MASSACHUSETTS INSTITUTE OF TECHNOLOGY Department of Electrical Engineering and Computer Science

6.014 Electrodynamics

	Issued:	February 6, 2002
Problem Set 1	Due in Recitation:	February 13, 2002

Suggested Reading: Course Notes, Week 1; Text: Sections 1.1-1.4

In general, each homework assignment covers the material presented in the preceding week's lectures and recitations.

Problem 1.1

A certain computer chip uses 4-volt logic and can drive an external 73-ohm load at that voltage through a DC blocking capacitor (the square wave is then ± 2 v). Its clock frequency is 2 GHz, enabling a 1-GHz square wave to be generated and modulated. If that external load is a non-reactive dipole antenna with radiation resistance $R_r = 73$ ohms and peak gain $G_o = 1.5$, then:

- a) What is the maximum RF power (watts) radiated?
- b) How many watts/ m^2 (maximum) are radiated 10 meters from the transmitter? Assume there are no obstacles in the path.
- c) What is the effective area A_e of an isotropic receiving antenna at 1 GHz?
- d) What is the maximum power that can be received at 1 GHz and 10-meter distance by this system? Assume we can ignore the small corrections needed because we are sending a square wave and not a sinusoid.
- e) What is the approximate maximum data rate (bps) that can be communicated to another digital device up to 10 meters away using this RF system?

Problem 1.2

The same 4-volt computer chip used in Problem 1.1 can alternatively drive a 73ohm LED that radiates 0.6-micron wavelength light isotropically with 10-percent efficiency. Assume a semiconductor photodetector having 90-percent quantum efficiency is used to receive the signals and that it is square, 1 cm on a side, and faces the transmitter at a distance of 10 meters.

- a) What is the maximum power radiated by the LED when 4 volts is applied?
- b) How many photons N per second are then radiated by this LED?

- c) What is the maximum number n of photons/sec detected by the detector at a distance of 10 meters?
- d) Approximately how many bits/sec (bps) can this system communicate at 10-meter separation if the rms fluctuations of the background photon flux within the passband of the narrowband filter covering the detector is ~10 photons/bit interval (i.e., if ~100 photons/bit are required)?
- e) Based in part on your answers to Problems 1.1 and 1.2, briefly explain which system probably has the most future for inter-device communications within a single room. Extra discussion is welcome, but not required.

Problem 1.3

Astronauts walking freely about on the moon (distance \cong 380,000 km) would like to have a 10-GHz portable phone to talk to Earth via a large dedicated Earth ground terminal at a nominal 2-kbps data rate. Assume a small transceiver is mounted on the astronaut's helmet and that its transmitting and receiving directionality is isotropic.

- a) As a practical matter, is the uplink (transmissions to the moon) or the downlink likely to pose the greater problem? To answer this question, briefly discuss quantitatively the major system parameters that will likely control system costs, i.e. transmitter power and antenna diameter. Consider the effects of isotropically radiating interfering terrestrial transmitters operating at the same frequency (we want our signal to be >10 times the same-band interference [we want SNR > 10]) and the fact that small sensitive receivers, 10-kW earth-based transmitters, and earth-based antennas with gains <10⁵ are relatively cheap. There are no "perfect answers" here, and common sense should help.
- b) Design an economic down-link (moon-to-earth) system to meet the astronauts' needs (assume the same system will also work for the up-link). For purposes of this problem, it suffices to find the antenna gain for the ground station and the required transmitter power. Note that transmitter powers above 10 kwatt are increasingly expensive (then costs increase ~linearly with power), as are antenna gains above 10^5 .(costs increase ~as G^{1.2}).

Problem 1.4

A certain uniform plane wave is characterized by $\overline{E}(\overline{r},t) = x^{4} \cos(10^{8}t + z) [v/m]$.

- a) What is the frequency f of this wave (Hz)?
- b) What is its direction of propagation?

- c) What is its wavelength $\lambda(m)$?
- d) If the medium in which it is propagating has permeability $\mu = \mu_0$, what then is its permittivity ϵ ?
- e) What is the magnetic field $\overline{H}(\overline{r}, t)$ for this wave?

Problem 1.5

- a) Let $\overline{A} = 2\hat{x} \cdot \hat{y} + \hat{z}$ and $\overline{B} = \hat{x} + 2\hat{y} + \hat{z}$. Find (i) $\overline{A} + \overline{B}$, (ii) $\overline{A} \bullet \overline{B}$, (iii) $\overline{A} \times \overline{B}$, (iv) magnitude of \overline{A} , and (v) the angle between \overline{A} and \overline{B} .
- b) Let $\underline{\overline{A}} = \hat{x} + (1-j)\hat{y} + j\hat{z}$, what is $\overline{a}(t)$ corresponding to this phasor?
- c) If $\overline{b}(t) = 2\hat{x}\cos(\omega t \pi/4) + \hat{y}\sin(\omega t + \pi/4)$, what is the corresponding phasor $\overline{\underline{B}}$?

