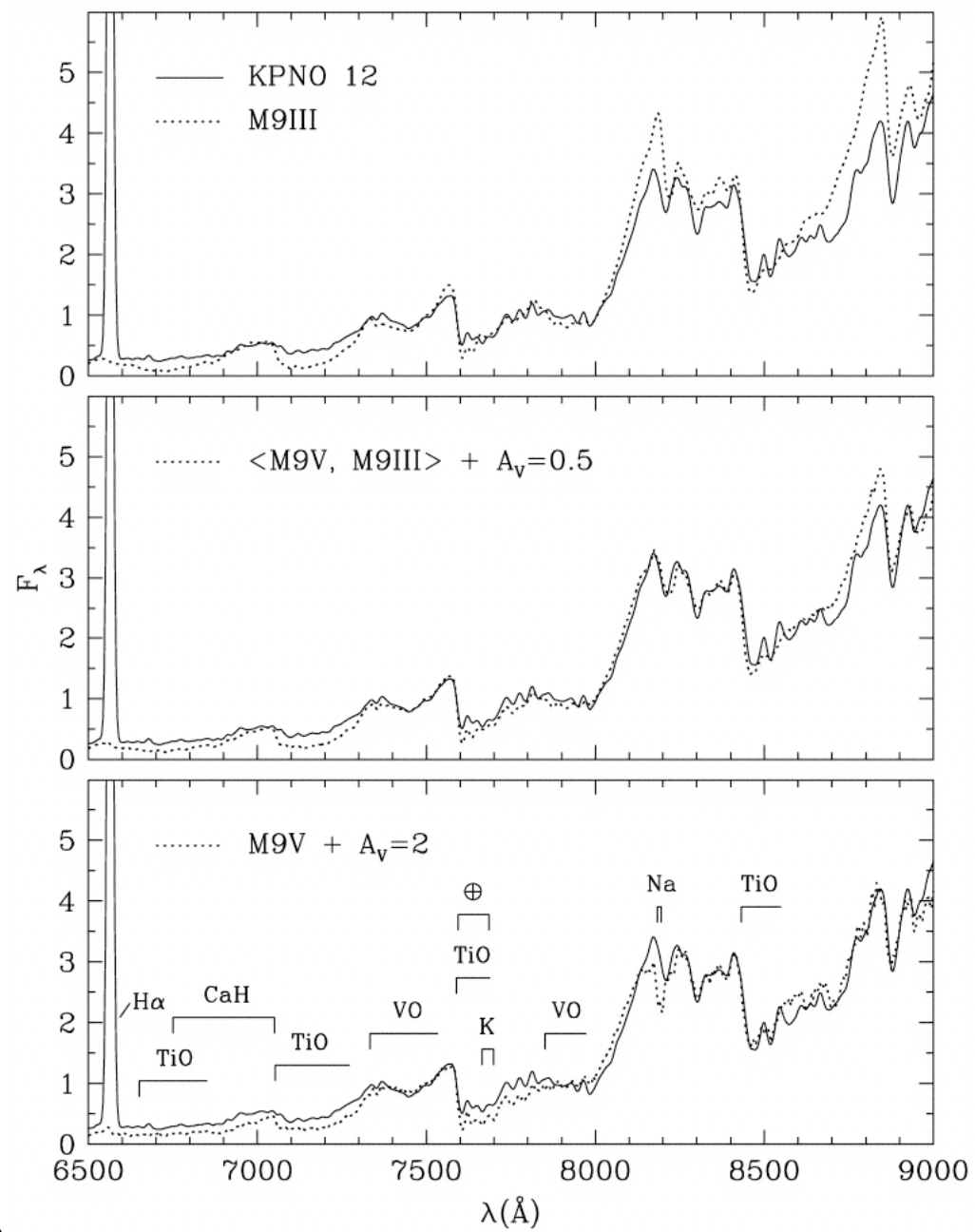
Burrows et al. (1997)

Effective Temperature (K)



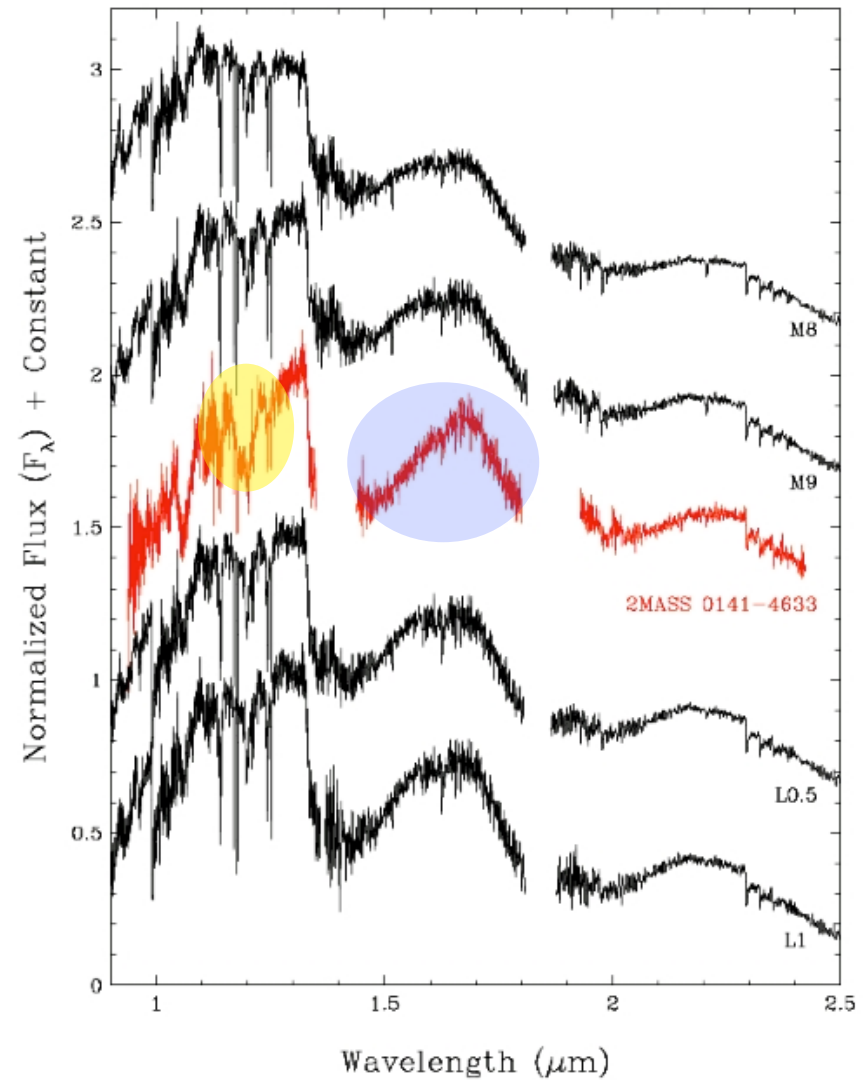
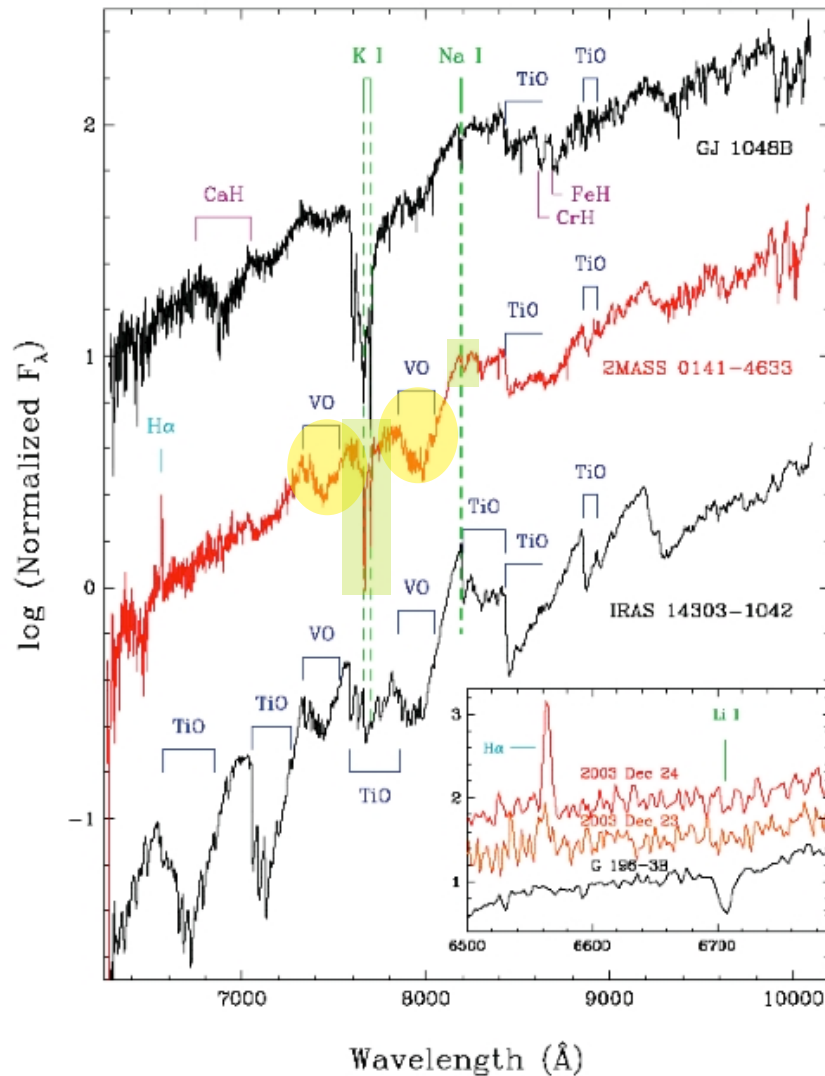
Lodders (1999)

$$P_{\text{ph}} \propto g/\kappa_R$$



Luhman et al. (2003)

2MASS 0141-4633: ~10 Myr and nearby

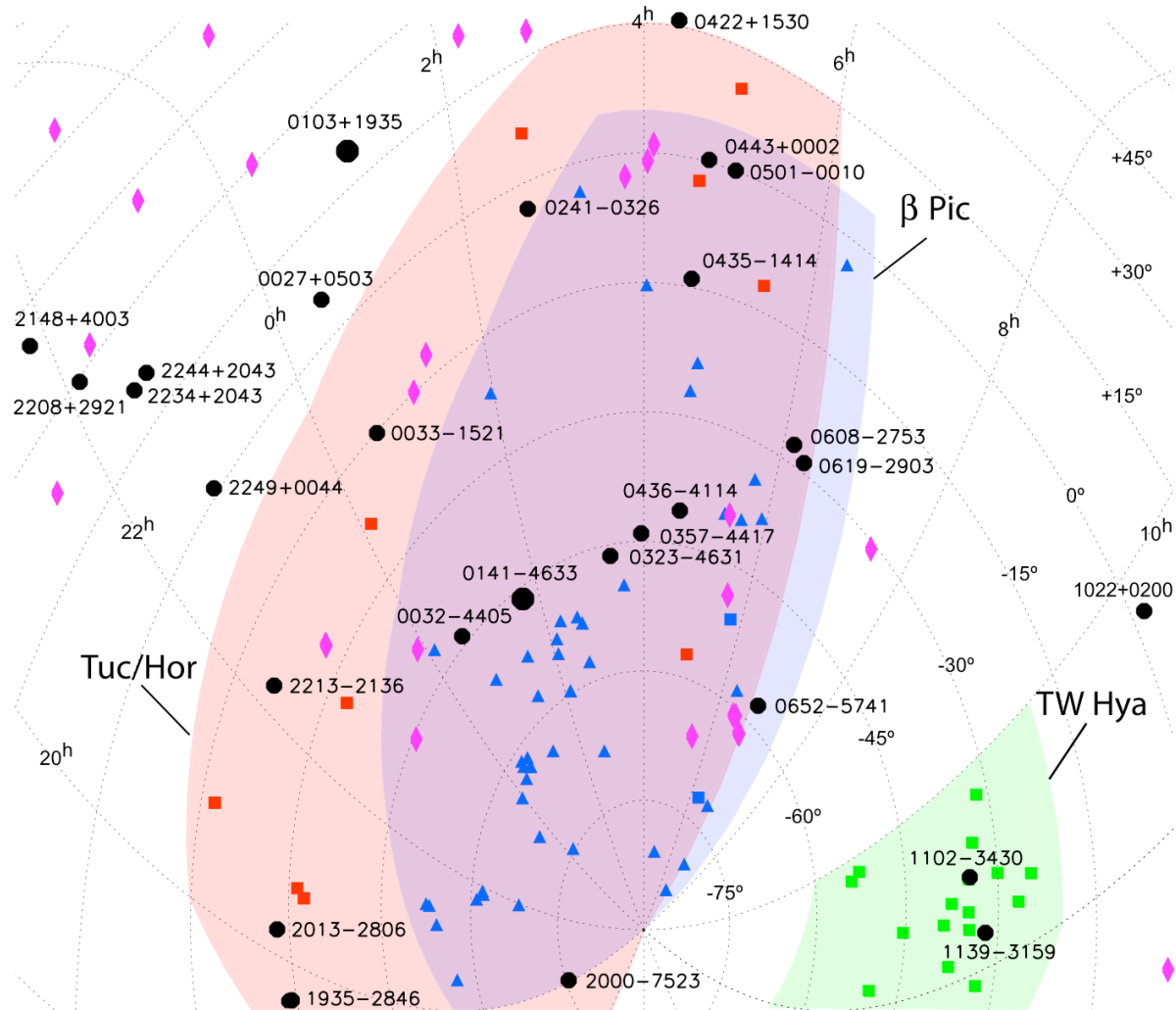


Kirkpatrick et al. (2006)

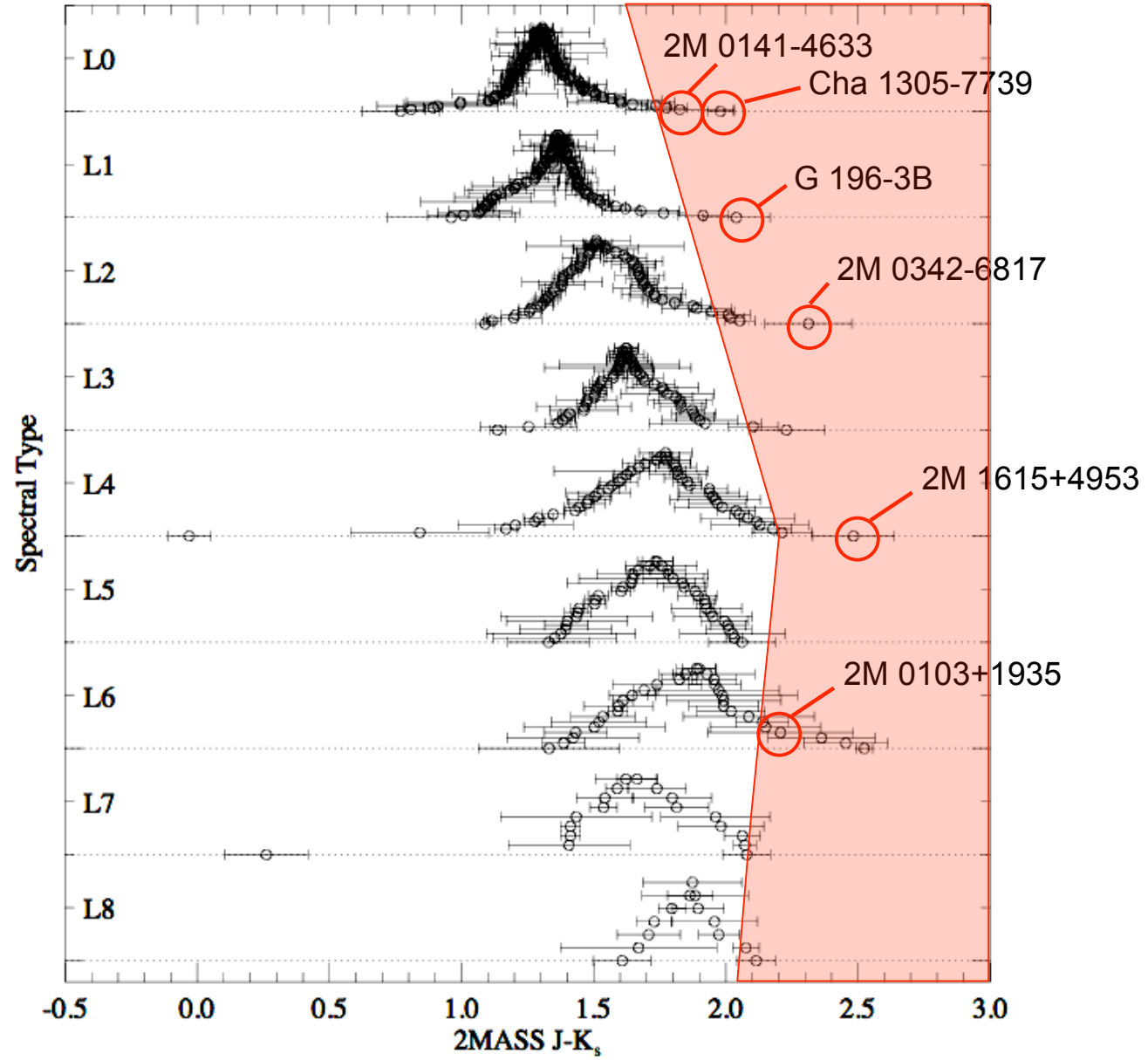
indicators of low surface gravity/youth:

