

# Statistical Properties of Exoplanets

## WHY?

- Compare to predicted outcomes of planet formation theories
- Assess prospects for future discoveries
- Place the Solar system in a larger context

# Statistical Properties of Exoplanets

**BEWARE!**

- Small number statistics
- Selection effects
- Trials factor (look at enough variables, and you *will* find correlations)

# The overall frequency of planets

From radial velocity surveys  
(nearby, mature FGK dwarfs)

- 1% have planets with  $M > 0.5 M_{\text{jup}}$ ,  
and  $P < 10$  days
- 6% have planets with  $M > 0.5 M_{\text{jup}}$ ,  
and  $P < 4$  years

# The overall frequency of planets

From microlensing surveys  
(distant M stars)

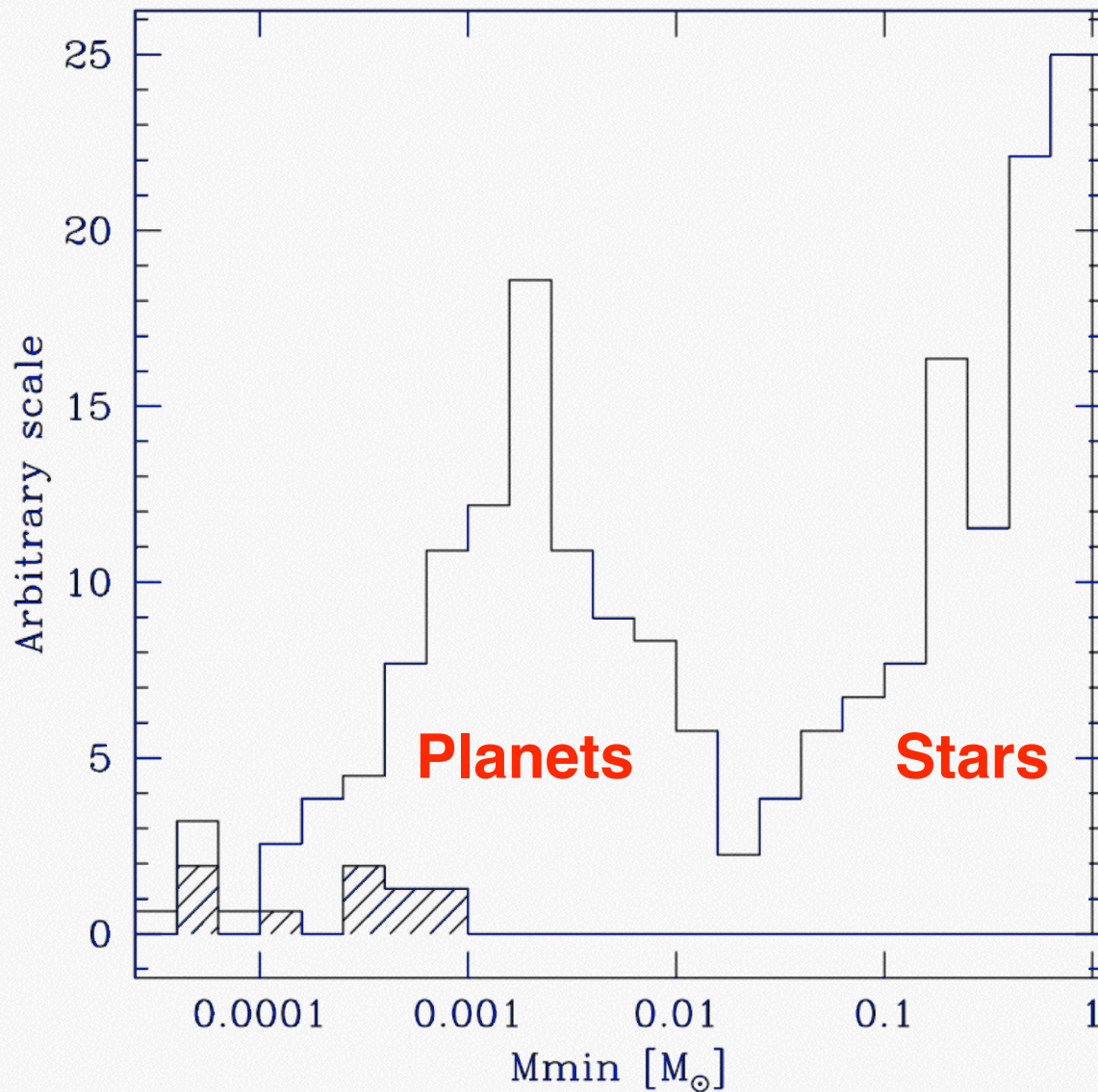
- $< 33\%$  have planets with  $M \sim M_{\text{jup}}$ ,  
and  $1.5 < a < 4 \text{ AU}$  Gaudi et al. (2002)
- $> 18\%$  have planets with  $M \sim M_{\text{nep}}$ ,  
and  $1.5 < a < 4 \text{ AU}$  ( $N = 2$ ) Gould et al. (2006)

# The overall frequency of planets

From transit surveys  
(distant GK dwarfs)