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Problem Set 1 Solutions

Available in Tutorials: February 14, 2002

Problem 1.1

a) Power =
$$v^2/R = 2^2/73 = 55$$
 mwatts

b)
$$P_r = G_t P_t / 4\pi r^2 = 1.5 \times 0.055 / 4\pi 10^2 = 6.57 \times 10^{-5} \text{ w/m}^2.$$

c) $A_e = G\lambda^2/4\pi = 1 \times (c/f)^2/4\pi = (3 \times 10^8/10^9)^2/4\pi = 0.0072 \text{ m}^2$ (if square, 8.8 cm on a side). Note that the gain over isotropic of an isotropic antenna is unity.

d)
$$P_{rec} = P_r A_e = 6.57 \times 10^{-5} \times 0.0072 = 4.73 \times 10^{-7} \text{ watts}$$

e) $R(bps) = P_{rec}/E_b \cong 4.73 \times 10^{-7}/4 \times 10^{-20} = 11.8$ Terabits/sec.] In practice, the limiting data rate might be a few orders of magnitude less because of radio frequency interference. Even so, the bandwidth required for the resulting data rate would still exceed the RF transmission frequency employed (1 GHz). Usually communications signals occupy bandwidths and communicate data rates that are less than 20 percent of the central frequency ($\Rightarrow \sim 200$ Mbps here). In practice we could reduce the transmitter power or the receiver sensitivity, and any remaining excess power would simply reduce further our low probability of error in the communicated bit stream.

Problem 1.2

- a) $P_{tmax} = v^2/R_r = 4^2/73 = 0.219$ watts. If we send half "one's" at this power level and half "zero's" at no power, then the average transmitted power is 0.109 watts.
- b) hf = photon energy, where f = c/ λ . P_{tmax} = 0.1×0.219 watts. Maximum photon flux N = P_{tmax}/hf = P_{tmax} λ /hc = 0.0219×0.6×10⁻⁶/(6.625×10⁻³⁴×3×10⁸) = 6.62×10¹⁶ photons/sec = N.
- c) If these are radiated isotropically, then we have $m = N/4\pi r^2$ photons sec⁻¹ m⁻² at radius r, of which ηA_{det} are detected, where $\eta = 0.9$ and $A_{det} = 10^{-4}$ here. Therefore $n = 0.9 \times 6.62 \times 10^{16} \times 10^{-4}/4\pi 10^2 = 4.74 \times 10^9$ detected photons/sec.
- d) If we assume 100 photons/detectable bit, then the maximum data rate is $\sim 5 \times 10^7$ bps.

Note that direct sunlight (~1 kw/m²) on the detector would add $P_{sun} = 0.1$ watt. The number of detected solar photons would be $0.9\times0.1/hf = 0.09 \ \lambda/hc = 0.09\times6\times10^{-7}/6.625\times10^{-34}\times3\times10^8 \cong 2.7\times10^{17}$. This number of solar photons could be reduced by a factor of 1000 by a narrowband optical filter at the receiver, would correspond to 3×10^{14} detected photons per second, all random. If we slice each second into 5×10^7 pieces, one piece per bit, then ~100 background photons would be detected in each such brief interval, and the rms variation in photon count per "zero" interval would be $(6\times10^6)^{0.5} = 2450$ photons/sec, large compared to the detection threshold of 100 photons per bit that we assumed. By setting the detection threshold for "one" near 10,000 photons, excursions of ~four standard deviations would be required to produce a bit error. Conservative design would therefore increase the detected photons required to represent a "one" to perhaps 20,000 or more, reducing the available data rate to ~250 kbps.

e) Since so much depends on manufacturing costs and unknown interference levels, it is difficult to be certain which mode, RF or optical, would be best for this application. However, at these data rates both RF and optical systems can be integrated on single chips, so manufacturing cost is less likely to control than is interference, and path blockage. In general, RF is more likely to fill a room through scattered papers, books, etc. than are optical signals, which are moreover vulnerable to direct sunlight if they operate in the visible. If we are guided by our answers to Problems 1.1 and 1.2, the preferred choice is simple—RF.

Problem 1.3

 $P_{rec} = P_t G_t G_r \lambda^2 / (4\pi r)^2$ Since we are using the same antennas in both directions, a) and λ and r are also the same in both directions, P_{rec} is simply a constant times P_t , where a one-watt transmitter (like a cell phone) is reasonable for the astronaut and 10 kW is reasonable for the ground transmitter. Thus this argument suggests that costs will be set by the needs of the downlink, which is disadvantaged by a factor of $\sim 10^4$. However the astronaut's isotropic antenna will pick up all N interfering transmitters on Earth; N could be millions if the same frequency band is used by individuals. For example, one billion people in view of the moon might produce 10 million simultaneous calls, of which one percent might be in our band (then N $\approx 10^5$ calls). If the interferers also radiate ~1 watt isotropically, then we need to have a ground station antenna that has gain in the direction of the astronaut that is $10(SNR) \times 10^5$ greater, or 10^6 . Since 10-kW transmitters and gains of 10^5 are cheap, we want to use such systems. They should suffice, however, because we have a 10^4 power advantage over a 1-watt interfering cell phone, plus a 10^5 antenna gain advantage, or 10^9 overall advantage, which is large compared to our assumed set of 10^6 1-watt isotropic interferers. Were this not the case, the uplink antenna gain might have to be increased substantially, making it the dominant cost element.