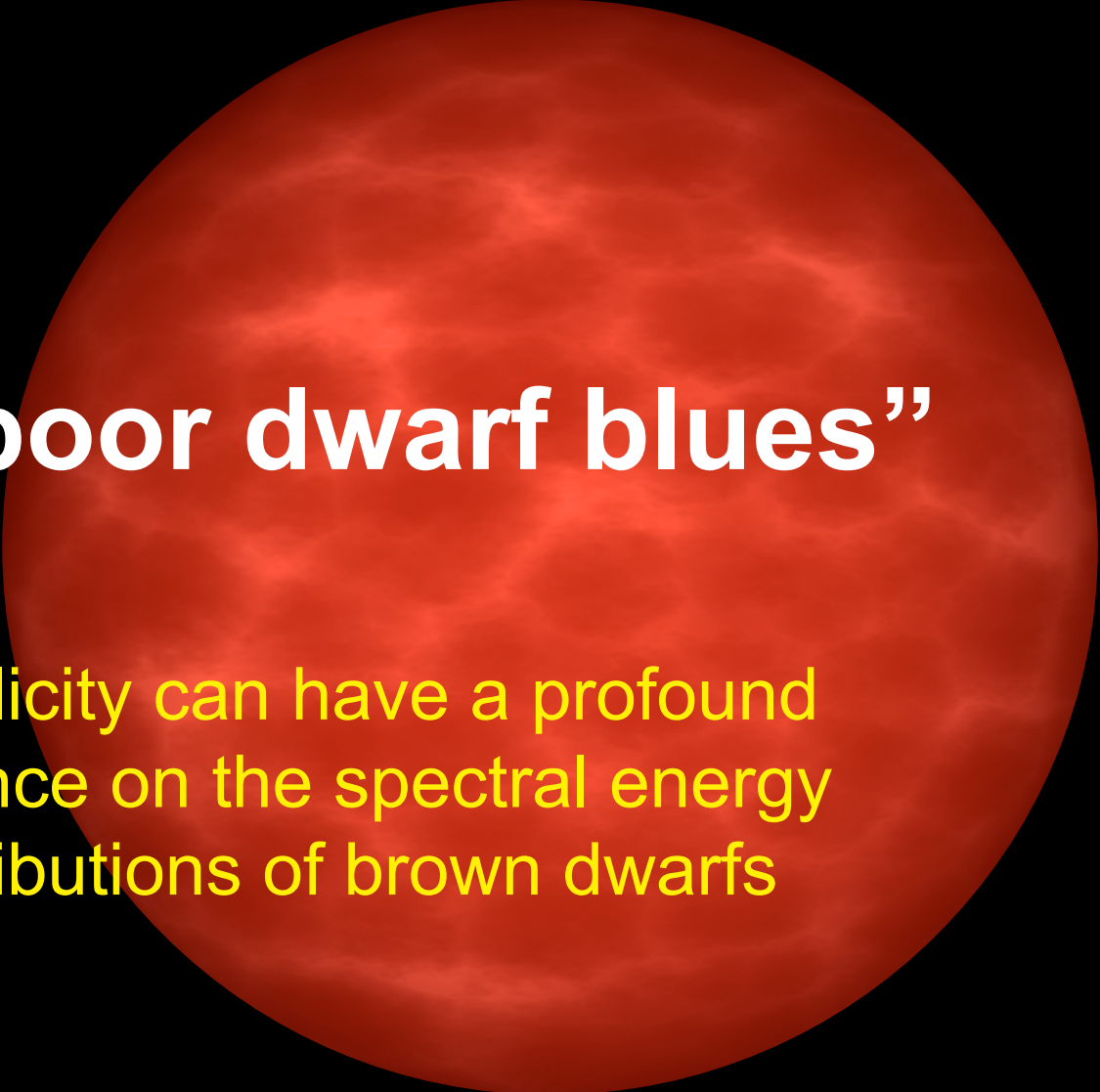
1. weak optical alkali lines (Na I, K I, Li I?)
2. enhanced metal oxide bands (VO, H₂O, CO)
3. a “peaky” H-band peak
4. reddened near-infrared SEDs

10-50 Myr brown dwarfs in the vicinity of the Sun



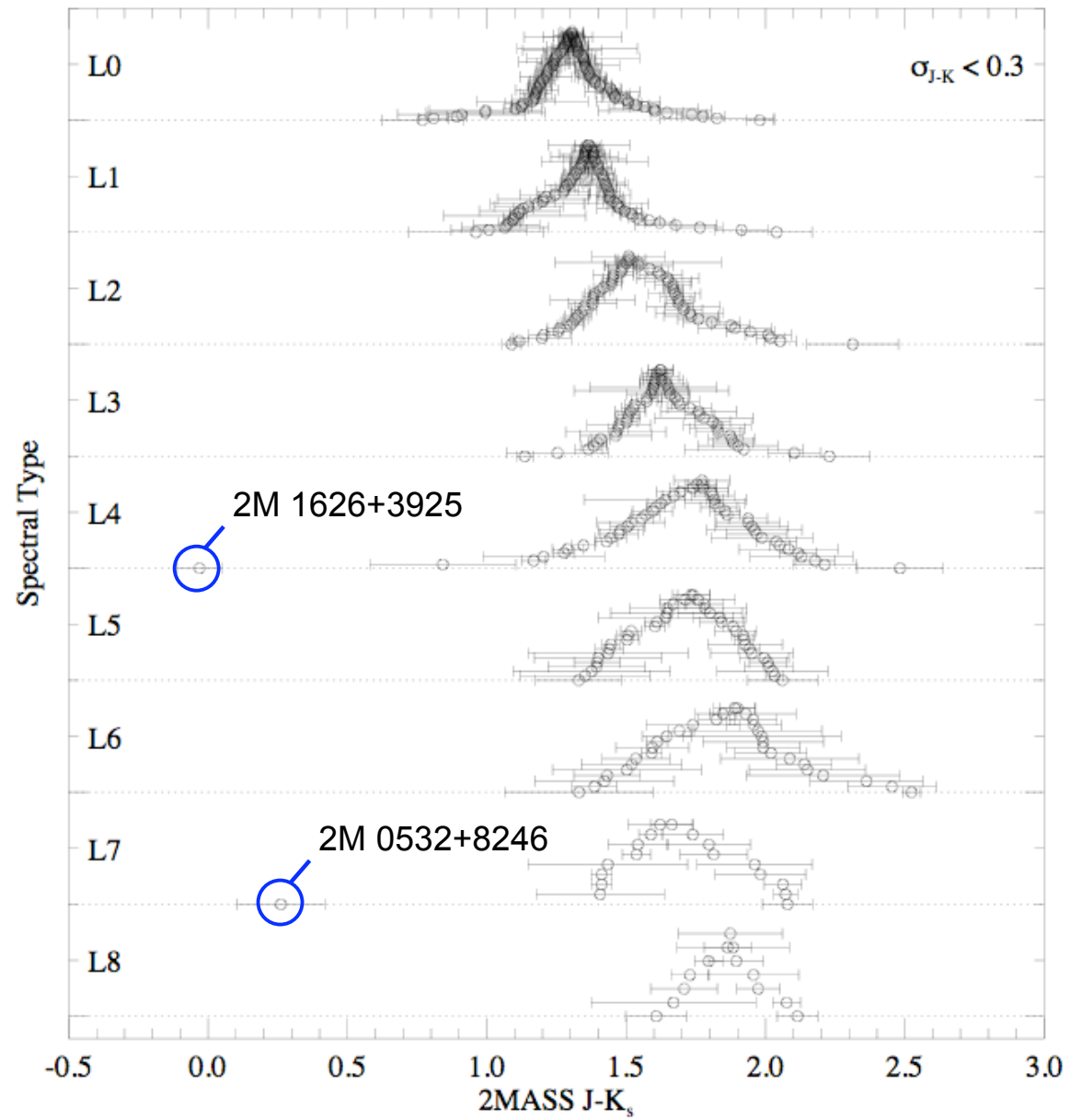
Cruz et al. (2006)

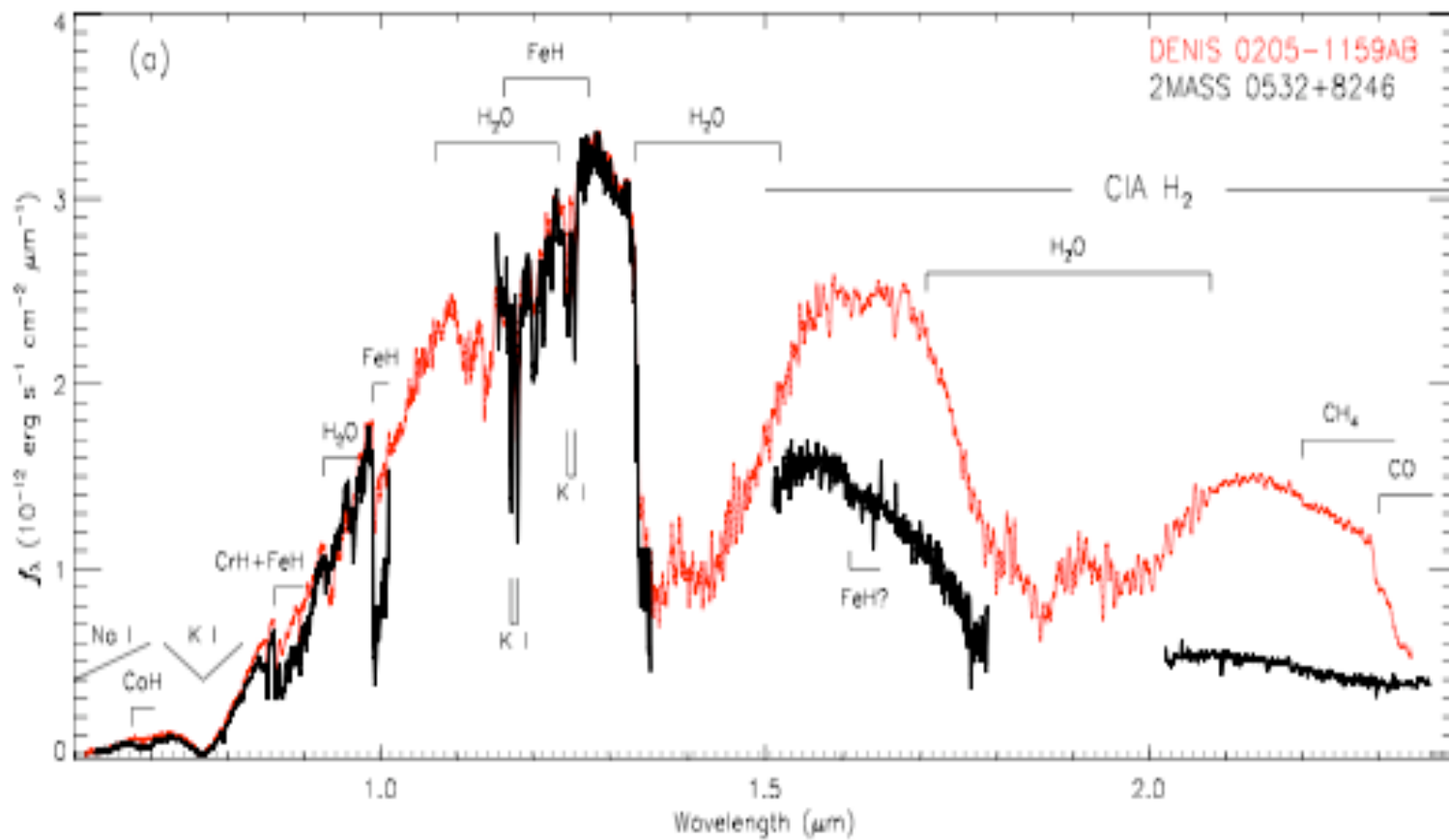




“the poor dwarf blues”

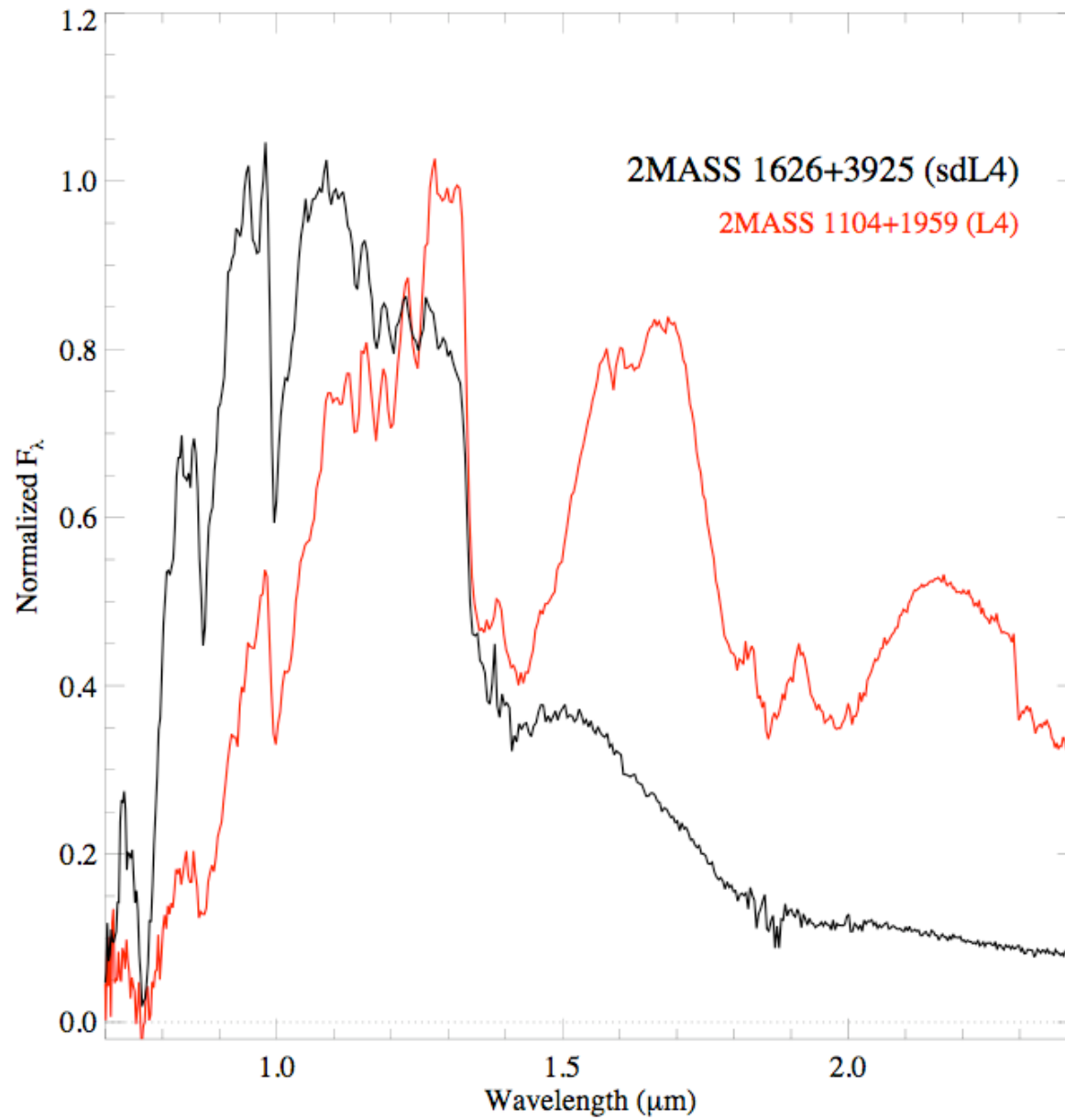
metallicity can have a profound
influence on the spectral energy
distributions of brown dwarfs