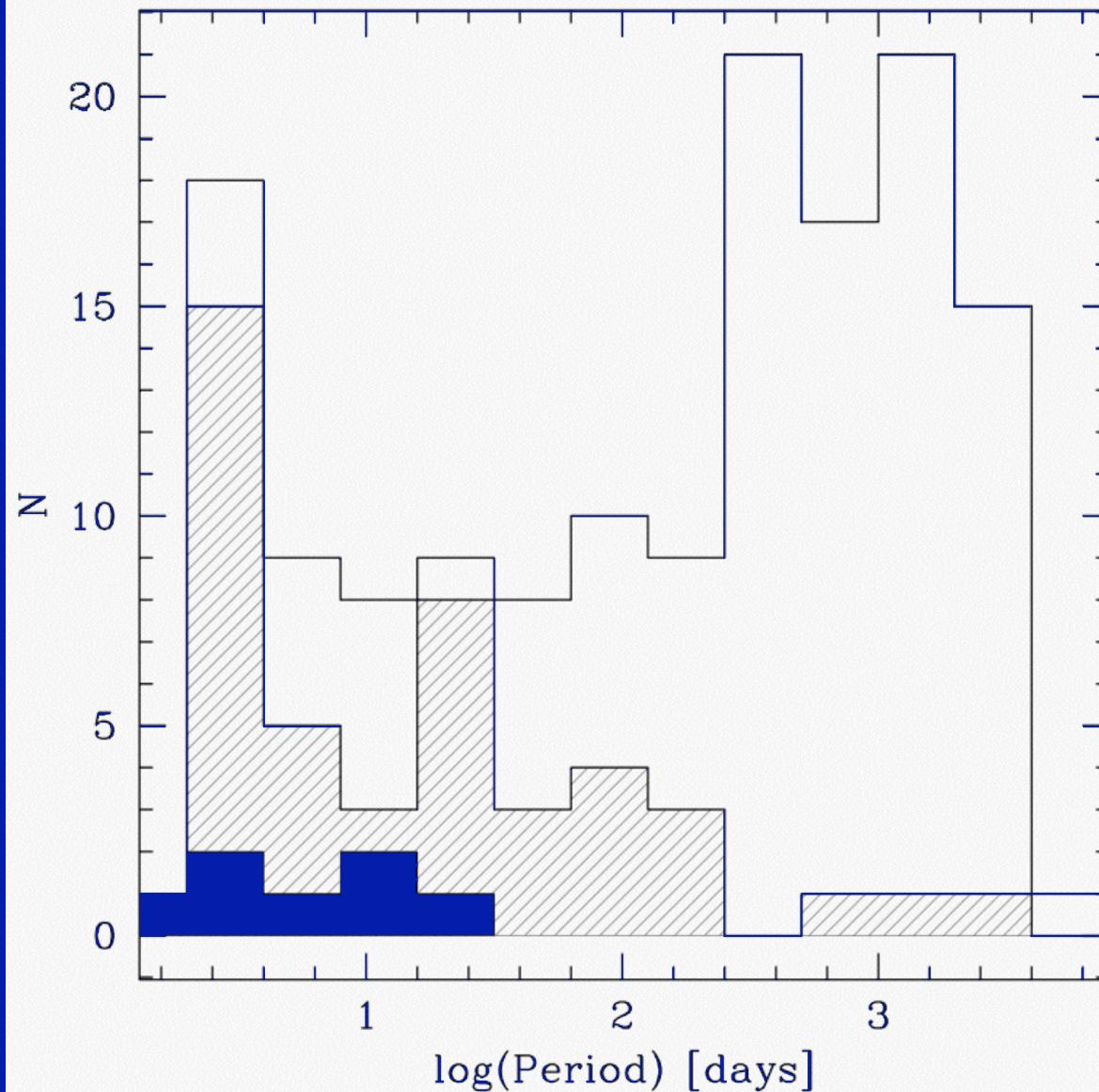
- ~ 0.5% have planets with  $R \sim R_{\text{jup}}$ ,  
and  $2.5 < P < 6$  days Gould et al. (2006)
- ~ 0.1% have planets with  $R \sim R_{\text{jup}}$ ,  
and  $P < 2.5$  days Gaudi et al. (2005)

# Planet masses



- Brown dwarf desert (20-60 M<sub>jup</sub>)
- $dN/dM \sim M^{-1}$
- Plentiful small planets?