b) $P_{rec} = P_t G_t G_r \lambda^2 / (4\pi r)^2$. We assume no two astronauts will talk to earth simultaneously in the same narrow band. If we use a 1-watt transmitter for the astronaut, r = 380,000 km, $G_t = 1$, $G_r = 10^5$, and $\lambda = 0.03$ m, then $P_{rec} = 1 \times 1 \times 10^5 (0.03)^2 / (4\pi 3.8 \times 10^8)^2 = 3.95 \times 10^{-18}$ watts, which should be compared to the received power requirement $RE_b = 2000 \times 4 \times 10^{-20} = 8 \times 10^{-17}$ watts. Since P_{rec} is ~20 times less than the required RE_b , either the ground station antenna gain must be increased by a factor of ~20; or the astronaut's transmitter must radiate >~20 watts, or some combined solution must be adopted. If we assume the antenna costs grow as $G^{1.2}$ and power costs increase linearly with power P_t , then the factor of 20 might be applied entirely to the astronauts' transmitter power (making it 20) watts), unless the baseline cost of that transmitter were larger than that of the antenna.

In practice the final optimized design might divide the "cost pain" among several elements. These element costs typically increase faster than linear with the critical parameters. For example, to optimize this system the ground station antenna gain might be increased a factor of 1.5, the astronaut's helmet antenna gain might be increased to 3.33 (requiring the astronaut to generally face the earth when phoning, the helmet power might be increased to 3 watts, the voice data rate (and quality) might be reduced (say to 1.5 kbps), the astronauts' battery weight might be increased a factor of 1.5, and the phone speaking time per charge might be reduced a factor of two.

Most engineering design proceeds approximately in this fashion. A nominal point design is analyzed, and then all system parameters, including performance, are incremented to produce a superior design using quantitative metrics that combine cost, performance, and risk.

Problem 1.4

a) Since
$$\omega = 10^9$$
, $f = \omega/2\pi \approx 15.9$ MHz.

- b) The argument $(10^8 t + z)$ remains constant as time increases if z goes negative. The wave is therefore propagating in the negative z direction.
- c) Setting the first time derivative of the argument $(10^9t + z)$ to zero (the argument doesn't change for a fixed point on the moving waveform), we find $dz/dt = 10^8$, less than the speed of light, so $\lambda = v/f = 10^8/1.59 \times 10^7 \cong 6.3$ m.

d)
$$v = (\epsilon \mu_0)^{-0.5} = 10^8 [\text{m/s}]$$
, whereas $(\epsilon_0 \mu_0)^{-0.5} \cong 3 \times 10^8$. Therefore $(\epsilon/\epsilon_0)^{0.5} = 3$, or $\epsilon = 9\epsilon_0 = 7.97 \times 10^{-11} [\text{F/m}]$

e) We can use Faraday's law, $\partial \overline{H}/\partial t = -(\nabla \times \overline{E})/\mu_o$, and then integrate, or we can take advantage of the fact that we have only a single uniform plane wave, which implies

 \overline{H} is in phase with \overline{E} , at right angles, and $1/\eta_o$ times greater, where $\eta_o = (\mu_o/\epsilon)^{0.5} = (1.257 \times 10^{-6}/7.97 \times 10^{-11})^{0.5} = 126$ ohms. Here, for fun, let's use Faraday's law.

$$\nabla \times \overline{\mathbf{E}} = \det \begin{vmatrix} \partial/\partial x & \partial/\partial y & \partial/\partial x \\ E_x & E_y & E_z \\ \hat{x} & \hat{y} & \hat{z} \end{vmatrix} = \hat{y} \partial \mathbf{E}_x / \partial z = -\hat{y} \sin(10^8 t + z), \text{ so}$$
$$\overline{\mathbf{H}} = \int \hat{y} \left[\sin(10^8 t + z) / \mu_0 \right] dt = -\hat{y} \cos(10^8 t + z) / 10^8 \mu_0 = \boxed{-\hat{y} \cos(10^8 t + z) / 126 \left[\text{A/m} \right]}$$

Problem 1.5

a) (i)
$$3\hat{x} + \hat{y} + 2\hat{z}$$
, (ii) $2 - 2 + 2 = 1$, (iii) $\overline{A} \times \overline{B} = \det \begin{vmatrix} A_x & A_y & A_z \\ B_x & B_y & B_z \\ \hat{x} & \hat{y} & \hat{z} \end{vmatrix}$
= $(-1 - 2)\hat{x} + (2 - 2)\hat{y} + (4 + 1)\hat{z} = -3\hat{x} + 5\hat{z}$

b)
$$\overline{\dot{a}(t) = (\hat{x} + \hat{y})\cos \omega t + (\hat{y} - \hat{z})\sin \omega t}$$

The real parts $\Rightarrow \cos \omega t$, imaginaries $\Rightarrow -\sin \omega t$

c)
$$\overline{\underline{B}} = -2j\hat{x} + \hat{y} = 2.24e^{j2.68}$$