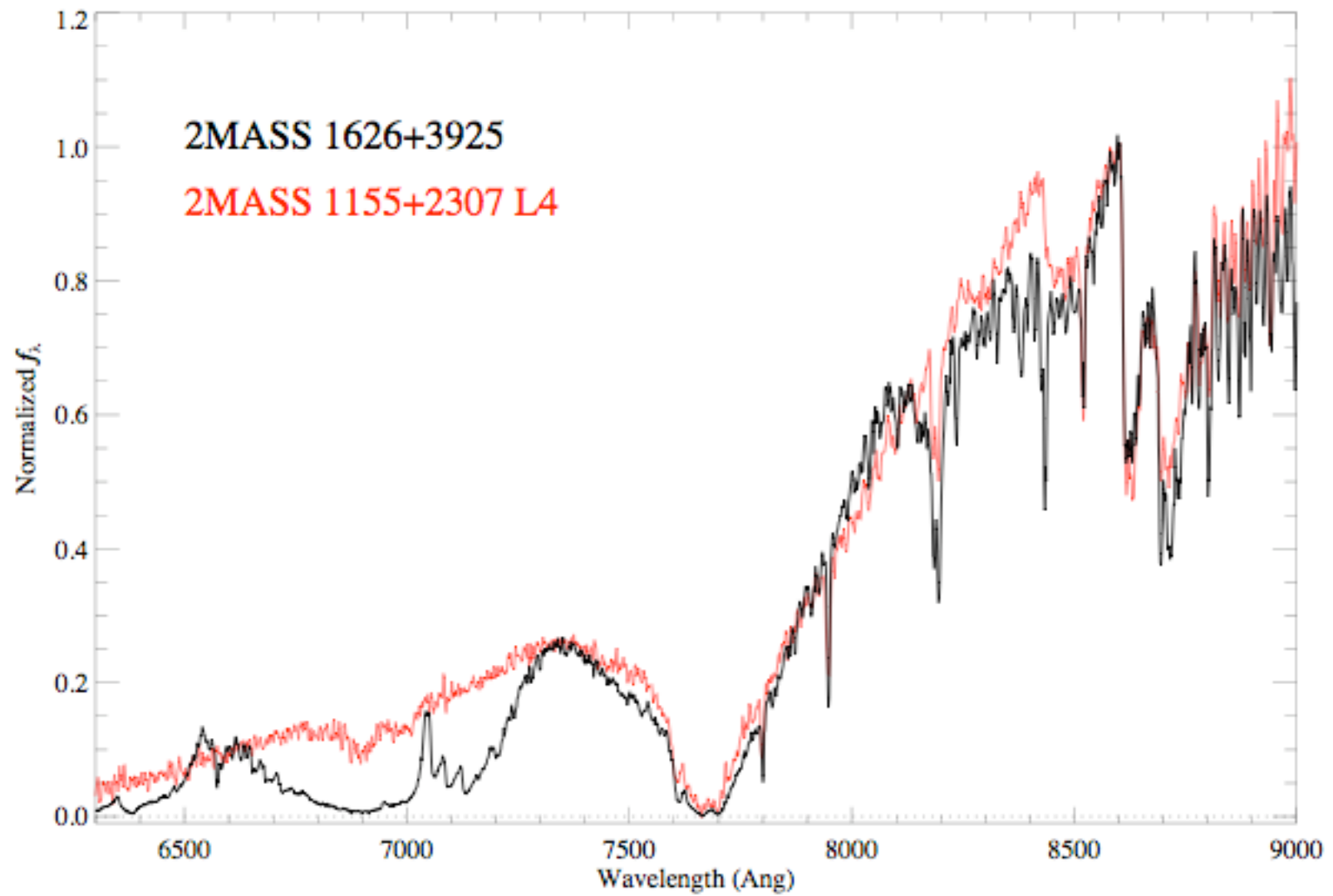


2MASS 0532+8246

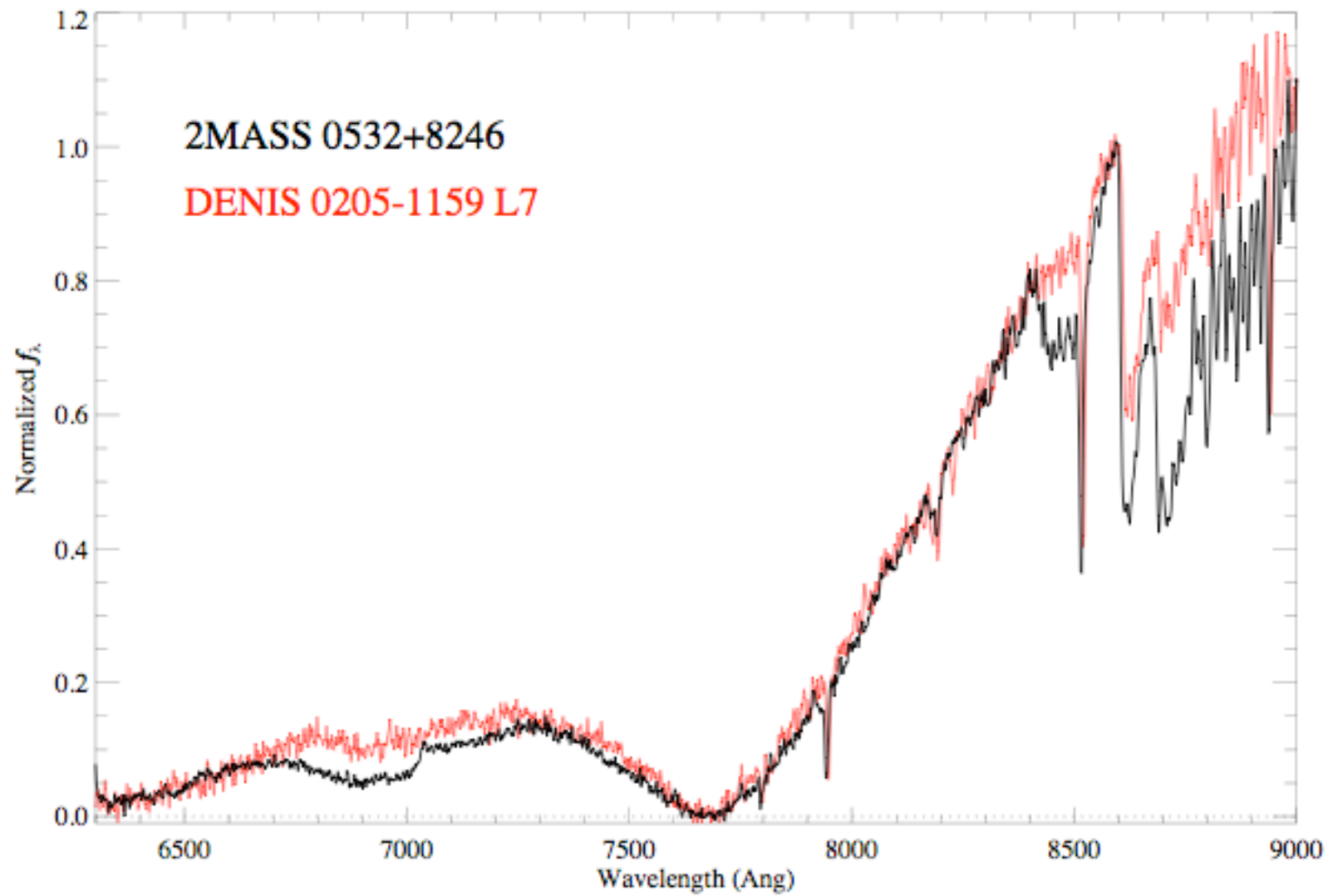
Burgasser et al. (2003)



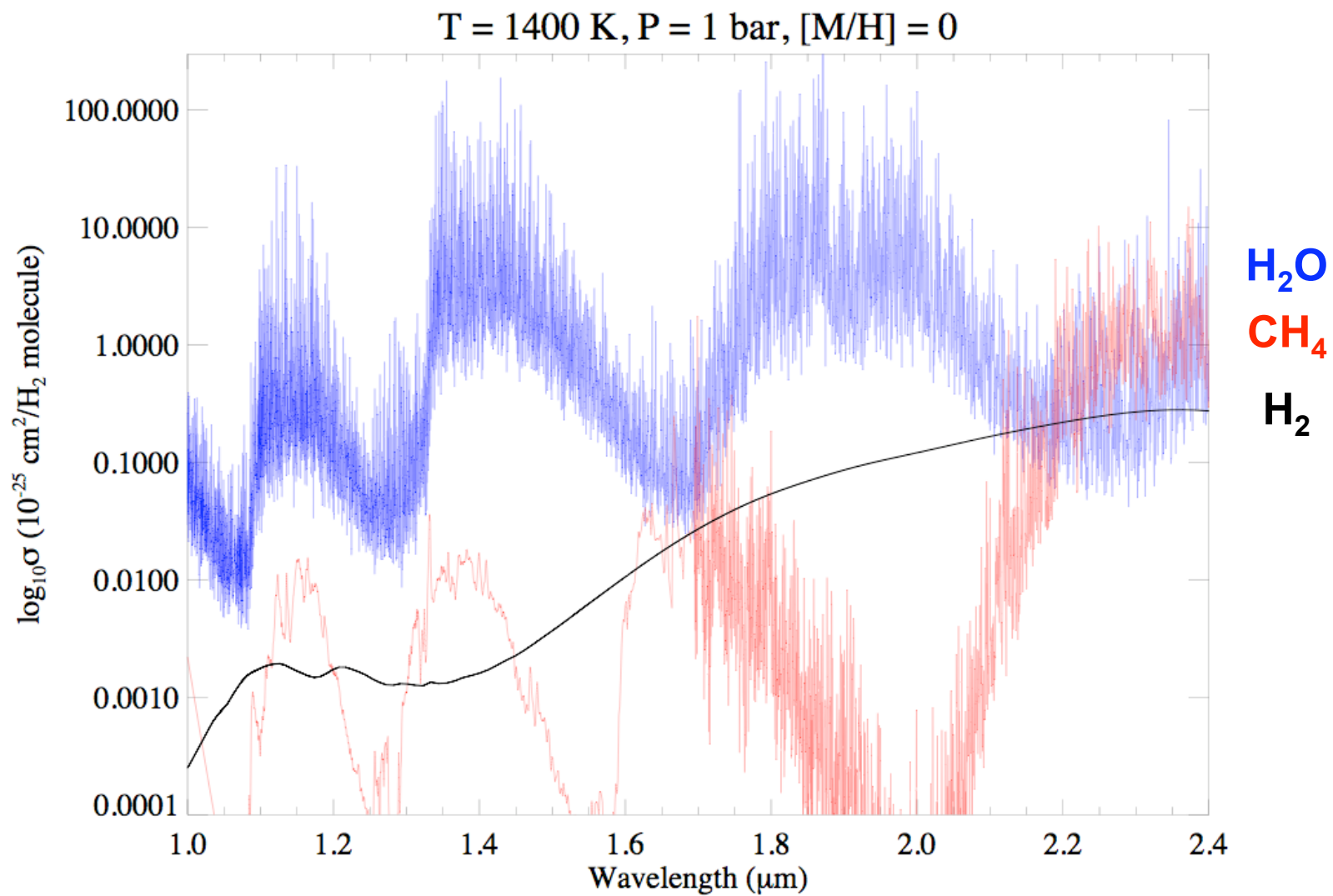
2MASS 1626+3925
Burgasser (2004)



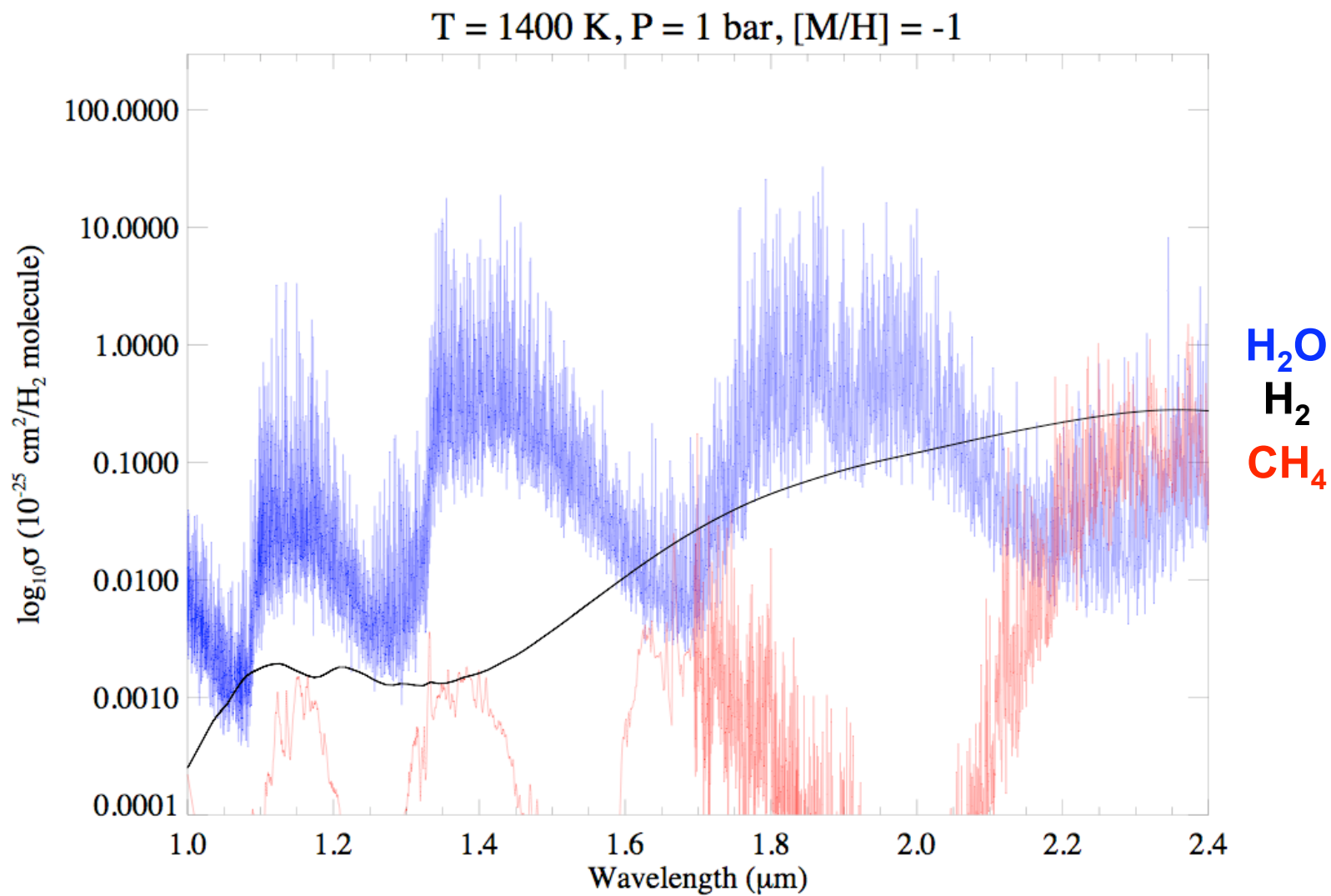
Burgasser, Kirkpatrick & Cruz (2006)



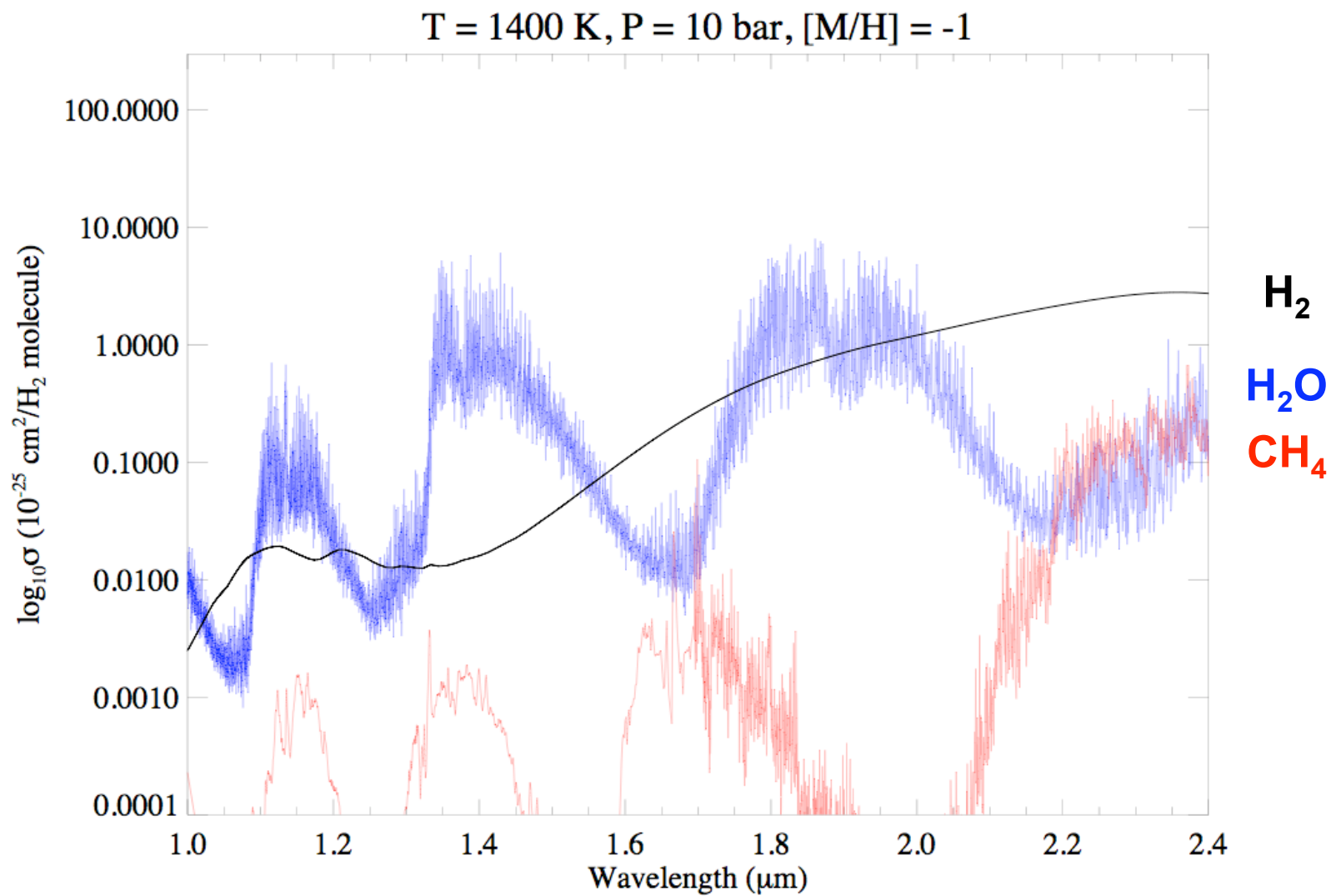
Burgasser, Kirkpatrick & Cruz (2006)