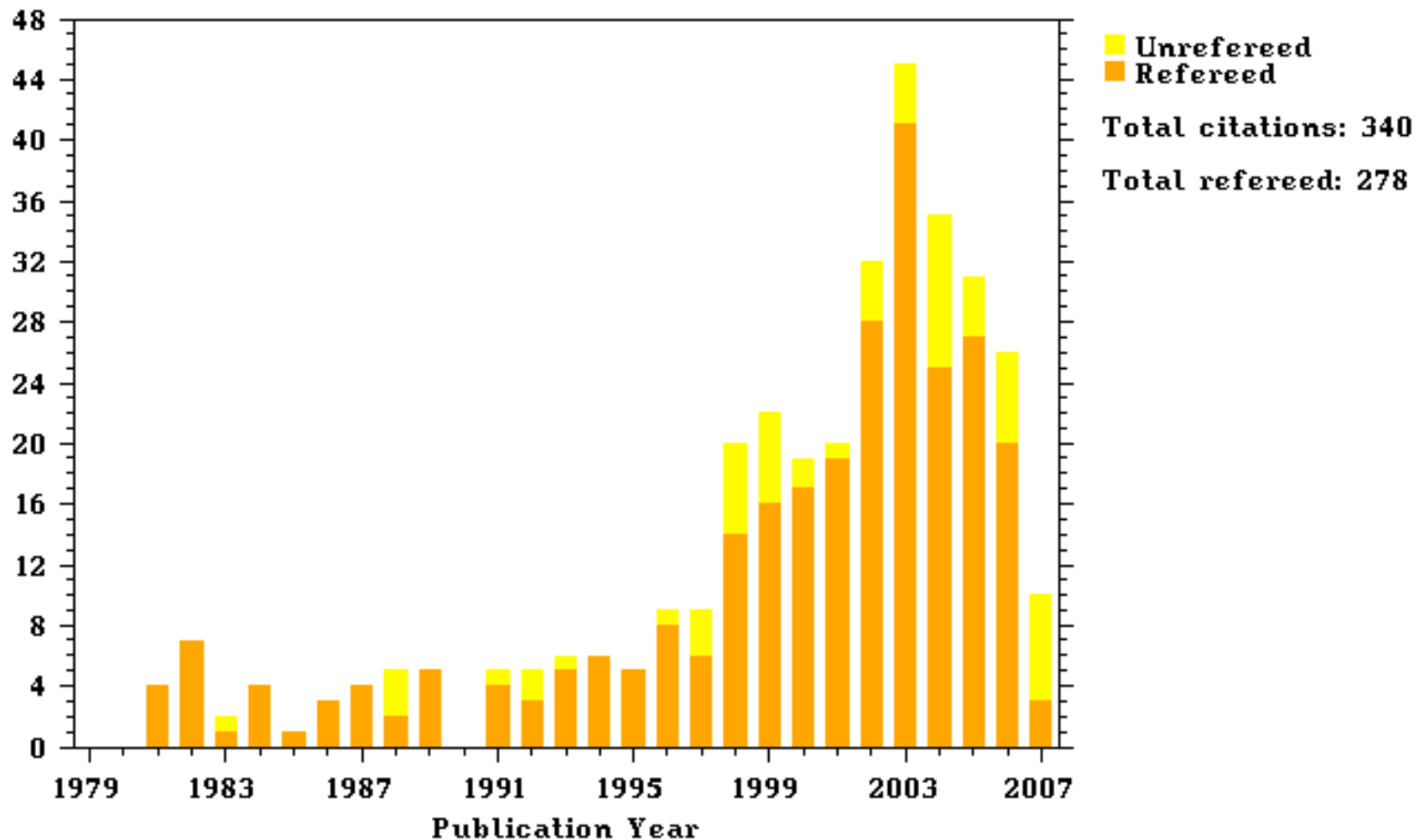
# Orbital periods



- “Pile-up” at  $P \sim 3$  days
- Very few with  $P < 3$  days
- Rising frequency with increasing  $P$  ?

# Goldreich & Tremaine (1980)

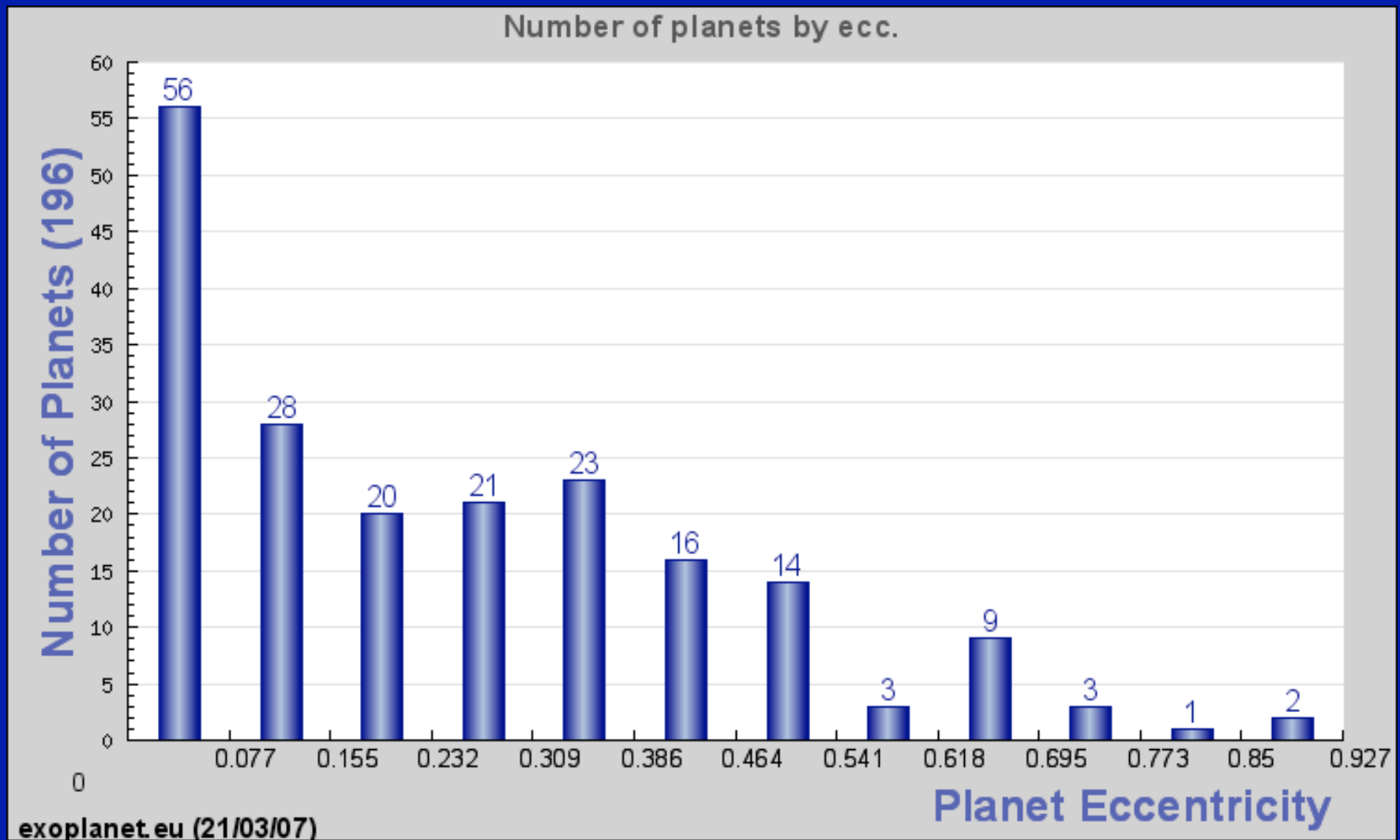
Citations/Publication Year for 1980ApJ...241..425G

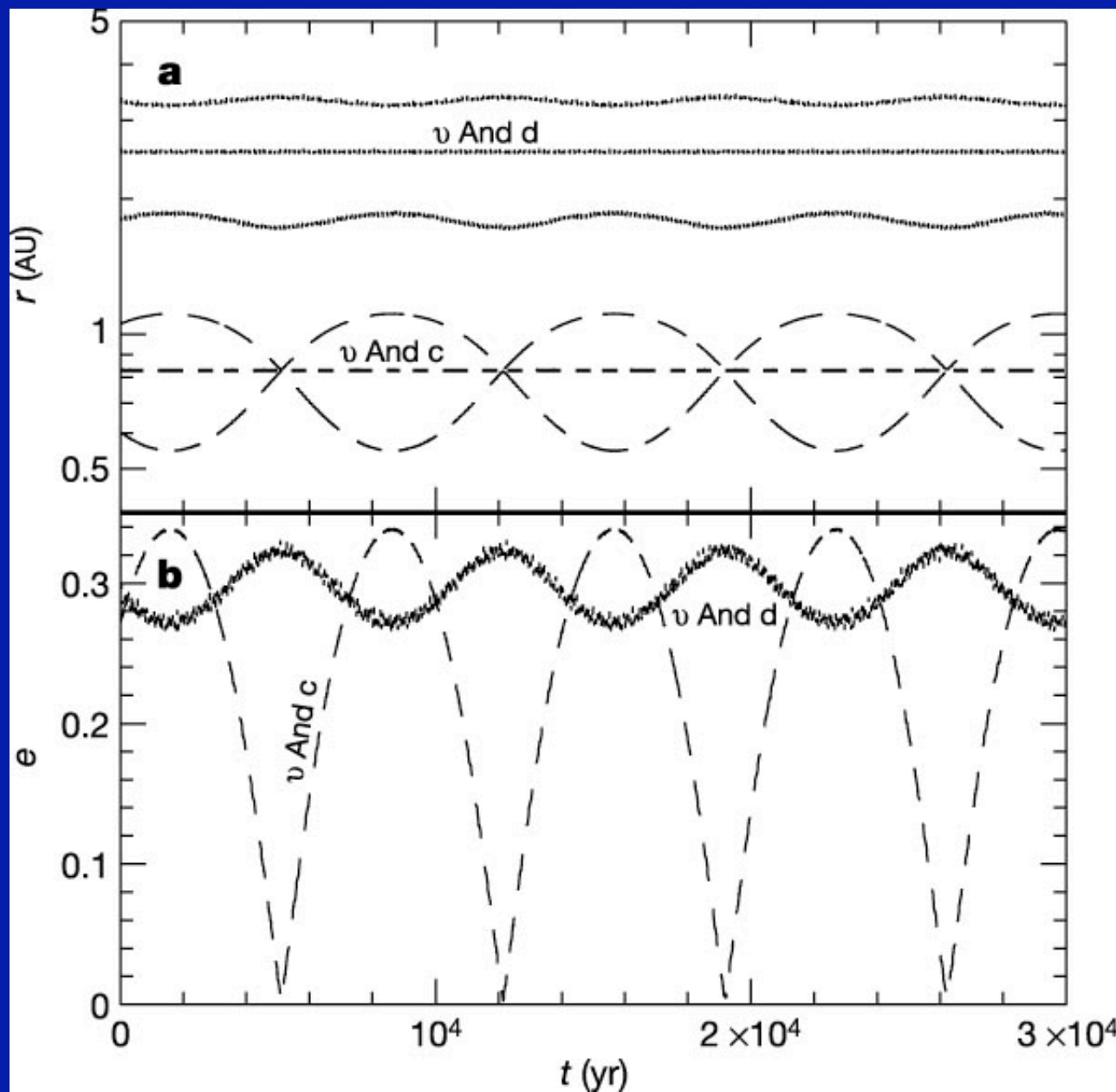


# Planetary Migration

- Tidal interaction with protoplanetary disk
  - Torques from spiral-density waves (Type I)
  - Clear gap; viscous evolution of disk (Type II)
  - Timescale problem
  - Halting problem
- Tidal interaction with planetesimal disk
- Planet-planet scattering

# Orbital eccentricities





## *v* And

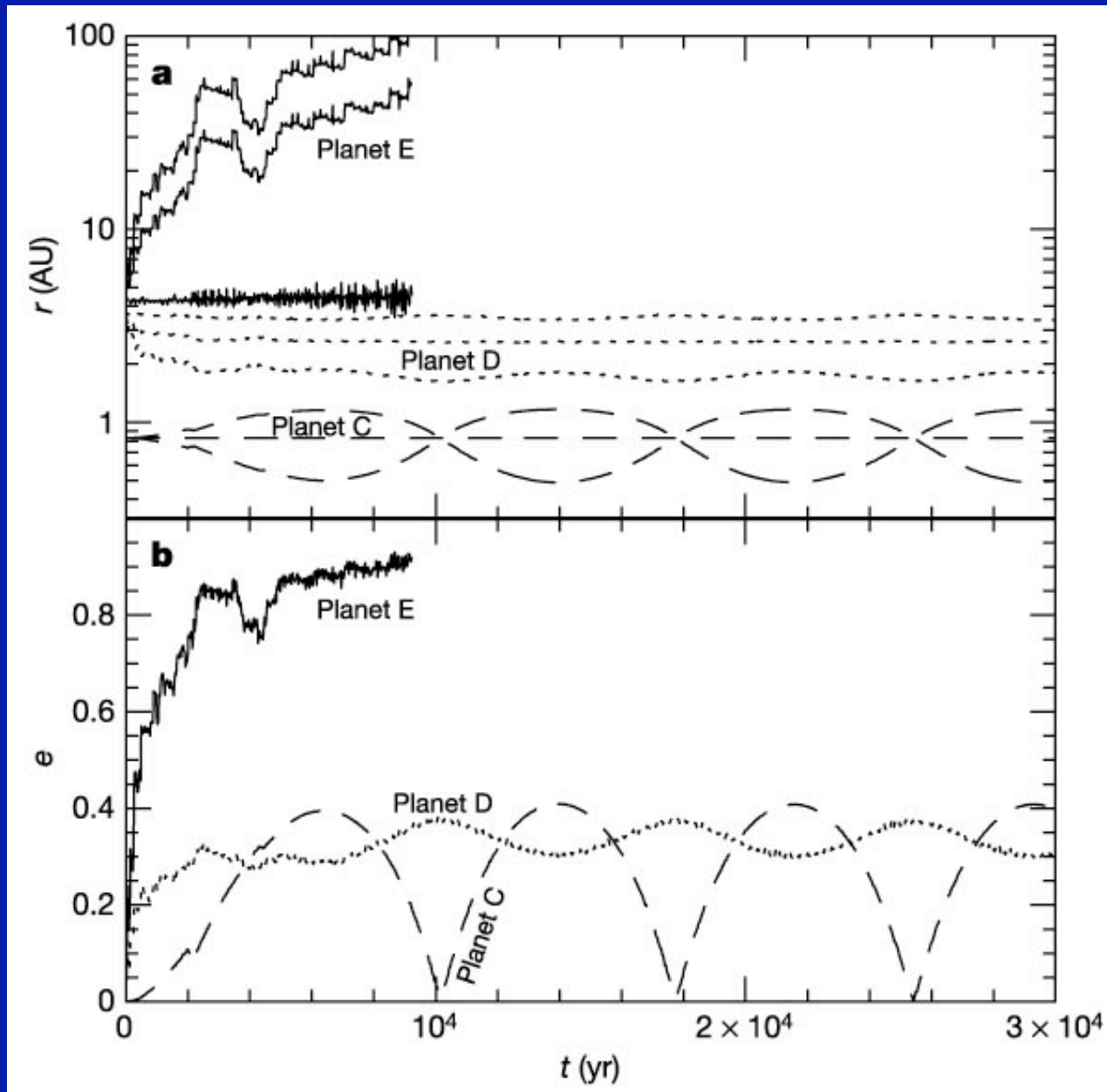
Evidence for  
planet-planet  
scattering