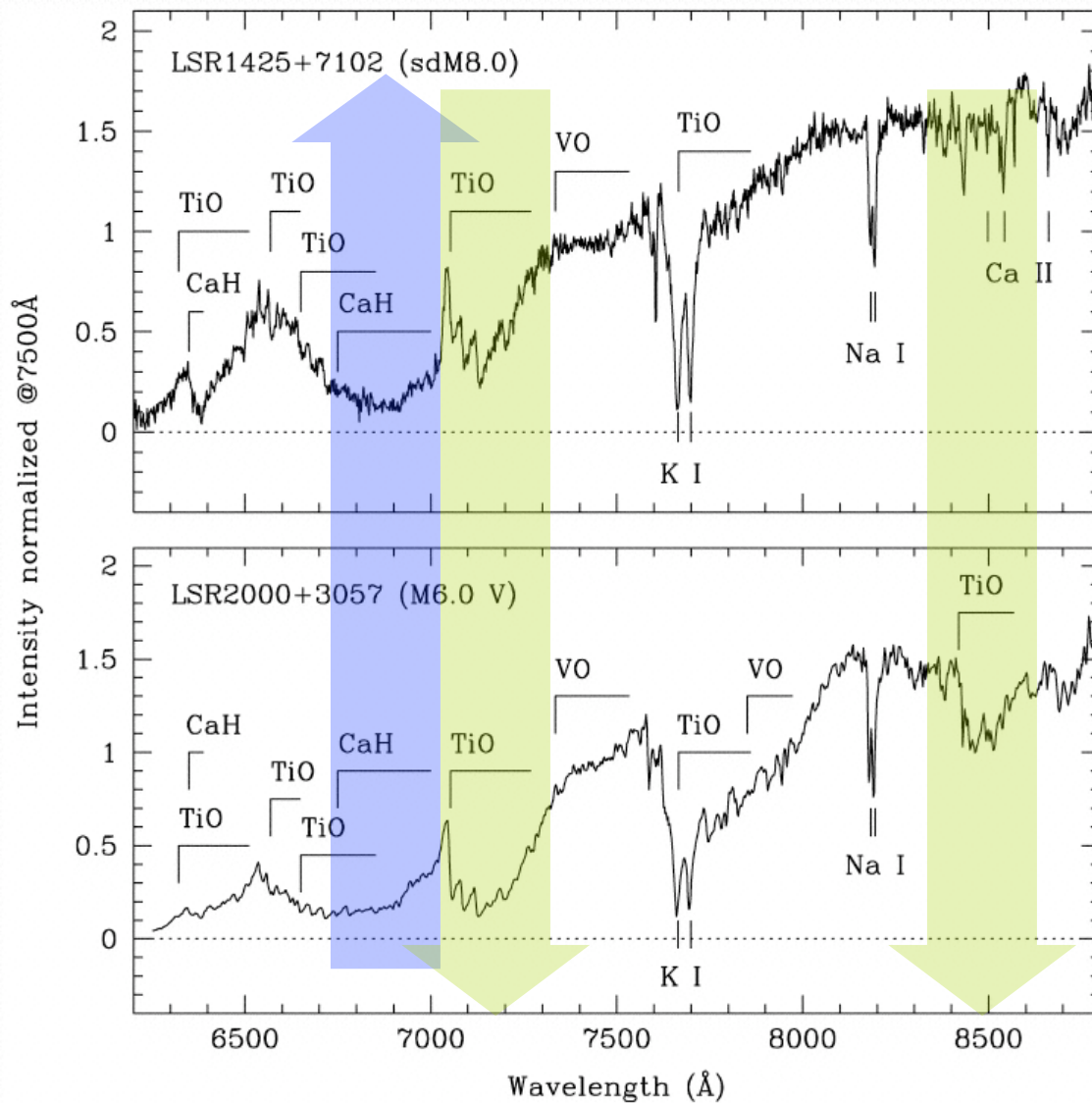
opacities from A. Burrows



opacities from A. Burrows



opacities from A. Burrows



M subdwarfs

strong metal hydrides
weak metal oxides

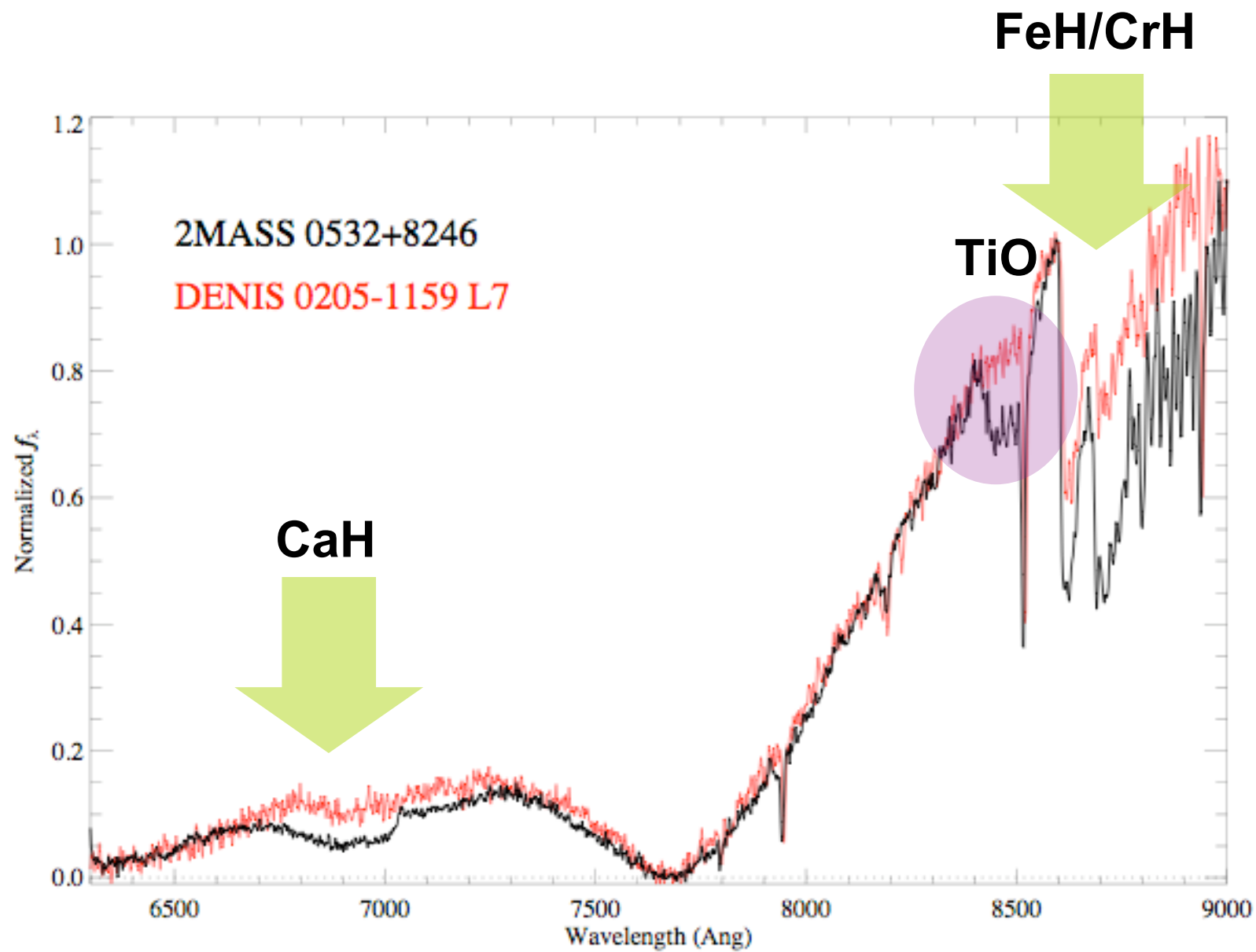
sdMs: $[M/H] \sim -1.2$

esdMs: $[M/H] \sim -2.0$

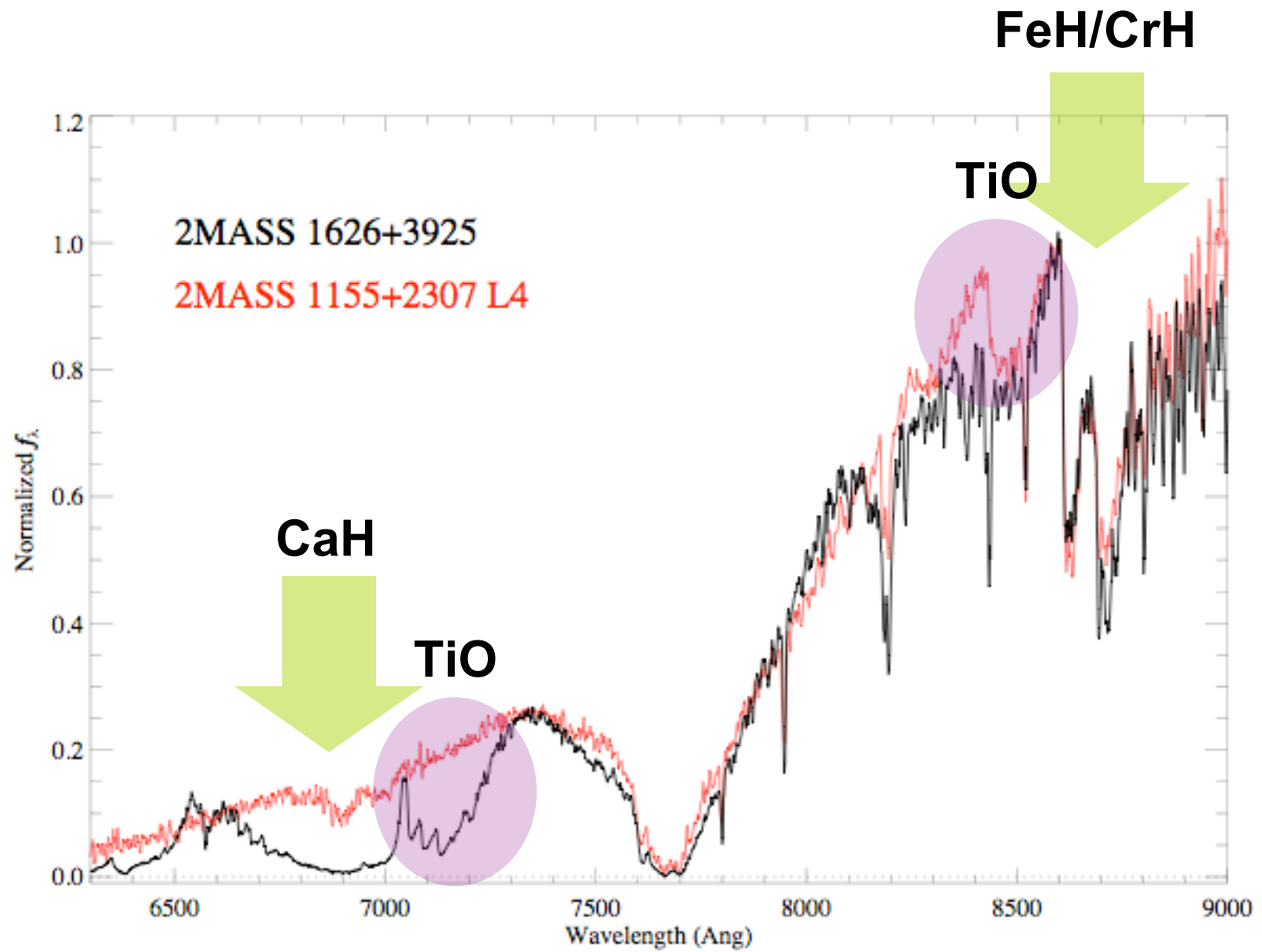
$V_{\text{LSR}} \sim -200 \text{ km/s}$

Lepine et al. (2003)

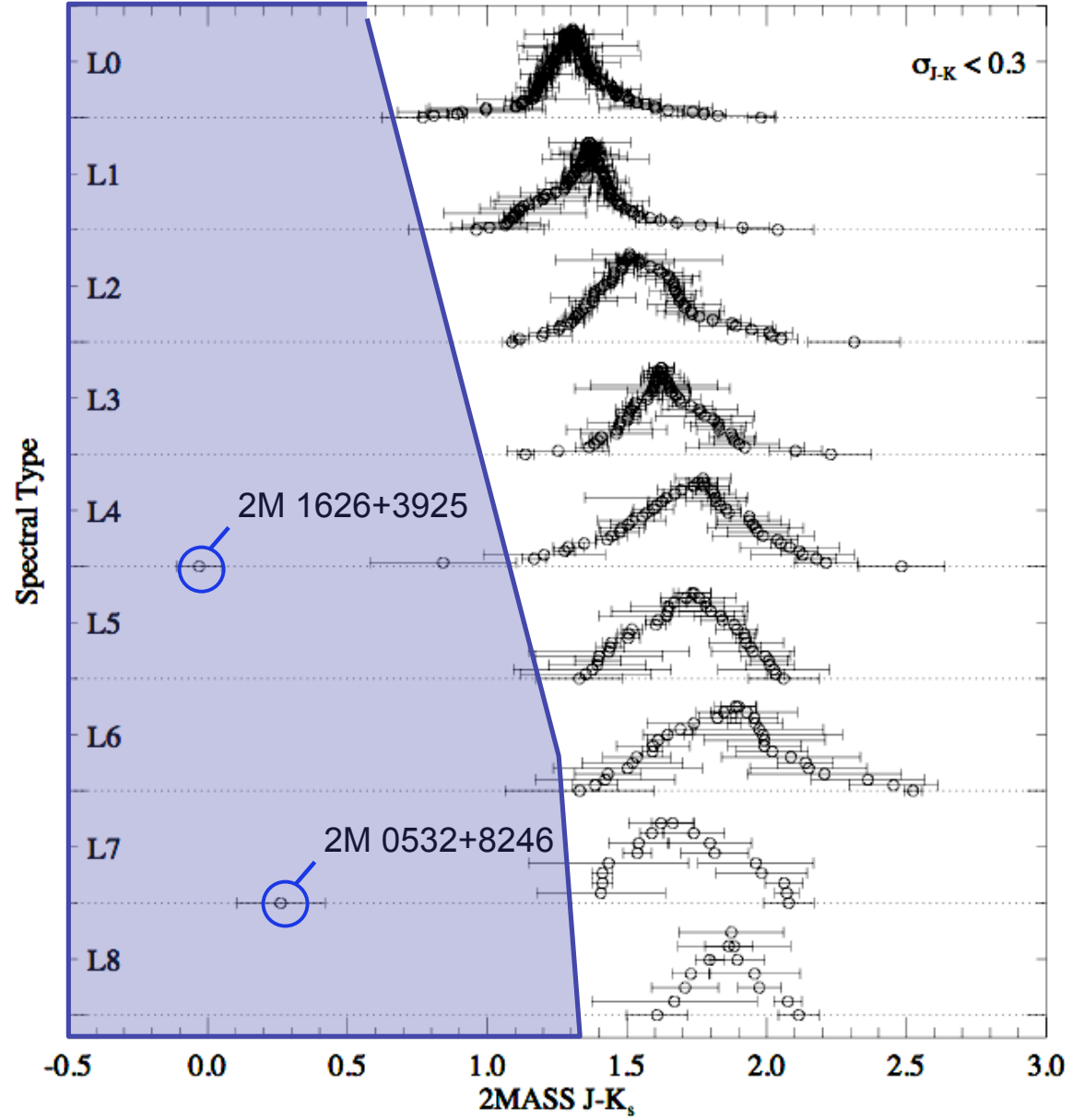
see also Gizis (1997); Lepine et al. (2004); Burgasser et al. (2006)