Ford, Lystad, &  
Rasio (2005)

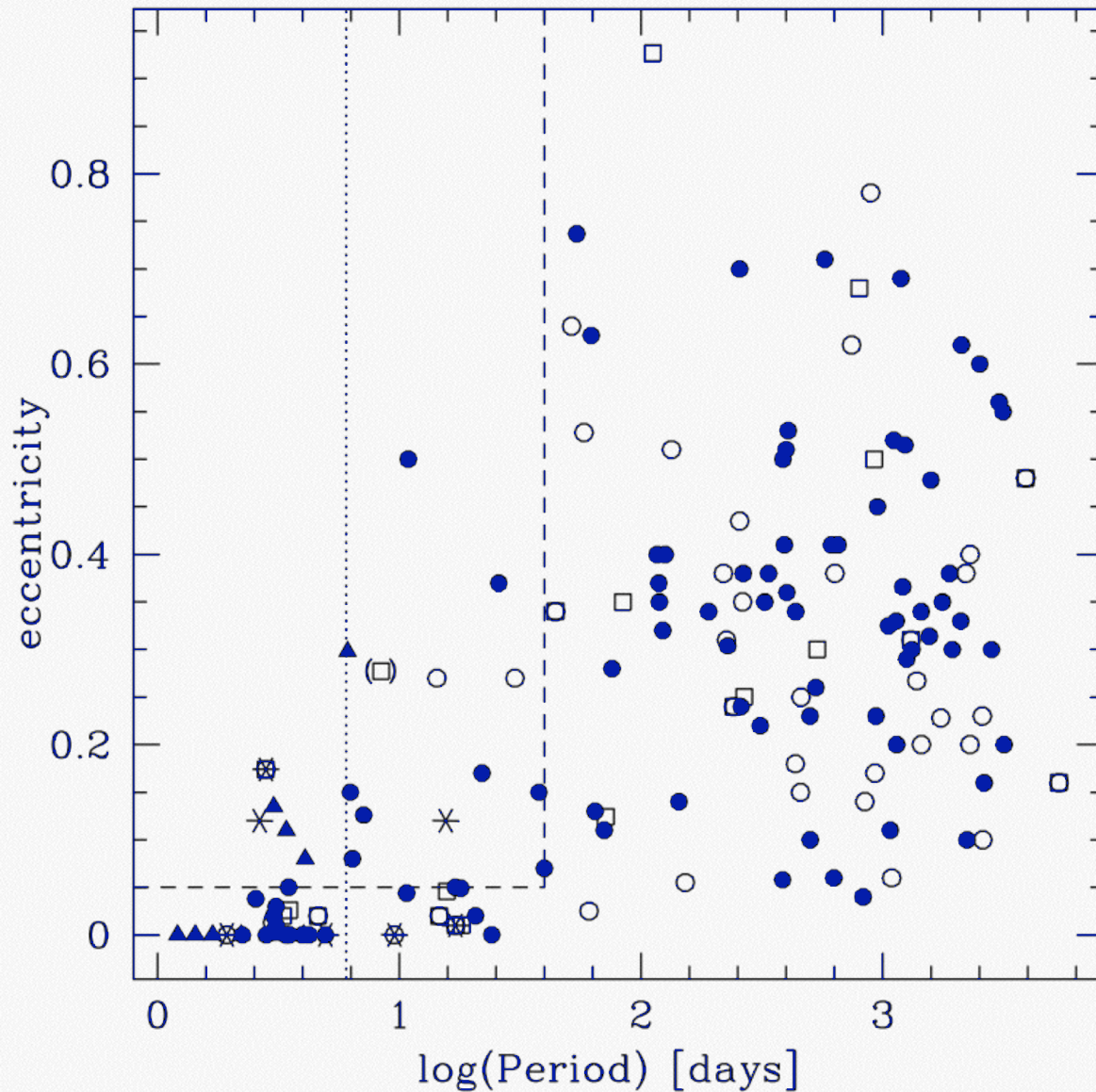
*v* And

Evidence for  
planet-planet  
scattering

Ford, Lystad, &  
Rasio (2005)