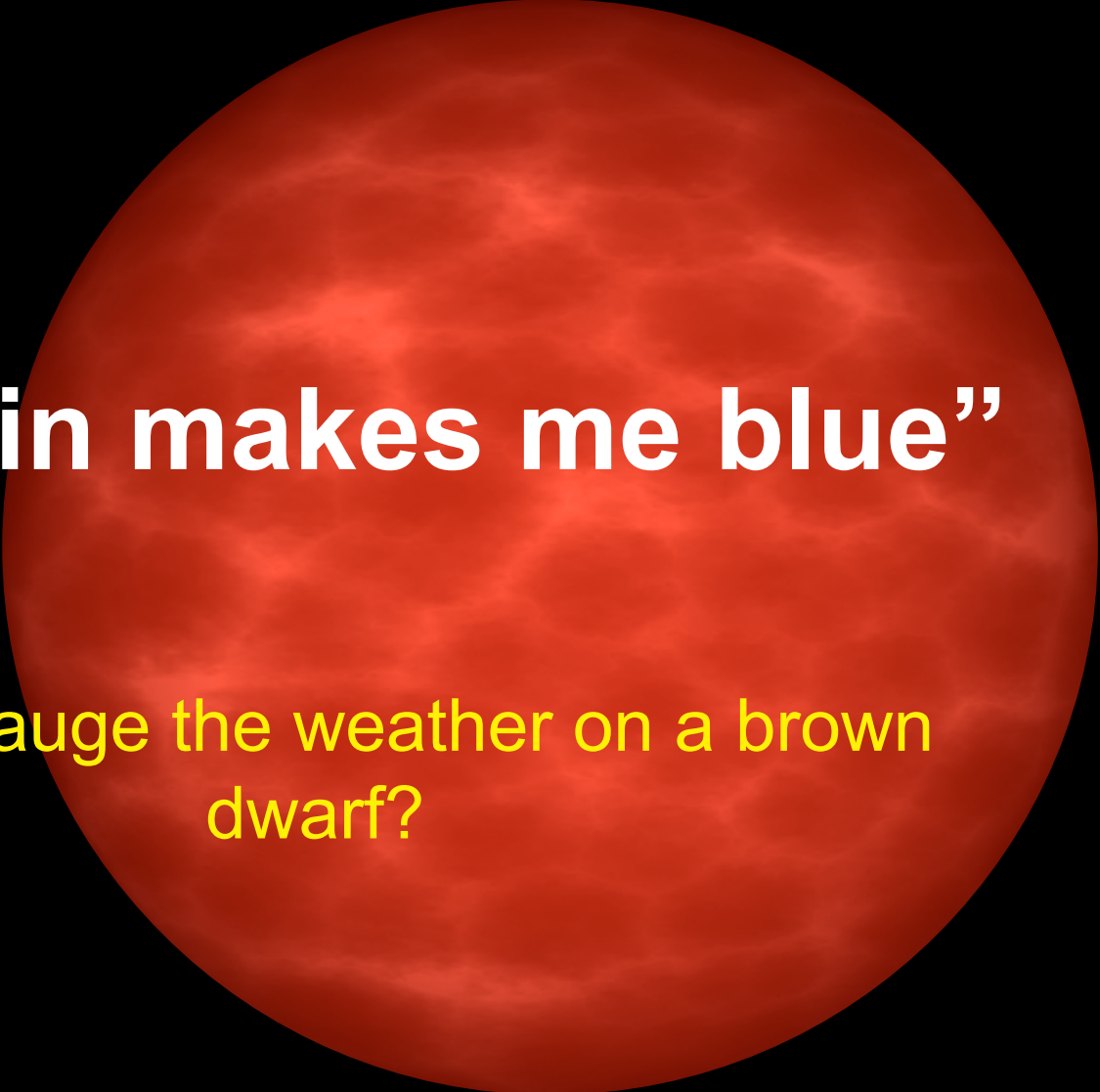


Burgasser, Kirkpatrick & Cruz (2006)



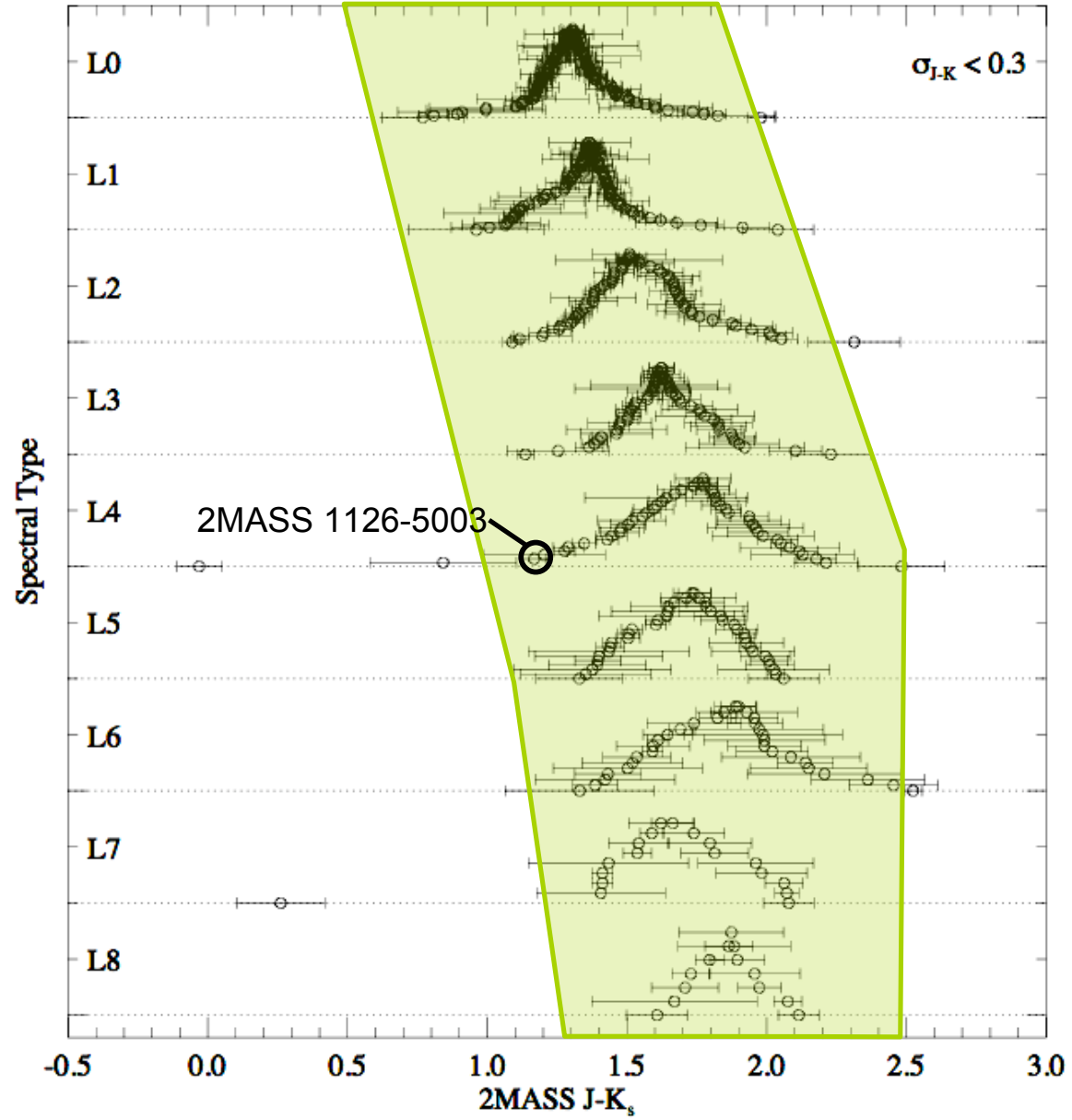
Burgasser, Kirkpatrick & Cruz (2006)

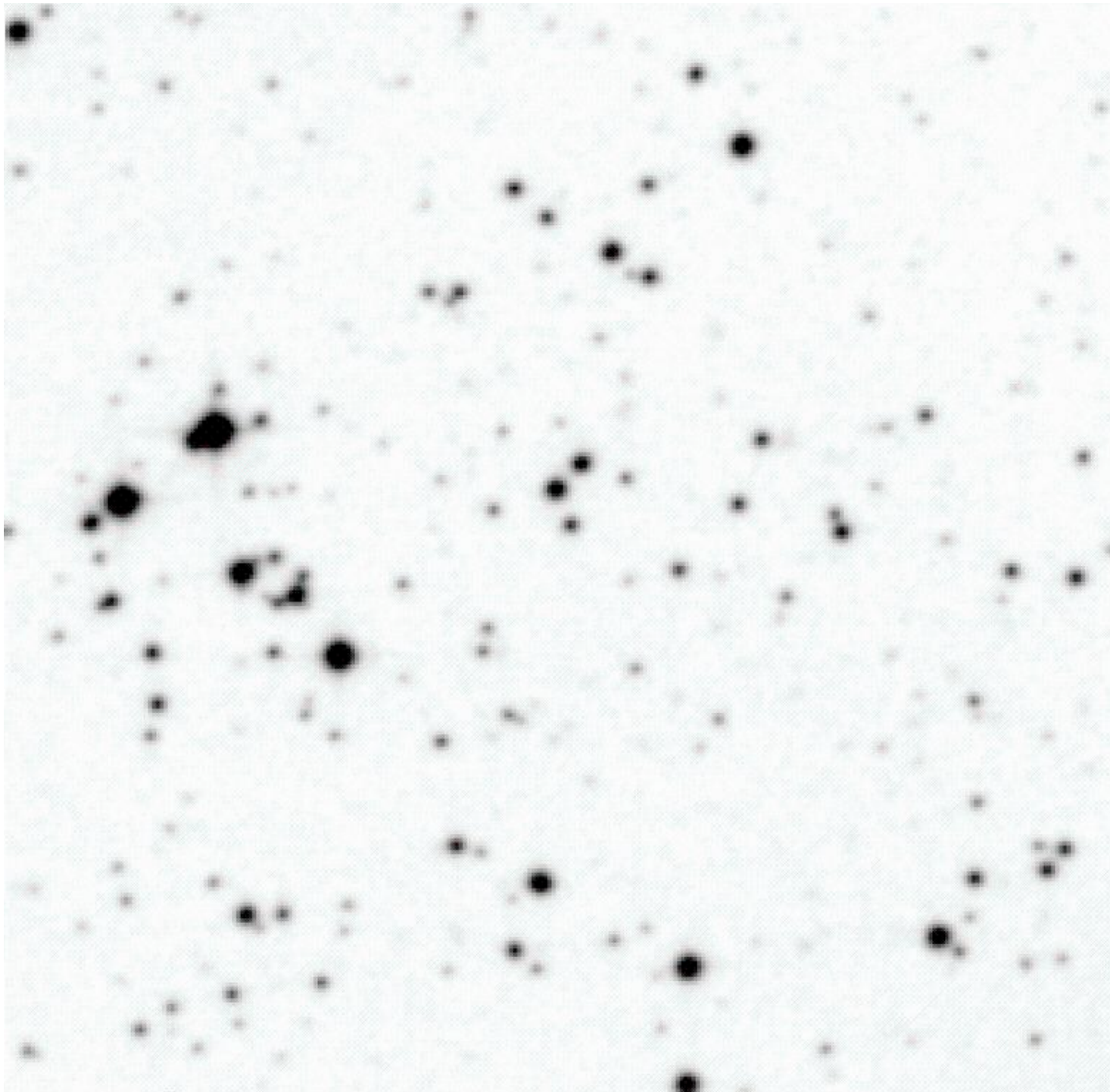




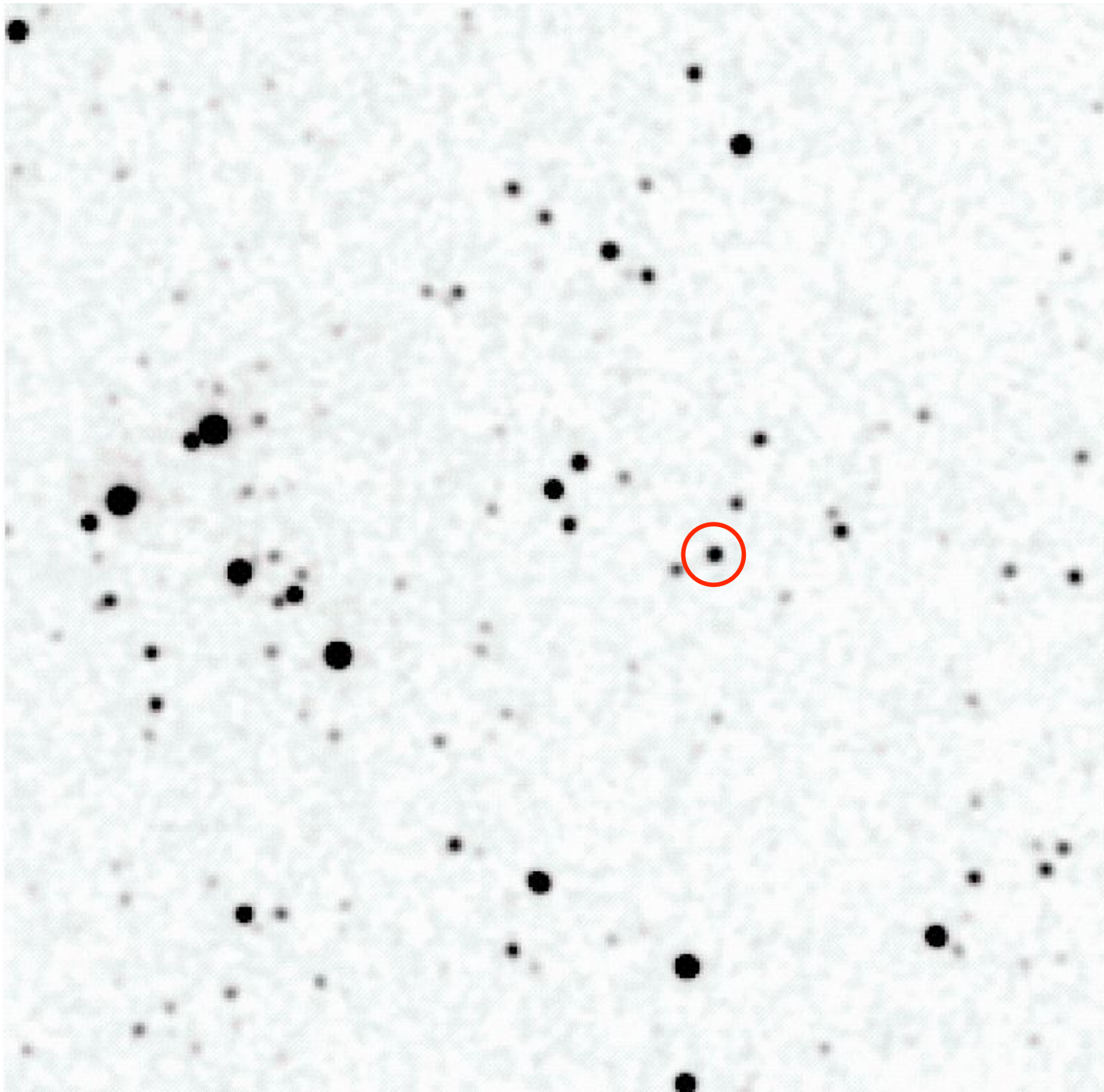
“the rain makes me blue”

can we gauge the weather on a brown
dwarf?

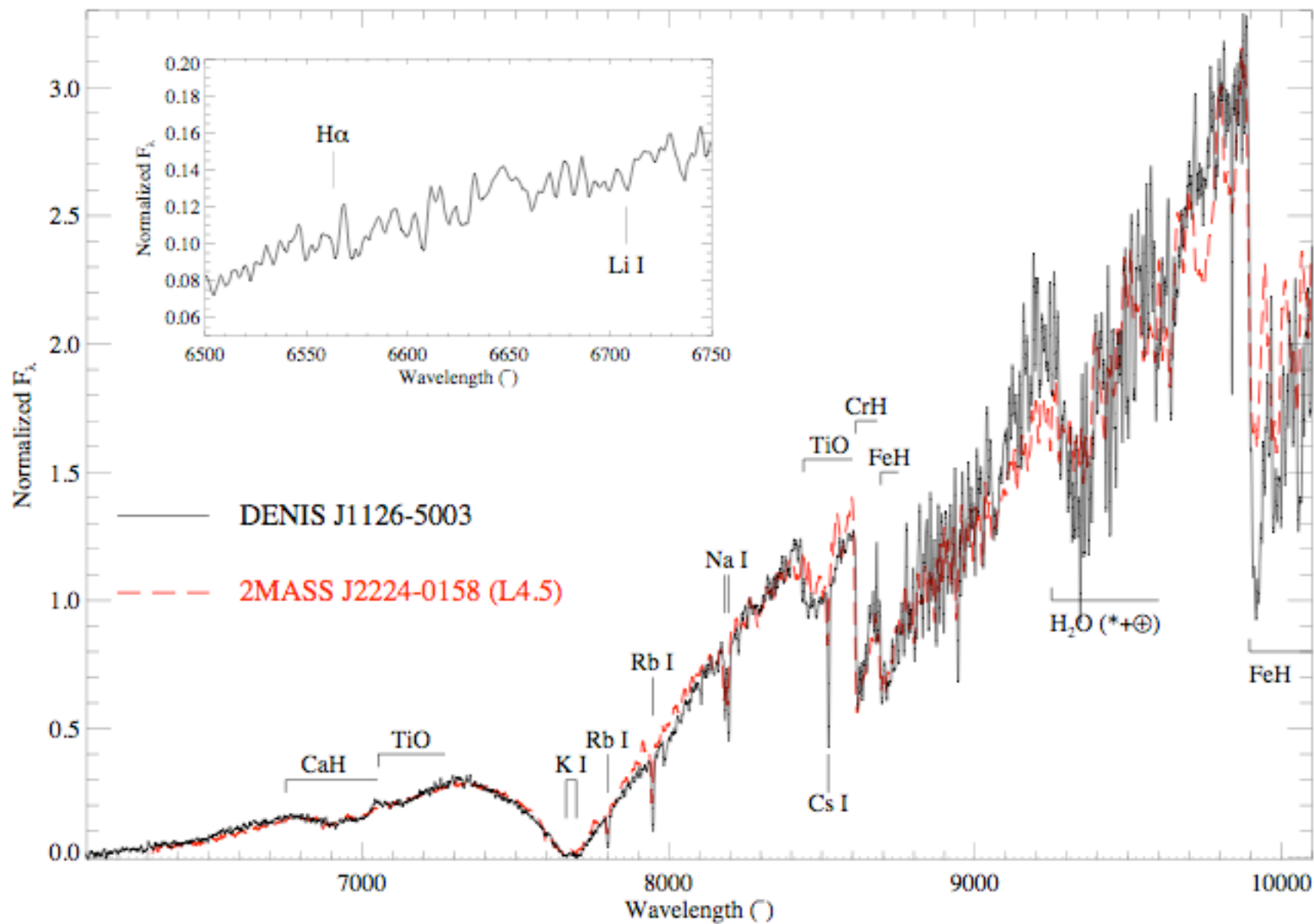


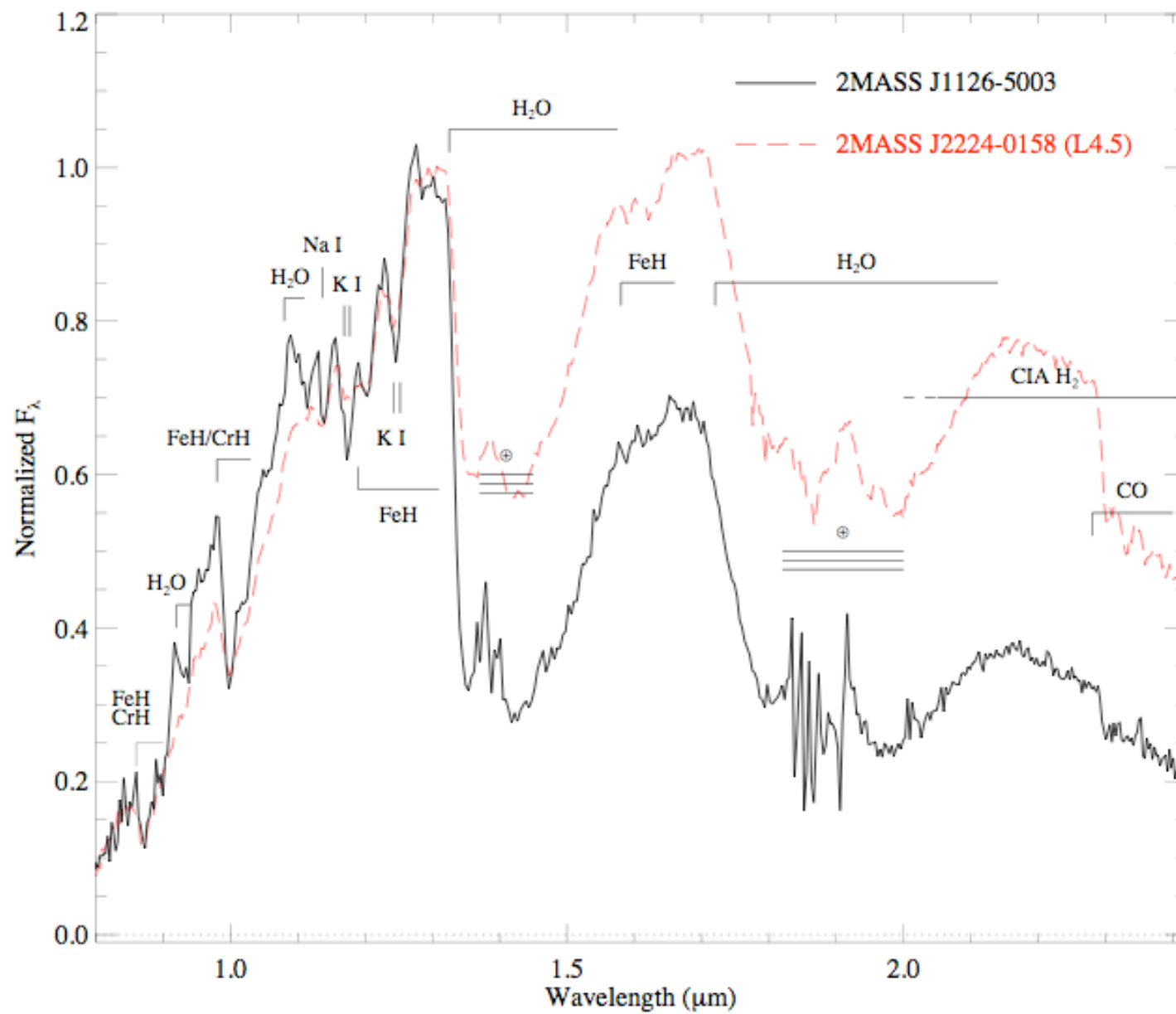


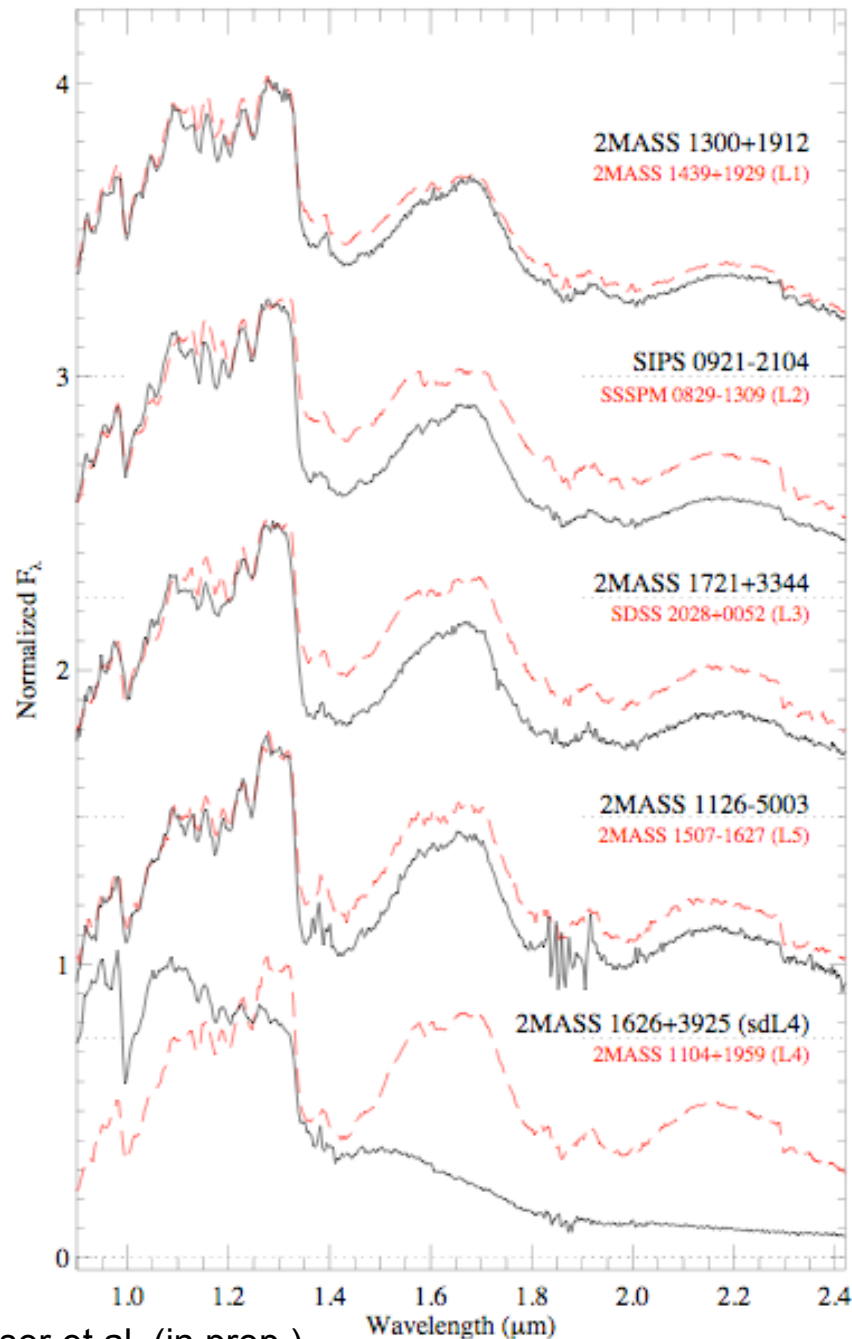
1985
SERC I_N



1999
2MASS J







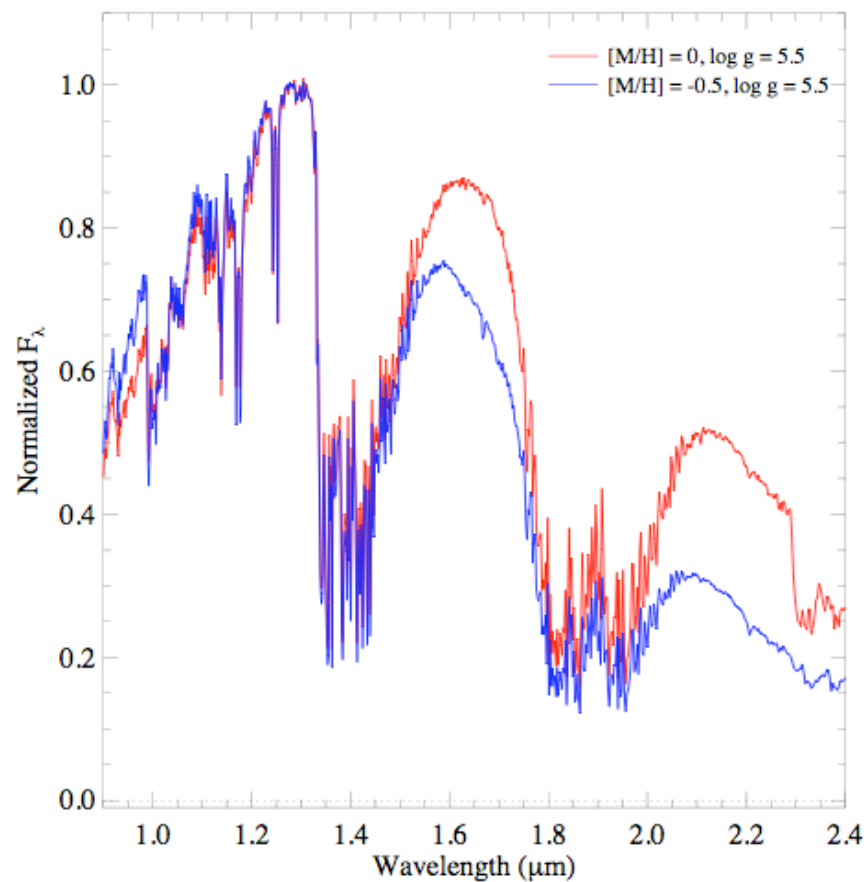
Burgasser et al. (in prep.)

blue L dwarfs

Cruz et al. (2003, 2006); Burgasser et al. (2004); Knapp et al. (2004); Chiu et al. (2006); Schmidt et al. (2007)

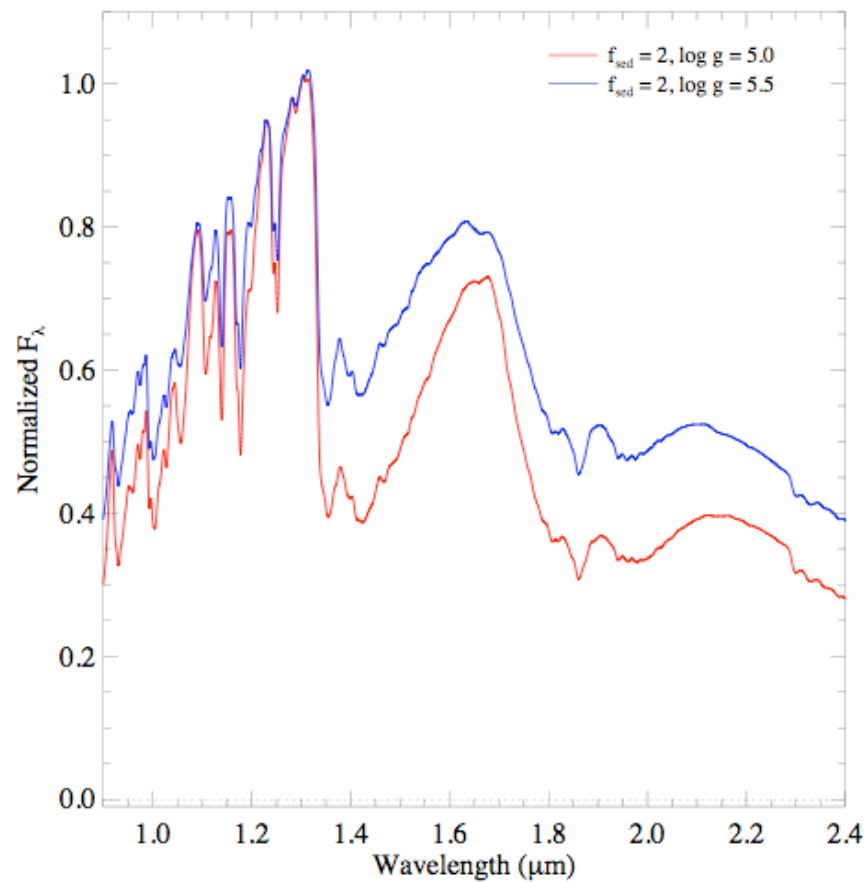
- J-K $\sim$ 0.3 mag too blue
- normal optical spectra
- deep H₂O absorption
- large proper motions

Lower metallicity?