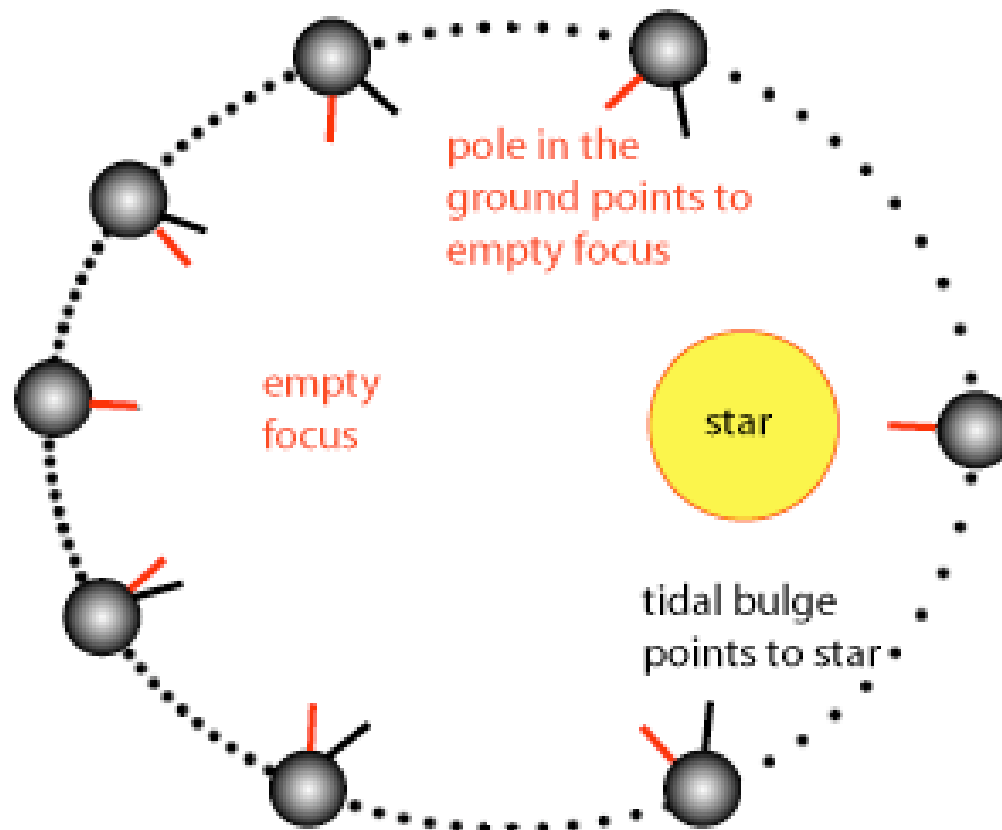


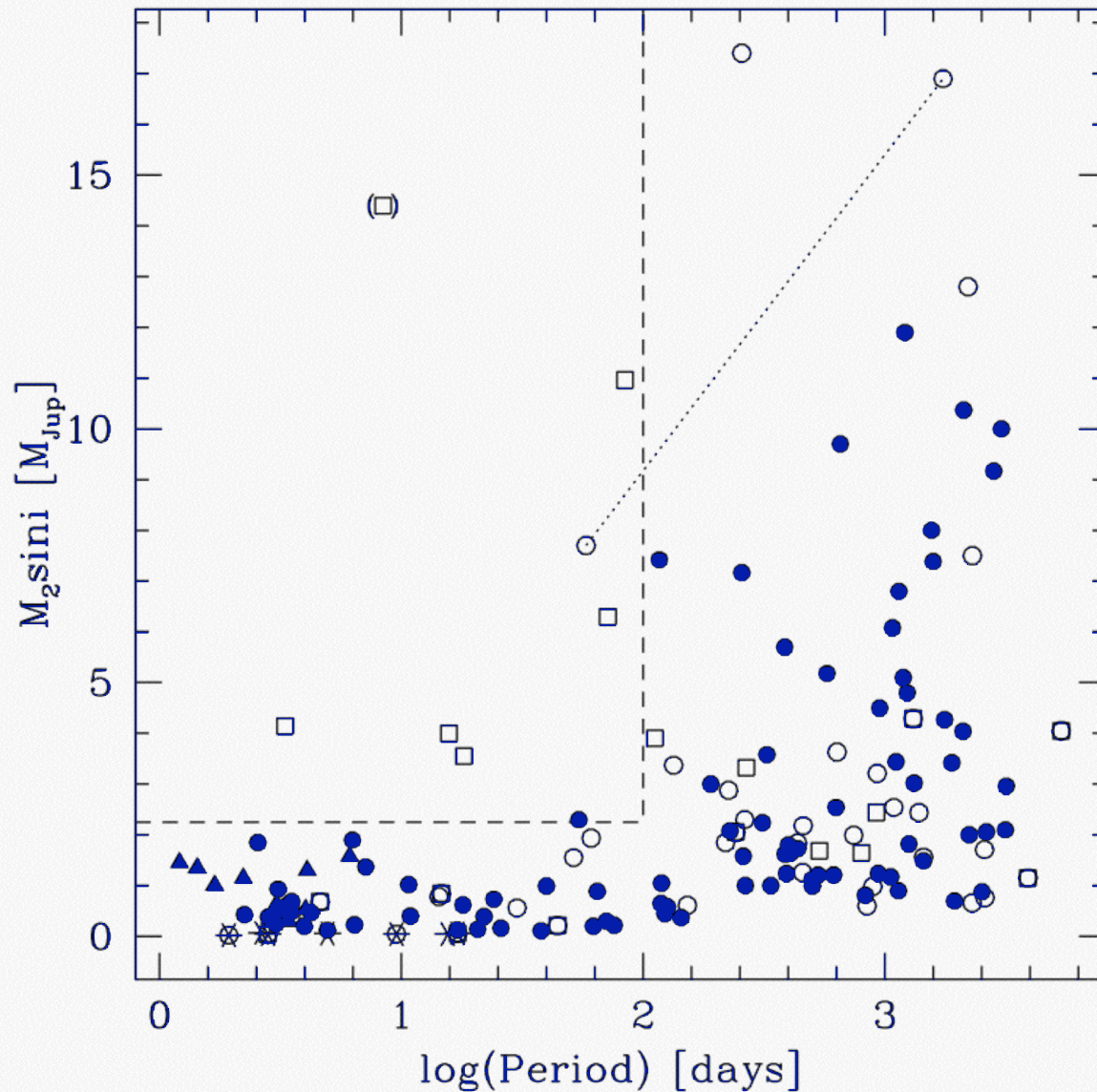
# Period / eccentricity



Circular orbits for  
 $P < 5$  days

# Tidal circularization

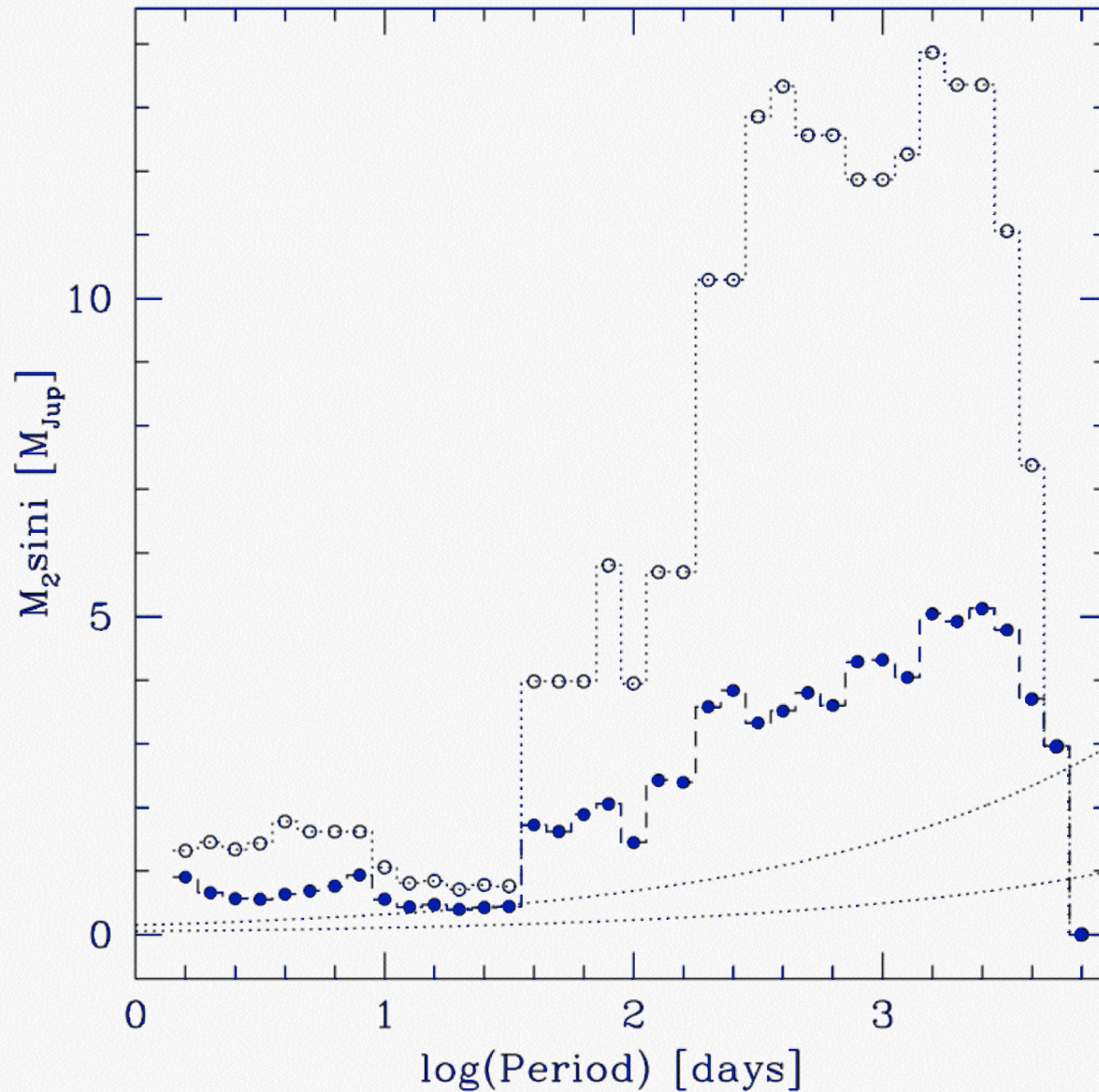