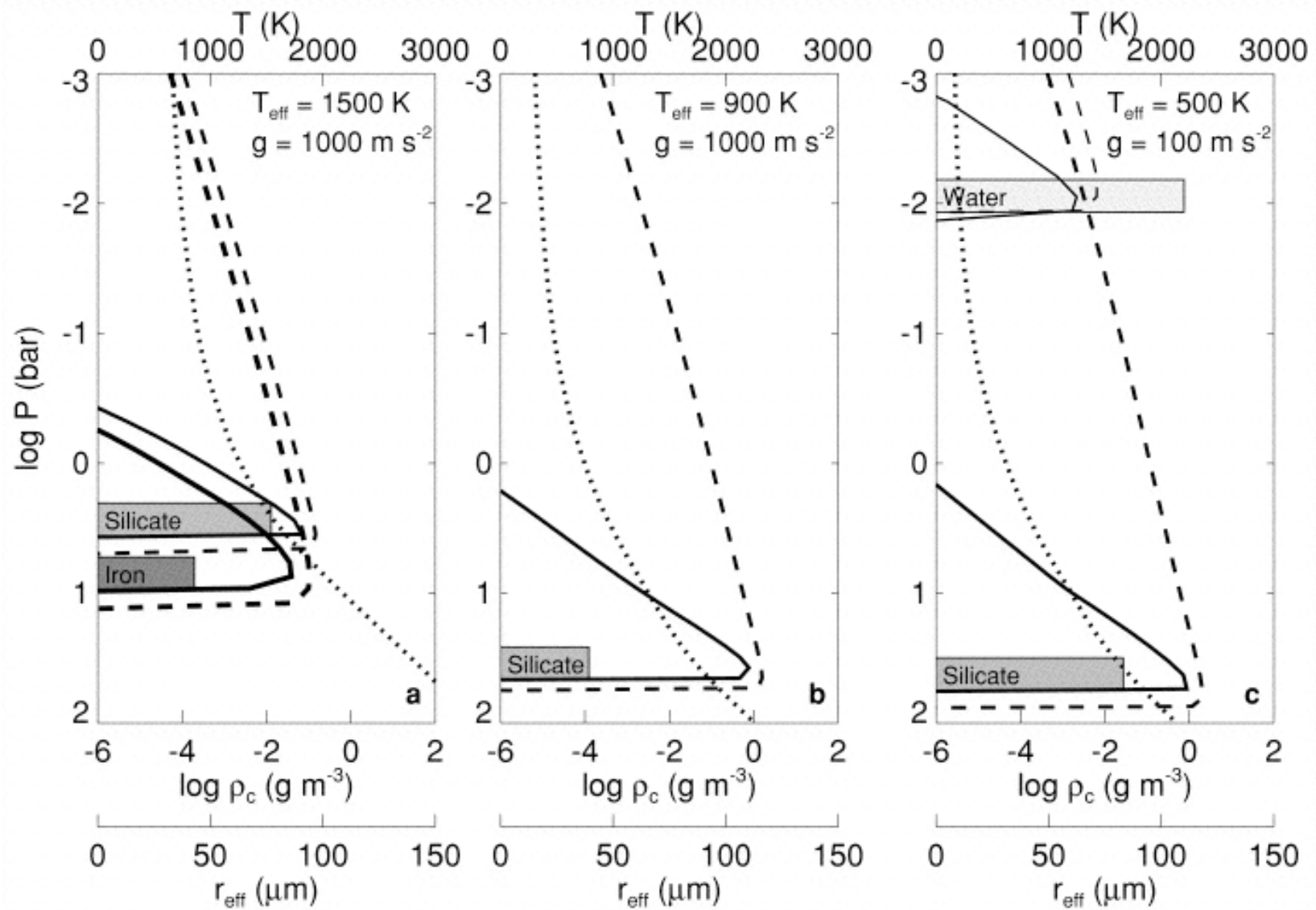


models from Burrows et al. (2006)

Higher surface gravity?

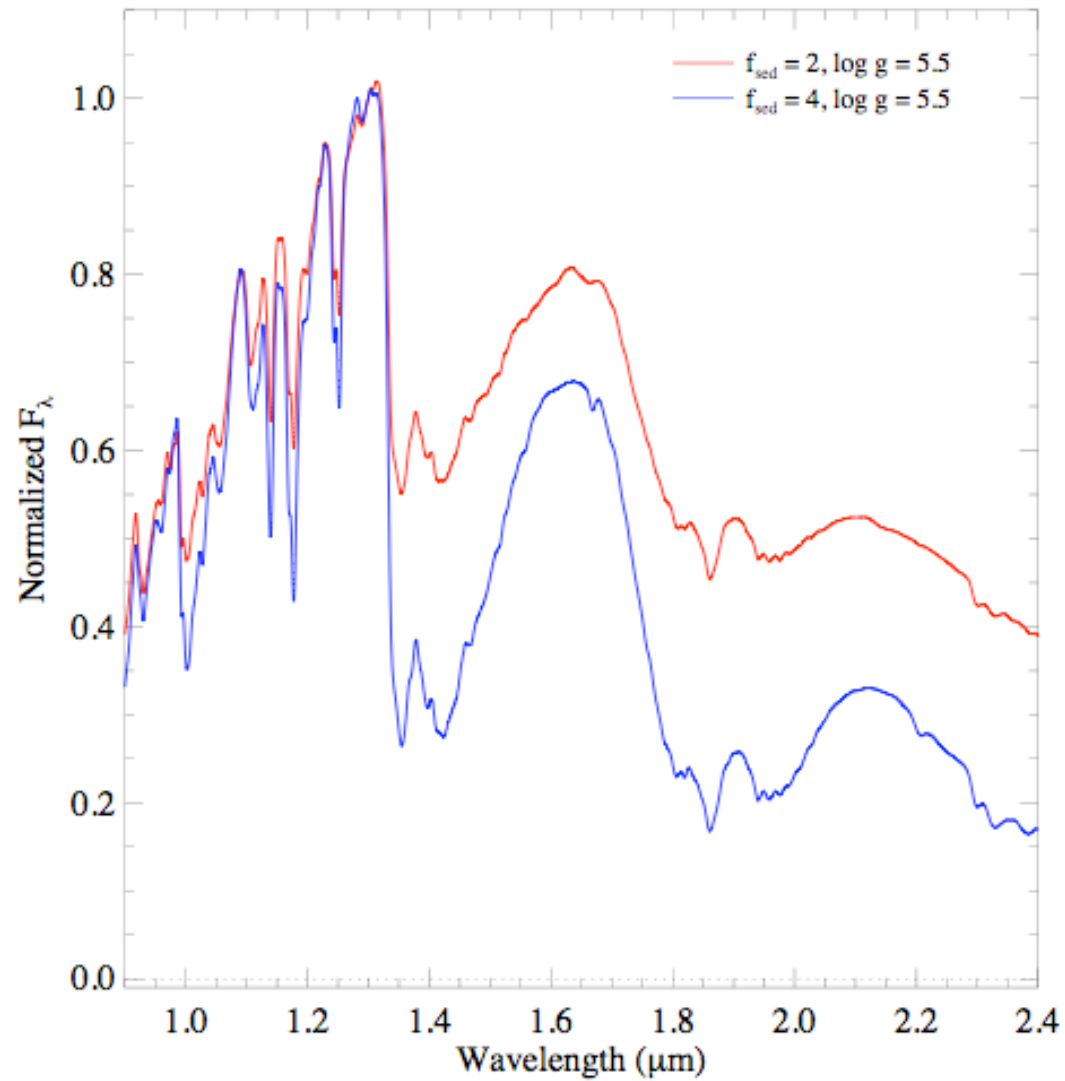


models from D. Saumon & M. Marley



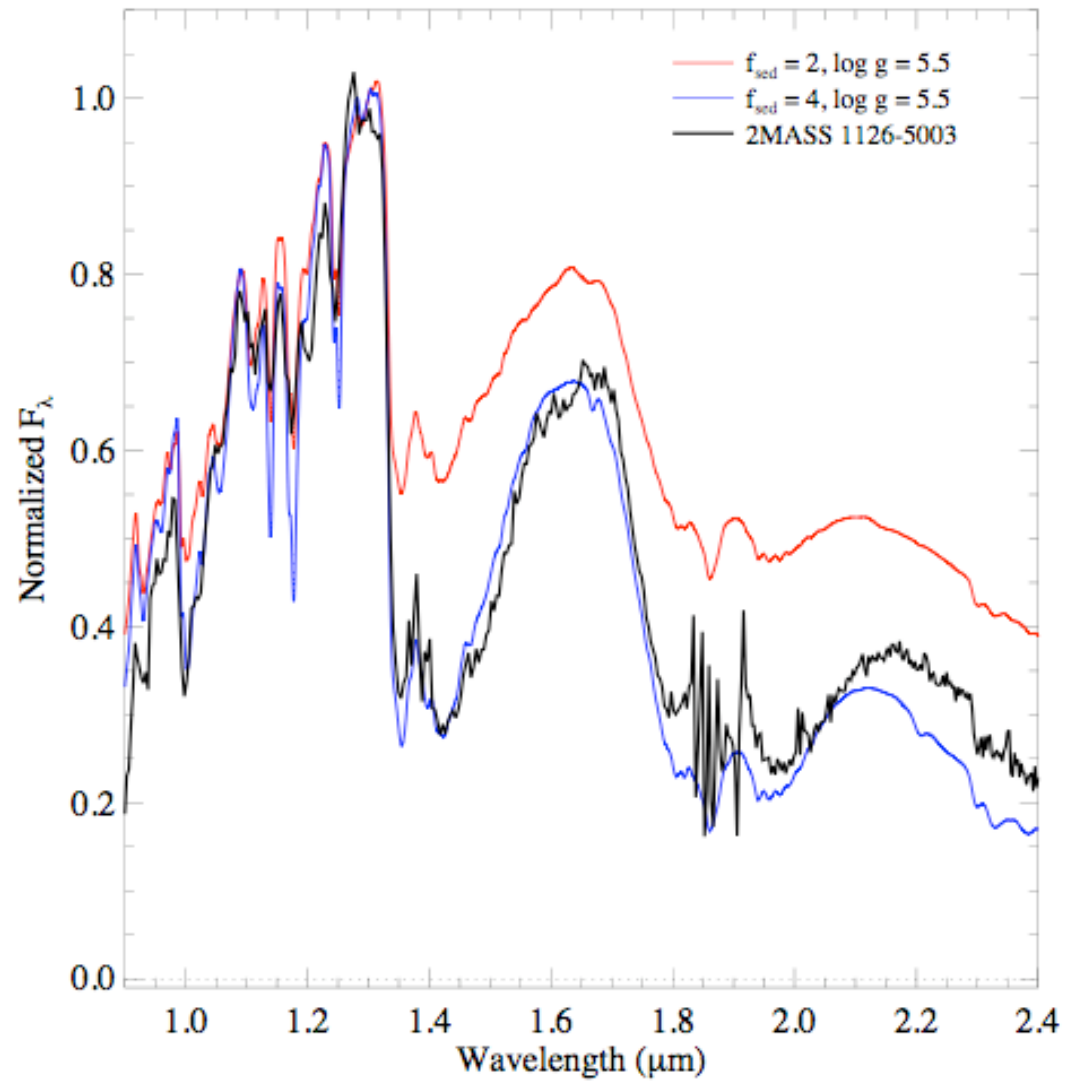
Ackerman & Marley (2001)

Thinner clouds?

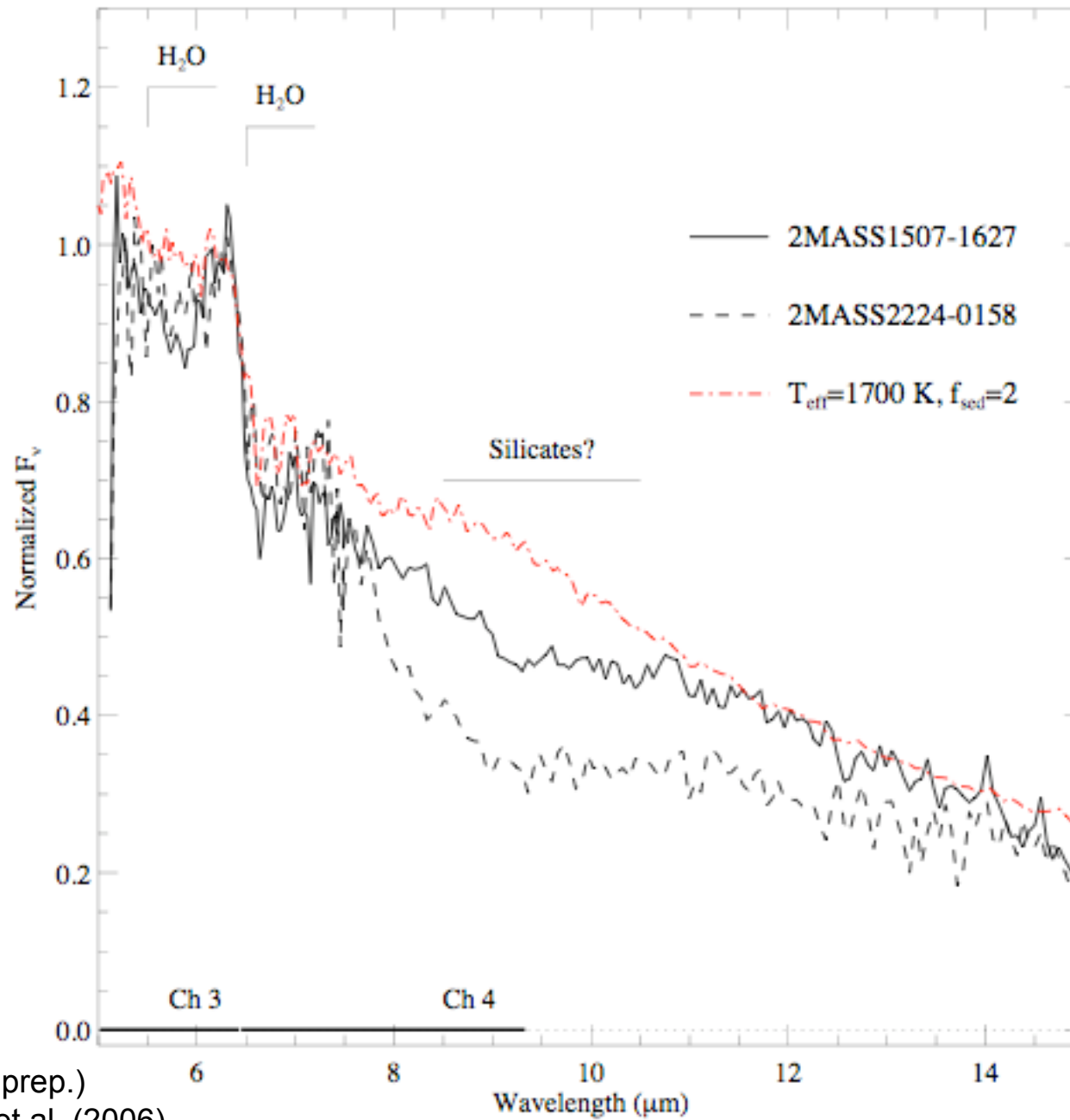


models from D. Saumon & M. Marley

Thinner clouds?



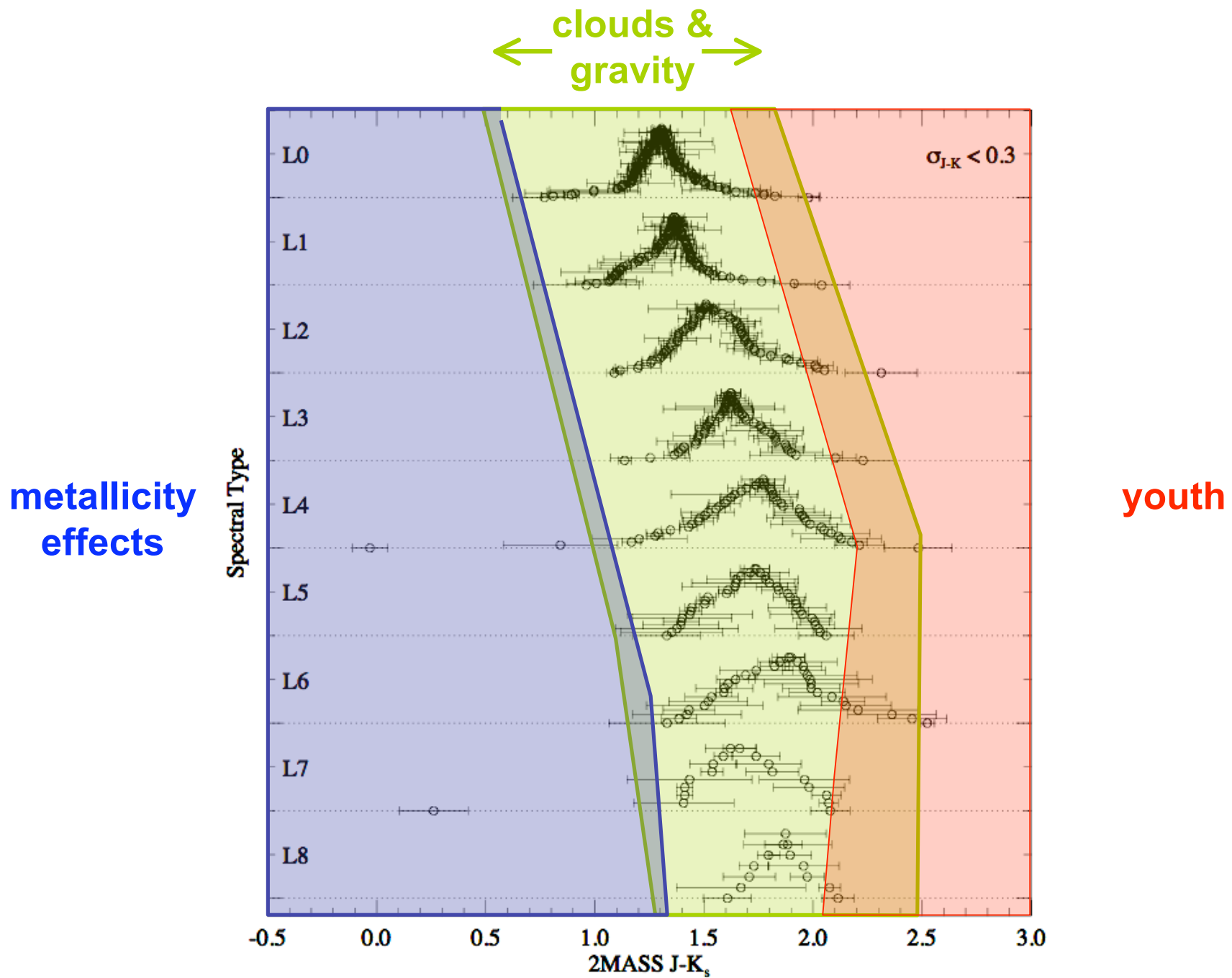
Burgasser et al. (in prep.)
see also Knapp et al. (2004)

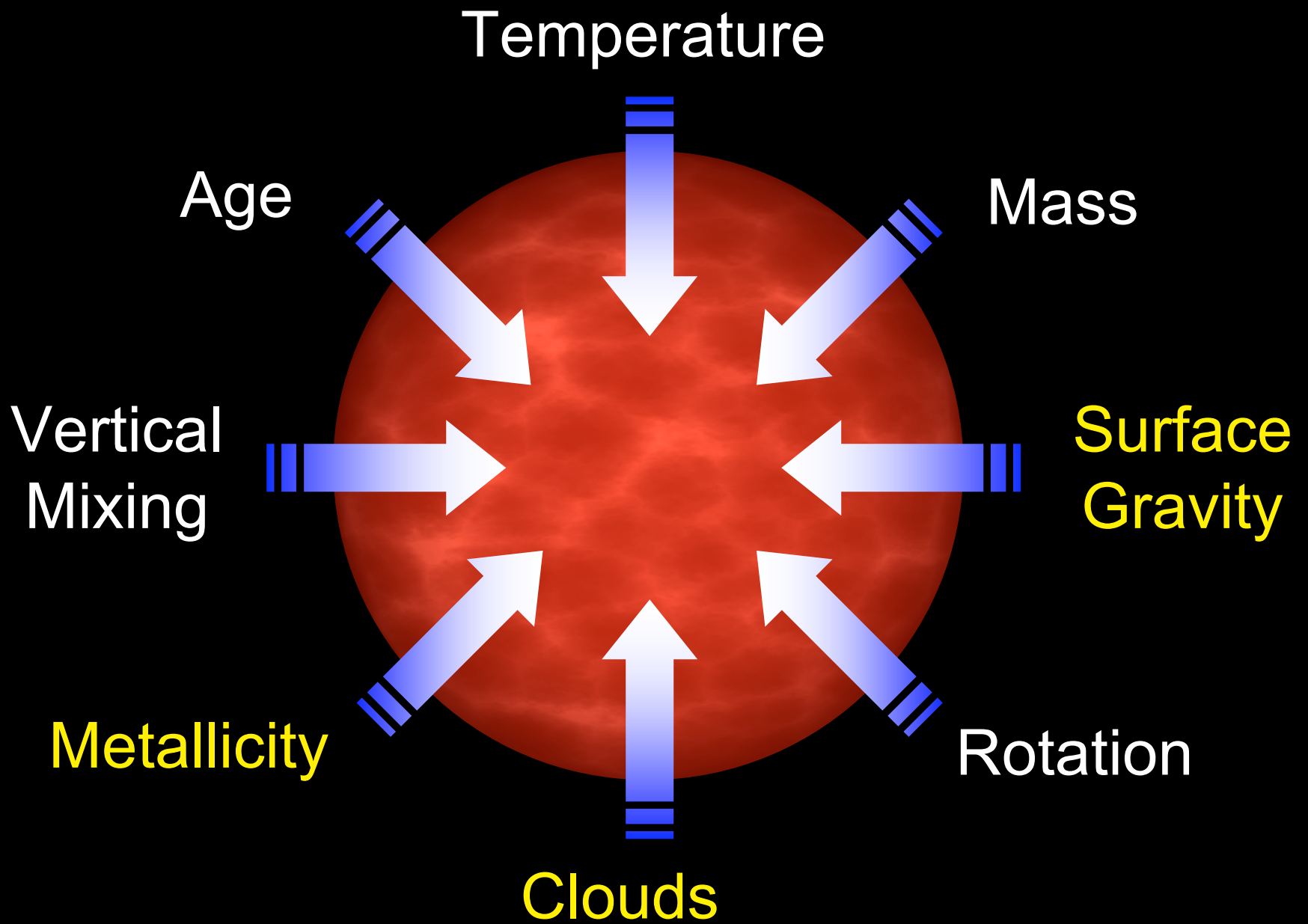


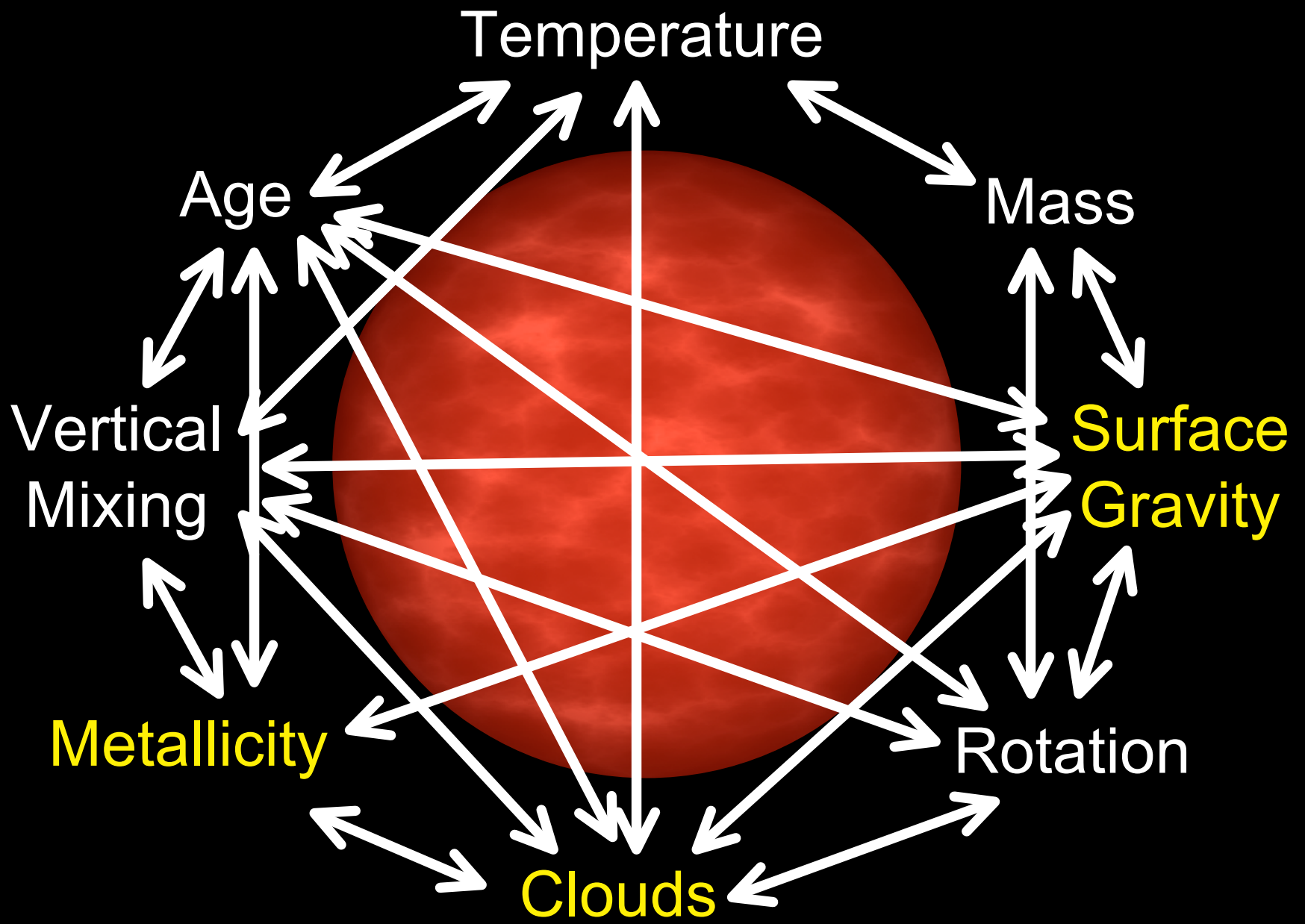
Burgasser et al. (in prep.)
 data from Cushing et al. (2006)

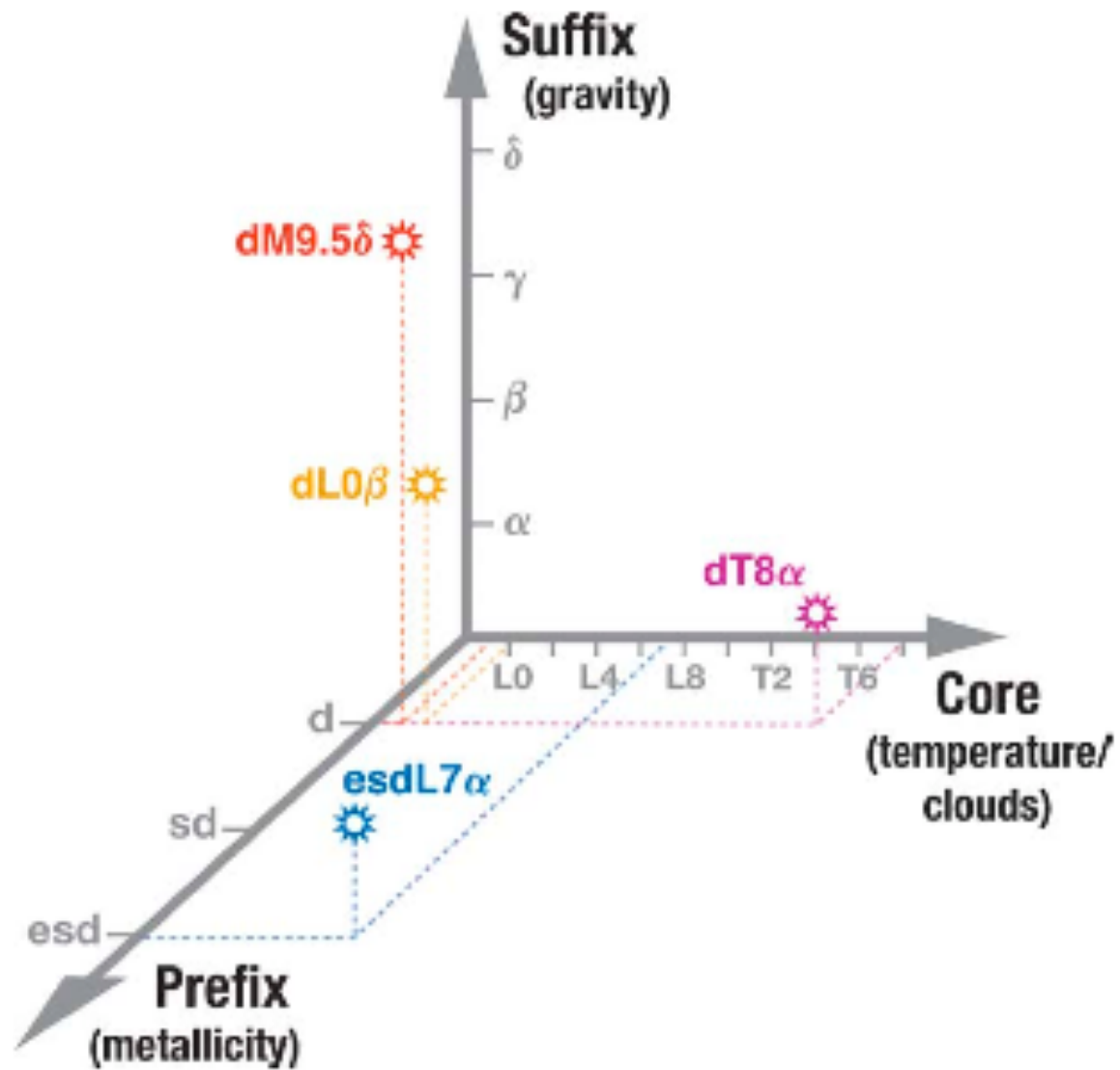


“round ‘em up”

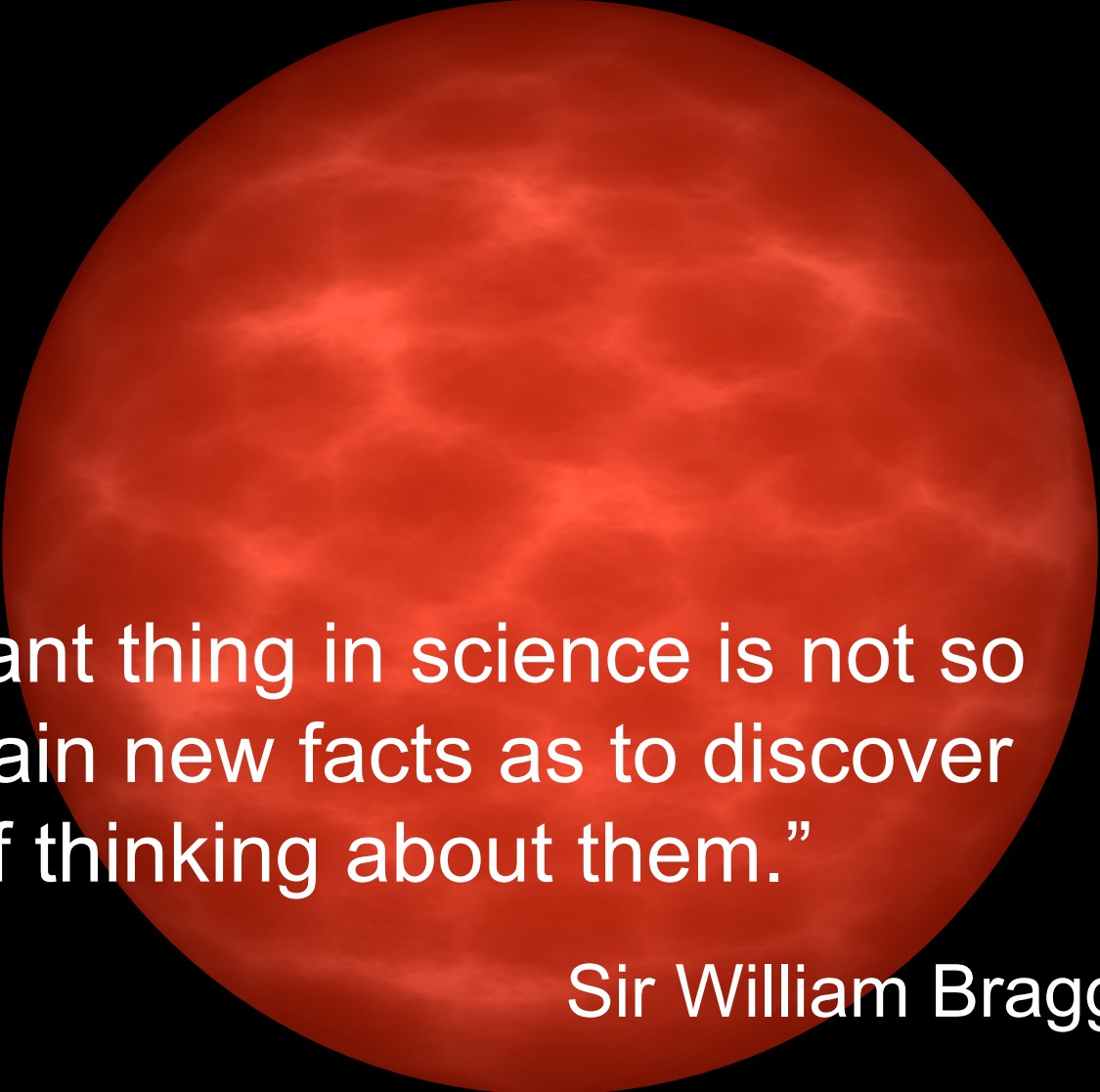








Kirkpatrick (2005)



“The important thing in science is not so much to obtain new facts as to discover new ways of thinking about them.”

Sir William Bragg