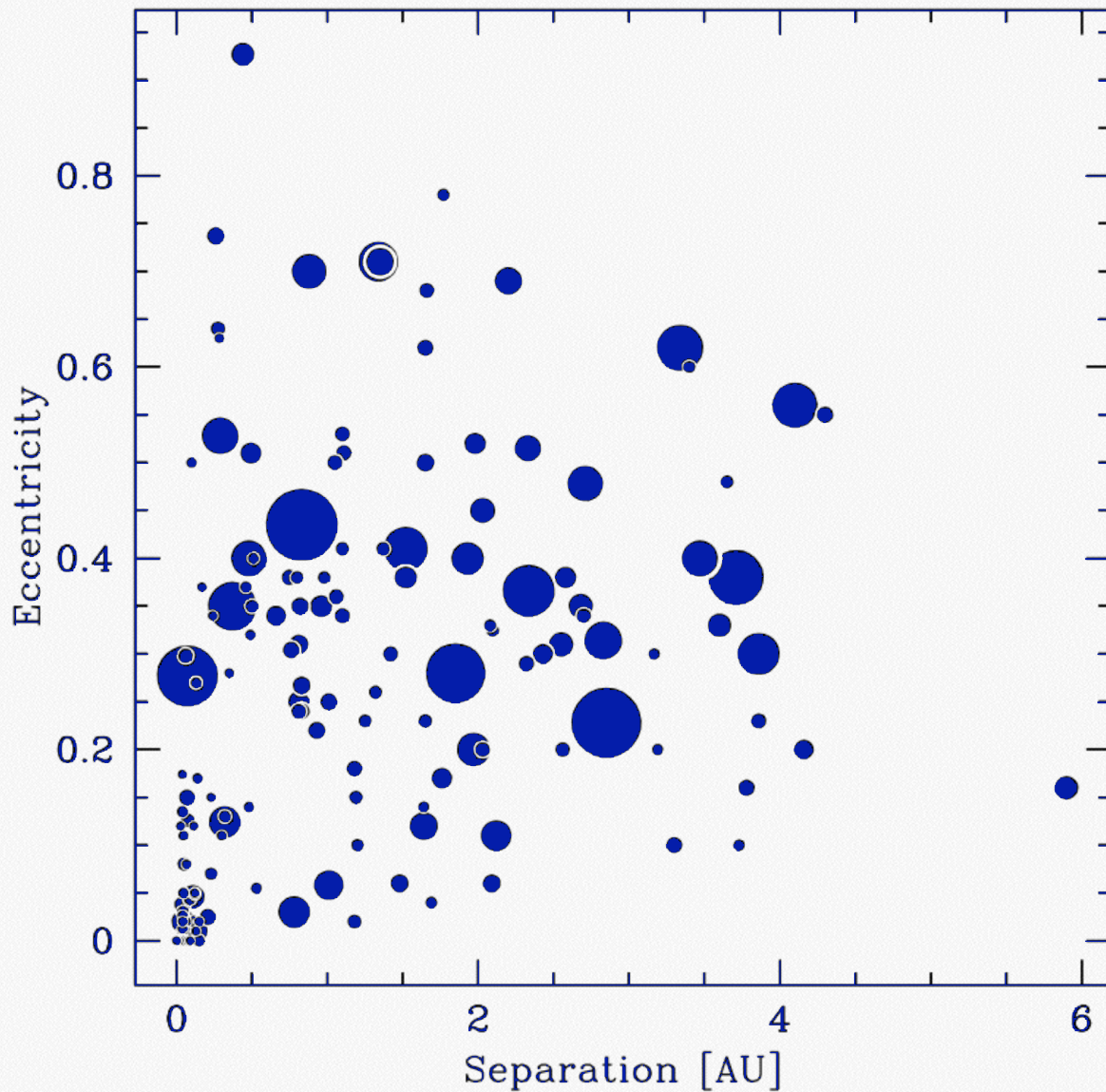
## Period / mass

- Shorter-period planets have smaller masses
- Exceptions are all in multiple star systems

# Period / mass

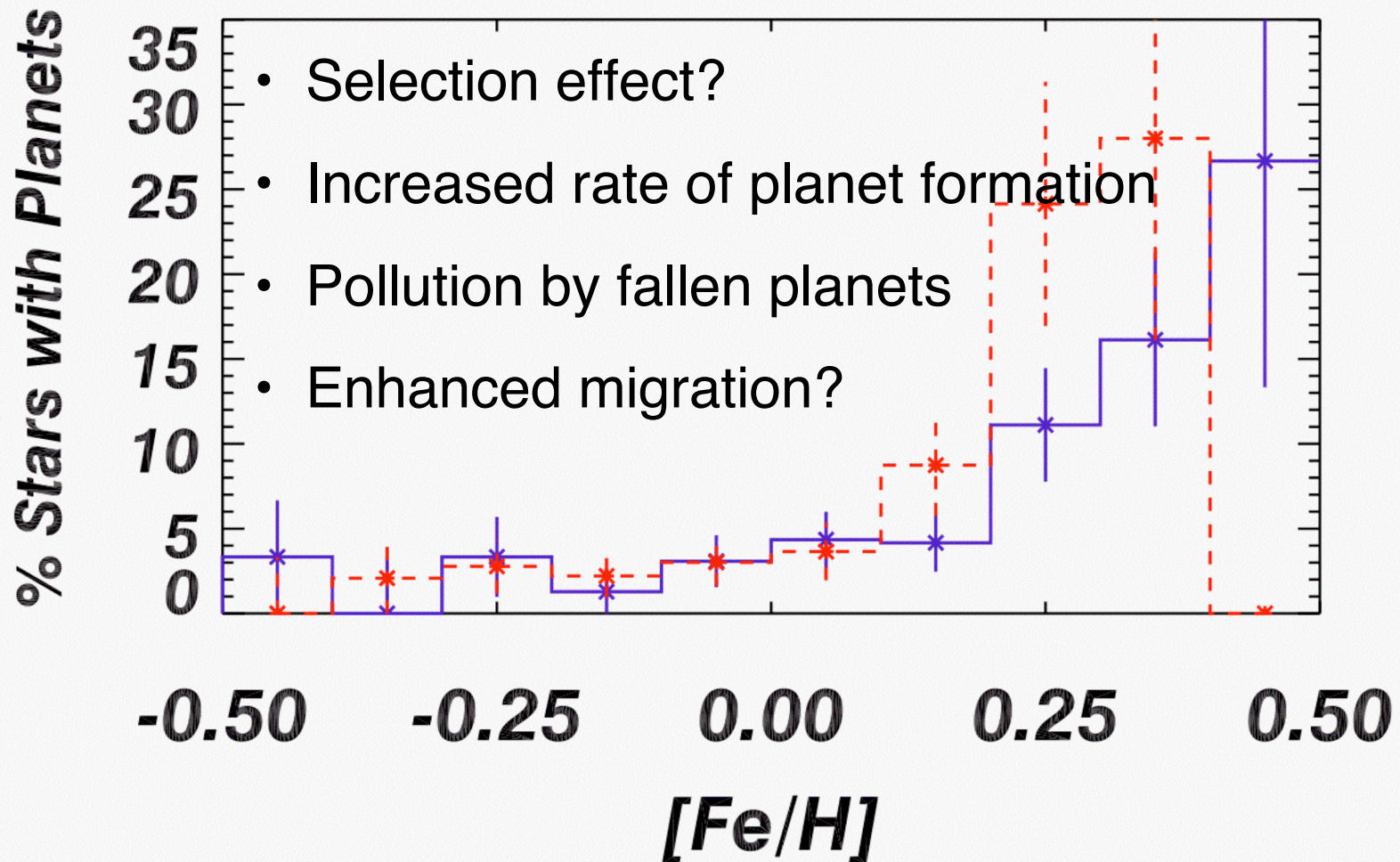


- Shorter-period planets have smaller masses
- Exceptions are all in multiple star systems



Period /  
mass /  
eccentricity

# Stellar metallicity



# Orbital period ratios

